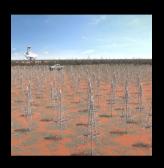


SKA Organization

- Australia, Canada, China, Germany, India, Italy, Netherlands, New Zealand, South Africa, Sweden, UK.
- Italy representative: Vettolani, Feretti
- China: agreement for 10 student/yr for 5 yrs to U. Western Australia to "learn and perform research in.. galaxy evolution, star formation.."

SKA – 1 & precursors



SKA1-LOW (MWA – Western Australia 2048 dipoles 80-300 MHz)



SKA1-SURVEY (ASKAP — Western Australia 36 12-m with Phased Array detectors 700 MHz to 1800 MHz)



SKA1-MID (MeerKAT – South Africa 64 12m antennas 900 MHz to 1670 MHz)

Update on timeline

- Re-baselining (== re-scoping!)
- Sep 2014 Engineering Meeting
- Input from Science Working Groups
- -> Science Review Panels
- Science and Engineering Advisory Committee (SEAC) -> SKA Board March 2015
- Critical Design Reviews in 2016 -> Construction starts

- OAB: Ghirlanda,
 Wolter (tra gli "editors")
- Bonnoli, Campana, Covino, D'Avanzo, Ghisellini, Guzzo, Melandri, Tagliaferri, Tavecchio, Trinchieri (tra gli "other contributors")
- Cosmology, HI, jets, GRBs

June 28, 2013

Italian SKA White Book

Editors: L. Feretti & I. Prandoni On behalf of the SKA-Italy WG



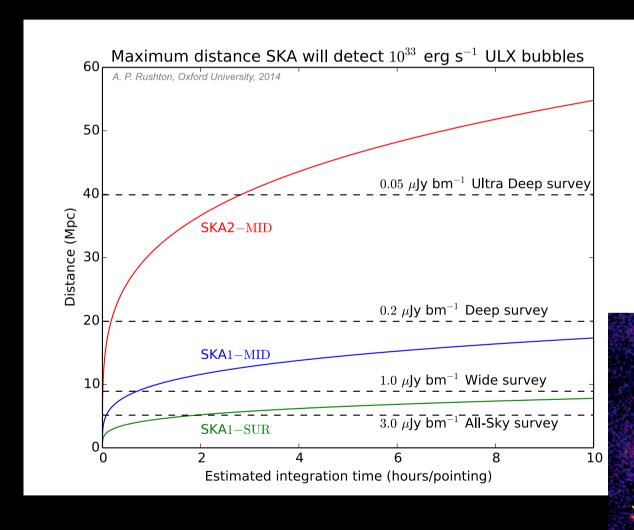




New Science Book (in press)

- Many chapters already uploaded on arXiv
- Radio Continuum: http://arxiv.org/html/1412.6942v1
 Revealing the Physics and Evolution of Galaxies and Galaxy Clusters with SKA Continuum Surveys by Prandoni & Seymour has 18 chapters (including: Radio investigation of Ultra-Luminous X-ray Sources in the SKA Era by Wolter et. al)
- Transients: *The SKA view of GRBs* by Burlon et al.
- Synergies etc: The connection between radio and high energy emission in black hole powered systems in the SKA era by Giroletti et al.
- <u>arXiv:1501.00420</u> Studies of Relativistic Jets in Active Galactic Nuclei with SKA Agudo et al.

SKA1 gives a complete look to the local Universe (up to 10-20Mpc)

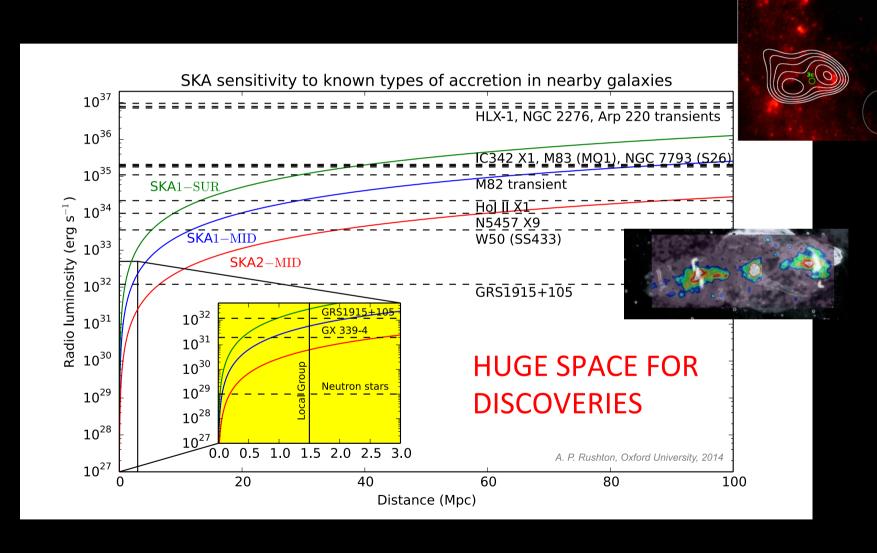


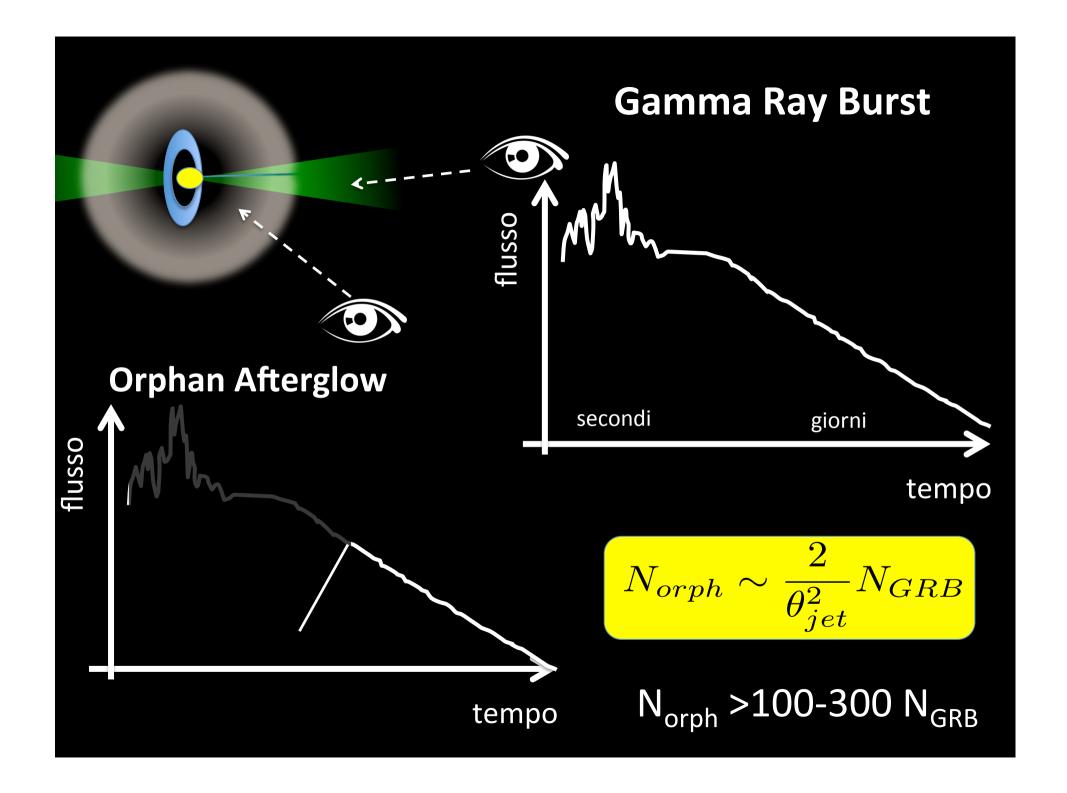
Extended structures: Bubbles, jets?

Core – point like

Transients

Limits for 1hr integration





Gamma Ray Bursts ←→ SKA



PB: Today only 30% of GRBs are detected in radio, WHY?

ANSW: due to radio faintness wrt current facilities and follow up strategy. SKA can follow up the entire GRB population (Ghirlanda et al. 2013 MNRAS, 425, 2543)



Search for signatures of Pop-III progenitors

- ATCA proposal nearly completed (PI D. Burlon)
- numerical simulatons on pop-III vs pop-II
- High energy properties of GRB-PopIII (F. Nappo)



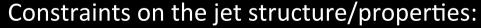
Radio observation of short GRBs:

ATCA proposal (PI - D. Burlon)





ANSW: Orphan Afterglows can be detected by forthcoming surveys (JVLA, ASKAP, SKA) 1.4 GHz @ F>0.1mJy \rightarrow 10⁻³ OA yr⁻¹ deg⁻² (Ghirlanda et al. 2014 PASA, 31, 22)





- JVLA survey proposal (PI R. Kunal)
- ALMA proposal for GAIA transients follow up (PI S. Covino)
- Multiwavength characterization of orphans

.. e inoltre

- Working Group sui transienti radio (P.I. R. Fender)
 - Fast Radio Bursts
 - Tidal Disruption Events
- Epoch of Reionization
- Continuum
- Cosmology
- Cradle of Life
- HI galaxy science
- Magnetism
- Pulsars
- Transients
- Proposals ? Collaboration for precursors?
- INAF PhD fellowships???