



NGS2: GeMS' new NGS WFS

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What is GeMS ?

HP 1 - A Fossil Relic in the Galactic Bulge





VVV - VISTA (4m) FWHM~0.8", Kam ~ 17





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NAC-CNAC

GSAOI+GeMS - Gemini-South (8m) FWHM~0.1", Kim ~ 20



Outlines

- Previous NGS module
 - design
 - limitations
- NGS2
 - design
 - benefits
- NGS2 preliminary on-sky performance



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GeMS NGS1 WFS



- Tiny, difficult to align fibre feed
- 4% optical throughput!

Credit: F. Rigaut

Several attempts to fix but little improvement.

* Note that this has nothing to do with pyramid sensing



















NGS1 WFS issues

- Impact Guide Star acquisition
 - Linear model
 - No distortion model
- Asterism distorted as it offsets
 - induce plate scale + rotation in science exposure
 - Guide star end up outside the probe





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NGS1 WFS display

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NGS1 WFS display

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NGS1 limitation

- Blind acquisition into tiny APD device
- Very low throughput (4%)
- Sky coverage very limited
 - V-band 15.5 under GREAT conditions
- Distortions unmapped



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NGS2 requirements

- TT sensing on up to 3 NGSs
- Use a focal plane array
- Frame rate up to 800 Hz
- No vignetting
- Rlim = 17 with noise better than 35mas
- Focus adjustement to minimize spot size
- No heat nor light input in canopus
- As simple as possible to integrate in GeMS



















NGS2 Nuvu

Camera spec

- 512x512 pixels
- 16um pixels
- EM gain up to 5000
- Low CIC 0.00058e-/px/fr
- RON ~300/EMgain
- Multi ROI capability
- pixel scale 263mas







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NGS2 Solution

- Focal plane sensor: Re-image system
 - Throughput exceed 75%
 - Solves limiting magnitude (sky coverage)
- Access to full field
 - FF imaging during acquisition simplifies ops
 - Field distortions can be mapped and include in window position model
- No moving mechanism across the field



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Non-sidereal observations

Neptune observation Guiding on Triton

Non-sidereal observations

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NGS2 observation: Science performance S20191018S0105.fits 4233x4233

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Conclusion

- NGS Works !
 - Being able to get 30mas TT on R=17.5
 - Even close loop on R=18.2 !
 - Major upgrade for GeMS Sky Coverage !
- Ultra-fast acquisition
 - Full Field imaging during acquisition simplifies ops
 - Field distortions can be mapped and include in window position model
- No moving mechanism across the field

