

# The Adaptive Optics Module of MAVIS: preliminary design and trade-offs

Valentina Viotto  
on behalf of the **MAVIS AOM** team

WFS in the VLT/ELT era IV Workshop 2019

# What is MAVIS?

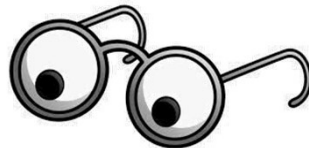


**MCAO**  
**Assisted**  
**Visible**  
**Imager (and)**

30"x30" corrected FoV

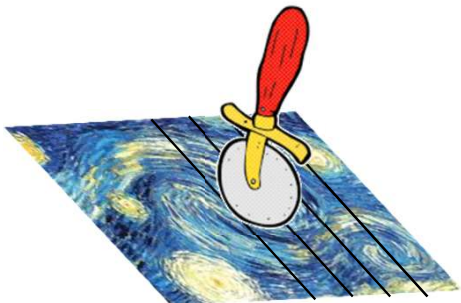
VRI

FoV 30"x30"  
pxscl 7.1mas



**Spectrograph**

Monolithic IFU  
3.6"x2.5"  
370-950 nm  
R 5000-10000



Australian  
National  
University



INAF  
ISTITUTO NAZIONALE  
DI ASTROFISICA  
NATIONAL INSTITUTE  
FOR ASTROPHYSICS



MACQUARIE  
University  
SYDNEY · AUSTRALIA



ONERA  
THE FRENCH AEROSPACE LAB

## MAVIS AOM core team:

G. Agapito, M. Aliverti, M. Bergomi, M. Bonaglia, L. Busoni, T. Fusco, G. Gausachs, D. Gratadour, D. Greggio, P. Haguenhauer, L. Marafatto, B. Neichel, E. Pinna, C. Plantet, K. Radhakrishnan, **F. Rigaut**, S. Stroebele, A. Vaccarella, D. Vassallo, V. Viotto

# Already existing AO equipment

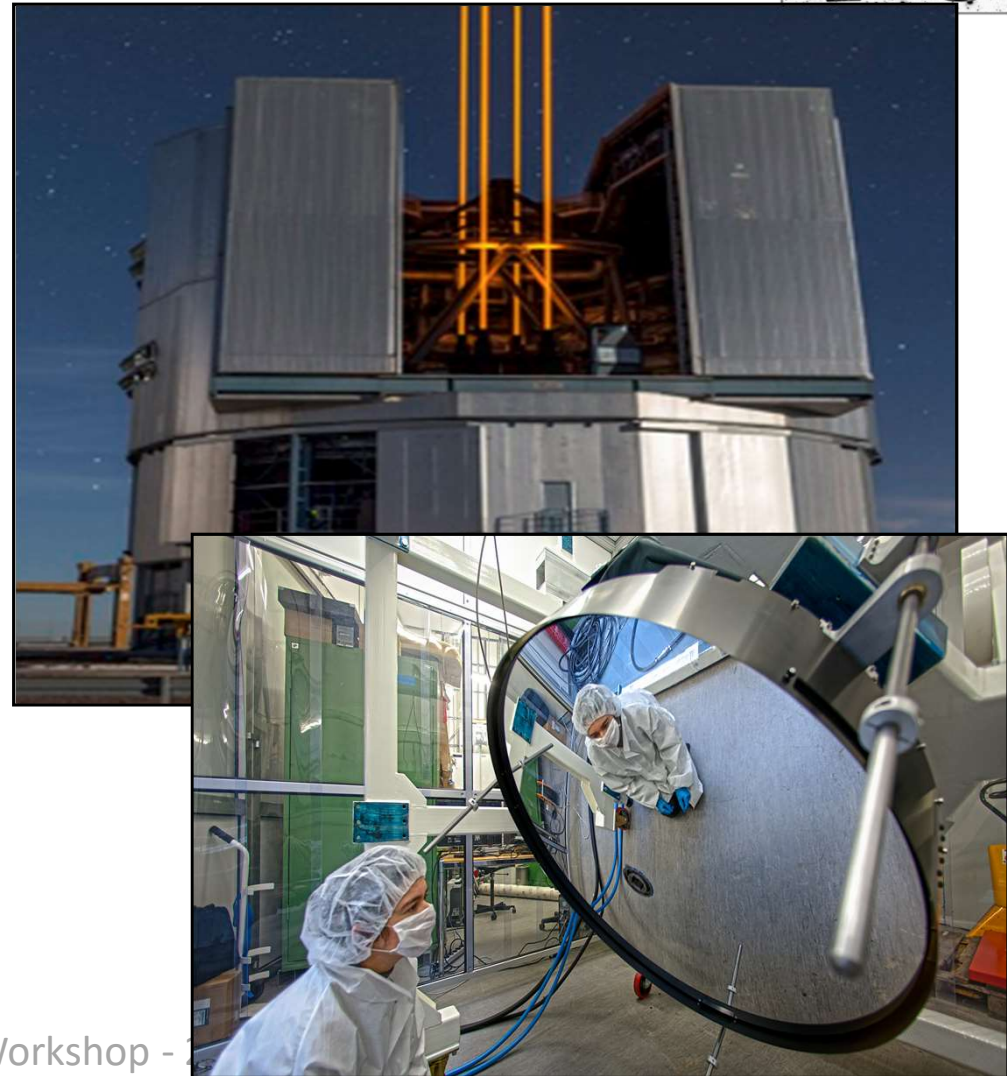


## Context:

### Adaptive Optics Facility completed

- Where: VLT UT4
- What: key technical components are
  - **Deformable Secondary Mirror** (1170 act)
  - **4 Laser Guide Stars** (routinely working)
- Serving:
  - MUSE: Optical IFU
  - HAWK-I: Wide-field IR imager
  - ERIS (from 2020): 1-5 $\mu$ m imager/IFU

**AOF full potential not exploited yet**



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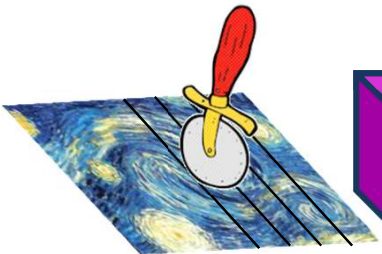
# AOM conceptual view

VLT Nas.

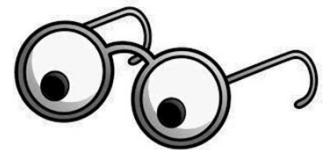
## AOM machine



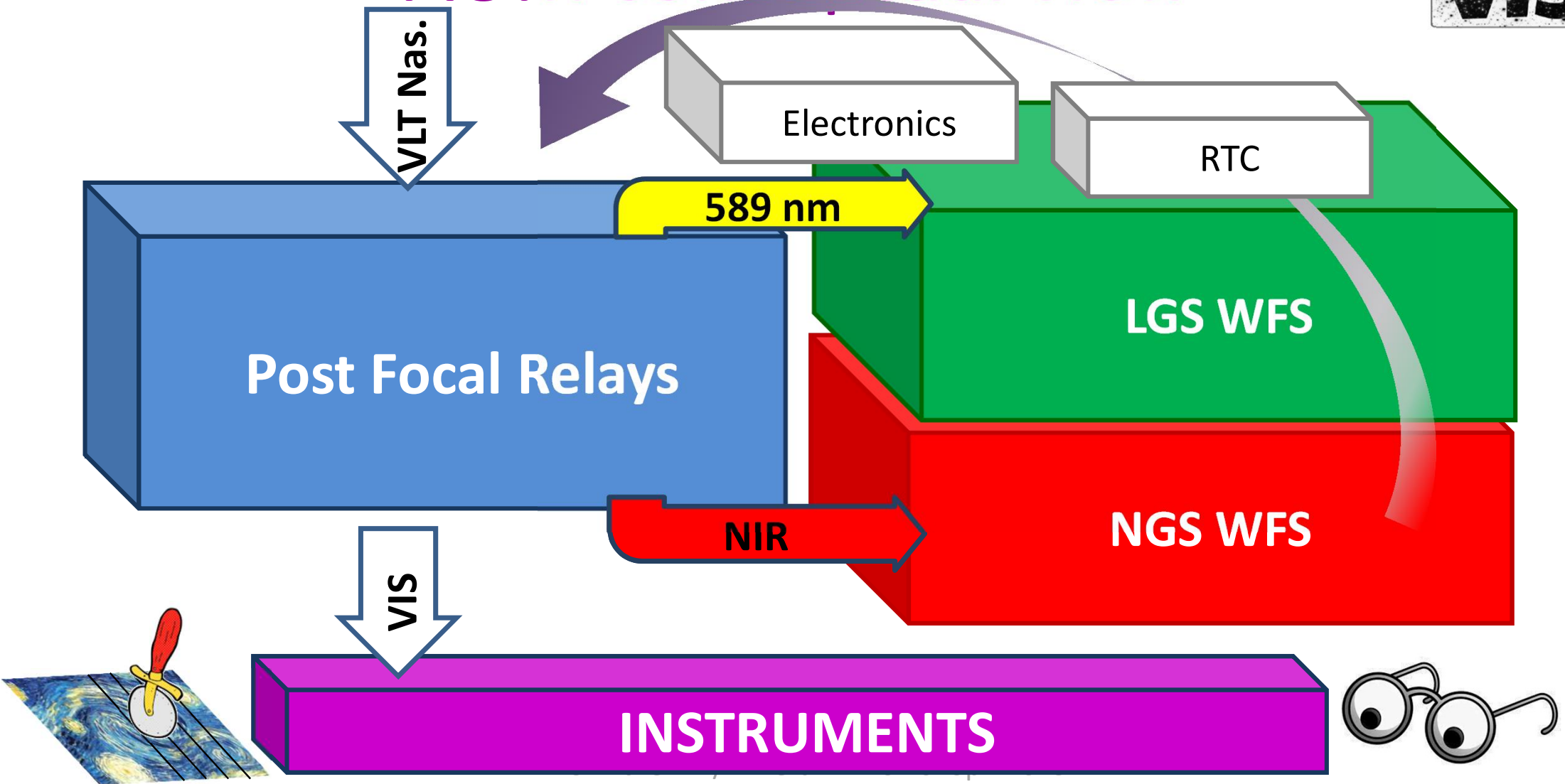
VIS



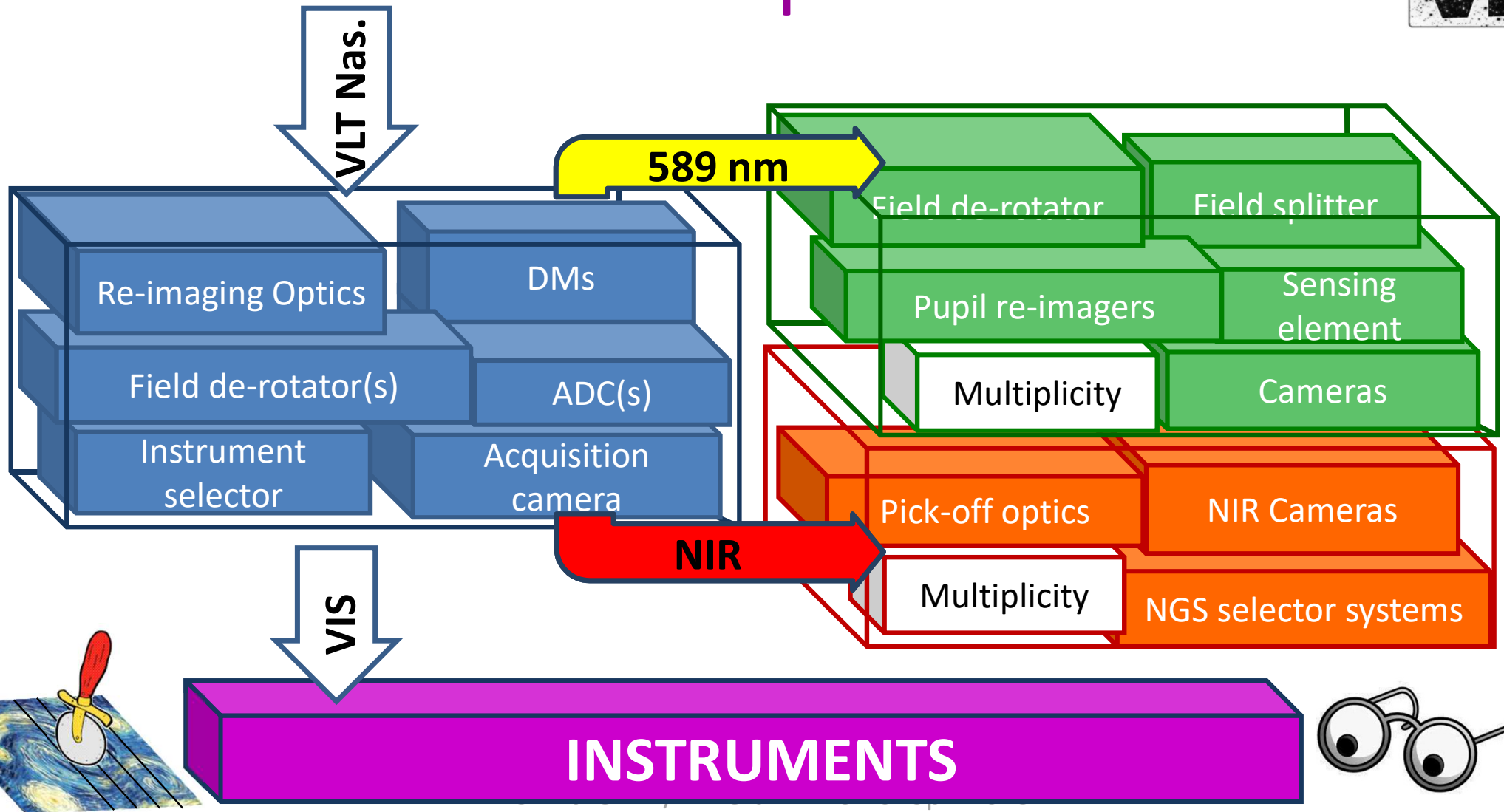
**INSTRUMENTS**



# AOM conceptual view

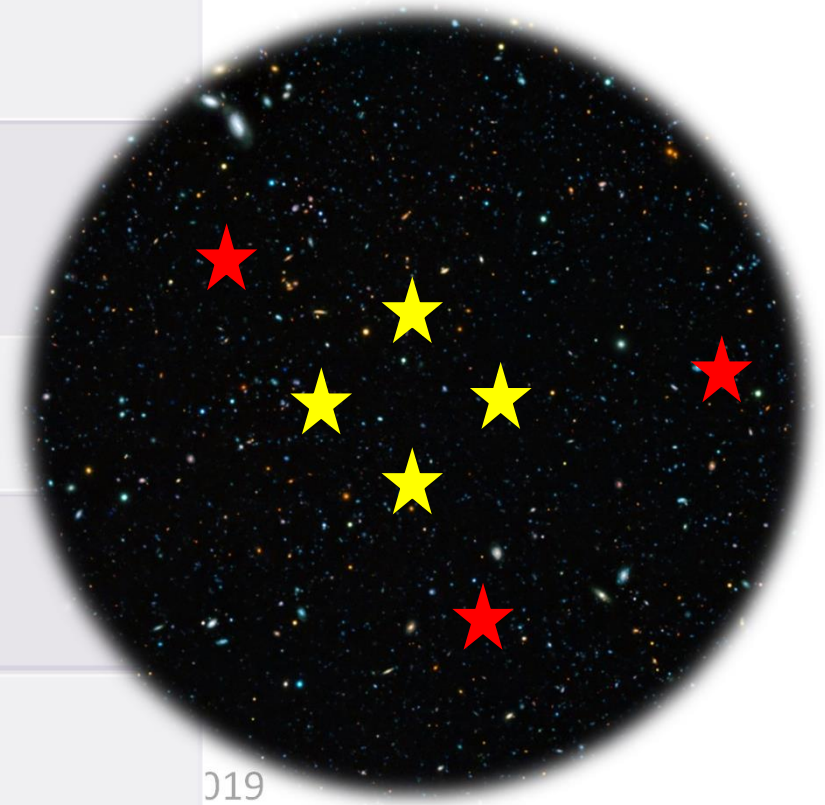


# AOM conceptual view



# AOM trade-offs

Trade-off	Questions
De-rotation scheme	<ul style="list-style-type: none"> <li>How to compensate NGS FoV, LGS FoV and SCI FoV apparent rotations</li> </ul>
PFR optical design	<ul style="list-style-type: none"> <li>Reflective vs Transmissive</li> <li>1 or 2 ADCs</li> </ul>
DM configuration	<ul style="list-style-type: none"> <li># of DMs</li> <li>Conj. altitude</li> <li>Pitch</li> </ul>
LGS configuration	<ul style="list-style-type: none"> <li># of LGSs</li> <li>Asterism</li> </ul>
NGS WFS trade-off	<ul style="list-style-type: none"> <li># of NGSs</li> <li>FoV accessibility</li> <li>Before/after DMs</li> </ul>
Wavelengths	<ul style="list-style-type: none"> <li>How blue?</li> <li>J-H band for NGS?</li> </ul>





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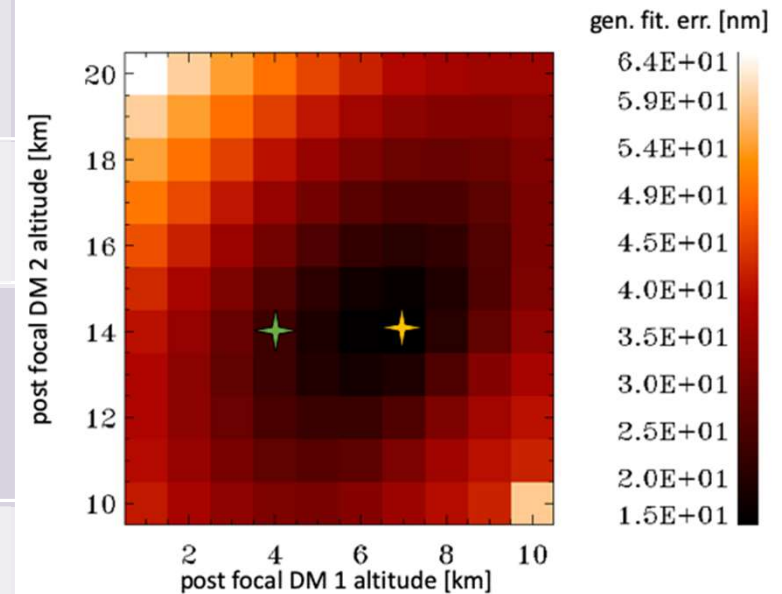
## Rationales:

- Quality
- Throughput
- MAIT
- Flexibility
- Modularity

# AOM trade-offs



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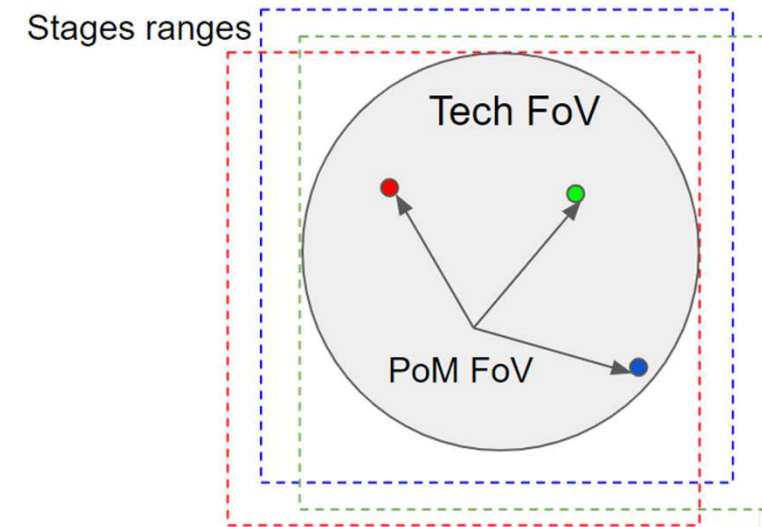
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*«One laser, two LGS for two WFS: technology, on-sky demonstration and first results»*  
-> P. Haguenauer tomorrow!!!



# AOM trade-offs

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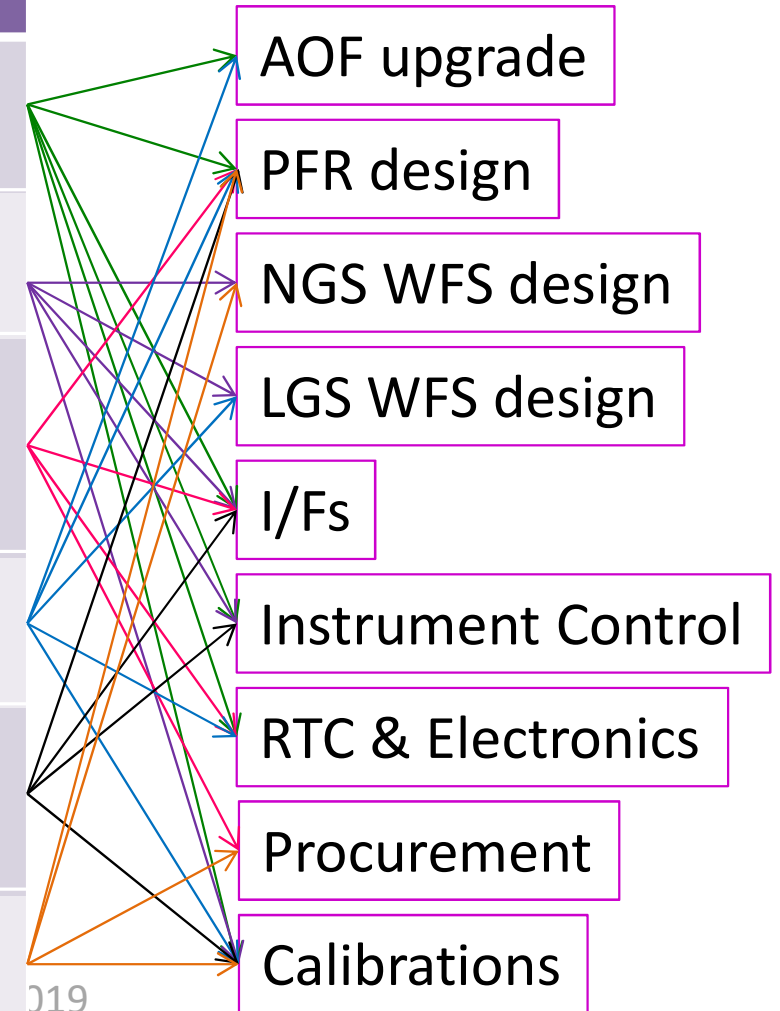


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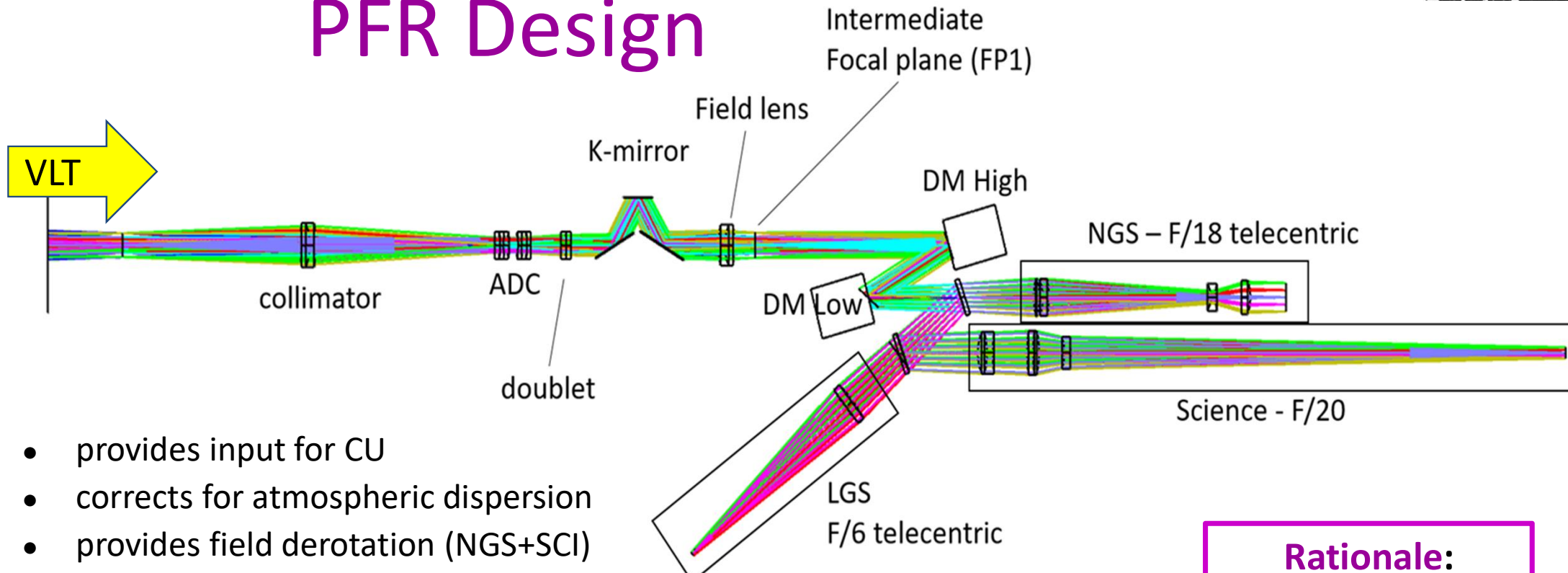
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# PFR Design



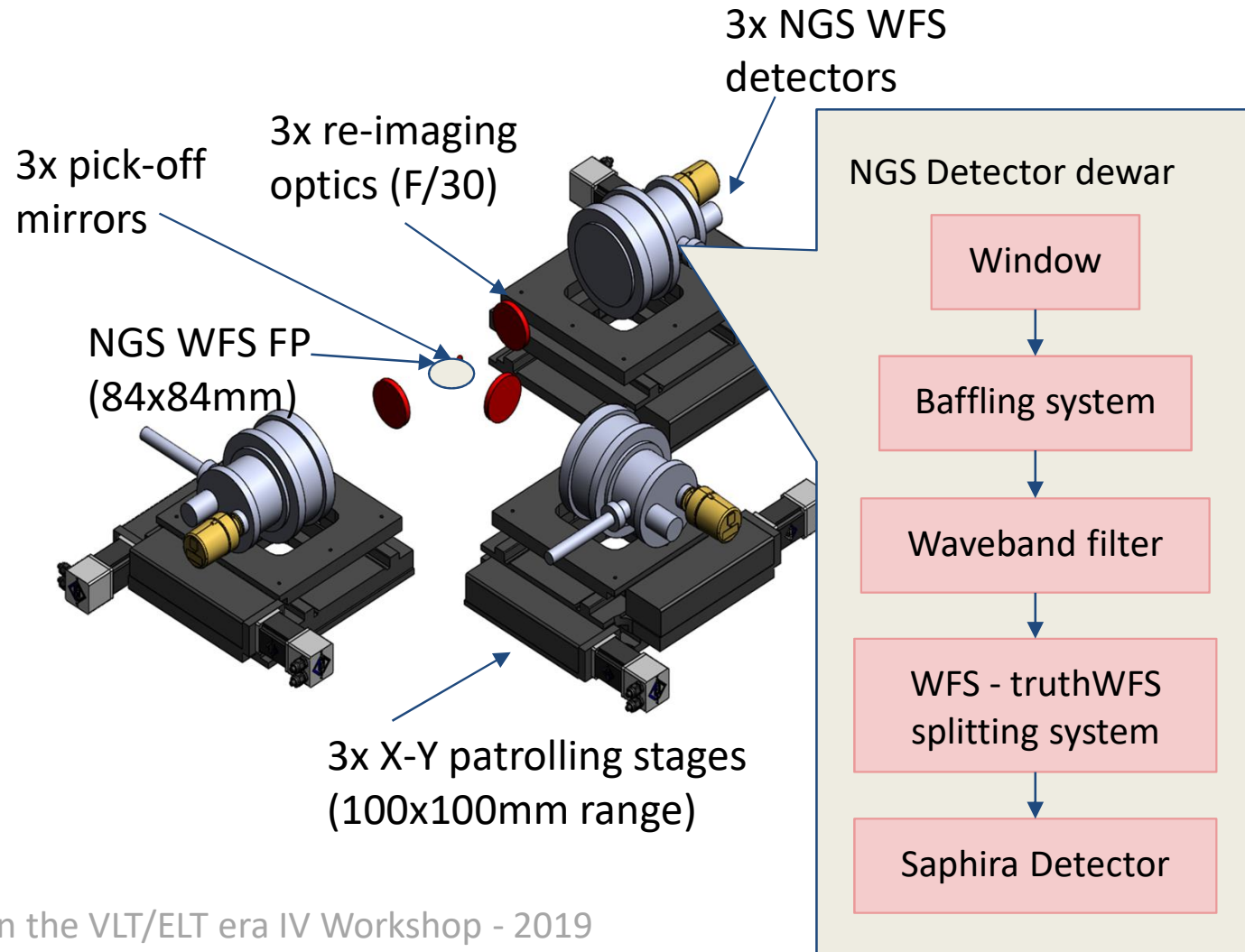
- provides input for CU
- corrects for atmospheric dispersion
- provides field derotation (NGS+SCI)
- re-images meta-pupils onto DMs
- feeds the NGS WFS with a 2 arcmin FoV at infinity (NIR band)
- provides means for stars acquisition [acquisition camera still missing]
- feeds the LGS WFS with the LGSes FoV at 90-230km altitude (Na line)
- delivers a 30" diameter FoV to the instruments (VIS band)
- provides at least 2 output ports and includes means to switch

## Rationale:

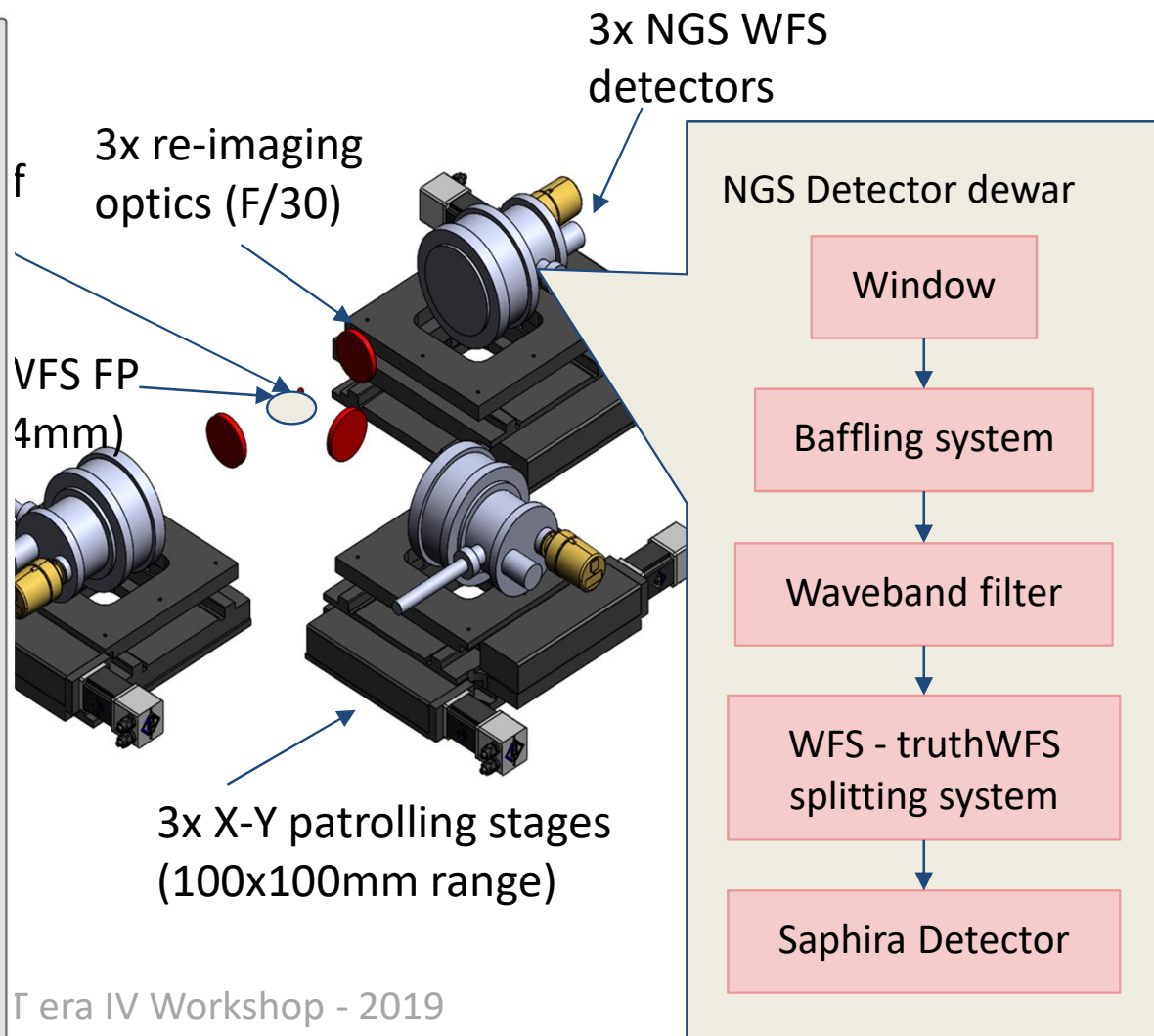
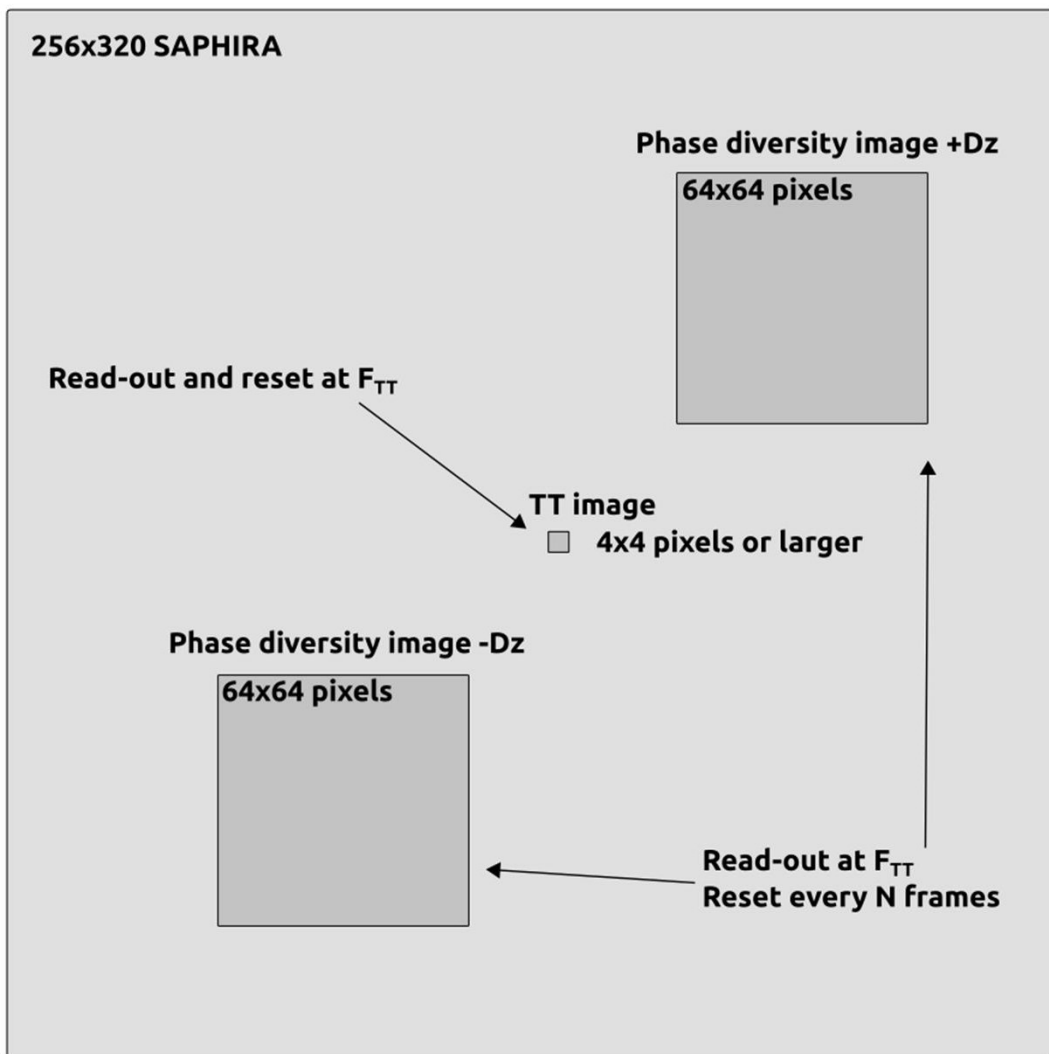
- Distortions
- Throughput
- Compactness
- Alignment
- Modularity

# NGS WFS Design

- **3 NGS**es can be sensed, in the J+H bands, at the same time
- 2 arcmin diameter unrestricted patrolling **FoV** (including SCI FoV)
- **1x1 subaperture** pupil sampling
- **30mas sampling** on the detector
- Tomographic **truth sensing** included



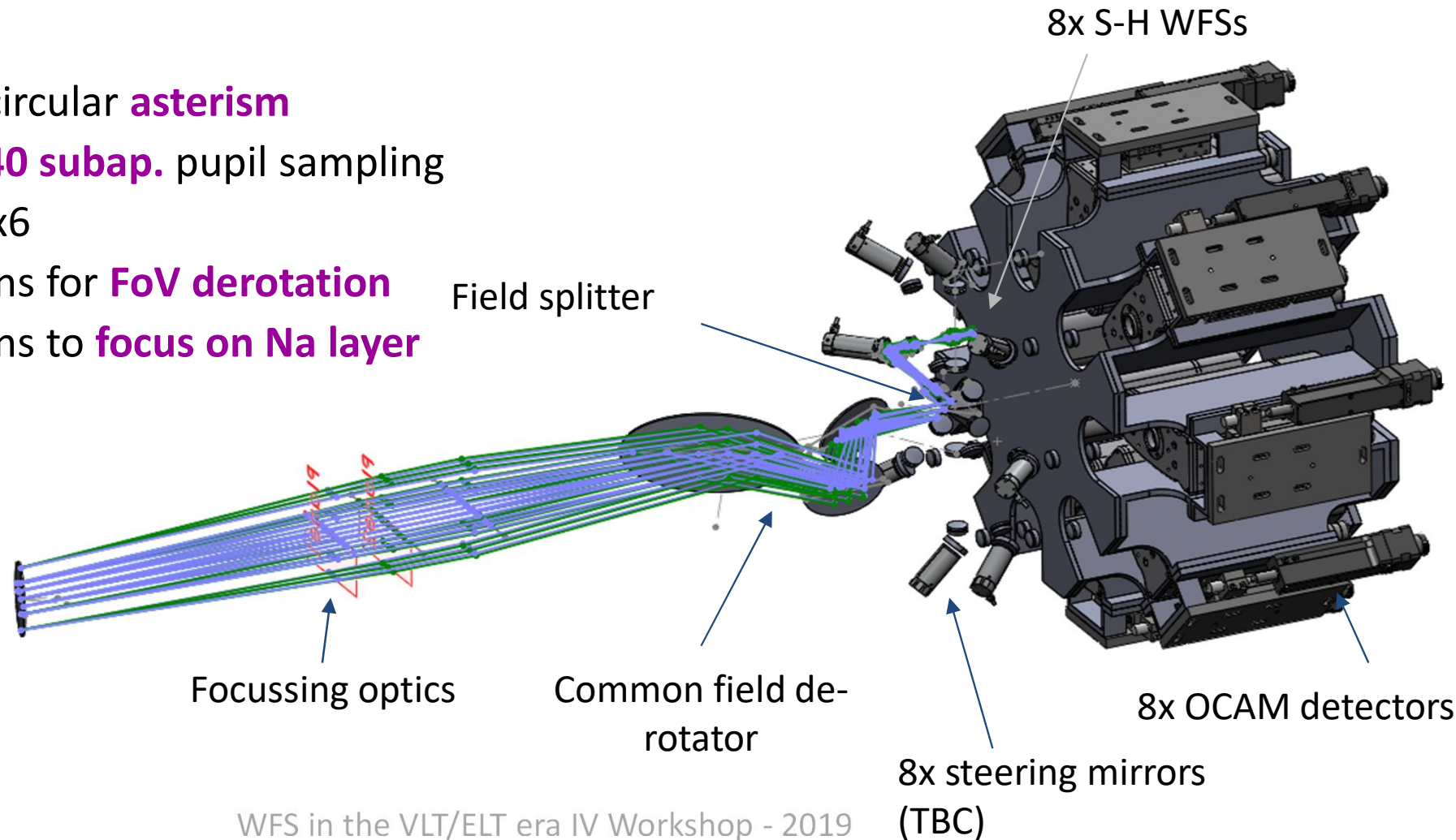
# NGS WFS Design



# LGS WFS Design



- 8 LGSes
- 17.5" radius circular **asterism**
- S-H WFS **40x40 subap.** pupil sampling
- #pix/subap 6x6
- includes means for **FoV derotation**
- includes means to **focus on Na layer** (elevation)



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# Simulations

## ESO TLR atmospheric profile:

$L_0$  [m]=25m

seeing (zenith) [arcsec] = 0.8

Na multi-peak profile

## Telescope:

Pupil = 8m with 16% obscuration

Zenith angle= 30deg

TFoV 120arcsec diameter

## DMs:

pitch [0.22,0.22,0.3]m

altitude [0,4,14] km

## Control:

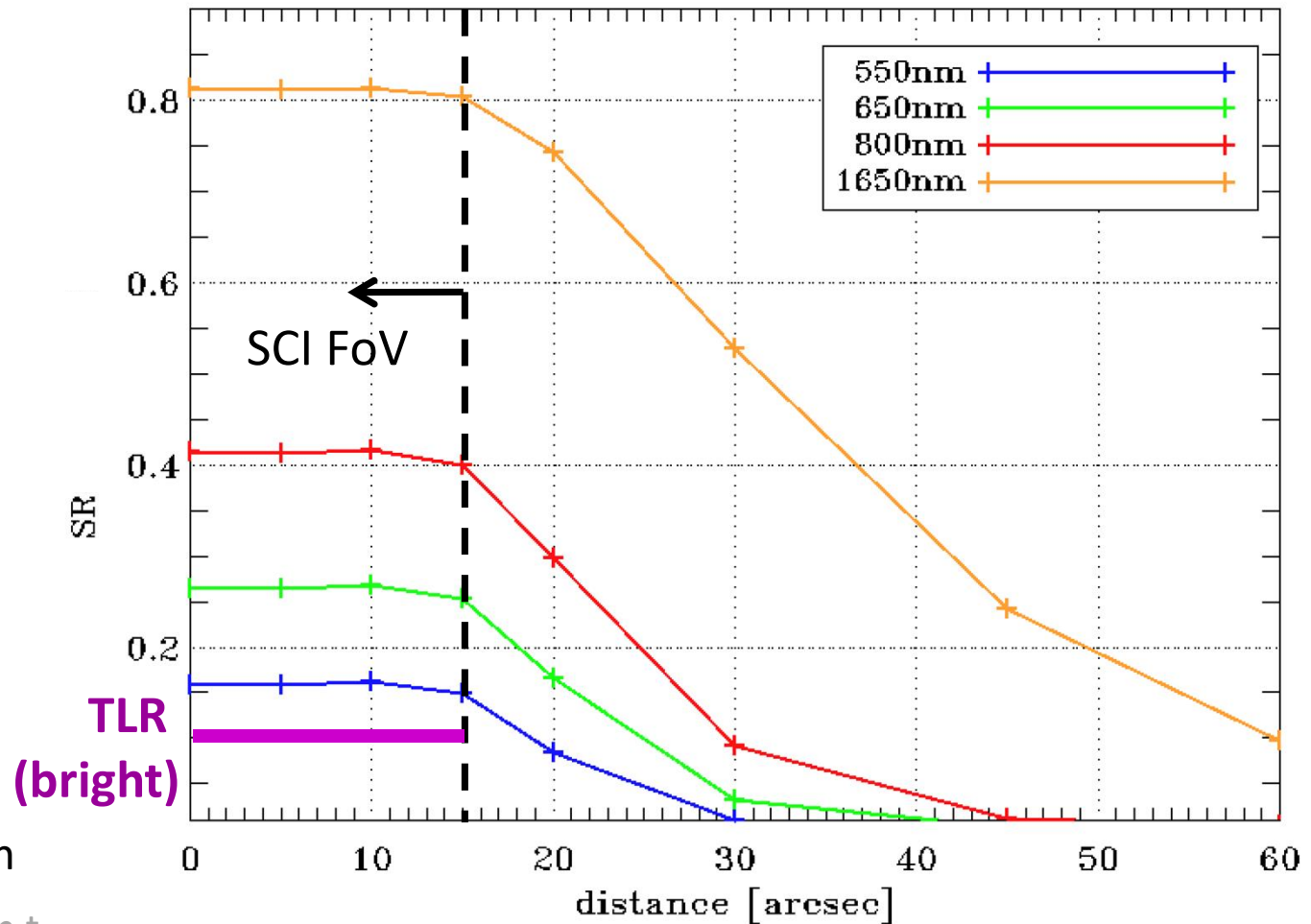
Rec layers @ Cn2 profile altitude.

Opt. FoV: 15arcsec radius

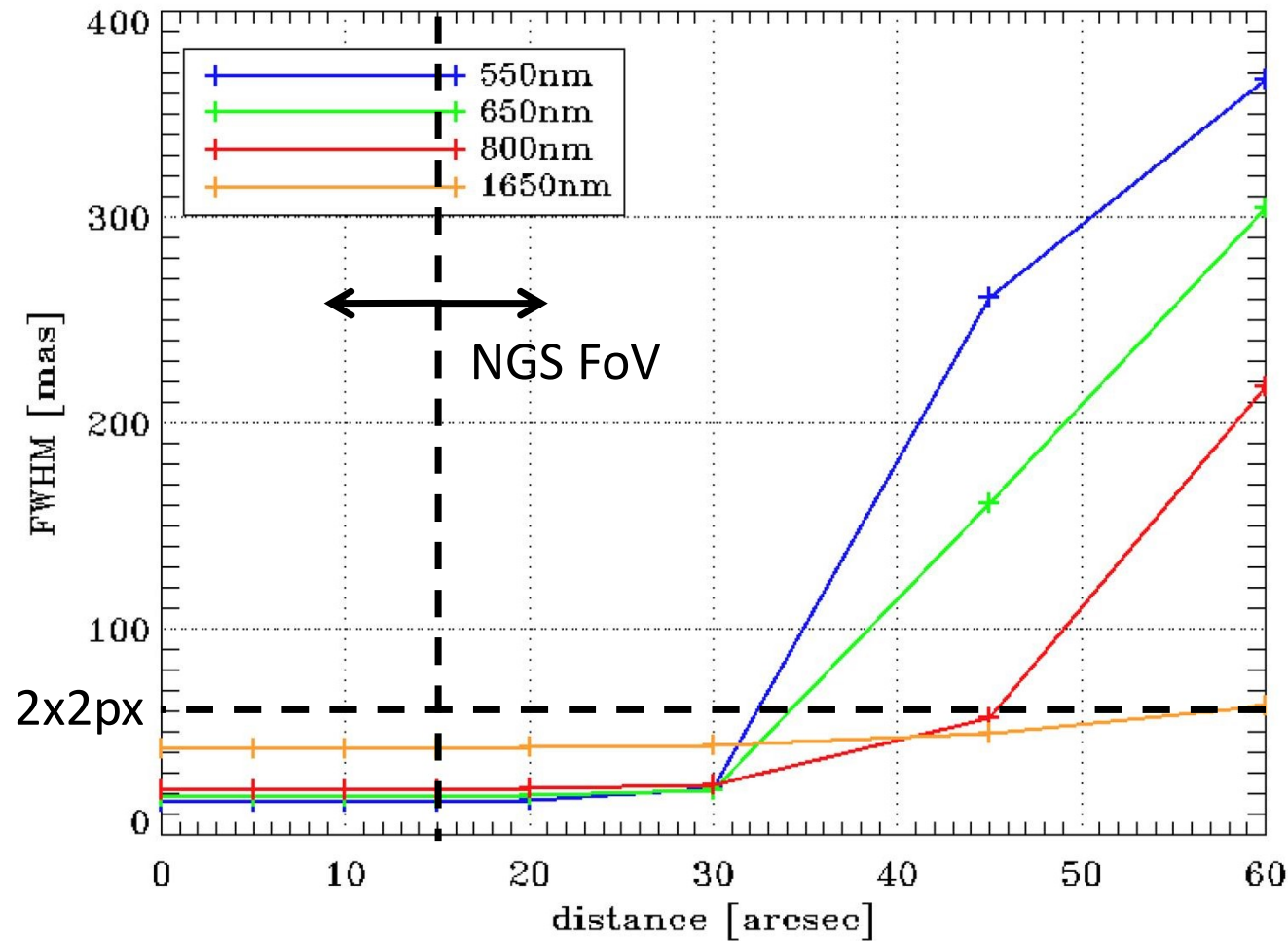
POLC

Integrator controllers with 0.3 gain

Delay = 2frames (2ms)



# Simulations - TFoV



## Sky Coverage criteria:

- Strehl Ratio
- Encircled Energy (info loss)
- FWHM (sensitivity)

# Simulations – Sky Coverage



## Assumption and terms

Tomographic error

(parameters: off-axis jitter + availability of stars )



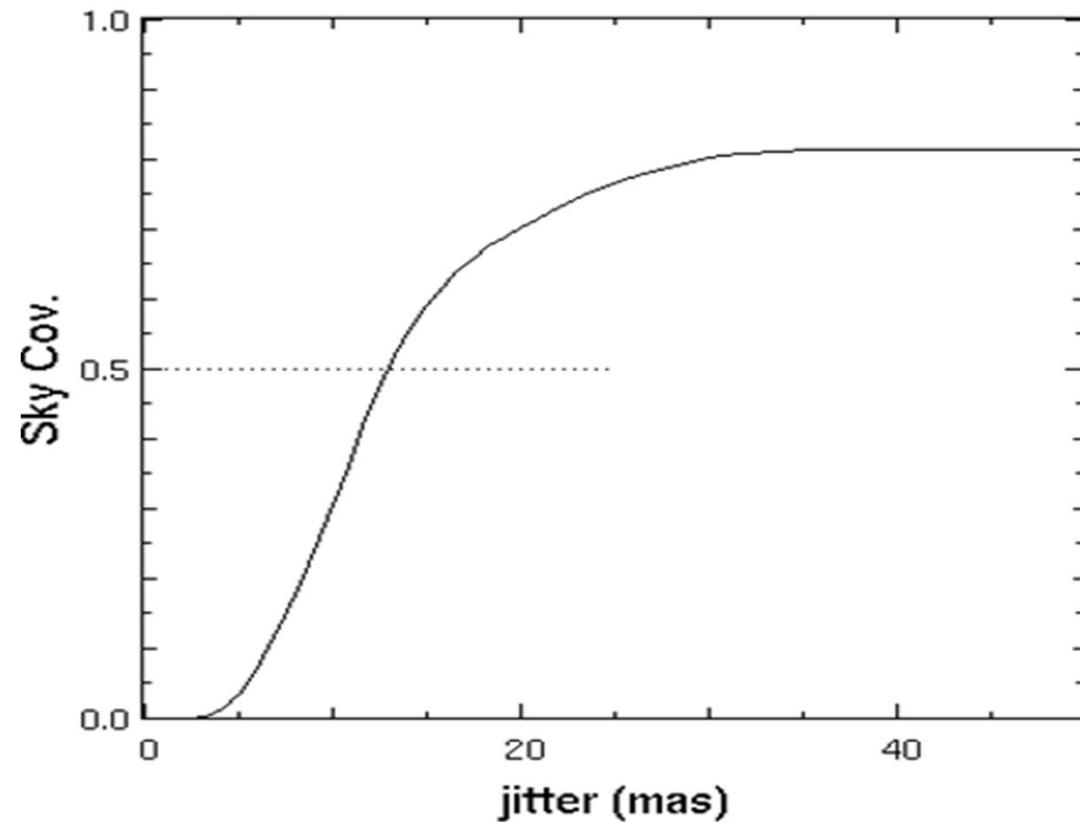
Windshake error

(parameter: frame rate)



Noise error

(parameter: ph/subap/frame)





# WFE Preliminary Budget

TLR-33: SR = 10% (15% goal) at V band (550 nm),  
~132 nm (120 nm goal).

Contributions	Nominal [nm]	Calibration residual (%)	After calibration [nm]
AO (no vibration, bright NGS)	114	100	114
NCPA	50	50	25
Manufacturing	29	100	29
Alignment	8	100	8
Thermal	TBD		
Image motion	TBD		
<b>Total</b>			<b>120.5 nm</b>



Loading work in progress

# Thank you



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