



**SISSA**

# Merging rates of compact binaries in galaxies

Lumen Boco

**Supervisors:**

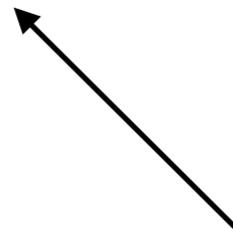
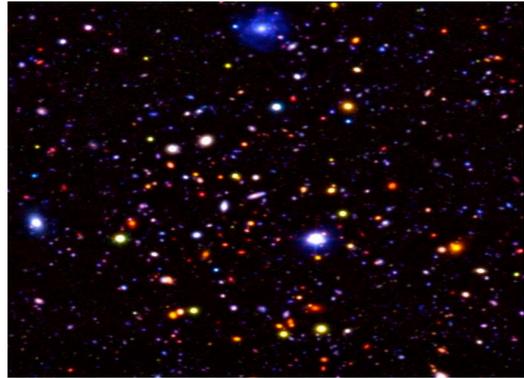
Andrea Lapi

Carlo Baccigalupi

Francesca Perrotta

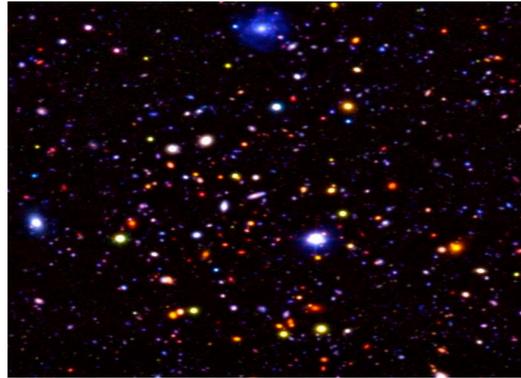
# **Merging rates of compact binaries in galaxies**

Galaxy statistics

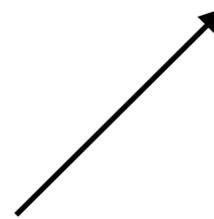
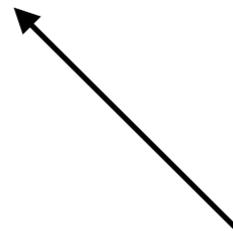


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Galaxy statistics

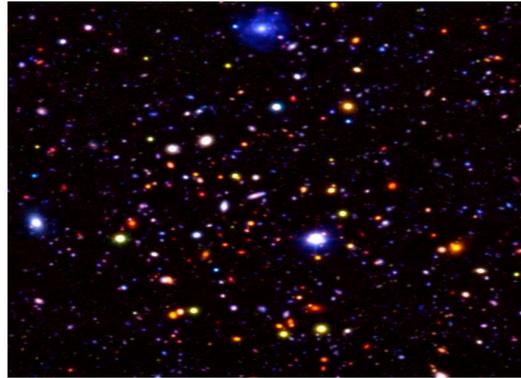


Single galaxy evolution



# Merging rates of compact binaries in galaxies

Galaxy statistics

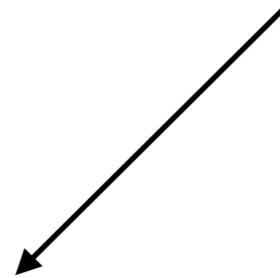
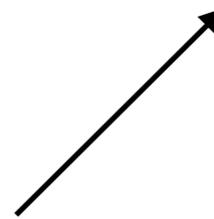
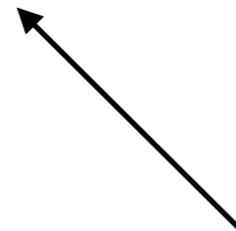


Single galaxy evolution

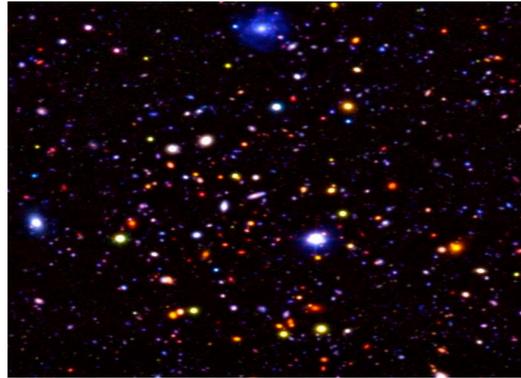


# Merging rates of compact binaries in galaxies

Stellar evolution



Galaxy statistics



Single galaxy evolution

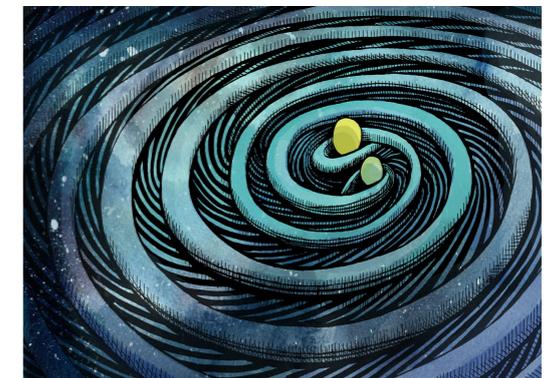


# Merging rates of compact binaries in galaxies

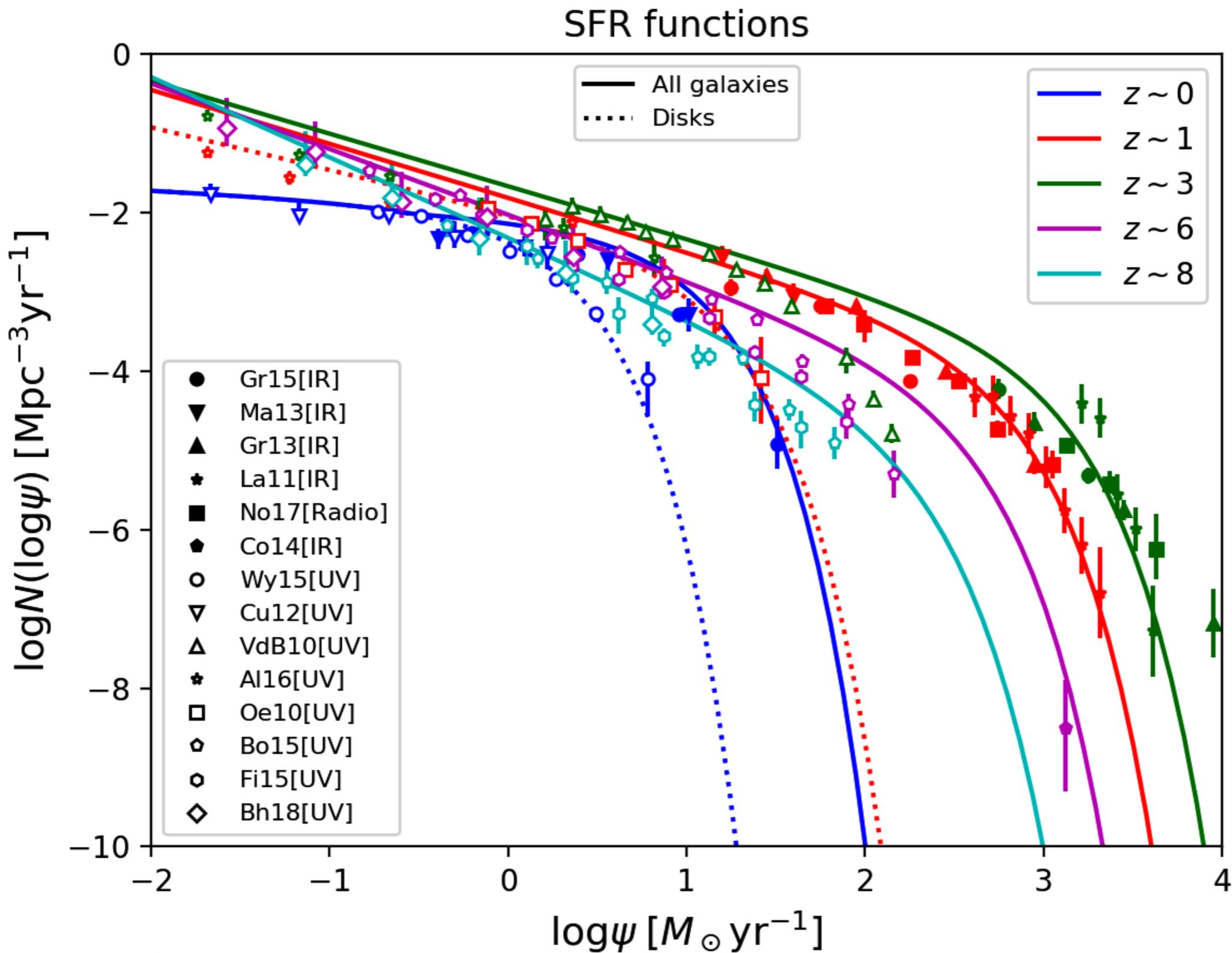
Stellar evolution



GW physics

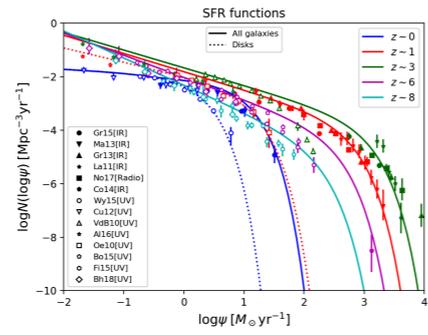


# Galaxy statistics



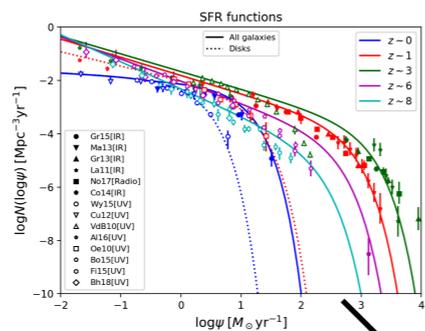
# Galaxy statistics

$$\int d \log \psi \psi$$

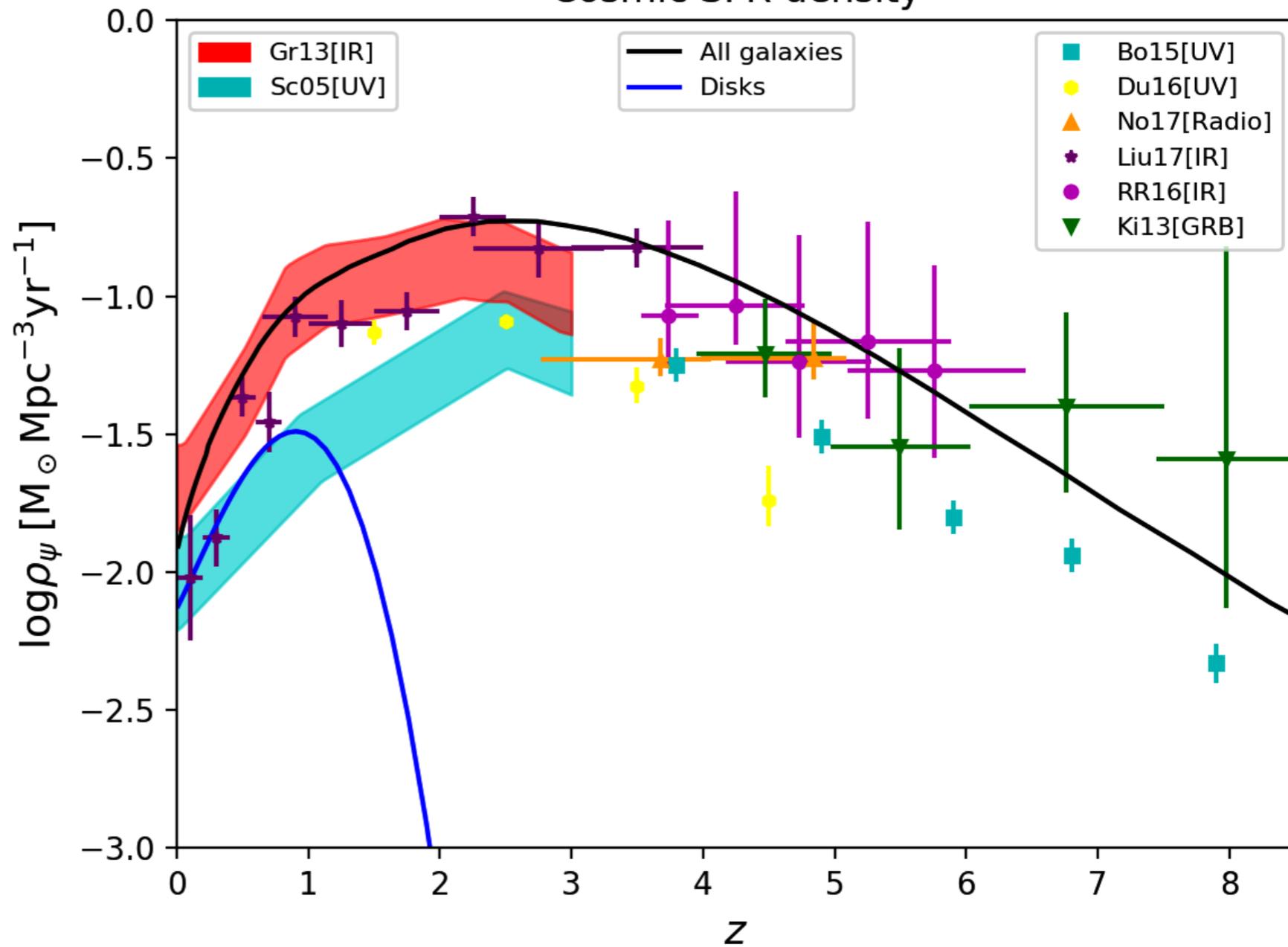


# Galaxy statistics

$$\int d \log \psi \psi$$



## Cosmic SFR density



# Galaxy evolution

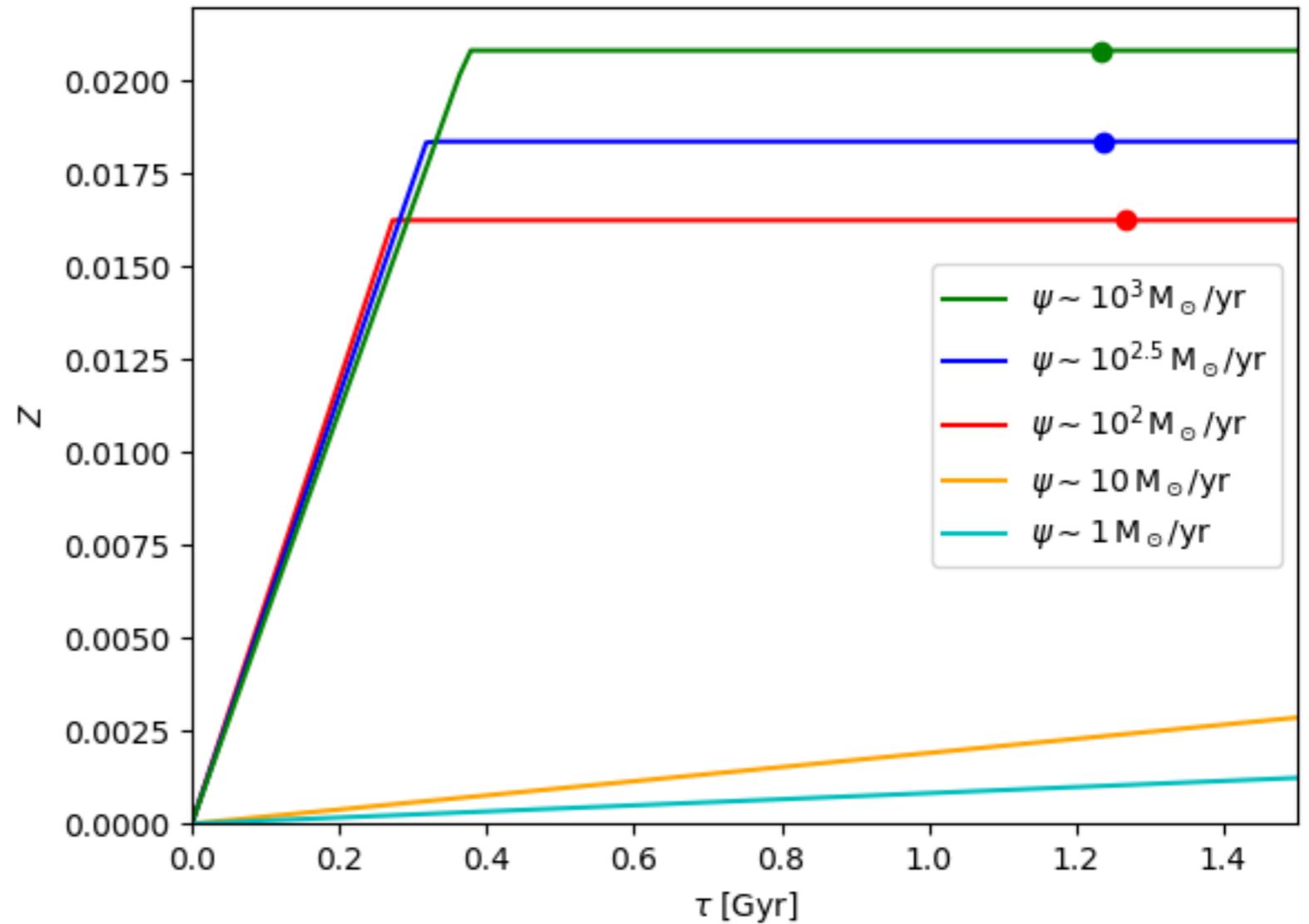
Chemical evolution

# Galaxy evolution

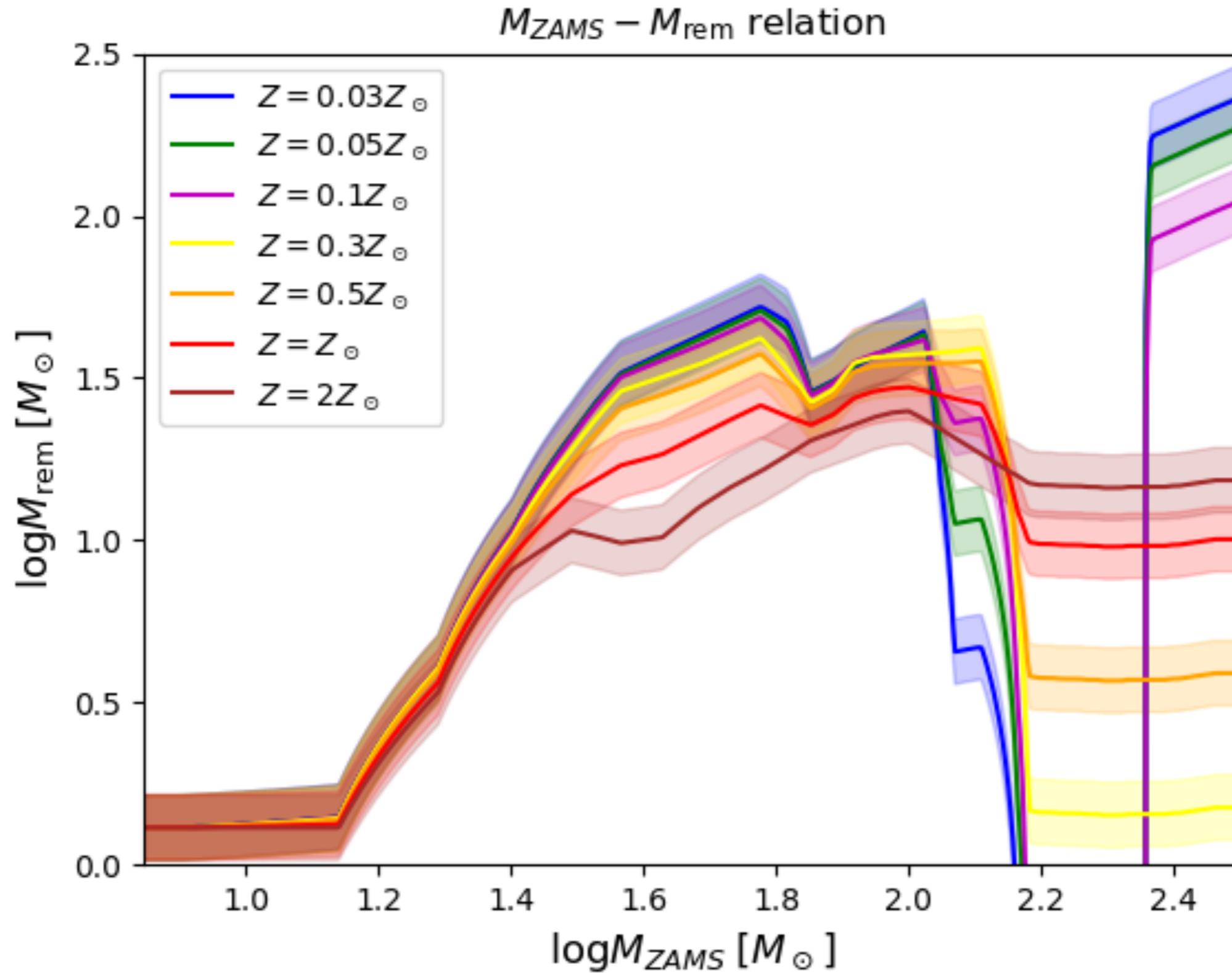
Chemical evolution

$$Z(\tau) = \begin{cases} Z_{\text{sat}} \frac{\tau}{\Delta\tau_{\psi}} & \frac{\tau}{\tau_{\psi}} \leq \Delta \\ Z_{\text{sat}} & \frac{\tau}{\tau_{\psi}} > \Delta \end{cases}$$

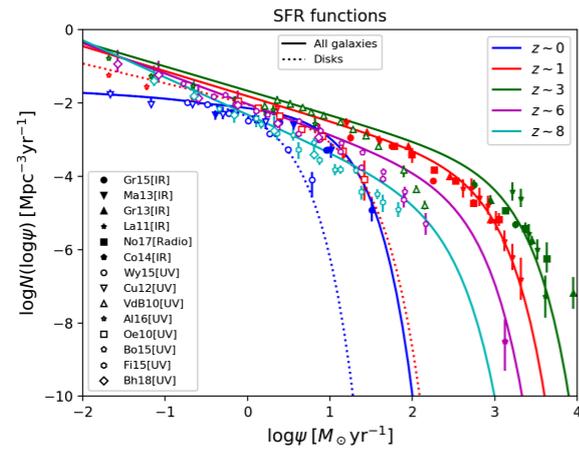
Metallicity evolution



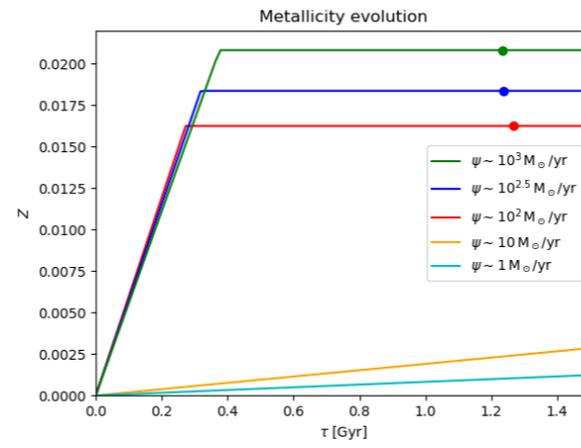
# Stellar evolution



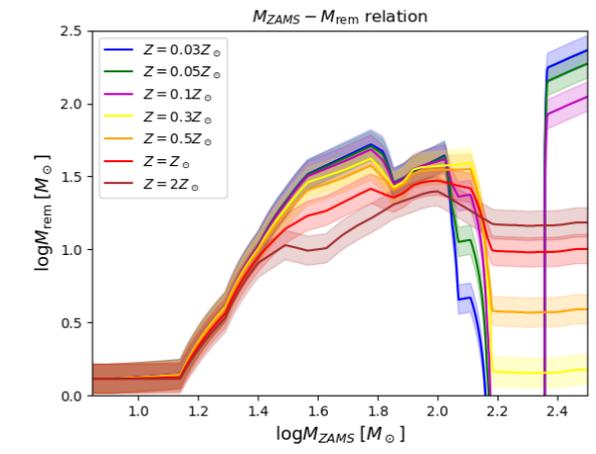
# Galaxy statistics



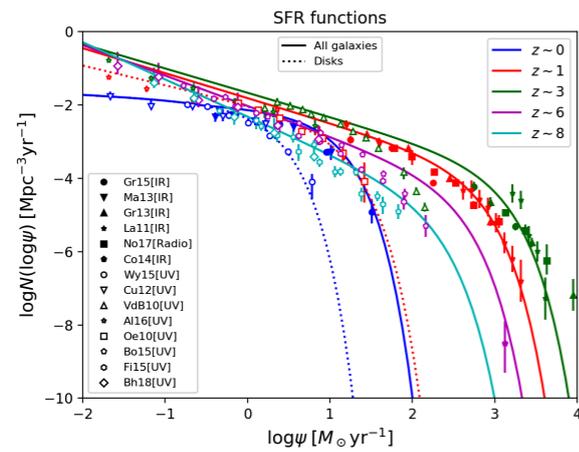
# Galaxy evolution



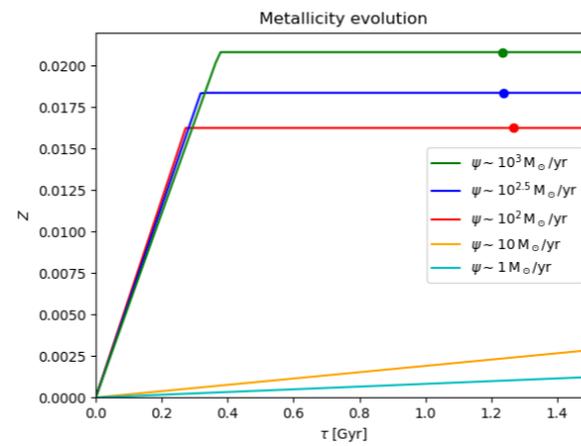
# Stellar evolution



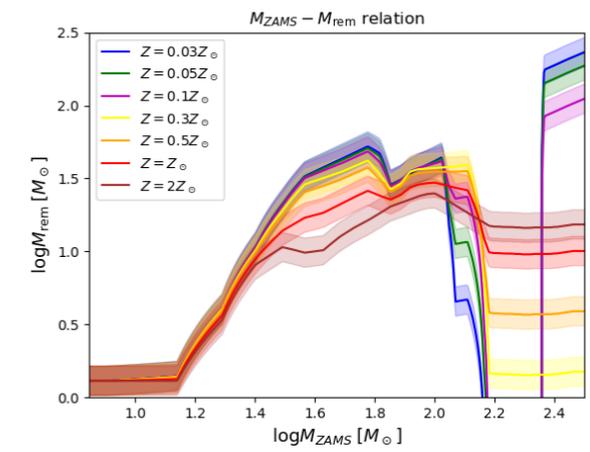
# Galaxy statistics



# Galaxy evolution



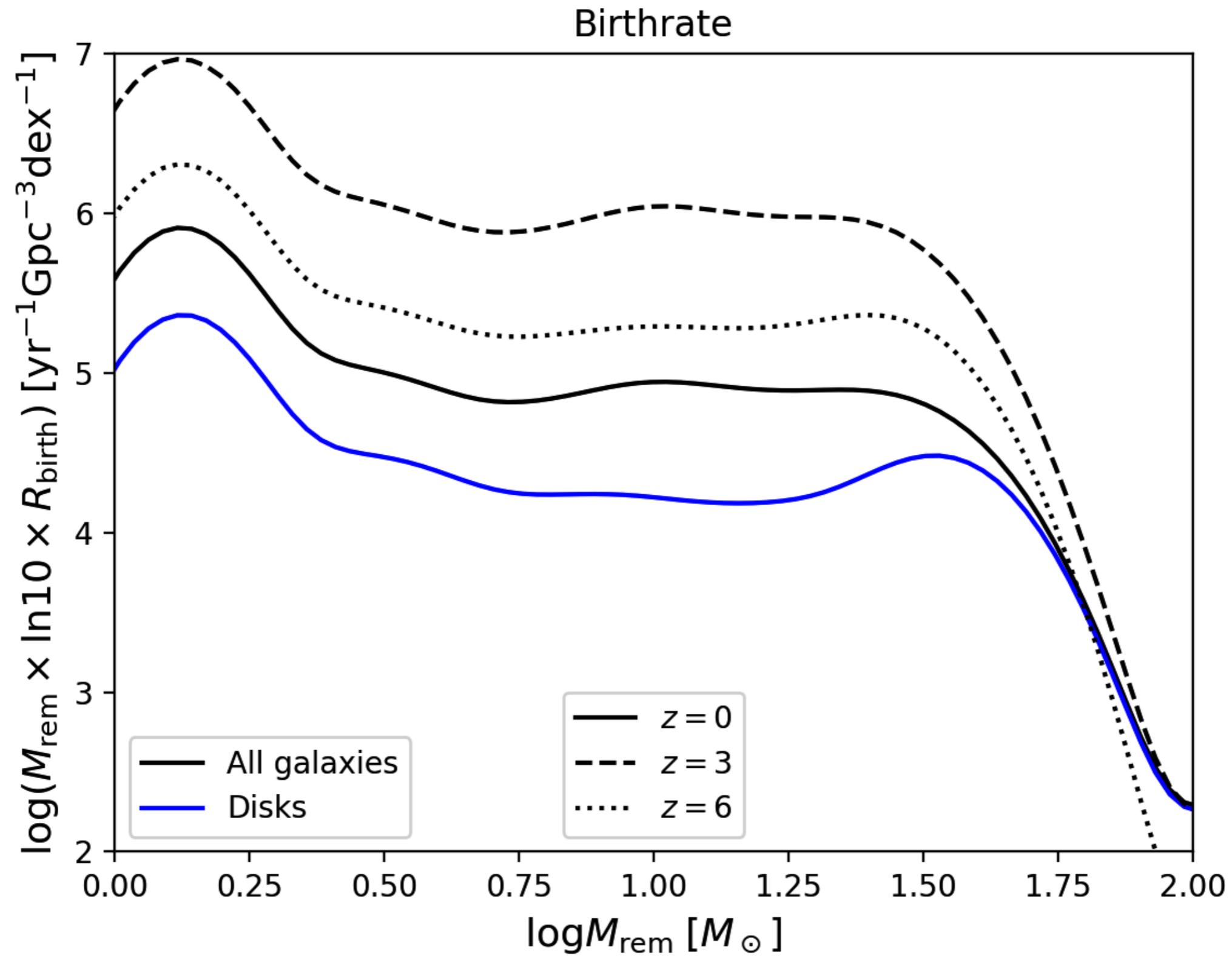
# Stellar evolution



## Cosmic birthrate of compact remnants

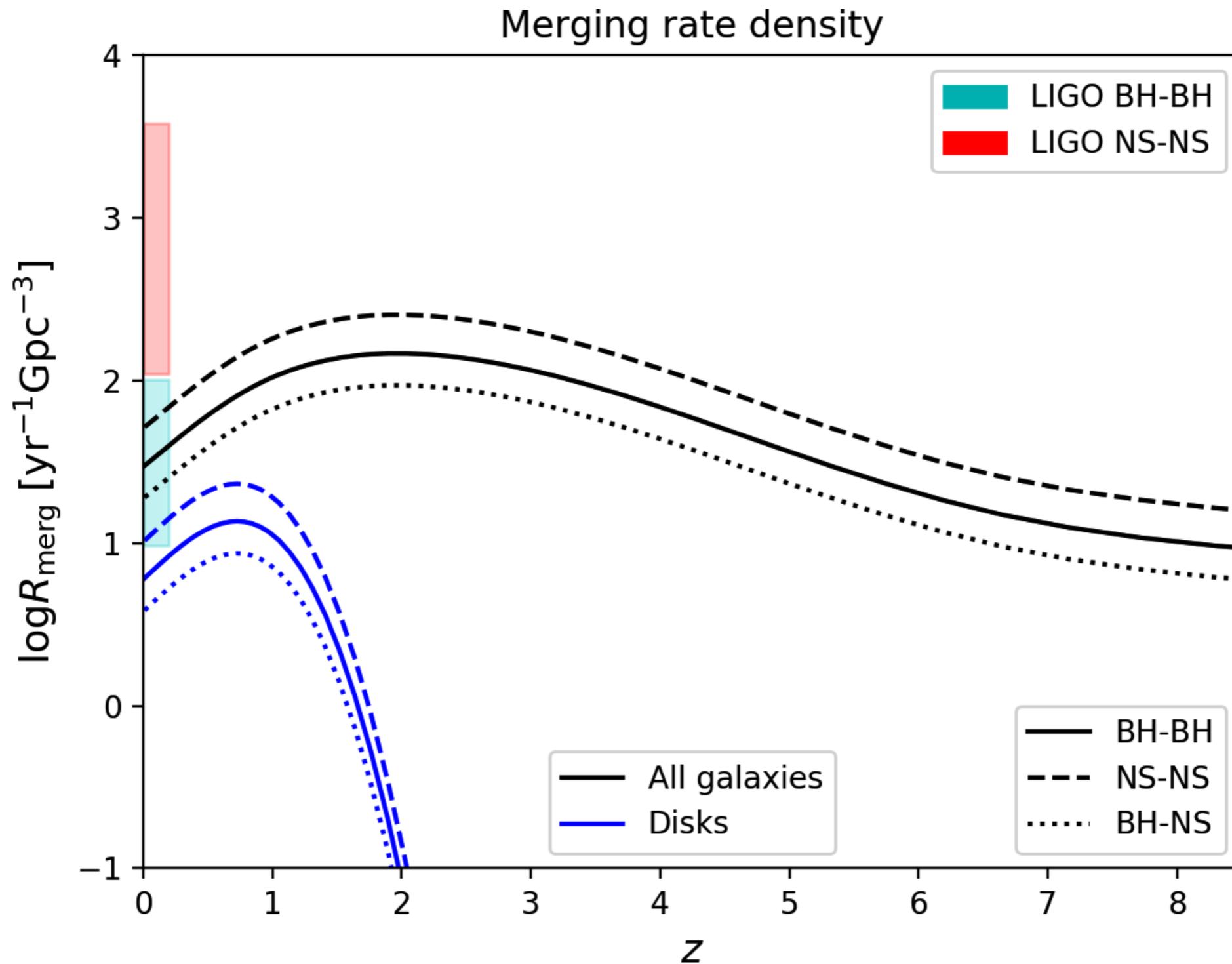
$$R_{\text{birth}}(m_{\bullet}, t) = \int d \log \psi \psi \frac{dN(\psi, t)}{d \log \psi dV} \int d \log Z \frac{dp}{d \log Z}(Z | \psi, t) \int dm_{\star} \phi(m_{\star}) \frac{dp}{dm_{\bullet}}(m_{\bullet} | m_{\star}, Z)$$

# Birth rate



# Merging rate

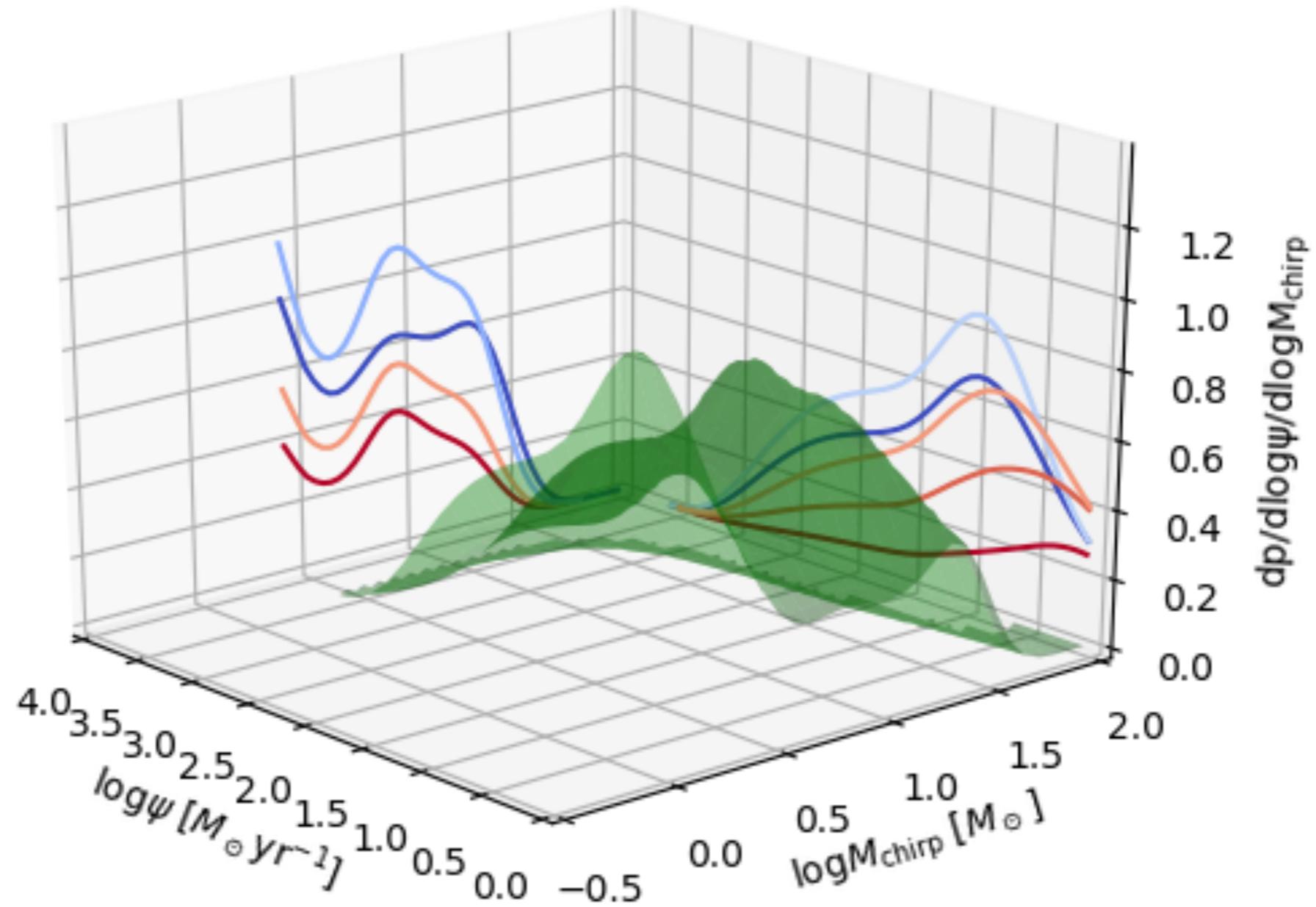
$$R_{\text{merge}}(t) = f_{\text{eff}} \int dm_{\bullet} \int dt_d \frac{dp}{dt_d} R_{\text{birth}}(m_{\bullet}, t - t_d)$$



# SFR-Chirp mass joint PDF

$$M_{\text{chirp}} \equiv \frac{(m_1 m_2)^{3/5}}{(m_1 + m_2)^{1/5}}$$

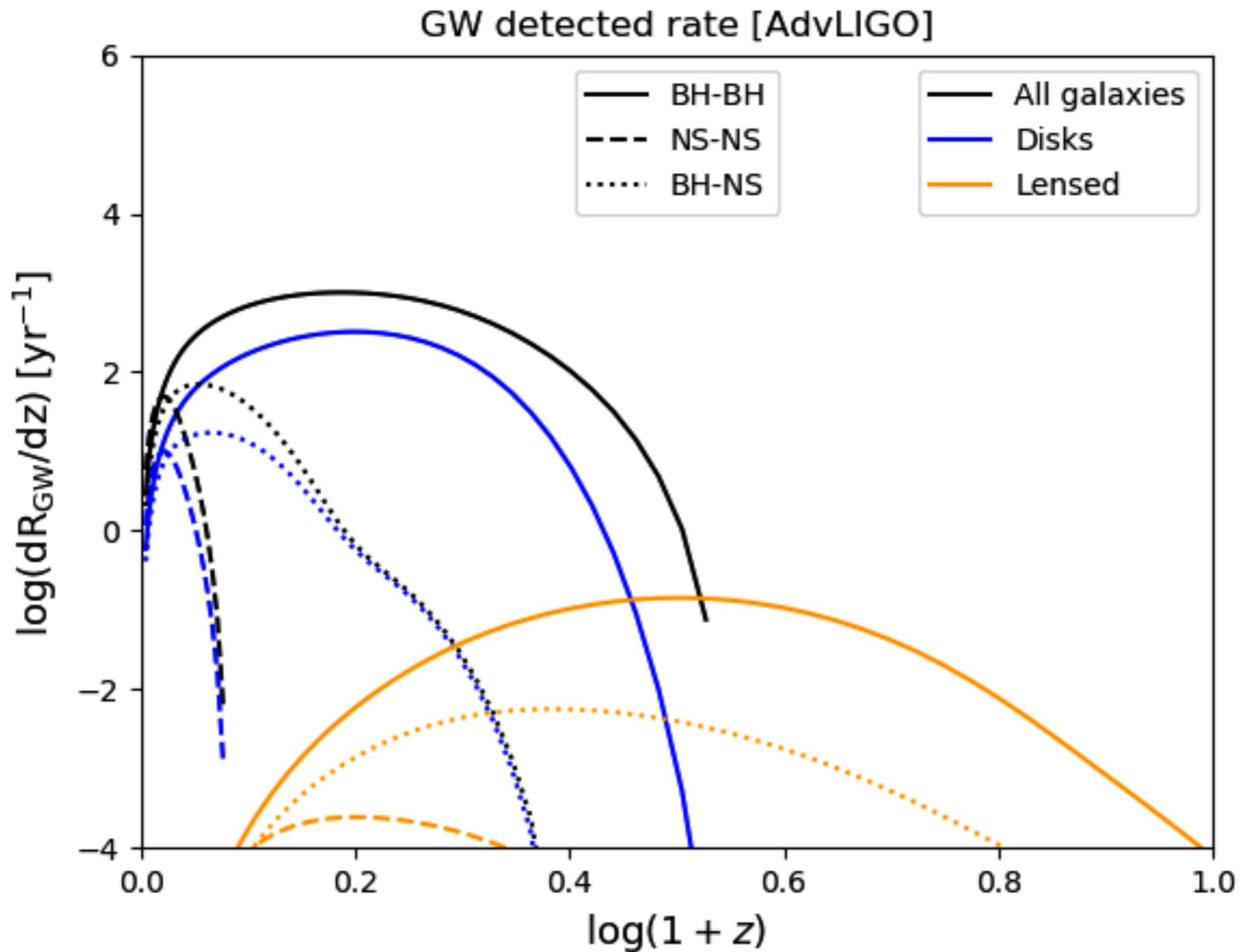
SFR-chirp mass joint PDF(BH-BH,z=0)



**Thank you for  
your attention!**

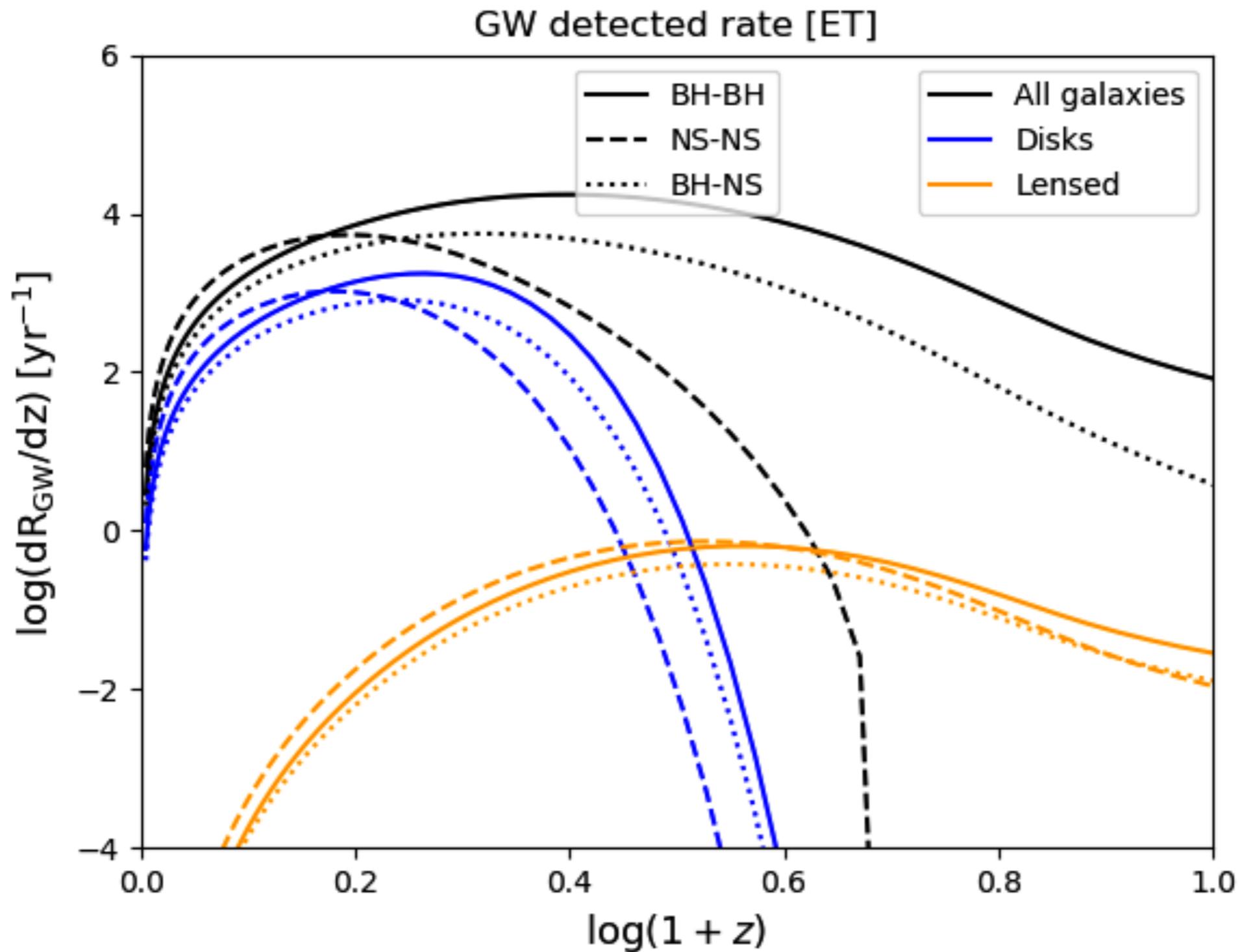
# Detection rates

$$\frac{d\dot{N}}{dz}(> \rho_0, z) = \frac{dV}{dz(1+z)} \int d\mathcal{M} R_{\text{merge}}(\mathcal{M}, z) \int_{\rho_0}^{\infty} d\rho P_{\rho}(\rho | \mathcal{M}, z)$$



# Detection rates

$$\frac{d\dot{N}}{dz}(> \rho_0, z) = \frac{dV}{dz(1+z)} \int d\mathcal{M} R_{\text{merge}}(\mathcal{M}, z) \int_{\rho_0}^{\infty} d\rho P_{\rho}(\rho | \mathcal{M}, z)$$



# Galaxy evolution

Star formation rate

Chemical evolution

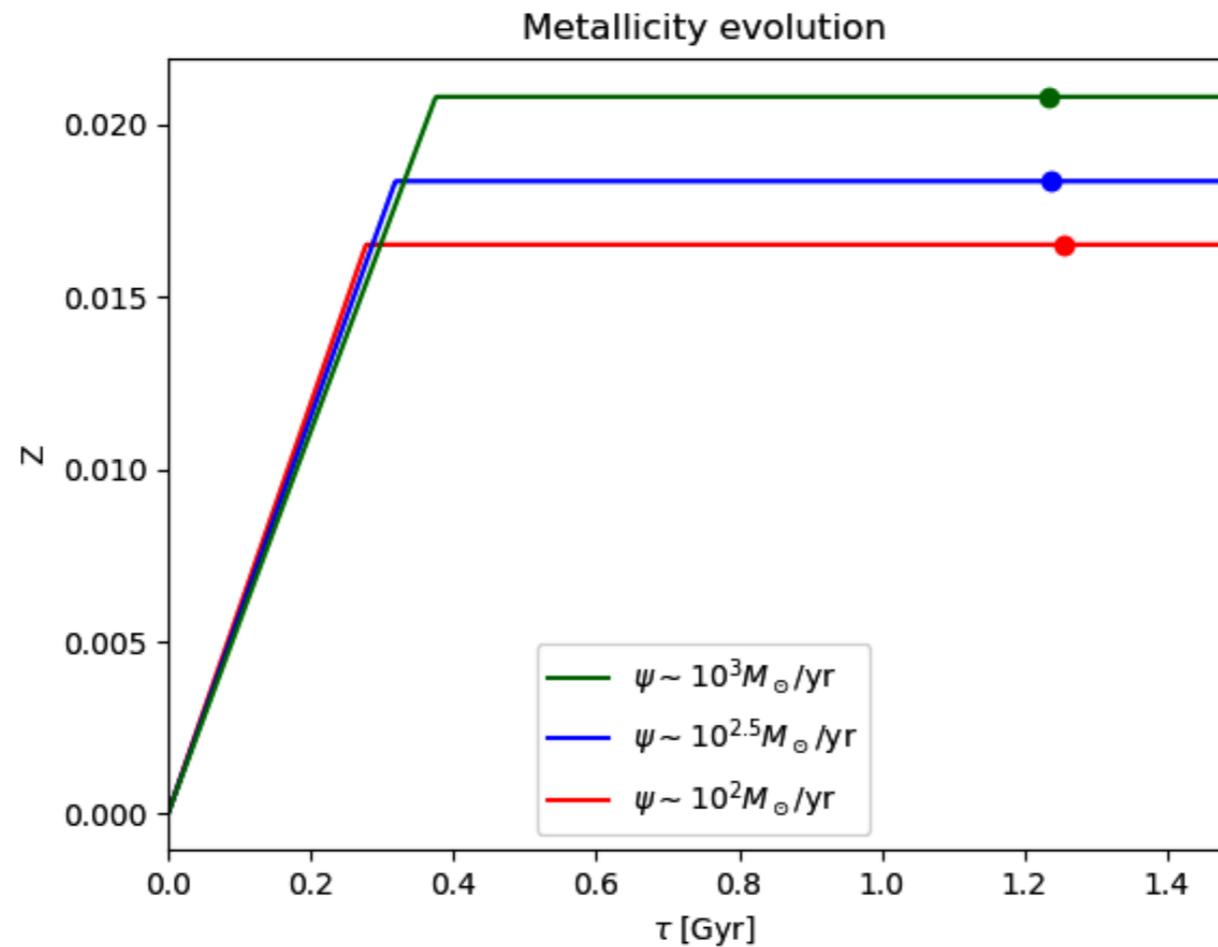
Spheroids

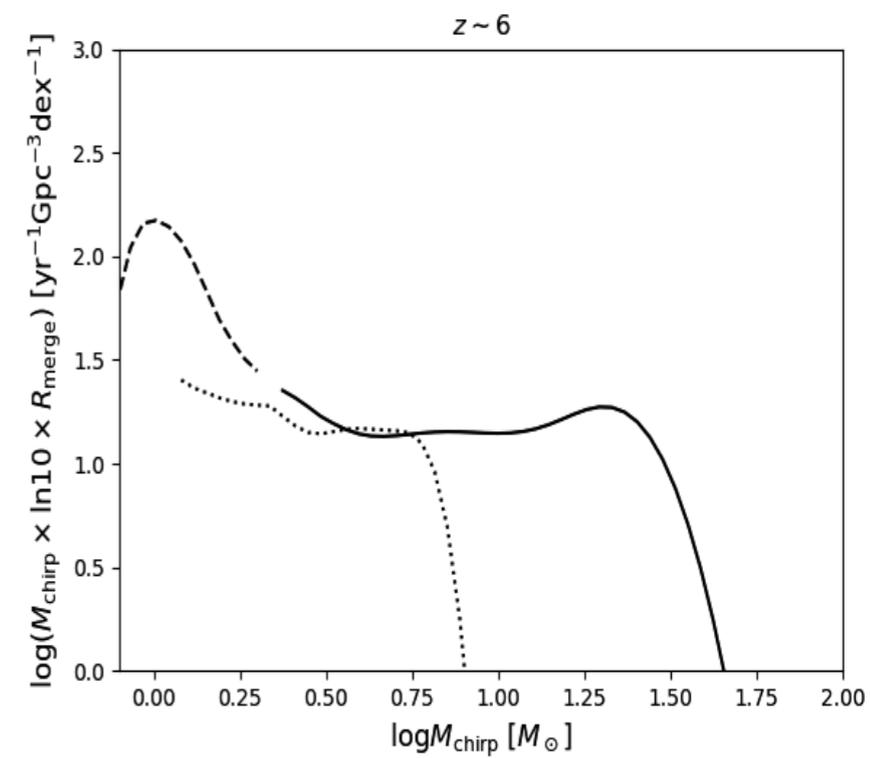
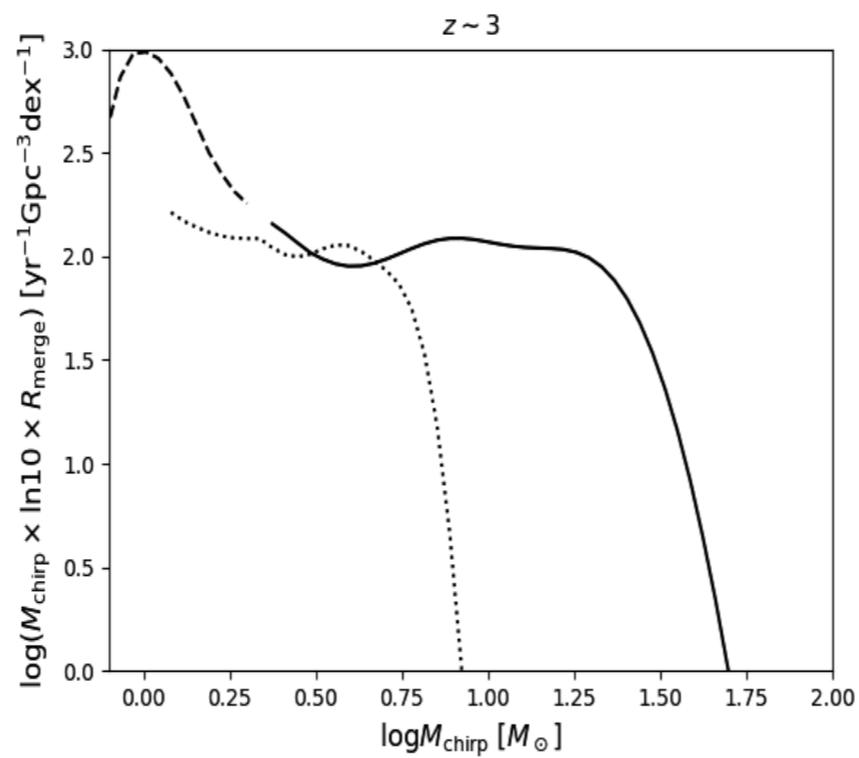
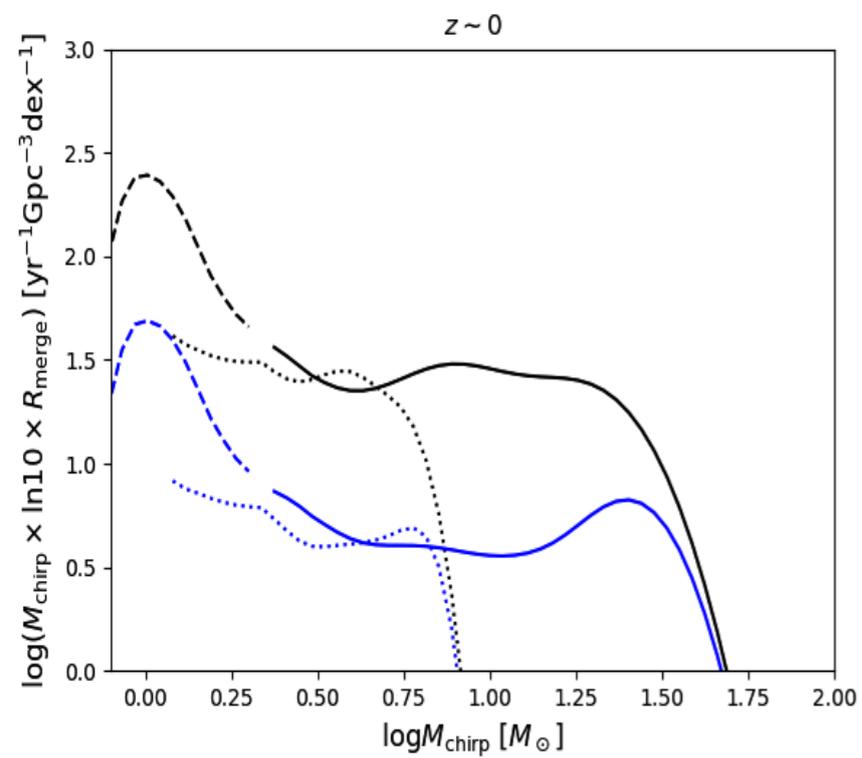
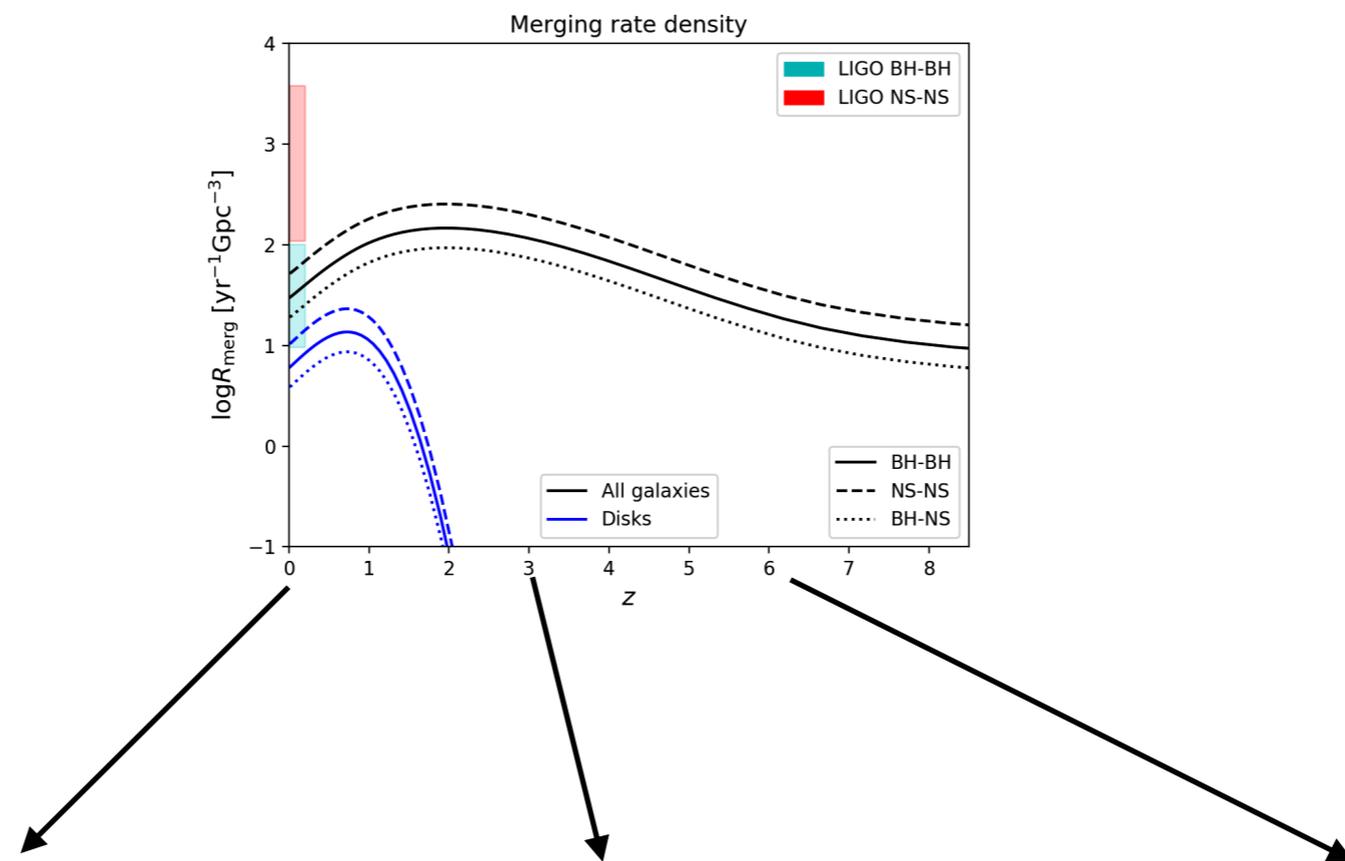
Disks

$$\psi(\tau) \propto \tau^k \Theta(\tau - \tau_\psi)$$

$$\psi(\tau) \propto e^{-\tau/\tau_\psi}$$

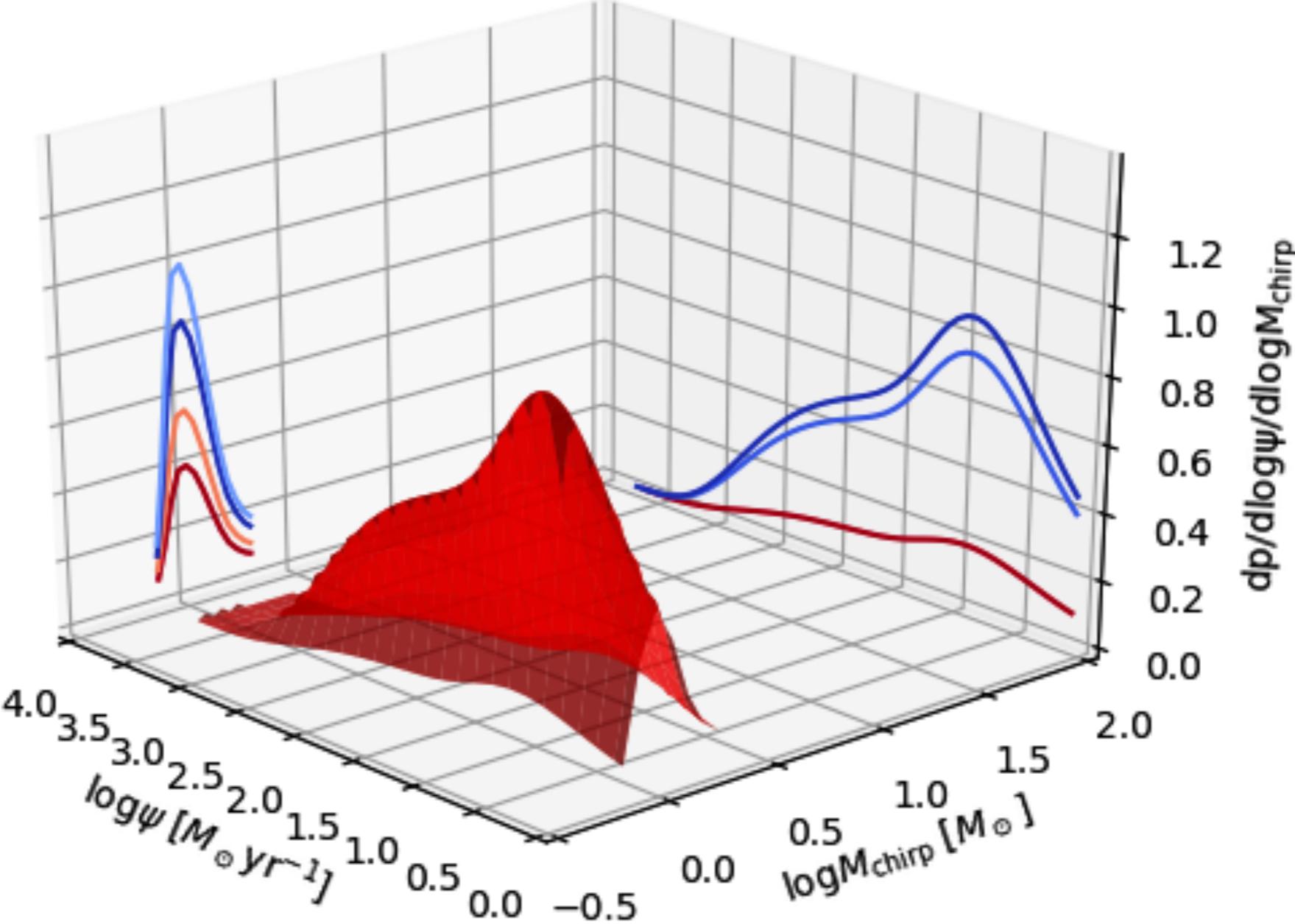
$$Z(\tau) = \begin{cases} Z_{\text{sat}} \frac{\tau}{\Delta\tau_\psi} & \frac{\tau}{\tau_\psi} \leq \Delta \\ Z_{\text{sat}} & \frac{\tau}{\tau_\psi} > \Delta \end{cases}$$





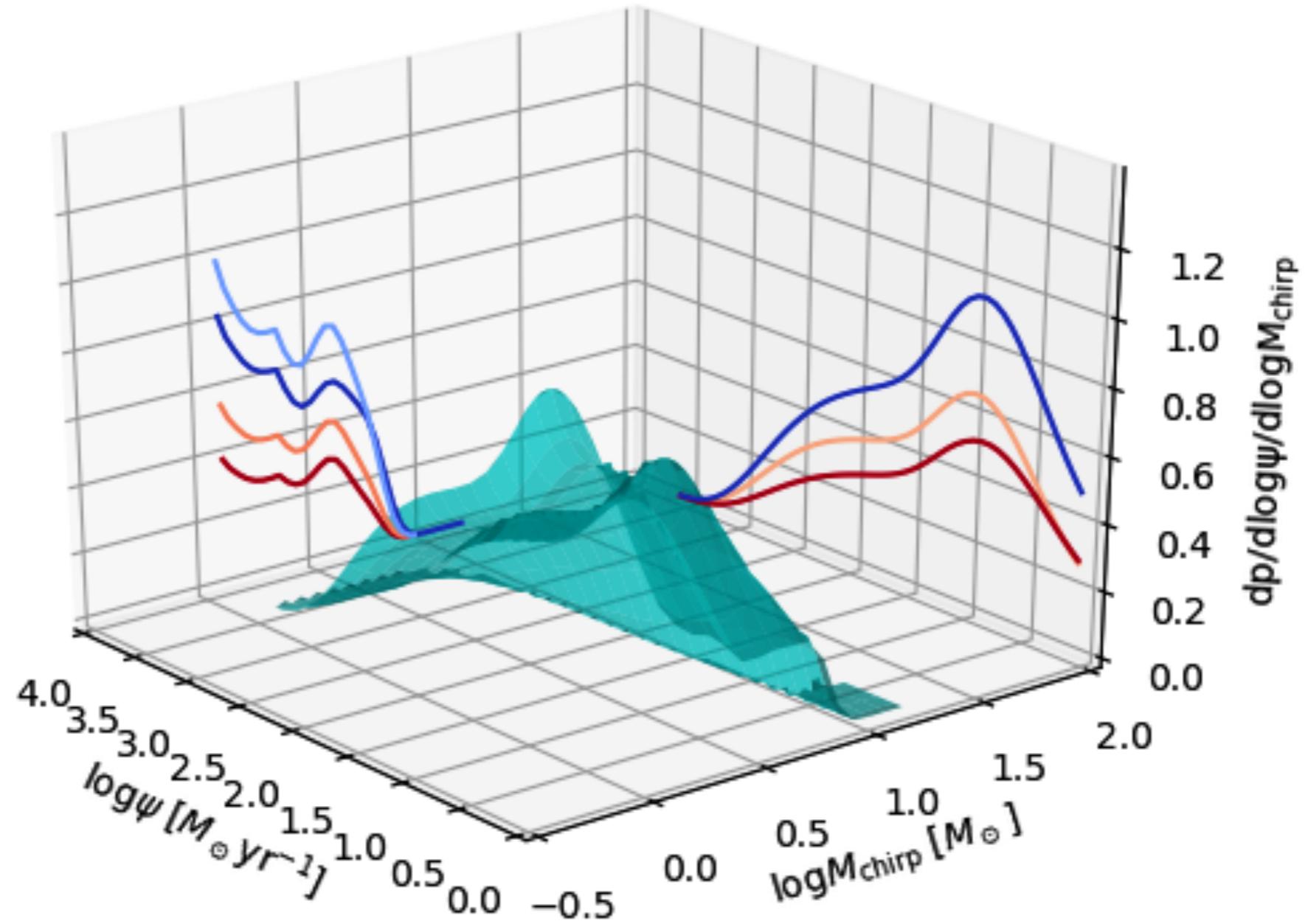
# SFR-Chirp mass joint PDF

SFR-chirp mass joint PDF(NS-NS,z=0)

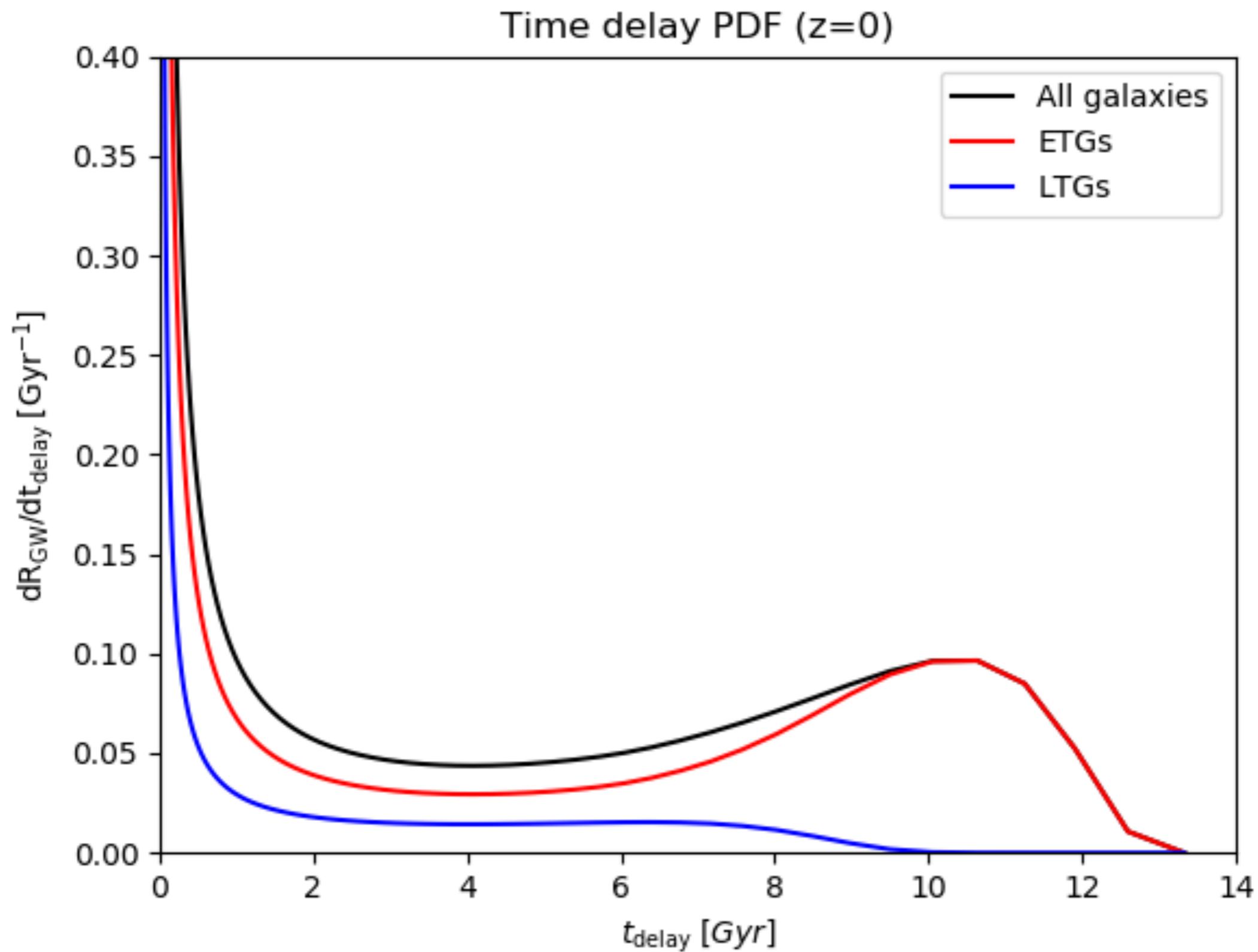


# SFR-Chirp mass joint PDF

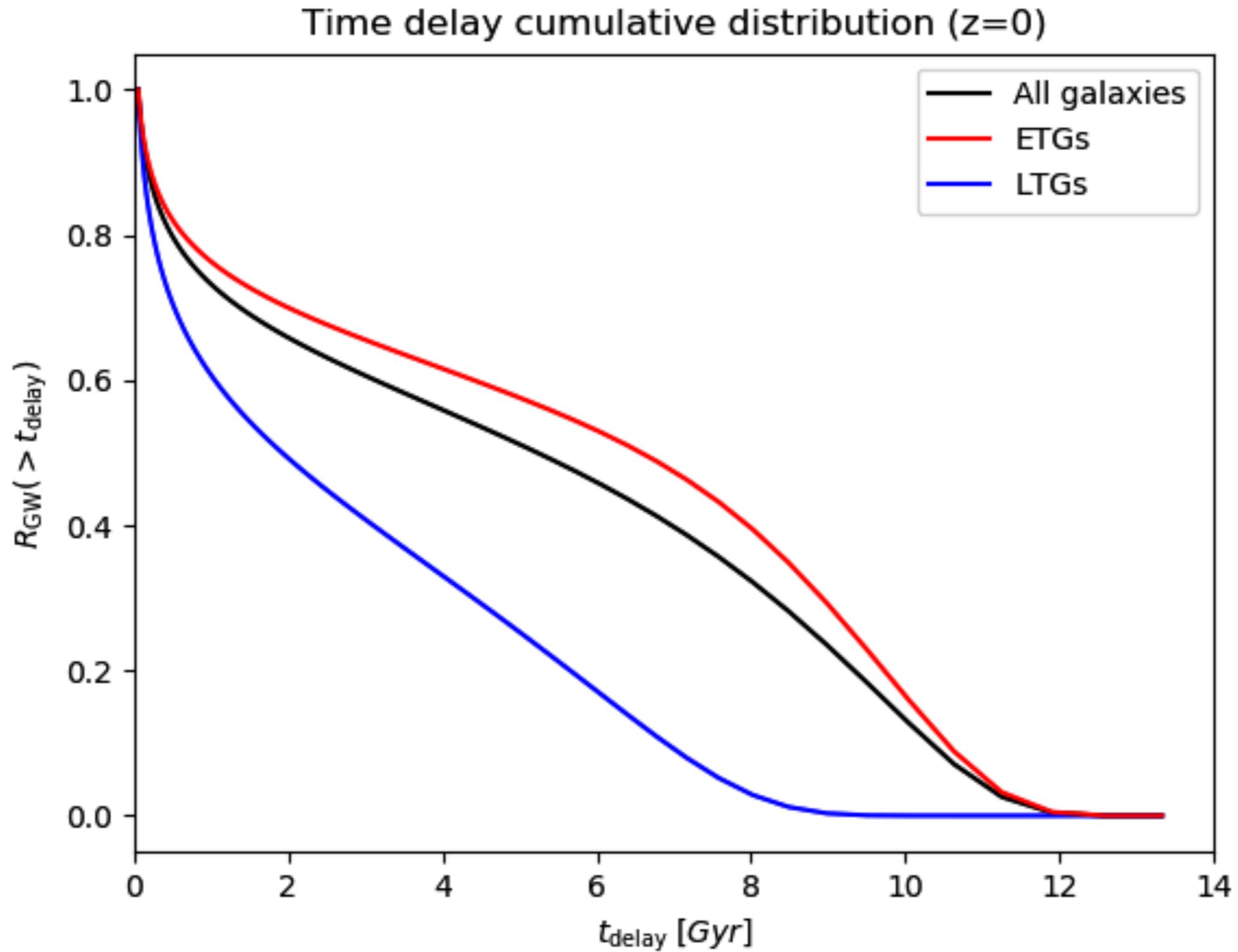
SFR-chirp mass joint PDF(BH-NS,z=0)

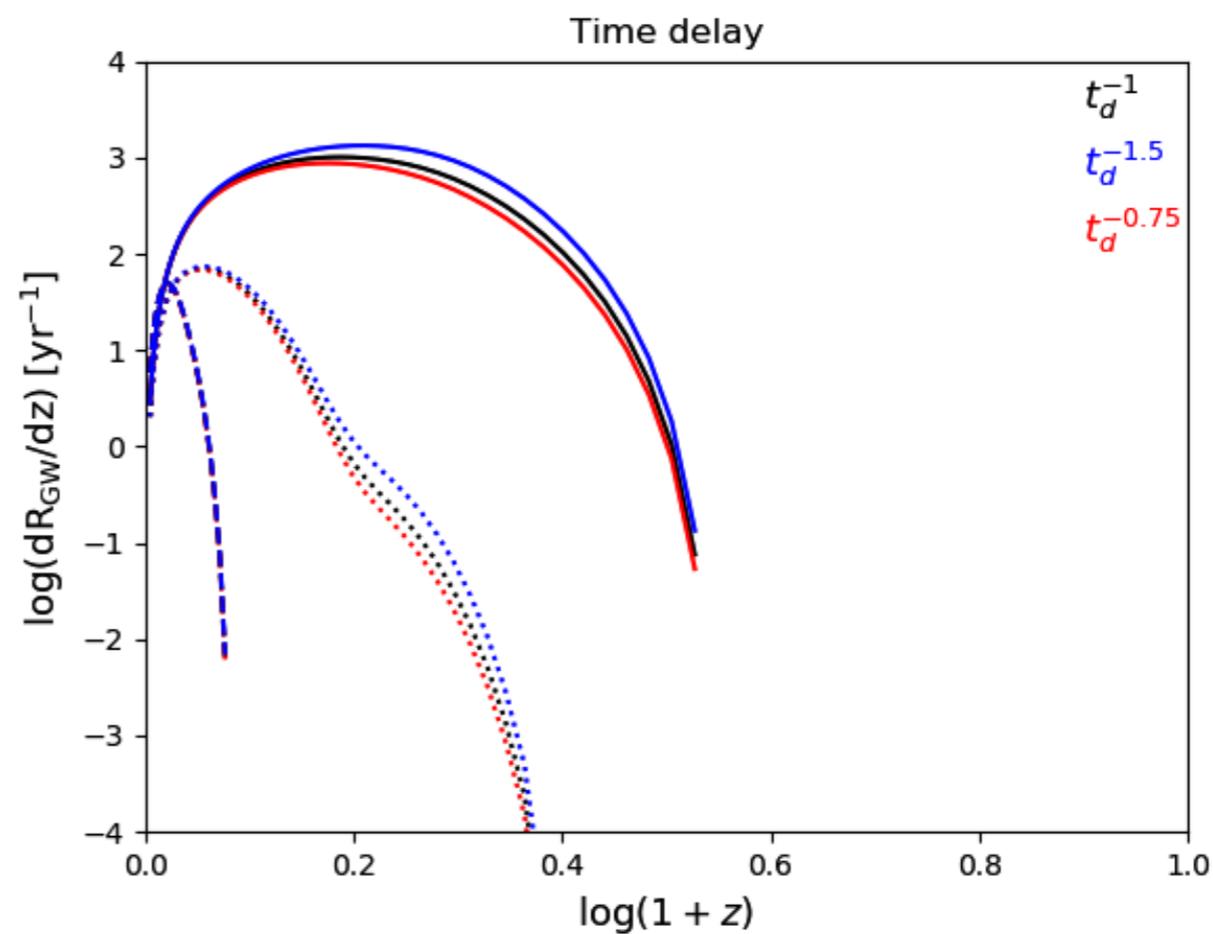
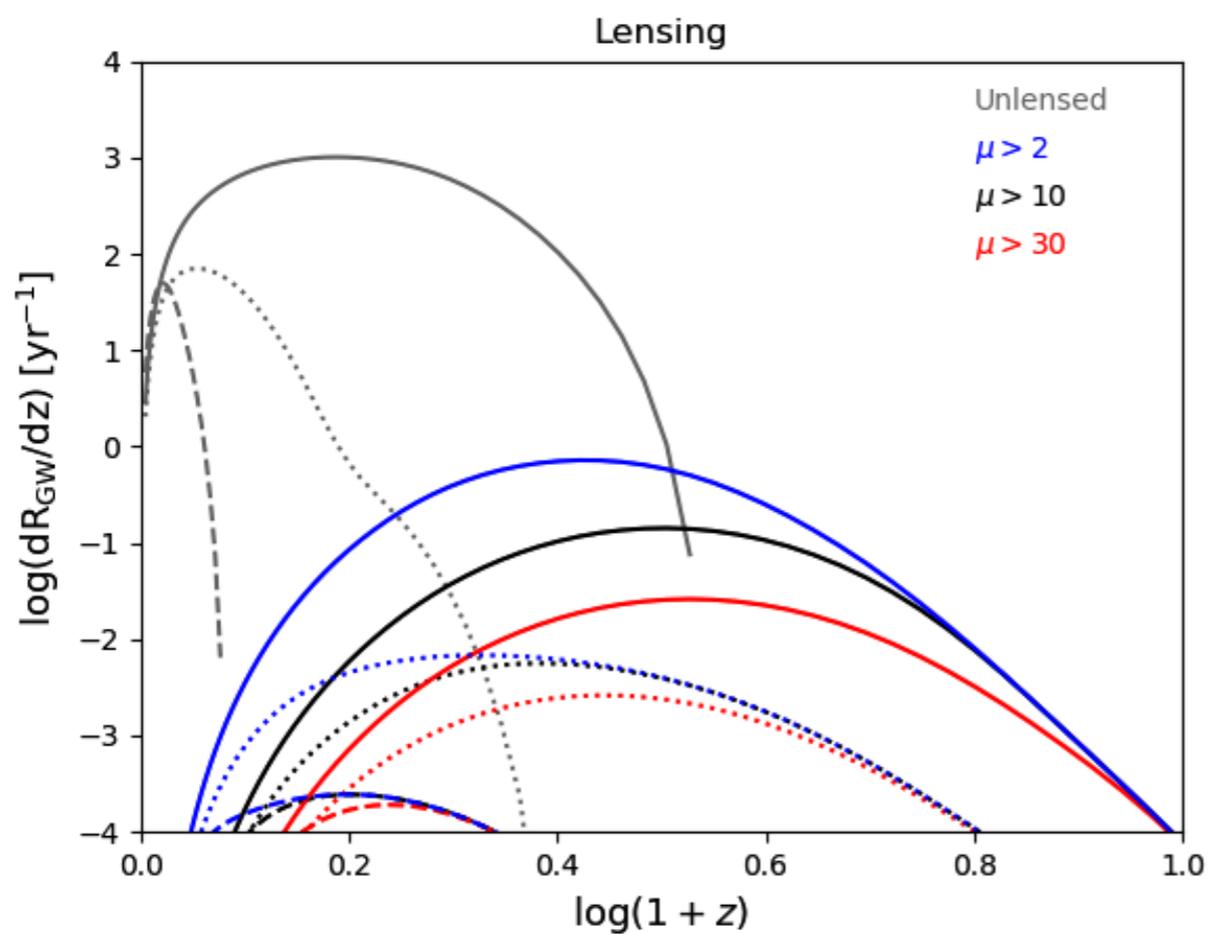
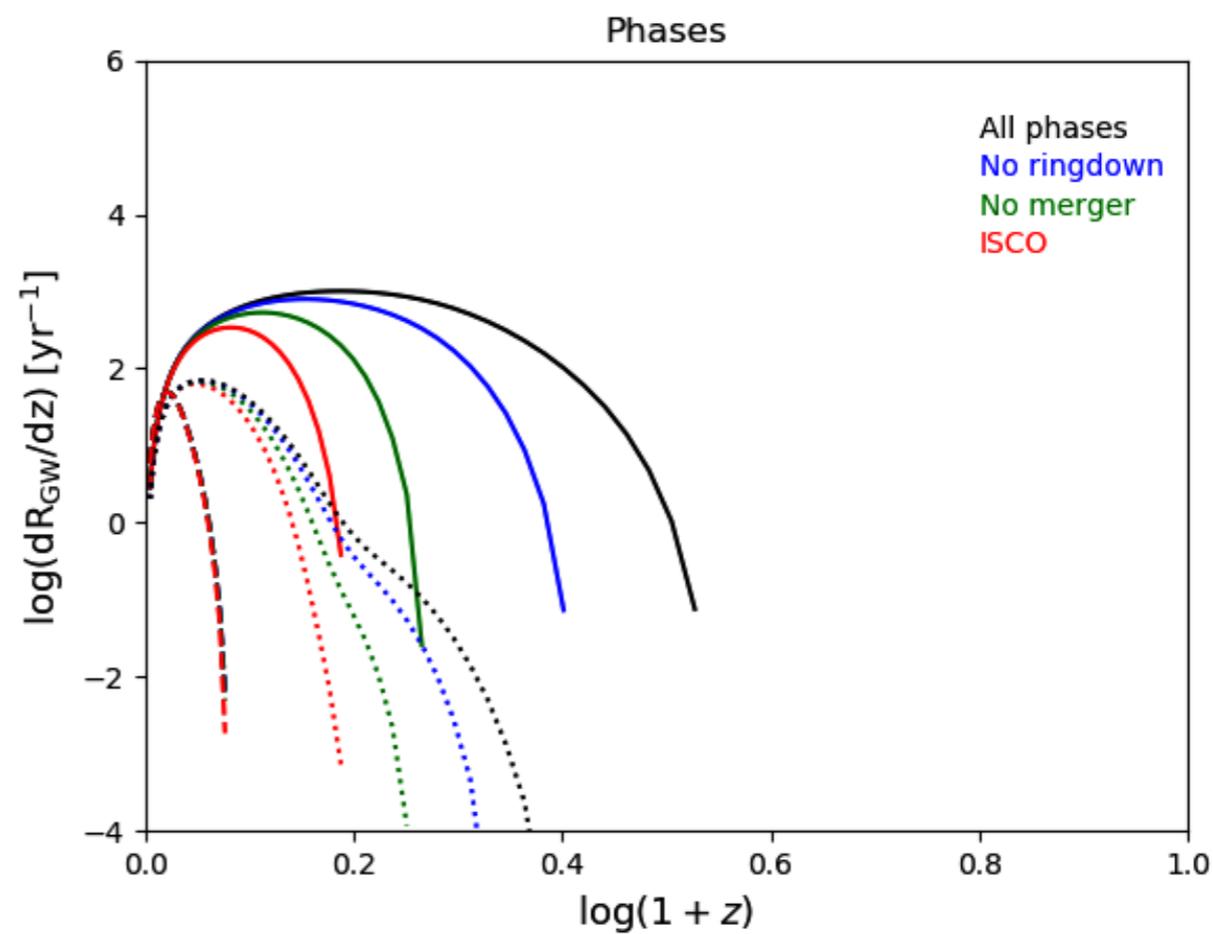
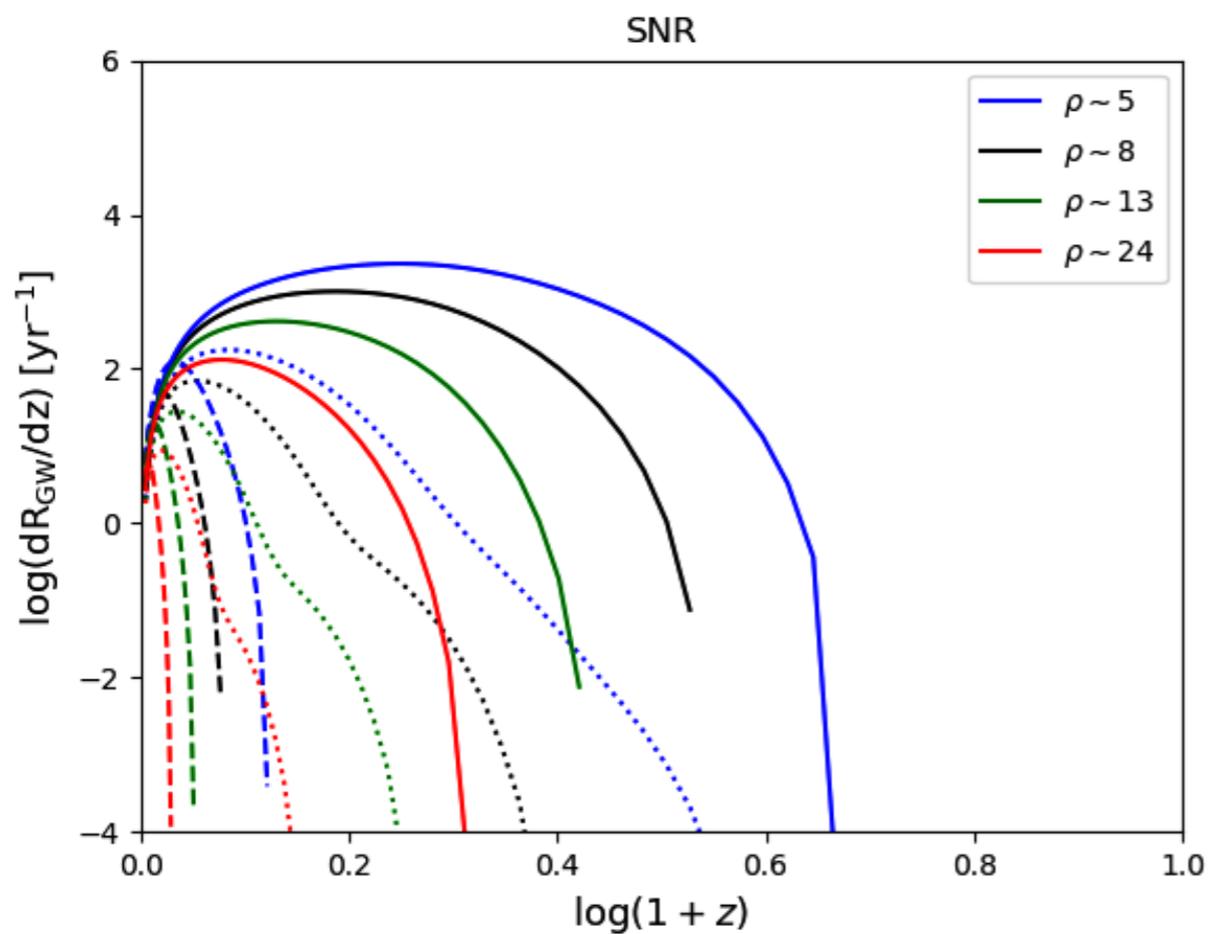


# Time delay distribution

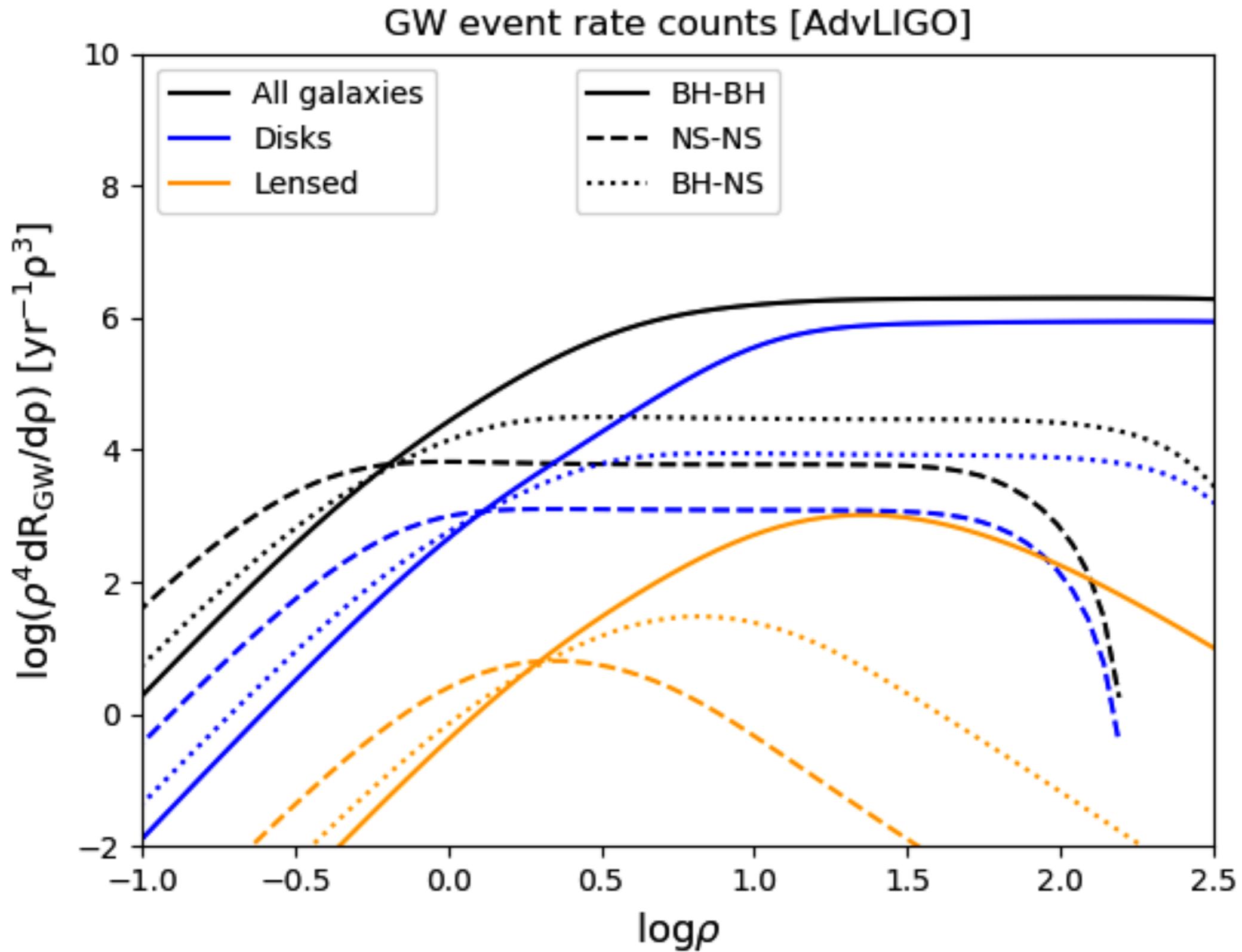


# Time delay distribution

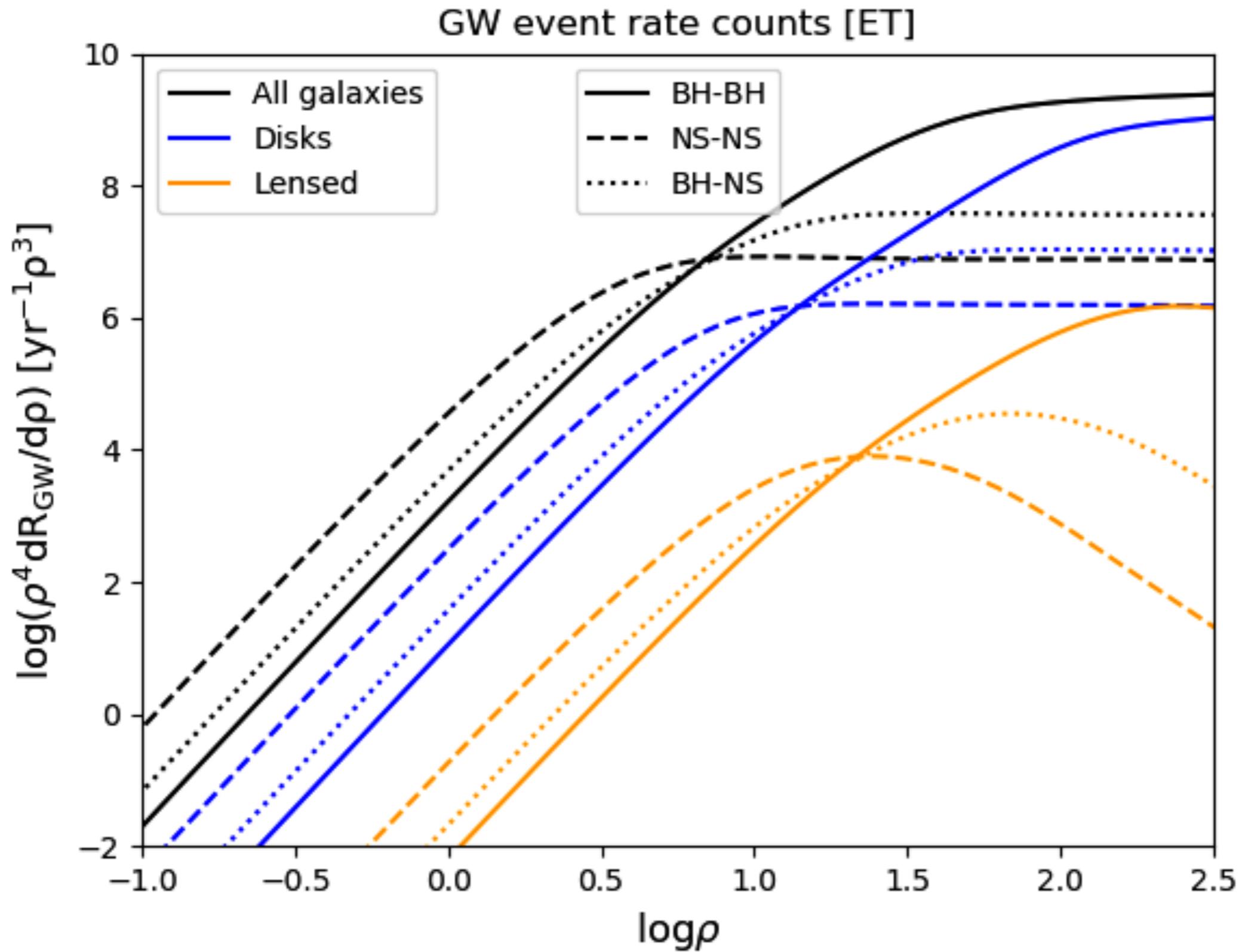




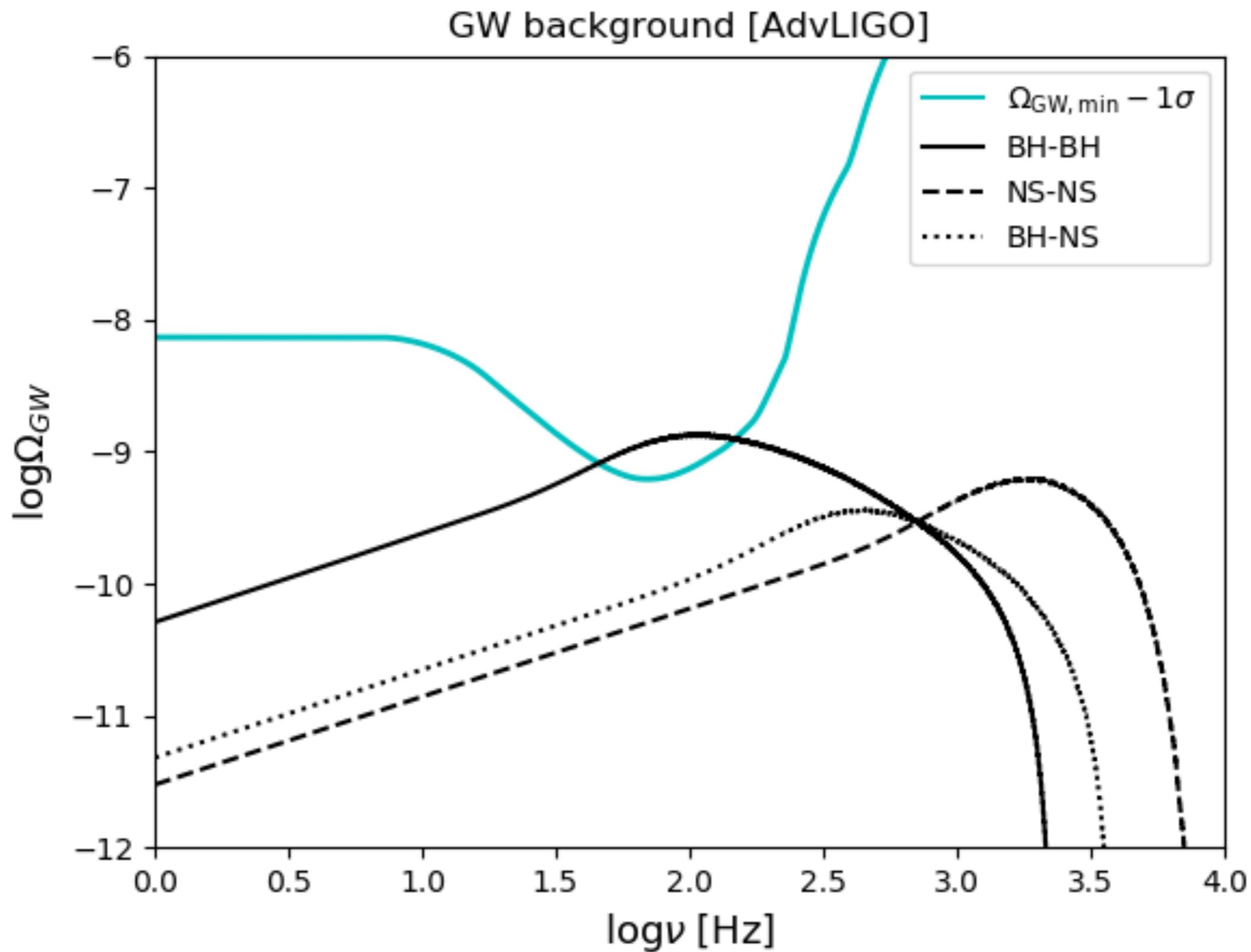
# GW number counts



# GW number counts



# GW background



# GW background

