LOFAR.IT



THE COMPUTING INFRASTRUCTURE

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INAF SCIENCE ARCHIVES & THE BIG DATA CHALLENGE 17-19 June 2019 Rome INAF-HEADQUARTERS

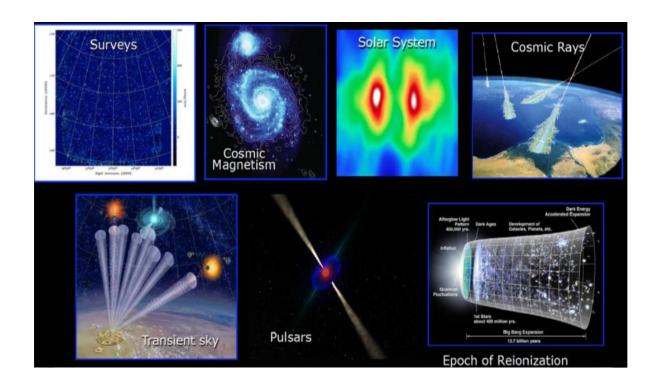


The International LOFAR Telescope

ILT consists of an **interferometric array of dipole antenna** stations, an it is distributed throughout **9 EU Countries**: NL, Germany, France, Italy, Poland, UK, Sweden, Ireland, Latvia.

Mainly based on the activities of 7 LOFAR Key Science Projects (KP)

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LOFAR - Radio Telescope

- **53 Stations** (**24 core (Exloo)**, 16 remote, 13 International)
- A LOFAR core station consists of 96 Low Band Antennas (LBAs), operating from 10 to 90MHz and 48 High Band Antenna (HBA) tiles that cover the frequency range from 110 to 250 MHz
- Remote stations in the Netherlands have the same number of HBA tiles, and LBAs
- International stations provide a single cluster of 96 HBA tiles and 96 LBAs (6 station in Germany, 3 in Poland, 1 in France, Ireland and UK, IT)





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LOFAR.IT: better late than...

Why were we interested in joining ILT?

LOFAR is the only instrument today to **produce imaging with a resolution of a few arcsec** (.... potentially < arcsec)

New scientific insight in low frequency radio Astronomy.

Potentially a large community is interested in LOFAR data.

Key Projects are the scientific groups that drive the evolution and technical knowhow in LOFAR data acquisition and analysis

SKA precursors/pathfinders drive frontier research

Offer to Italian researcher the possibility to improve knowledge on LOFAR data acquisition and analysis, to build a community of researchers.... ready for SKA.

LOFAR is the biggest SKA pathfinder (SKA-Low)



- Build a LOFAR 2.0 station in Medicina (2021-2022)
- Build a LOFAR data analysis infrastructure
- Implement a technical and scientific collaboration with ASTRON
- Develop a community that is able to work with LOFAR data (for science and technology)
- Participation of Italian community to **Key Projects** (surveys in particular)

Participation to the KPs, LOFAR guarantee time

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LOFAR.IT Board and WGs

Consortium for the participation to International LOFAR Telescope

Board: Gianfranco Brunetti(Coordinator INAF-IRA) Ugo Becciani(INAF-OA Catania) Segretario, Federica Govoni(INAF-OA Cagliari, UTG II), Francesco Massaro(UniTo), Jader Monari(INAF-IRA), Roberto Scaramella(INAF-OA Roma)

- Science Advisory Committee: Andrea Ferrara (Chair), Matteo Murgia, Mauro Messerotti, Grazia Umana, Gianni Bernardi, Ettore Carretti, Isabella Prandoni, Laura Pentericci, Marta Burgay, Rossella Cassano, Andrea Chiavassa(UniTO)
- Technological joint WG ASTRON-INAF: Established on March 2018 . Leaded by Astron. Primary objective: joint development of RCU for LOFAR 2.0 and eventually LBA2.0.
- Data WG: Giuliano Taffoni(INAF-OA Trieste) Chair, Alessandro Costa (INAF-OA Catania), Francesco Bedosti (INAF-IRA), Cristina Knapic (INAF-OA Trieste), Manuela Magliocchetti (INAF-IAPS Roma), Annalisa Bonafede (UniBo, Associata INAF IRA)



Jun 2017 -- Feb 2018: Negotiation with ILT board

Oct 2017 -- Working group to define a RoadMap to join ILT

March 2018 -- The roadmap is presented and approved by INAF

Apr 2018 -- ILT Board approves LOFAR-IT as new member

Aug 2018 -- Call for proposal (observing time) & Call for KP

Oct 2018 – LOFAR.IT computing infrastructure v1

Nov 2018 -- Call for proposal (observing time) & Call for KP + ideas

Jun 2019 -- First Italian LOFAR school



KP involvement (rapidly evolving with time):

Survey KP \rightarrow 17 full members + 1 Executive Body member (<u>3rd contributing country</u>)

Magnetism KP \rightarrow 4 members + 1 core member

Transient KP → 3 members

Solar KP \rightarrow 2 members

Proposals submission: about 90 h requested/cycle

(last 3 cycles)

Science:

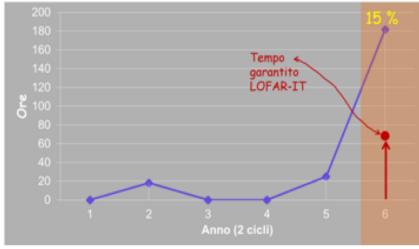
2018: 17 papers (2 IT PI) 2019: more than 25 (4-6 IT PI)

Technology:

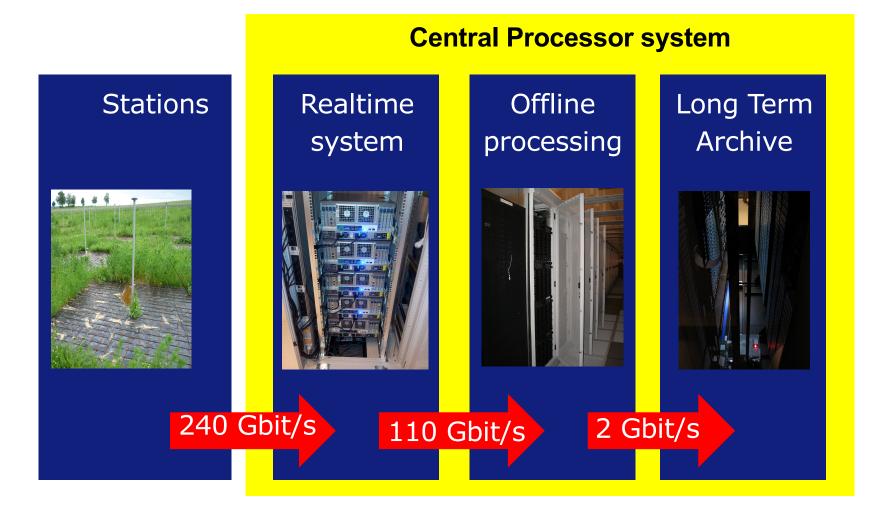
Build a distributed data reduction and analysis infrastructure and user support.

Involvement in LOFAR 2.0 RCU

pipeline and software for calibration and imaging optimization/development



LOFAR - Computing Model - Data flow



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LOFAR COMPUTING MODEL. Central Processor System:The Correlator

- 9 Dell T620 nodes
 - Dual Intel Xeon E5-26xx
 - 2 Nvidia Tesla K10
 - 2 Dual port 10GbE
 - 2 FDR Infiniband HCA





COBALT (COrrelator and Beamformer Application for the LOFAR Telescope)

Both the F-stage (Fourier transfom) and the X-stage (cross-correlation) are implemented in GPUs

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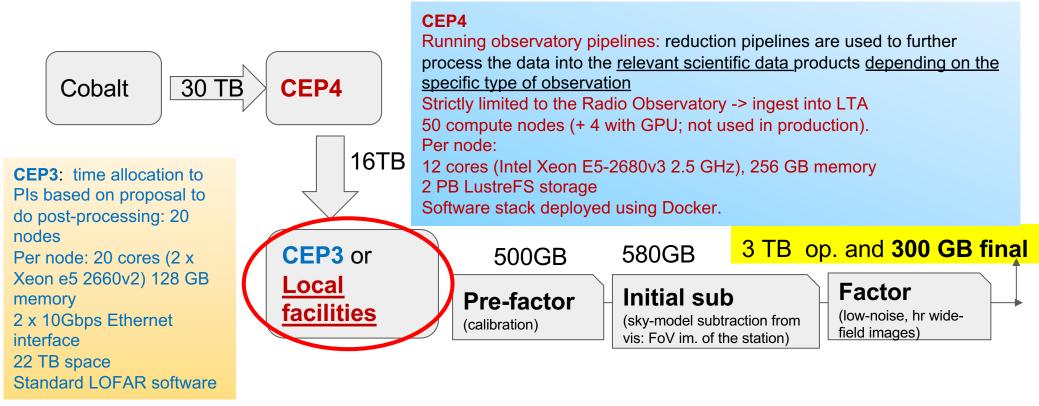
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LOFAR COMPUTING MODEL. Central Processor System:Post-Processing

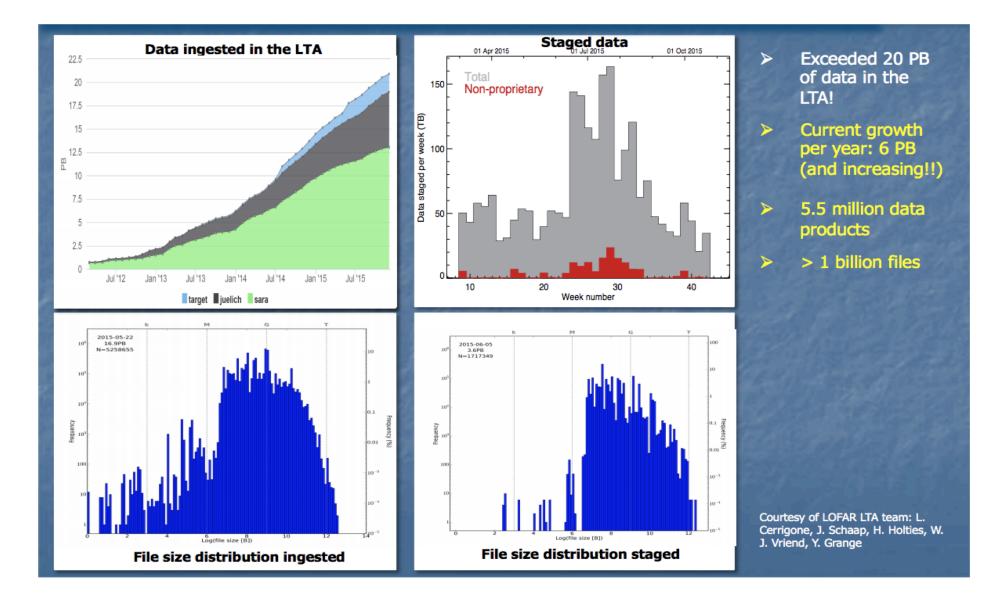
Two central processing (CEP) clusters in Groningen (i.e. near the correlator). Pipelines use locallydeveloped generic framework.

Distributed system built using a <u>co-design approach</u> (we know the algorithms and we design the HW)





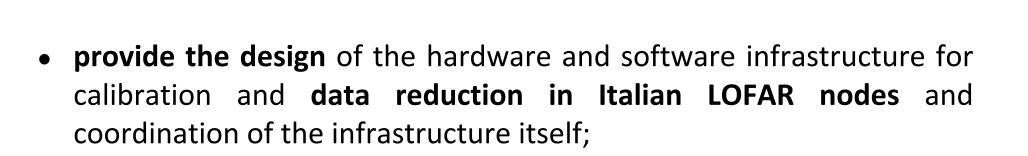
Technological Challenges: LTA



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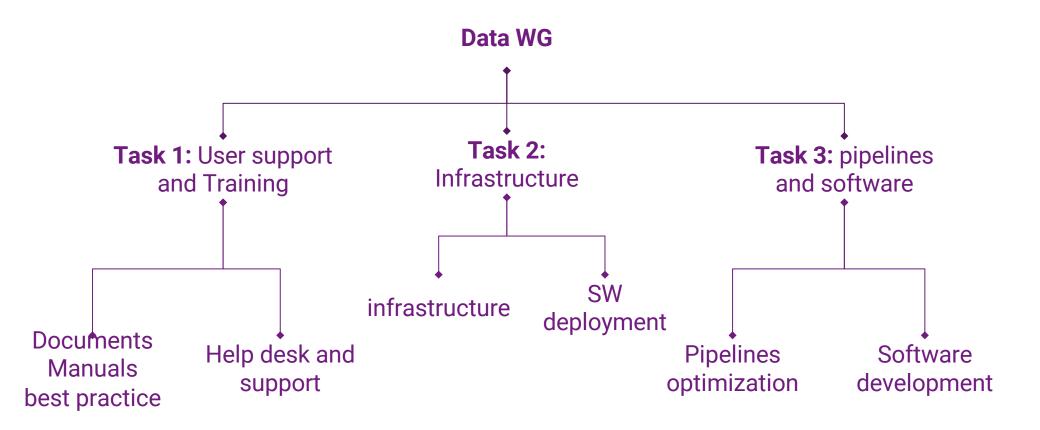


LOFAR.IT: Data Working Group

- coordinate the **installation**, **configuration** and **management** of specific software and pipelines for the reduction of LOFAR data;
- provide technical support to users belonging to LOFAR IT through testing, verification, optimization and development of pipelines for LOFAR data reduction;
- collaborate with LOFAR developers for further code testing and optimization/parallelization of codes and data reduction pipelines (e.g DDFacet pipeline);

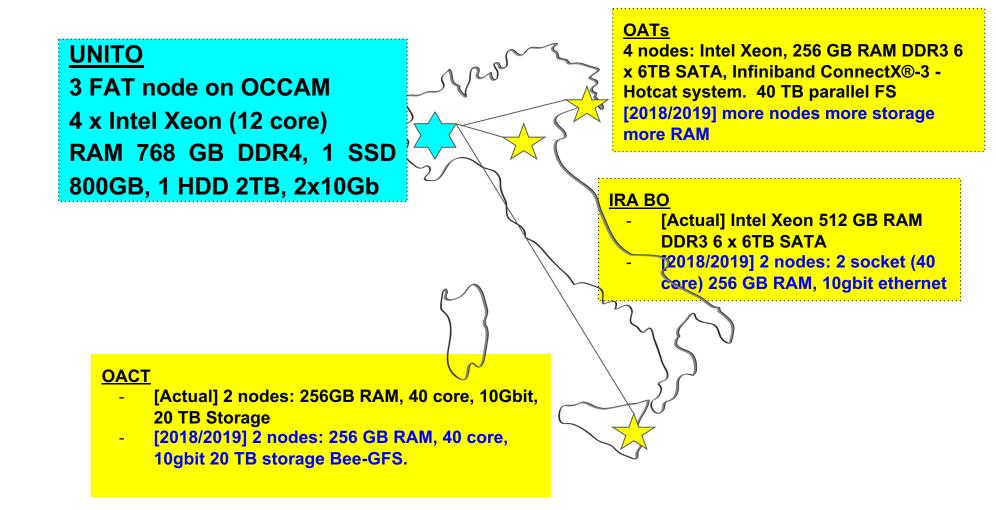
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LOFAR.IT e-Infrastructure





How to build the infrastructure

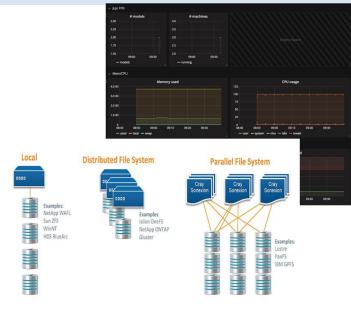
Monitoring and Operations (telgraf+grafana)

Simplified access to resources and Uniform use: containers and schedulers

Role of parallel filesystem to share data between nodes and to speed up IO

Identify or train "support" groups for data reduction and for HW/SW

Produce Manuals and HowTOs



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Software and data access

Software containerisation

- Platform independence: Build it once, run it anywhere
- Resource efficiency and density
- Effective isolation and resource sharing
- Operational simplicity and minimal overhead
- Orchestration

HPC Integration: Singularity containers

• parallel computing (MPI)

Container repository

- Versioning
- Fine grained Access control
- Usability
- Integrated with user software.







Pre-factor does the first calibration of LOFAR data

phase 1.a: flag - calibration - transfer of solutions to the target and initial calibration of the target - averaging. Computing time: 3 - 4 days, RAM (core) at least 64GB.

phase 1.b: high and low resolution images and models for auto-calibration. Computing time: 4 -5 days, RAM (core) at least 64GB.

Factor produces low-noise, high-resolution wide-field images from LOFAR HBA data. Computing time: 1 month

DDF alternative direction dependent pipeline for SK (different versions according to the user capabilities). Multithreaded and faster.

more....





LOFAR data reduction tests

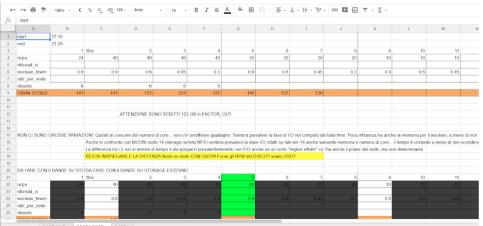
Tested **Factor** pipeline on OATS and OACT sites employing Singularity containers on single node.

Factor parameters

- 6 bands (2 MHz)
- #cpu@OACT: from 6 to 24 and from 10 to 40

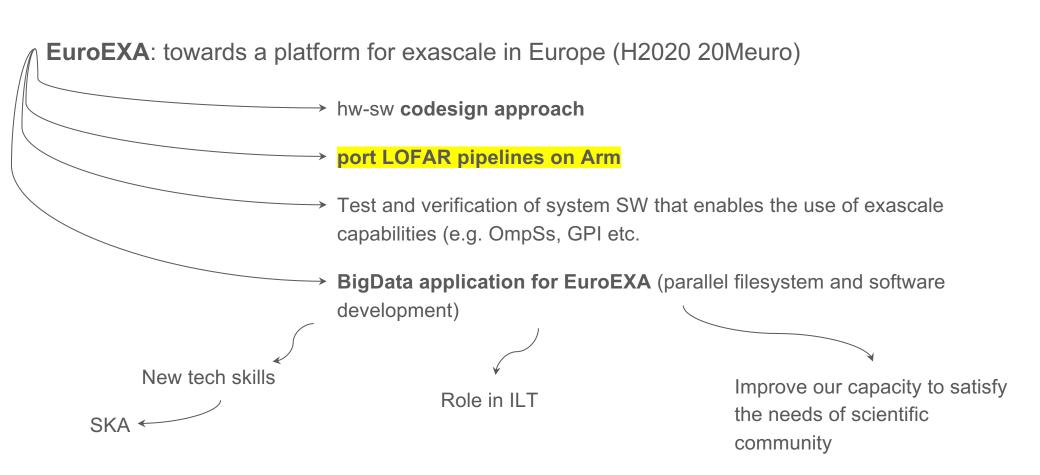
#cpu@OATs: from 3 to 40

- wsclean_mem@OACT: from 30 to 90%
- wsclean_mem@OATs: from 30 to 90%



Conclusions: the pipeline is **I/O time dependent** therefore a parallel File Systems should be employed (e.g. BeeGFS). Nodes with at least **10 cores and 256 GB RAM.**





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