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# Study of Alfvénic vortex shedding past a cylindrical obstacle

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# Introduction

 Fluid flow around a bluff obstacle may lead to periodical and alternate formation and disconnection of vortices from its each side (vortex shedding) and further formation of a chain of vortices (Kármán vortex street) – e.g. White (2006)

$$St = \frac{d \cdot f}{v}$$

- <u>Gruczecki et al. (2010)</u> vortex shedding has been widely studied in hydrodynamic conditions, but is less well understood in magnetohydrodynamic (MHD) conditions
- It has been studied in MHD conditions by a number of numerical simulations, mainly in 2D e.g. <u>Singha et al. (2006)</u>, <u>Dousset and Pothérat (2008)</u>, <u>Gruczecki et al. (2010)</u>...
  <u>Karampelas and Van Doorsselaere (2021)</u> MHD simulations in 3D



# Introduction

- E.g. Lee et al. (2015), Samanta et al. (2019) observations suggest the possibility of its occurrence in the Sun's atmosphere
- <u>Nakariakov et al. (2009)</u> it is a possible mechanism for excitation of kink mode oscillations in coronal loops:





# Numerical model

- Vortex shedding around a cylindrical obstacle was studied in MHD conditions in 3D using the numerical code Lare3d
- A parametric study was performed for different values of magnetic field (0, 5, 10 and 15 G) perpendicular to the plasma flow plane
- Gravitational field was not considered, ideal MHD equations were used





#### • Time evolution of density – $B_y = 0$





### Results

• Time evolution of density –  $B_y = 15 \text{ G}$ 





### Results

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• Density distribution at the last timestep 1000 (1120 s) – comparation





## Results

#### • Relative density change – comparation







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# Results

• Values of period obtained from the time evolution of  $\Delta \rho / \rho_0$  using wavelet analysis (<u>Torrence and Compo, 1998</u>)

B <sub>y</sub>	0	5 G	10 G	15 G
Р	12.701	10.830	9.428	8.736



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# Conclusion

- The presence of magnetic field perpendicular to the flow plane:
  - makes the vortices denser
  - makes the vortex shedding frequency increase
  - causes higher periodical density changes
- In the future we intend to use stronger magnetic field inside the obstacle and extend the problem on gravitational field to approximate it more to the solar corona



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