



16th  
European  
Solar  
Physics  
Meeting

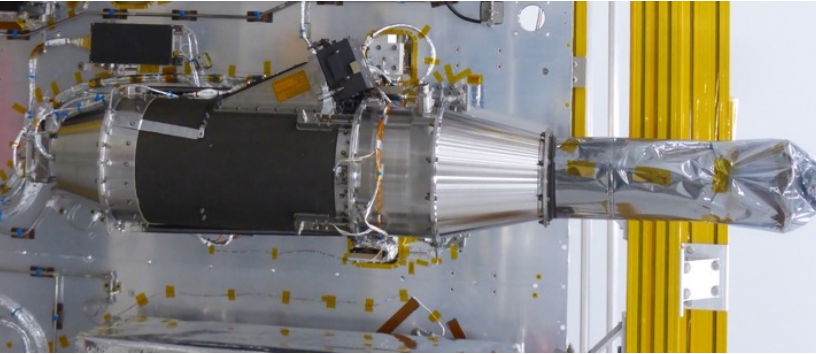
## Metis coronagraph aboard Solar Orbiter: instrument performance on the eve of the nominal mission phase

*Marco Romoli Univ. di Firenze, Italy  
and Metis Team*

The Metis coronagraph is one of the remote sensing instruments of the Solar Orbiter mission launched on February 10, 2020. Metis is an innovative externally occulted coronagraph designed to observe the solar corona in an annular field of view from  $1.6^\circ$  to  $2.9^\circ$  simultaneously for the first time in the broadband linearly polarised visible light and in the narrow-band H I 121.6 nm ( $\text{Ly}\alpha$ ). Metis has been characterised during the three-month long commissioning phase and throughout the cruise phase that will end after the Earth fly-by on November 27, 2021, official start of the nominal mission phase. In this work the Metis performance and main scientific achievements will be reviewed.



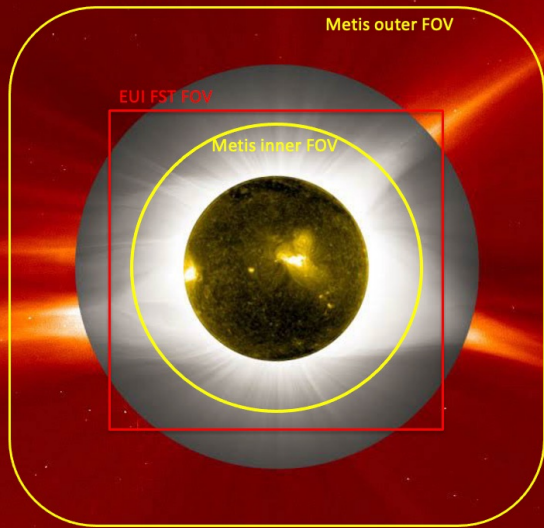
# Metis coronagraph on-board Solar Orbiter



Metis is an imaging **externally occulted** all-reflecting **coronagraph** designed to provide:

- **Full Imaging** of the extended corona ( $1.7 - \sim 9 R_{\odot}$ ) in
  - **UV Ly-alpha** ( $121.6 \pm 10$  nm), and
  - **visible light** (580-640nm) in **total and polarized brightness**
- **Density distribution** in corona of  $H^0$ , and  $e^-$
- **Global Maps of solar wind outflow** ( $H^0$ )
- **Large scale dynamics** of  $H^0$ , and  $e^-$  in CMEs

Polarized VL imaging @ 580 - 640 nm  
 UV HI Ly  $\alpha$  imaging @  $121.6 \pm 10$  nm  
 FoV ( $1.6^\circ \cdot 2.9^\circ$  annular,  $1.7 - 3.0 R_{\odot}$ ) @ 0.28 AU



Spatial resolution  $\leq 4000$  km ( $20''$ ) @ 0.28 AU  
 Time resolution  $\geq 1$  sec  
 Simultaneous VL and UV imaging





# Mission Profile & Metis Coronagraphic Observations

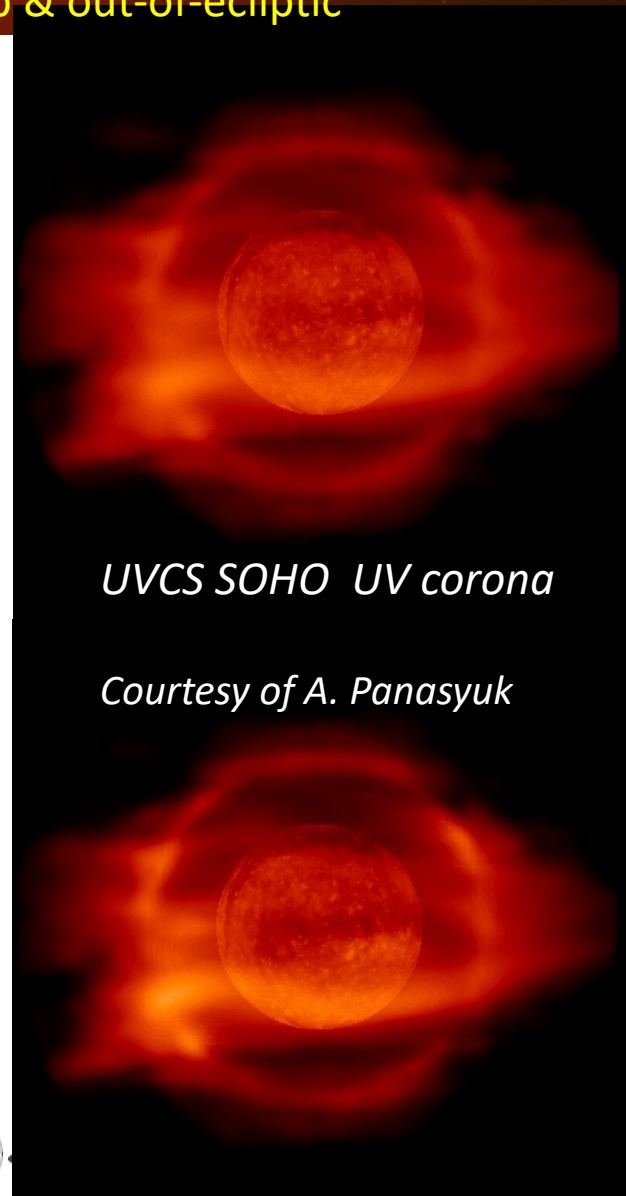
## Metis - first coronagraph pointing to the Sun close-up & out-of-ecliptic

**Close to the Sun 0.28 AU** (min perihelion) fine structure of wind plasma in corona in extended latitude & longitude ranges

**Out of the ecliptic ~ 34°**  
access to longitudinal structure of corona, solar wind and magnetic flux tubes channeling outflows

**Reduced rotation relative to the Sun 7.7°/d**  
intrinsic evolution of solar wind and of coronal density inhomogeneities due to reduced rotation effect at the limb

**Out of the geocorona**  
best UV coronal seeing conditions



*UVCS SOHO UV corona*

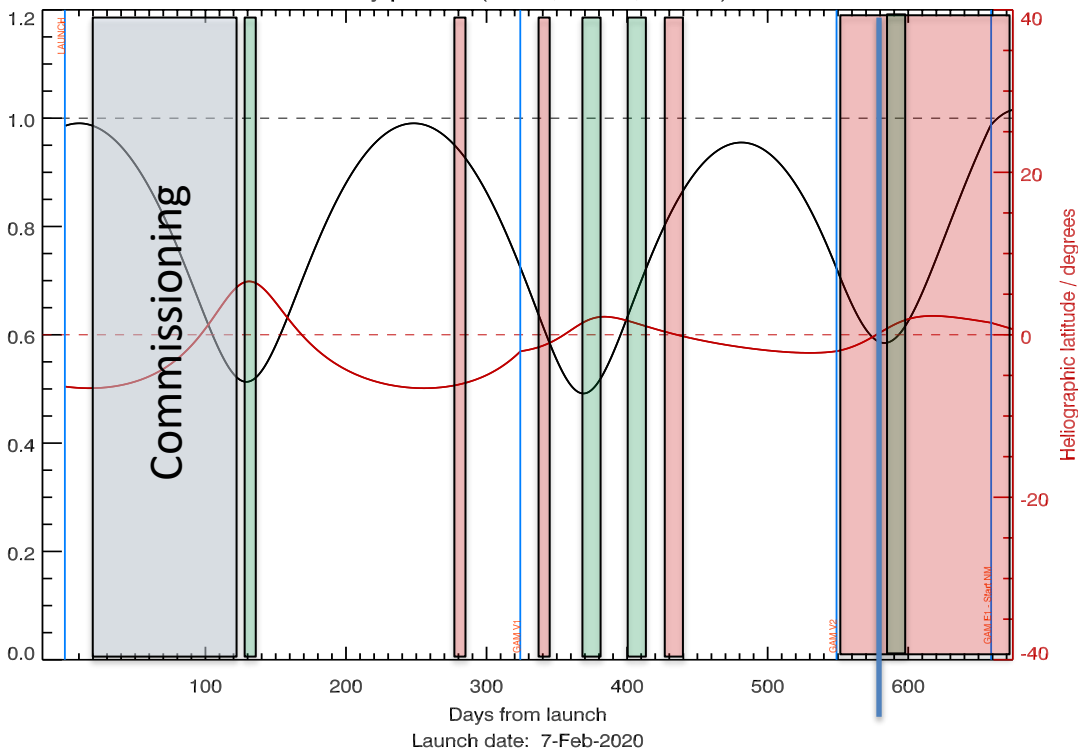
*Courtesy of A. Panasyuk*





# Cruise Phase

Early phases (LEOP, NECP, Cruise)



- Star radiometric calibrations
- Observing modes testing
- Several CMEs observed since January 2021
- Density fluctuation test
- Synoptic program test
- Stray light and polarizer tests with S/C rolls and slews
- Optimization of the internal occulter position





# Metis Performance Status



## Completed activities:

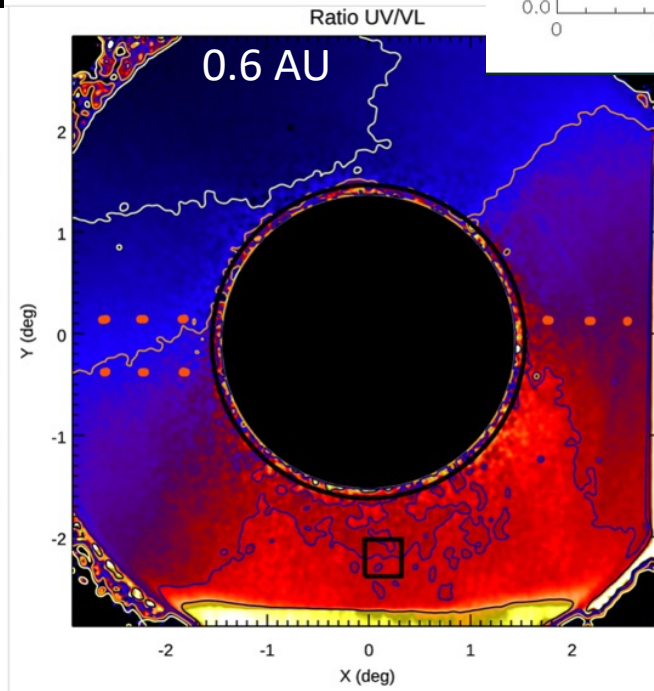
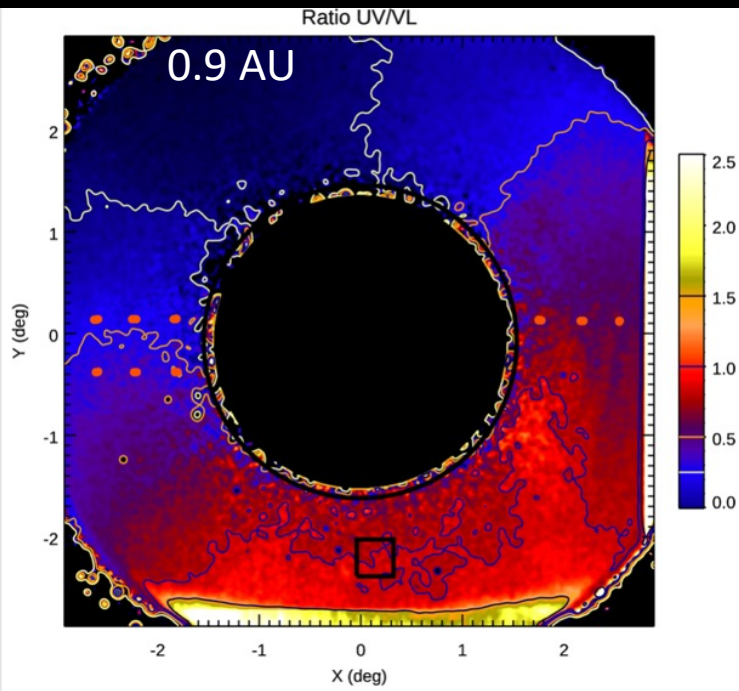
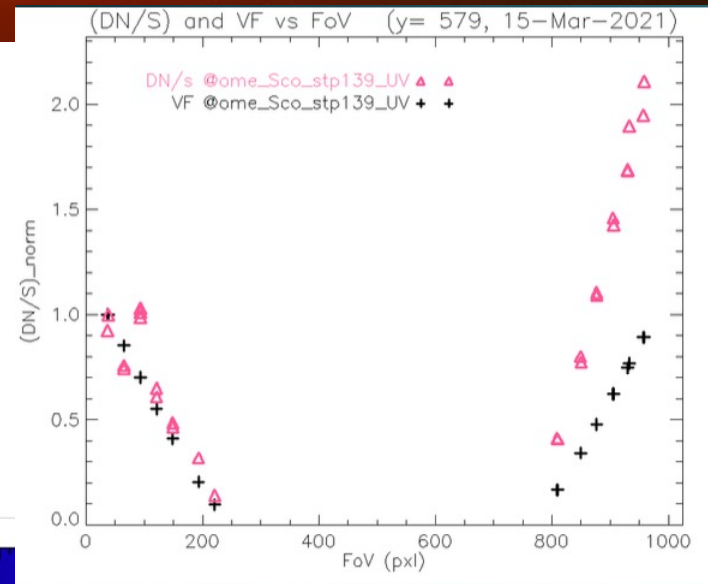
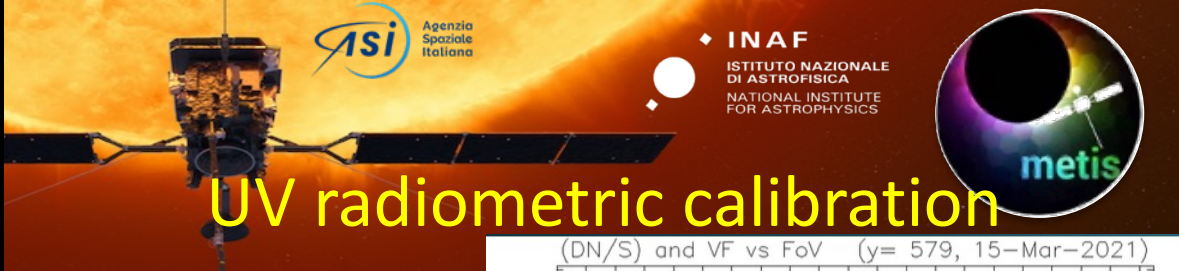
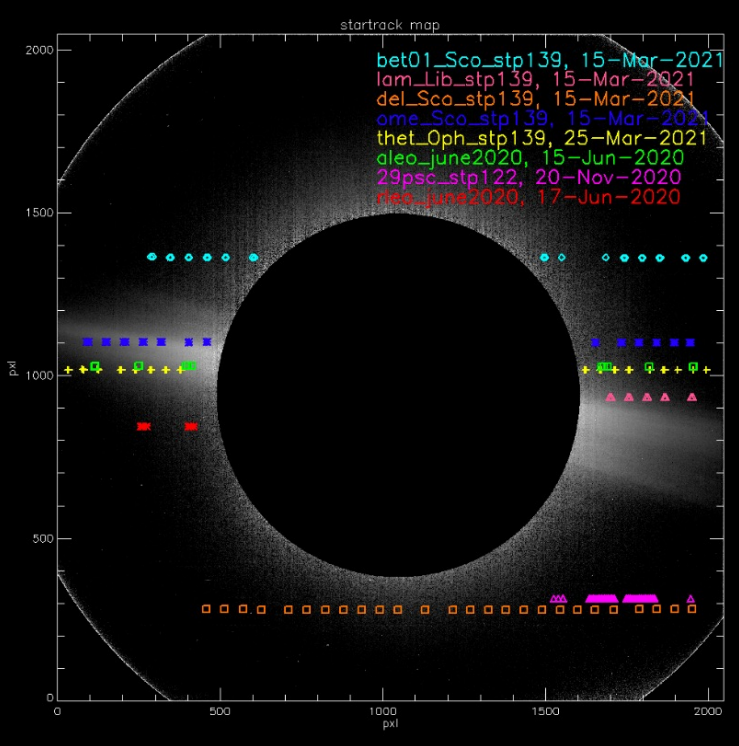
- **IO alignment** (Coarse and fine)
- Dark/Bias subtraction
- **Spatial resolution** (preliminary: need more star observations)
- **Polarimeter characterization**
- **Metis pointing verification**
- **Vignetting function** (Preliminary vignetting function for both VL and UV available)
- **VL radiometric calibration** (preliminary: confirms VL on-ground efficiency within uncertainties)
- **Stray light characterization**

## On-going activities:

- Perihelion tests on IO alignment and stray light (Feb-Mar 2022)
- Radiometric calibration (UV in progress – On-ground calibration temporarily used)

The full analysis is needed to complete the pipeline to L2 (within NMP)



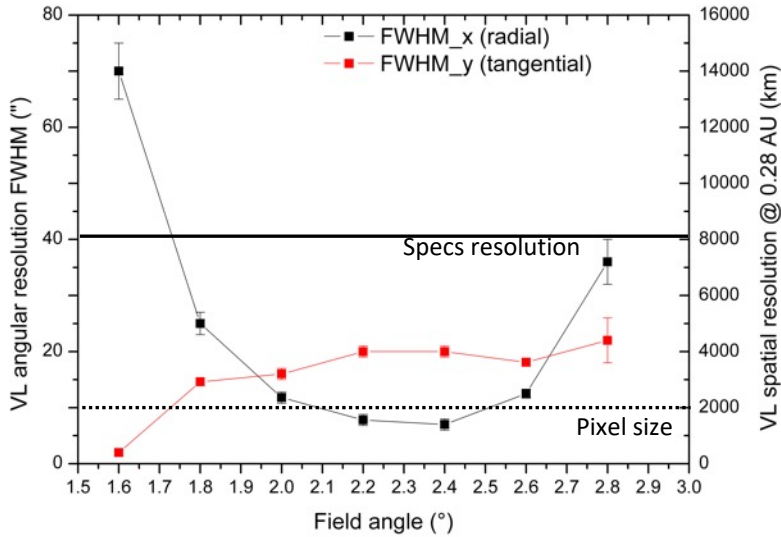




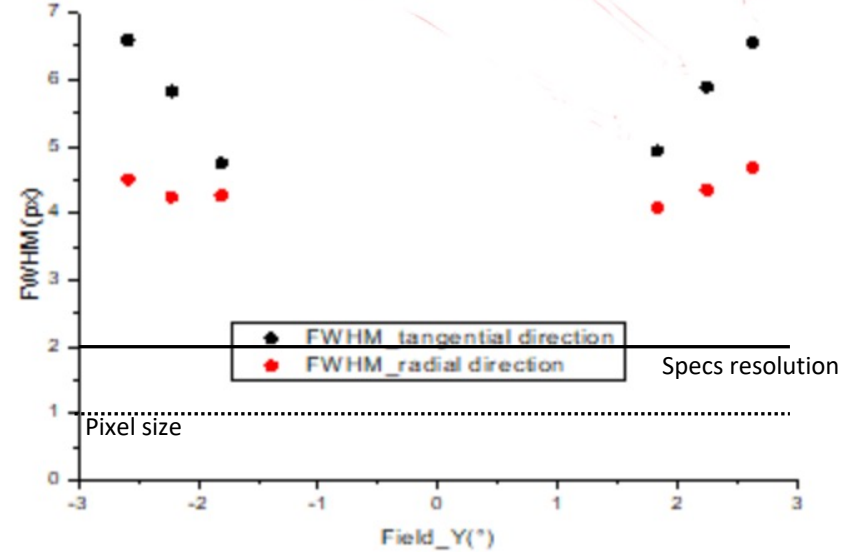
# METIS spatial resolution



## VL



## UV



	<b>Specs</b>
<b>Spatial Plate Scale (detector pixel)</b>	10 arcsec (VL) 2000 km @0.28AU 20 arcsec (UV)
<b>Spatial resolution</b>	40 arcsec (VL and UV analog)

Spatial resolution checked in-flight  
With Star observations





# Nominal Mission Phase



Starts in December 2021 – Metis will be **continuously** operating with upgraded telemetry:

- **Synoptic programs**
- **Increased Sun center pointing opportunities**
- **CME watch**
- **Targets of opportunity:**
  - Stars observations for radiometric and inter-instrument calibrations;
  - coordinated observations with PSP, SOHO, STEREO A and future solar missions
  - coordinated observations with ground based observatories
- **Density fluctuations**







# Nominal Mission Phase



Data will be available 3 months after download and processing as:

- **Level 2:** calibrated data (physical units).  
*Corrections for bias, dark current, flat-field, and vignetting, exposure normalisation, pointing, and radiometric calibration are applied. Stokes parameters, total/polarised-brightness images, polarisation angle/fraction.*
- *All the available orbital and attitude information is used and coordinates expressed in scientific coordinate systems (WCS).*
- **Level 3:** science data derived from L2 data,  
*Movies, Carrington maps; and data obtained after scientific analysis, i.e., electron-density maps, solar-wind outflow velocity maps.*

**Level 2 data release will start with the Nominal Mission (> Dec 2021)**



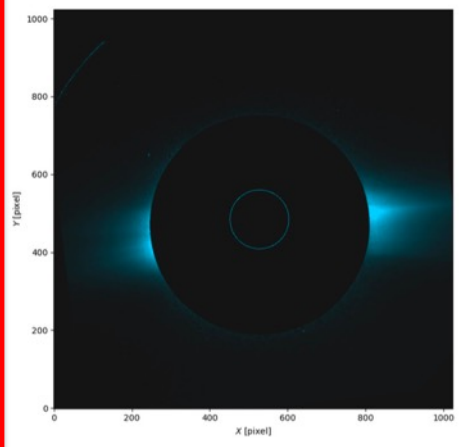


# Metis data

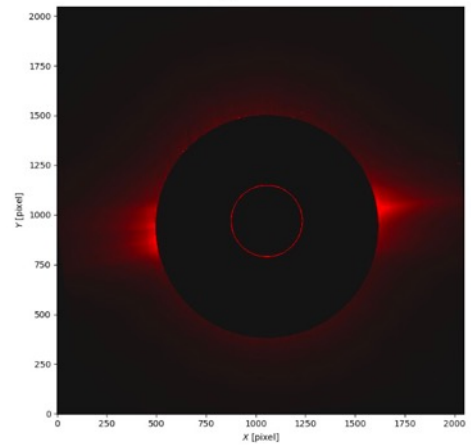
Level 2

Level 3

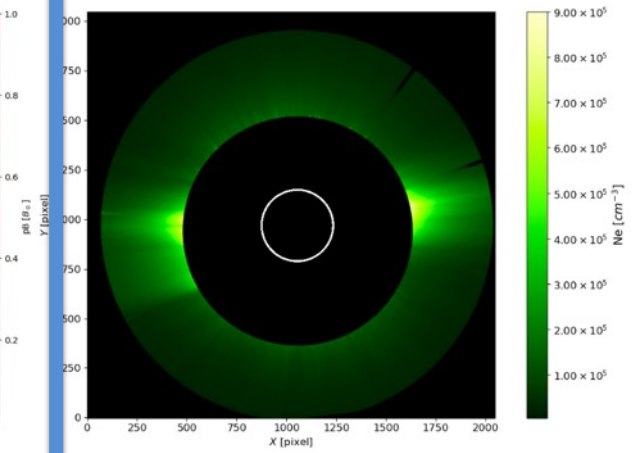
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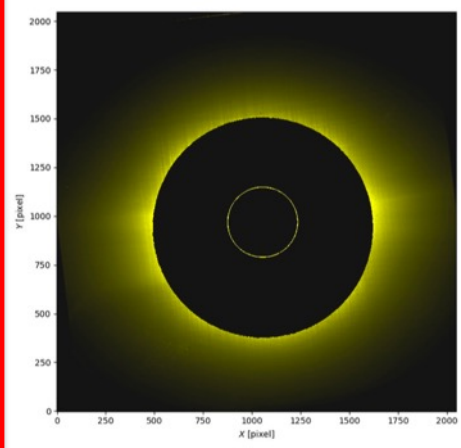
pB



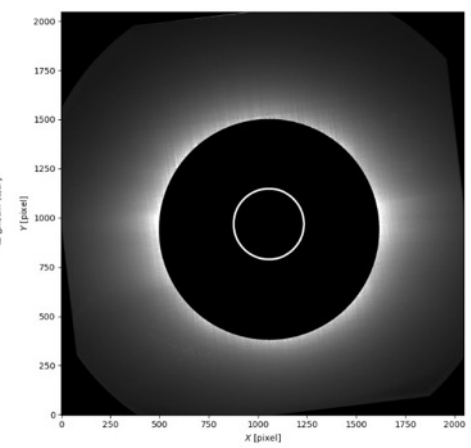
Ne



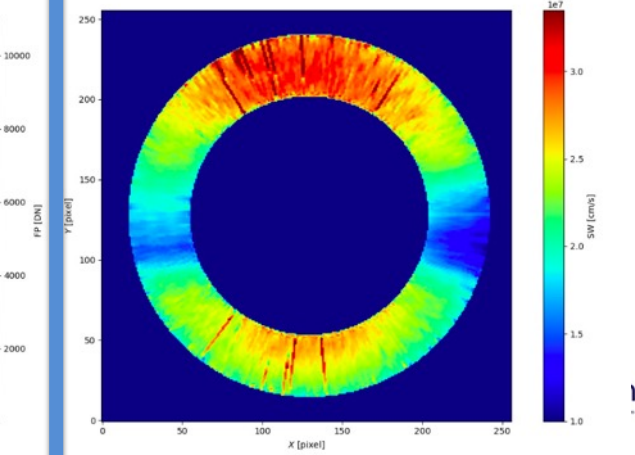
tB

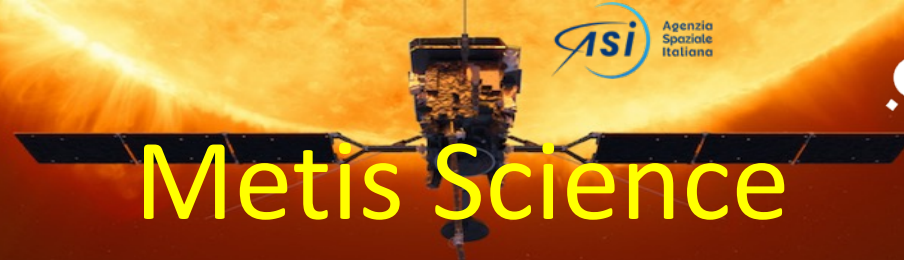


FP



SW



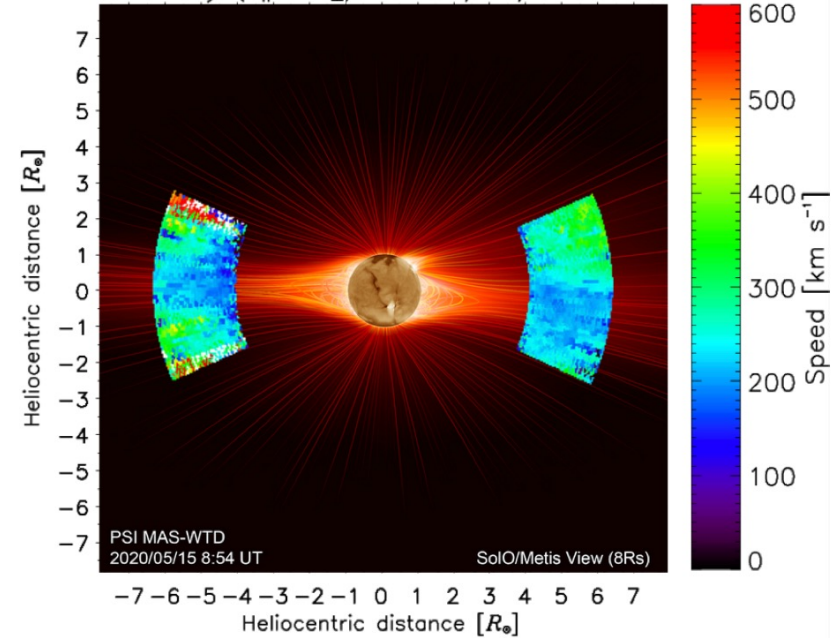


# Metis Science

## First paper on Metis data:

- Romoli et al., *First light observations of the solar wind in the outer corona with the Metis coronagraph*, A&A 2021
- Grimani et al., *Cosmic-ray flux predictions and observations for and with Metis on board Solar Orbiter*, A&A 2021
- Telloni et al., *Exploring the Solar Wind from its Source on the Corona into the Inner Heliosphere during the First Solar Orbiter – Parker Solar Probe Quadrature*, ApJL ,submitted
- Andretta et al., *The first Coronal Mass Ejection observed in both visible-light and FUV HiLy- $\alpha$  channels of the Metis Coronagraph on-board SolarOrbiter*, A&A, to be submitted

Outflow velocity ( $T_{||} = T_{\perp}$ ) 2020/05/15 11.39.25 UT

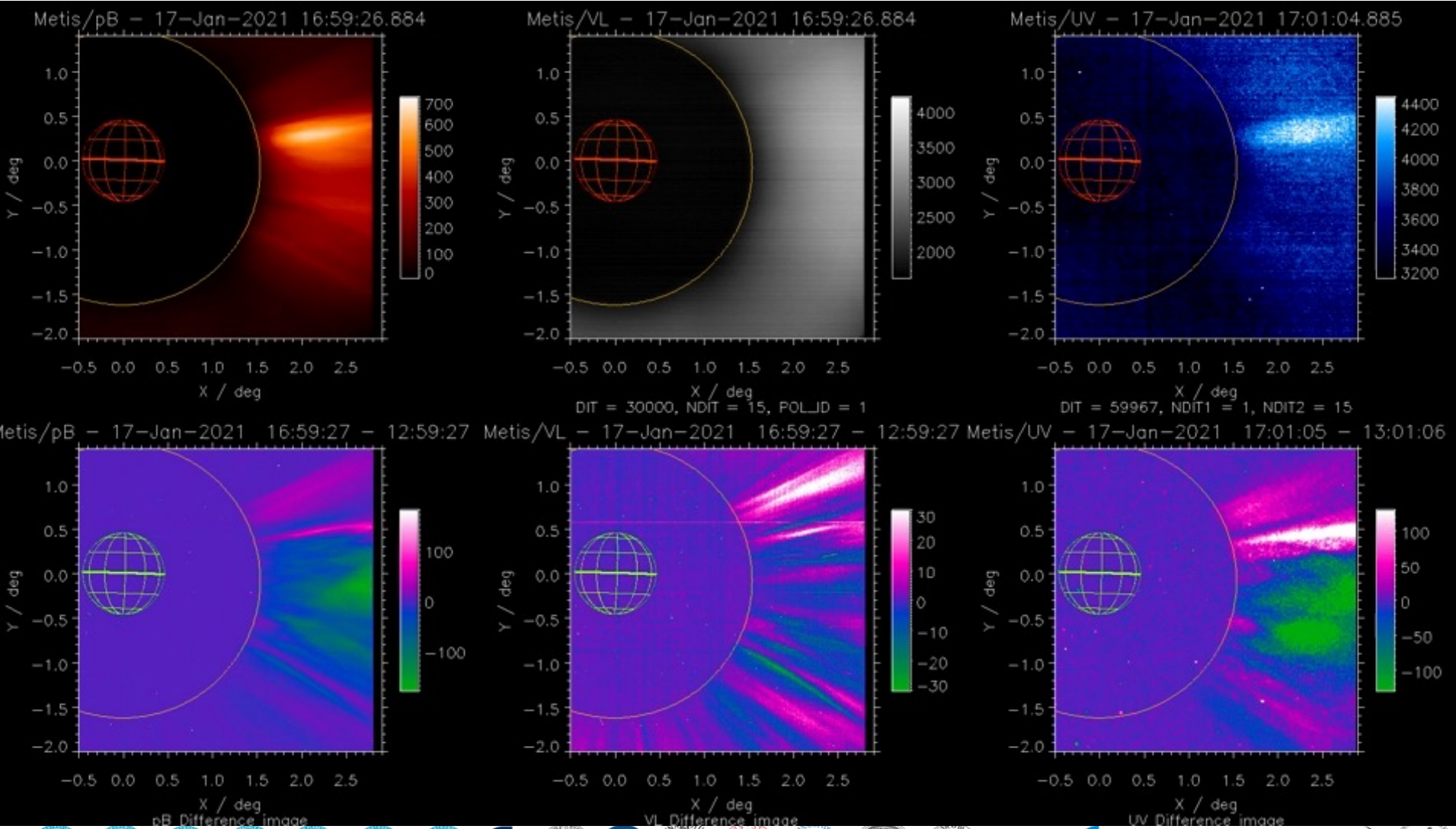


Many contribution from Metis in this ESPM-16 meeting





# First CME detection in VL and UV



# Science topics working groups

## TT1 - Wind diagnostics (R. Susino)

- a. Electron density  
(and electron temperature) (S. Fineschi)
- b. Hydrogen density (J.C. Vial)
- c. Wind velocity  
with Doppler dimming (R. Susino)

## TT2 - F-corona (F. Landini)

## TT3 - Combined synoptics (L. Teriaca)

## TT4 - Helium Diagnostics (V. Andretta)

## TT5 - Image enhancements (F. Frassetto)

## TT13 - Cosmic Rays (C. Grimani)

## TT14 - Sun grazing comets and other solar system bodies (V. Da Deppo)

## TT6 - Solar Wind (D. Telloni)

## TT7 - Large scale magnetic configuration and evolution, Streamers and pseudo-streamers (L. Strachan)

## TT8 - CMEs, prominence eruptions and blobs (P. Heinzel)

## TT12 - Modelling of CME propagation/evolution in corona and solar wind in connection with space weather (A. Bemporad)

## TT9 - Coronal shocks, particle acceleration (G. Zimbardo)

## TT10 - Plasma density fluctuations and waves (G. Nisticò)

## TT11 - Flux emergence, magnetic field reconnection, coronal heating, flares (F. Reale)

Synergies with several space missions:

SOHO, STEREO, SDO, PSP, Proba3, ASO-S, Aditya, UVSC, PUNCH, CODEX, Solar C,  
and Ground based telescopes



Thank you





# METIS Design Overview

