

INTRODUCTION

- Sunspot areas are strongly correlated with other solar activity indices as the Group Number or Sunspot Number
- Measures of the sunspot areas in historical drawings are very interesting in order to analyze the past solar activity
- Analysis is carried out by using computer programs that binarize the sunspot images

Sunspot area (in msh) corrected by foreshortening in the plane-parallel approximation

$$A_M = \frac{10^6}{2\pi R^2} \frac{A_S}{\cos \rho}$$

A_S → sunspot area measured directly onto the image

R → radius of the solar disk (in pixels)

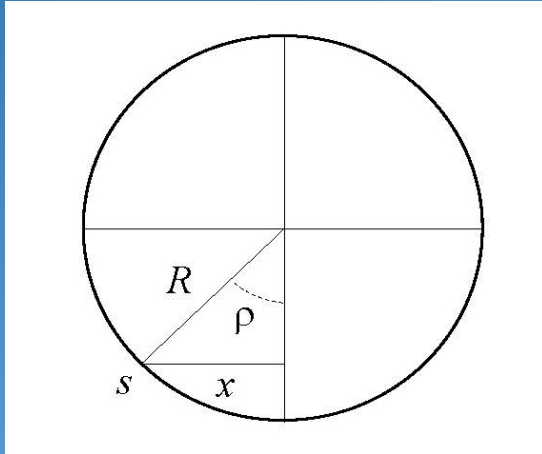
ρ → angle between the direction of the centre of the solar disk and the direction of the sunspot

On the uncertainties of the sunspot area measurements

F. Sánchez-Bajo and J.M. Vaquero
Universidad de Extremadura, Spain

UNCERTAINTIES

1. Uncertainty associated to the sunspot position $\rightarrow \sigma(A_M) = A_M \tan \rho \frac{\sigma(x)}{R}$



2. Uncertainty associated to the resolution of the image $\rightarrow \sigma(A_M) = A_M \frac{2\sigma(R)}{R}$

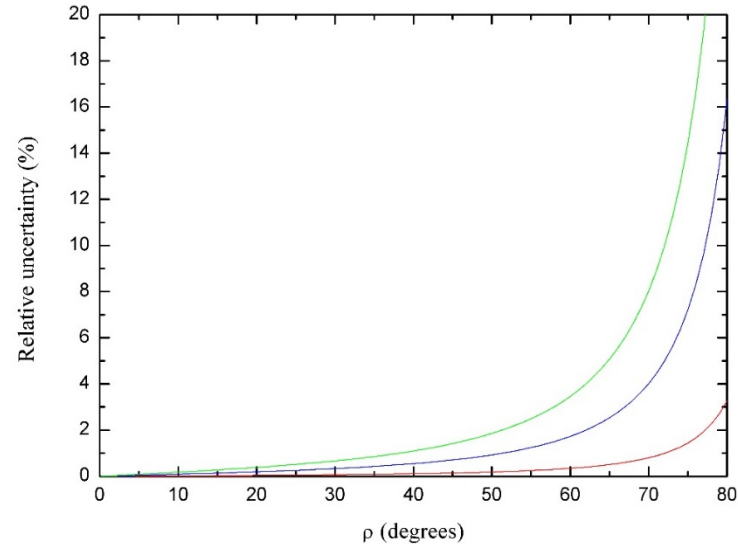
3. Uncertainty associated to the error in the determination of the sunspot area (binarization) $\rightarrow \sigma(A_M) = A_M \frac{\sigma(A_S)}{A_S}$

On the uncertainties of the sunspot area measurements

F. Sánchez-Bajo and J.M. Vaquero
Universidad de Extremadura, Spain

UNCERTAINTY DUE TO THE SUNSPOT POSITION

The relative sunspot area error increases with $\rho \rightarrow$ Very important in the analysis of historical drawings (inaccuracies in the sunspot positions)



Relative uncertainty of the sunspot area as function of the angle ρ for an error in x of 0.1 % (red line), 0.5 % (blue line) and 1 % (green line). The percentages are referred to the radius of the solar disk.

On the uncertainties of the sunspot area measurements

F. Sánchez-Bajo and J.M. Vaquero
Universidad de Extremadura, Spain

UNCERTAINTY DUE TO THE BINARIZATION PROCESS (I)

Binarization → transformation of an 8-bits image (256 gray levels from 0 to 255) to a black and white image (black = 0 / white = 1)

- **Binarization threshold** t ($0 < t < 1$)
- **Cut-off intensity** $c \rightarrow c = minval + t(maxval - minval)$

If intensity $> c$, then it is assigned the value 1 (white). Otherwise, it is assigned the value 0 (black)
Sunspot area (uncorrected by foreshortening) → sum of all black pixels

- **The sunspot area depends critically on the binarization threshold**
- **Optimal t value is unknown** → binarization threshold freely selected by the user

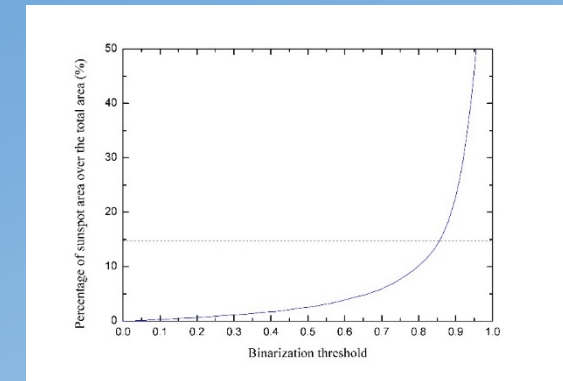
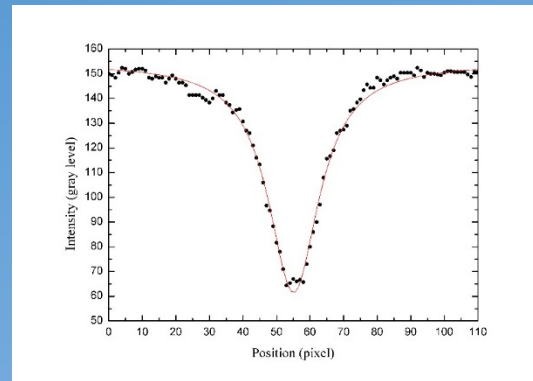
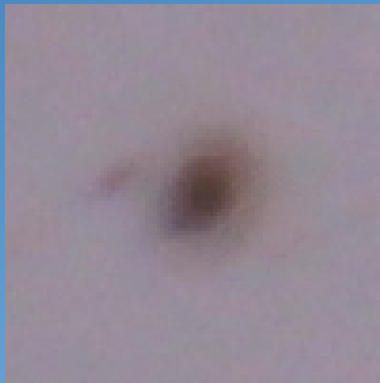
On the uncertainties of the sunspot area measurements

F. Sánchez-Bajo and J.M. Vaquero
Universidad de Extremadura, Spain

UNCERTAINTY DUE TO THE BINARIZATION PROCESS (II)

Estimation of the best binarization threshold values:

- ❖ Selection of a real sunspot (a) from a digital image (8-bits) and measuring the intensity along the sunspot
- ❖ Fit a Cauchy function to the intensity data (b). This Cauchy function represents a simulated sunspot
- ❖ Introduction of Gaussian noise to simulate a real sunspot, with fluctuations in the intensity
- ❖ Binarization of the noise image
- ❖ Comparison the number of black pixels in the binarized image with those of the original simulated image, using a convenient reference value to define the sunspot limits (c)



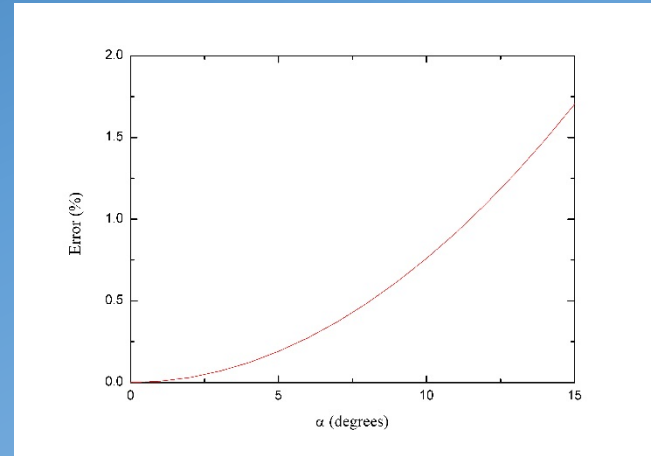
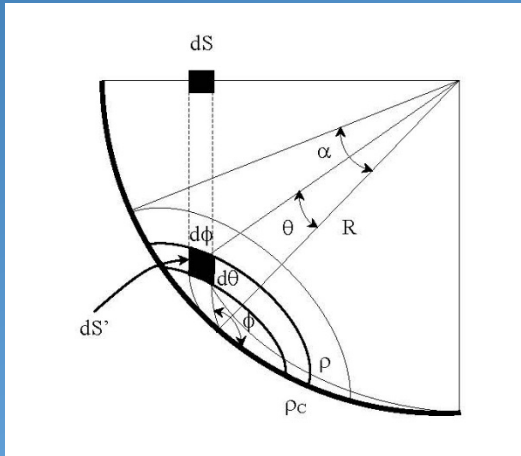
On the uncertainties of the sunspot area measurements

F. Sánchez-Bajo and J.M. Vaquero
Universidad de Extremadura, Spain

ANALYTICAL SUNSPOT AREA FOR CIRCULAR SHAPE

The size effects are ignored in the plane-parallel approximations

- Sunspot area of a circular sunspot with radius $\alpha \rightarrow S' = 2\pi R^2(1 - \cos \alpha)$
- Projected sunspot area (measured over the image) $\rightarrow S = \int dS = \int dS' \cos \rho = R^2 \int \cos \rho \sin \theta d\theta d\varphi = 0.5 \pi R^2 \cos \rho_c (1 - \cos 2\alpha)$ $\rho_c =$ angle between the direction of the centre of the sunspot and the direction of the centre of the disk



The error due to the use of the classical plane-parallel approximation is very small excepting for large sunspots

On the uncertainties of the sunspot area measurements

F. Sánchez-Bajo and J.M. Vaquero
Universidad de Extremadura, Spain

ANALYTICAL SUNSPOT AREA FOR ELLIPTICAL SHAPE

The spherical ellipse is the translation of the plane ellipse to a sphere, with semi-major angle α and semi-minor angle β

➤ Sunspot area of the elliptical sunspot $\rightarrow S' = 2\pi R^2 - 4R^2 \sqrt{\frac{A^2-1}{A^2(1+C^2)}} \Pi \left[\frac{C^2}{1+C^2}, \frac{\pi}{2} - \gamma \right]$

$\Pi \rightarrow$ Elliptic integral of the third kind

$$A = \frac{1}{\sin \alpha} \quad B = \frac{1}{\sin \beta} \quad C = \sqrt{\frac{B^2 - A^2}{A^2 - 1}} \quad D = \sqrt{\frac{B^2 - A^2}{A^2}} \quad \gamma = \sin^{-1} \sqrt{\frac{1 + D^2}{1 + C^2}}$$

➤ Projected sunspot area (measured over the image) $\rightarrow S = \int dS = \int dS' \cos \rho = R^2 \int \cos \rho \sin \theta d\theta d\varphi = \pi R^2 \cos \rho_c \sin \alpha \sin \beta = \pi ab \cos \rho_c$ (projected plane ellipse)

$$a = R \sin \alpha$$

$$b = R \sin \beta$$

On the uncertainties of the sunspot area measurements

F. Sánchez-Bajo and J.M. Vaquero
Universidad de Extremadura, Spain