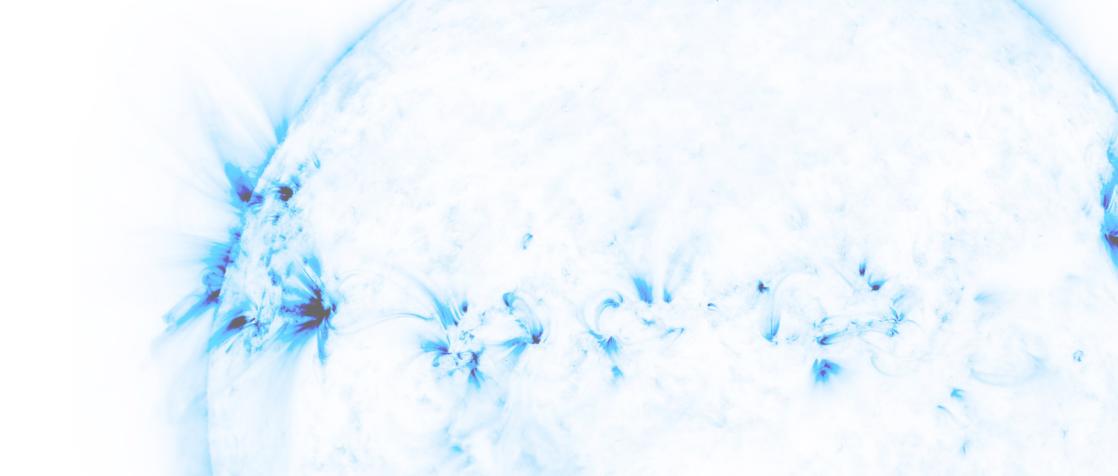


The University of Dublin

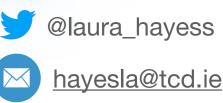




Solar Flare Impacts on Lower Ionosphere

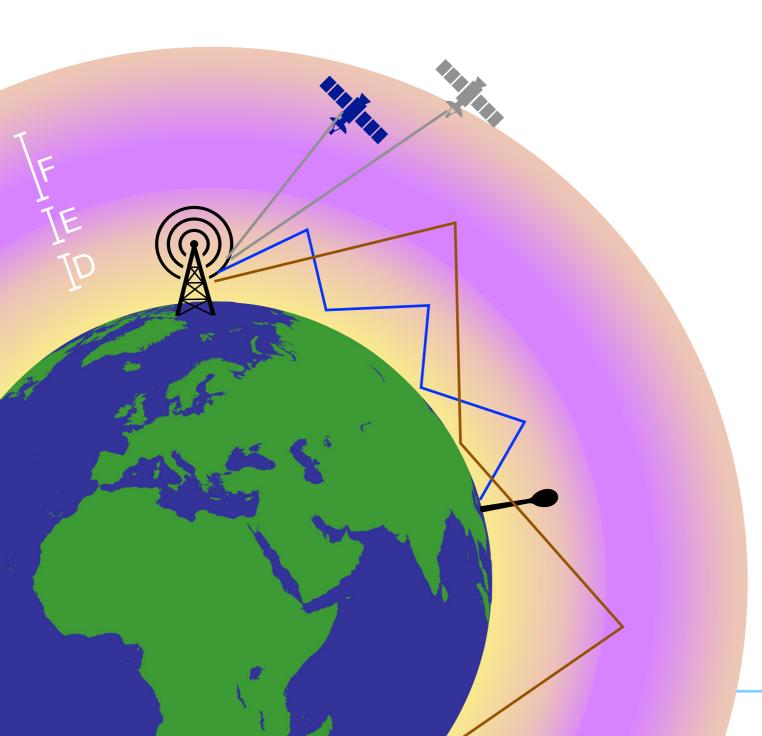
A statistical study

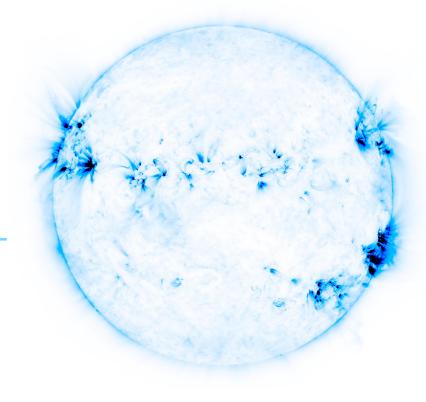
Laura A. Hayes Oscar O'Hara, Sophie Murray & Peter Gallagher



ESPM-16 Sept 7 2021

Earth's lonosphere Linked to the Sun





- Ionosphere plays a governing role in radio communications (e.g. GPS, HF comms)
- Critical frequency in which wave will reflect is dependent on the electron density N_e

 $f_{critical} \propto \sqrt{N_e}$

Earth's lonosphere

Linked to the Sun

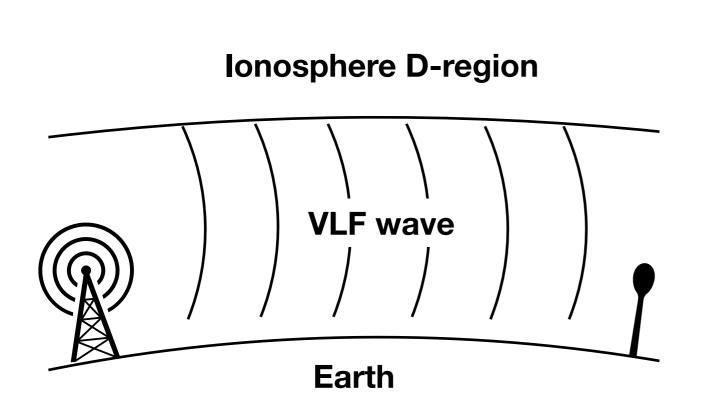
 Increased fluxes induce significant photoionisation

 Substantially increases electron density -> X-rays penetrate down to lowest lying Dregion impact HF radio comms

Increased e⁻ density

X-rays and

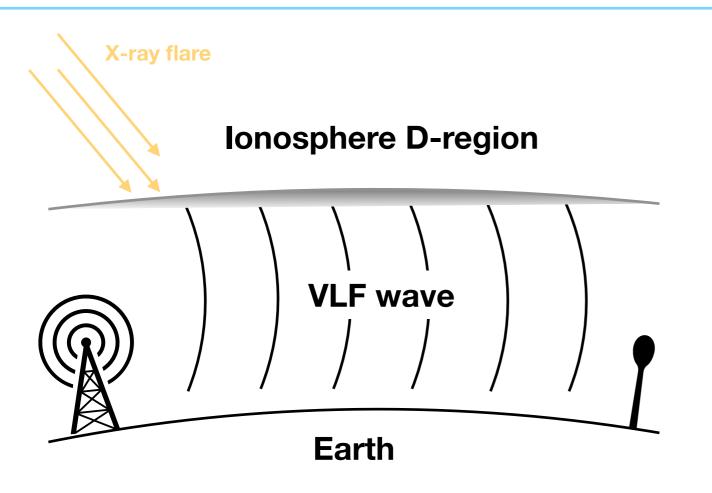
Solar flare impacts on D-region Lowest lying region



 D-region formed during the day by Lyman-α (121.6nm) acting on neutral Nitric Oxide (minor constituent)

- Very Low Frequency (VLF 3-30kHz) waves propagate in wave-guide
- Large (1MW) Navy transmitters are used for global submarine comms!

Solar flare impacts on D-region Lowest lying region

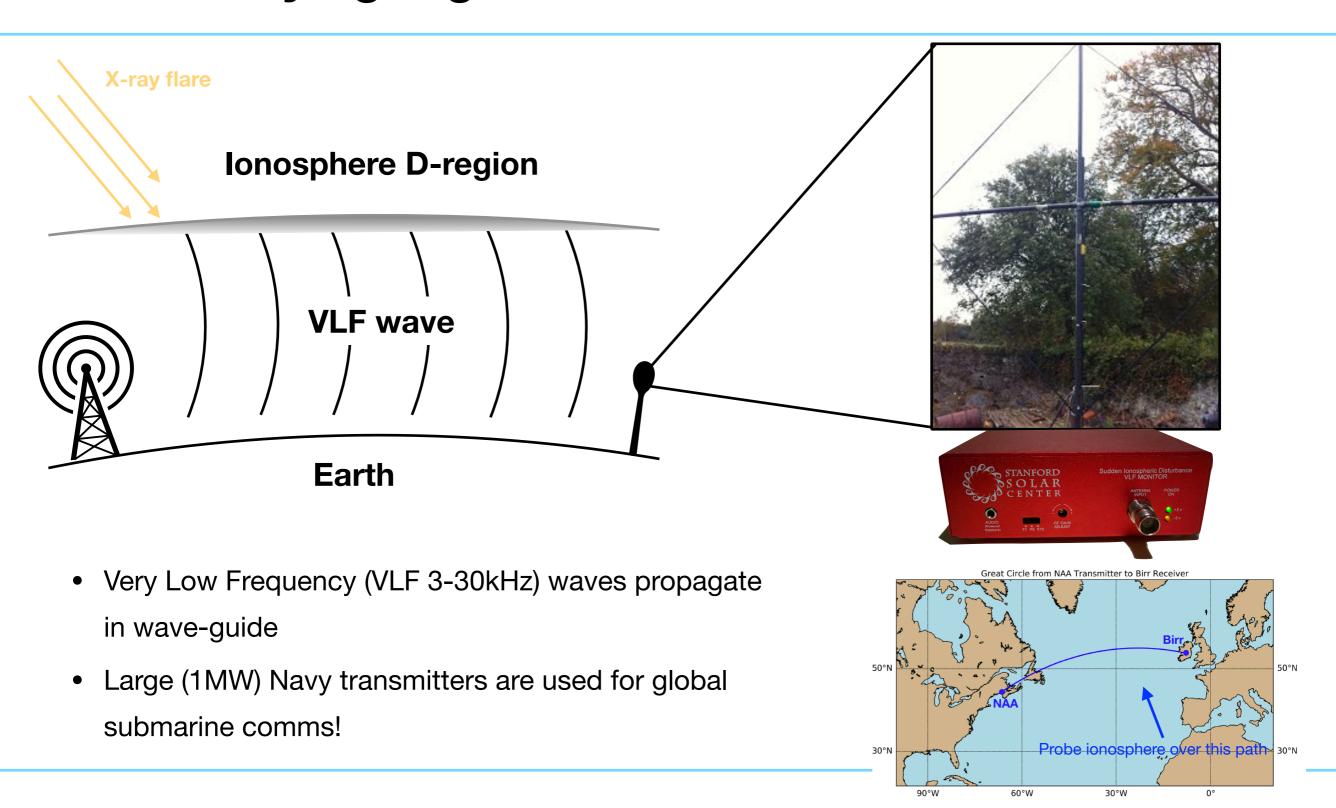


- Very Low Frequency (VLF 3-30kHz) waves propagate in wave-guide
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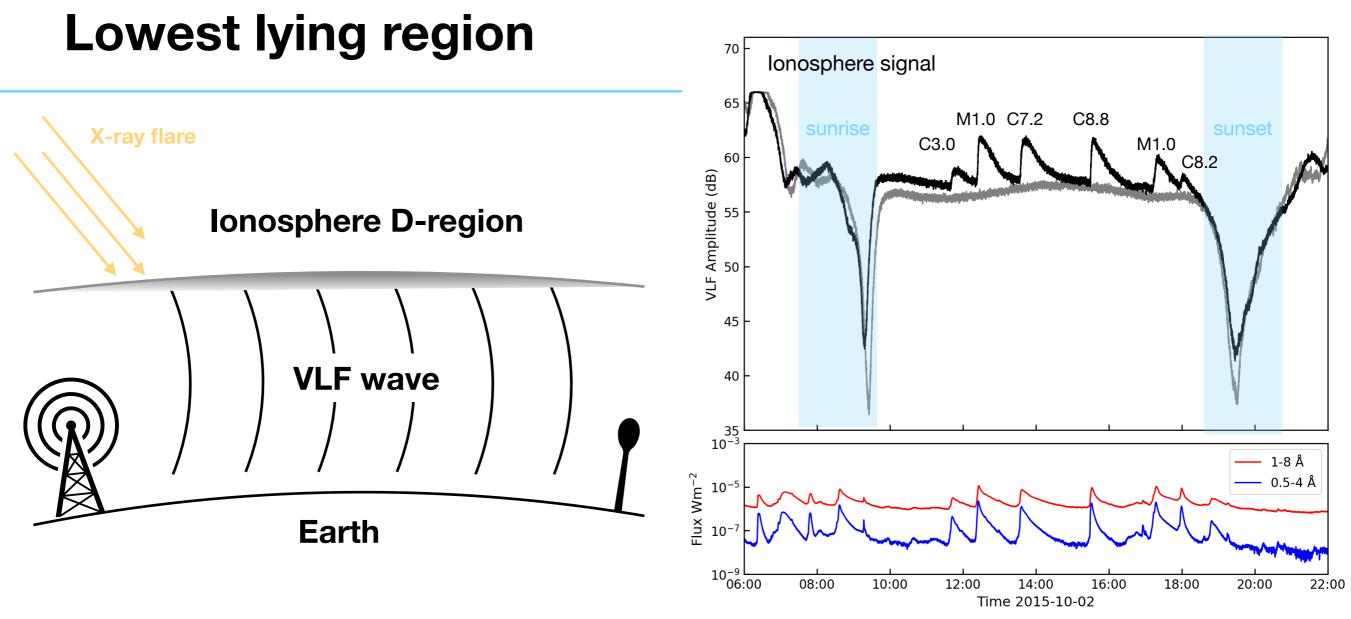
- D-region formed during the day by Lyman-α (121.6nm) acting on neutral Nitric Oxide (minor constituent)
- During flare X-rays <1nm ionise N₂ and O₂ (dominant particles)

Significantly changes electron density = changes propagation conditions for VLF waves

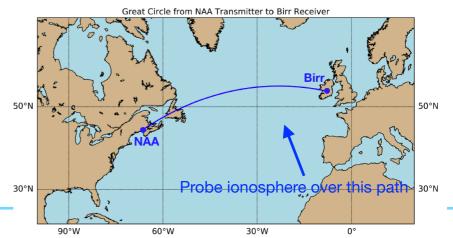
Solar flare impacts on D-region Lowest lying region



Solar flare impacts on D-region

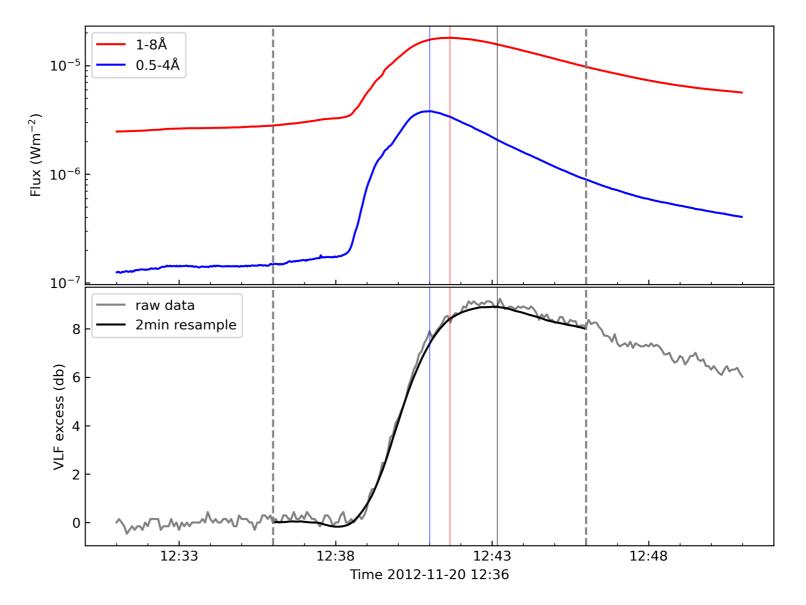


- Very Low Frequency (VLF 3-30kHz) waves propagate in wave-guide
- Large (1MW) Navy transmitters are used for global submarine comms!



Overview

- VLF at Birr has been recording since 2012 2018
- Total of 342 flares of X,
 M, and C class (missed many ()) but still one of
 the largest statistical
 studies of VLF/flare
 analysis



- peak X-ray fluxes
- peak VLF amplitude (excess from background)
- time delays between X-ray and VLF

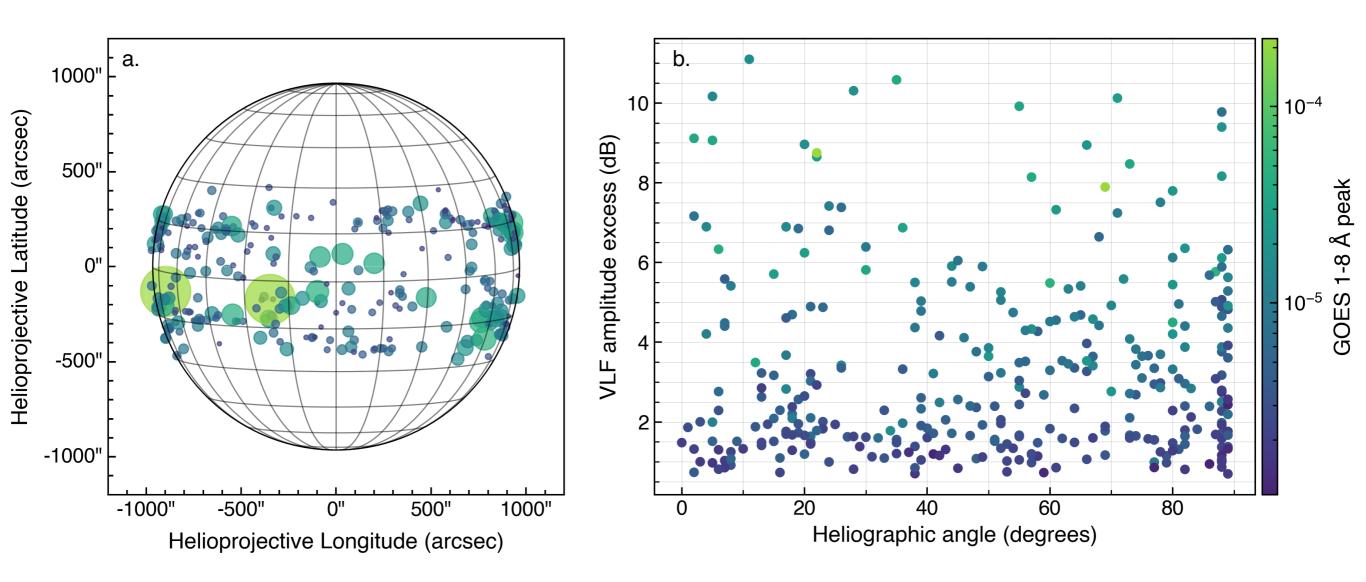
Flare flux and VLF amplitude

1-8Å 0.5-4Å b. a. 10.0 **10**⁻⁴ 1-8 Å X-ray peak flux (Wm⁻² VLF amplitude excess (dB) ' 10⁻⁵ 1.0 10⁻⁶ 10⁻⁵ 10-5 10-4 10^{-7} 10^{-4} 10^{-1} 1-8 Å X-ray peak (background sub) (Wm⁻²) 0.5-4 Å X-ray peak (background sub) (Wm⁻²)

Correlation coefficients of 0.8

Larger X-ray flux -> larger the ionospheric response

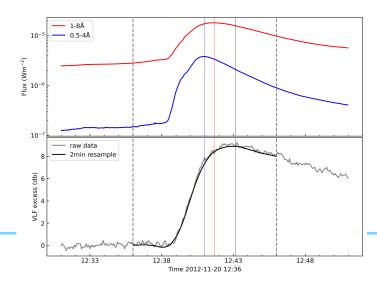
Flare location and ionospheric response

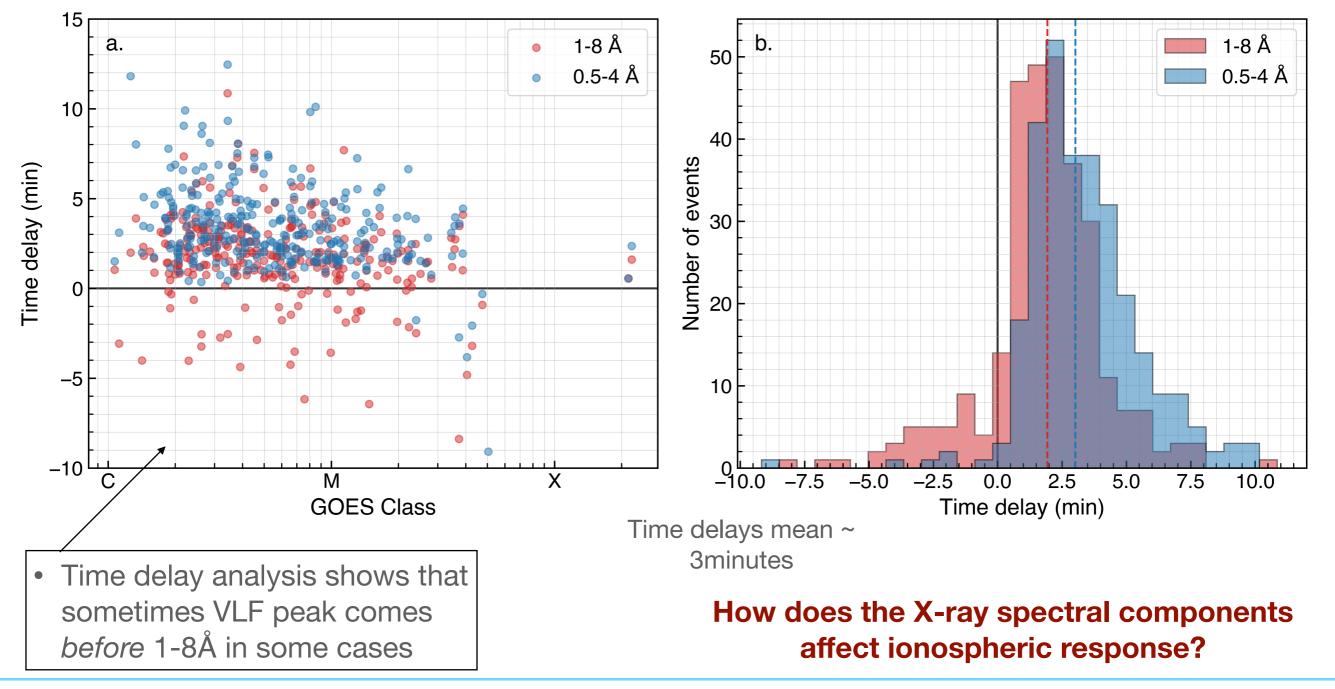


No relationship between flare location on disk and D-region impacts

Rules out Lyman-α being a major factor (centre to limb variation)

Time delay analysis





Statistical Flare Analysis Conclusions

- Solar flare X-ray and EUV emission has a direct impact on Earth's lower ionosphere the lowest lying D-region is most impacted.
- Statistical study of >300 flares shows larger soft X-ray flux, larger the ionospheric response.
- Unlike upper regions of the ionosphere, there is no dependence on the Dregion response to the location of the flare on disk.
- Time-delay analysis reveals that harder X-rays also have a significant impact and that the full spectral components of the flare should be considered in future works.

Much work still to be done regarding spectral components of flares and their impacts!