



Statistical study of Stream Interaction Regions at 1 and 1.5 AU

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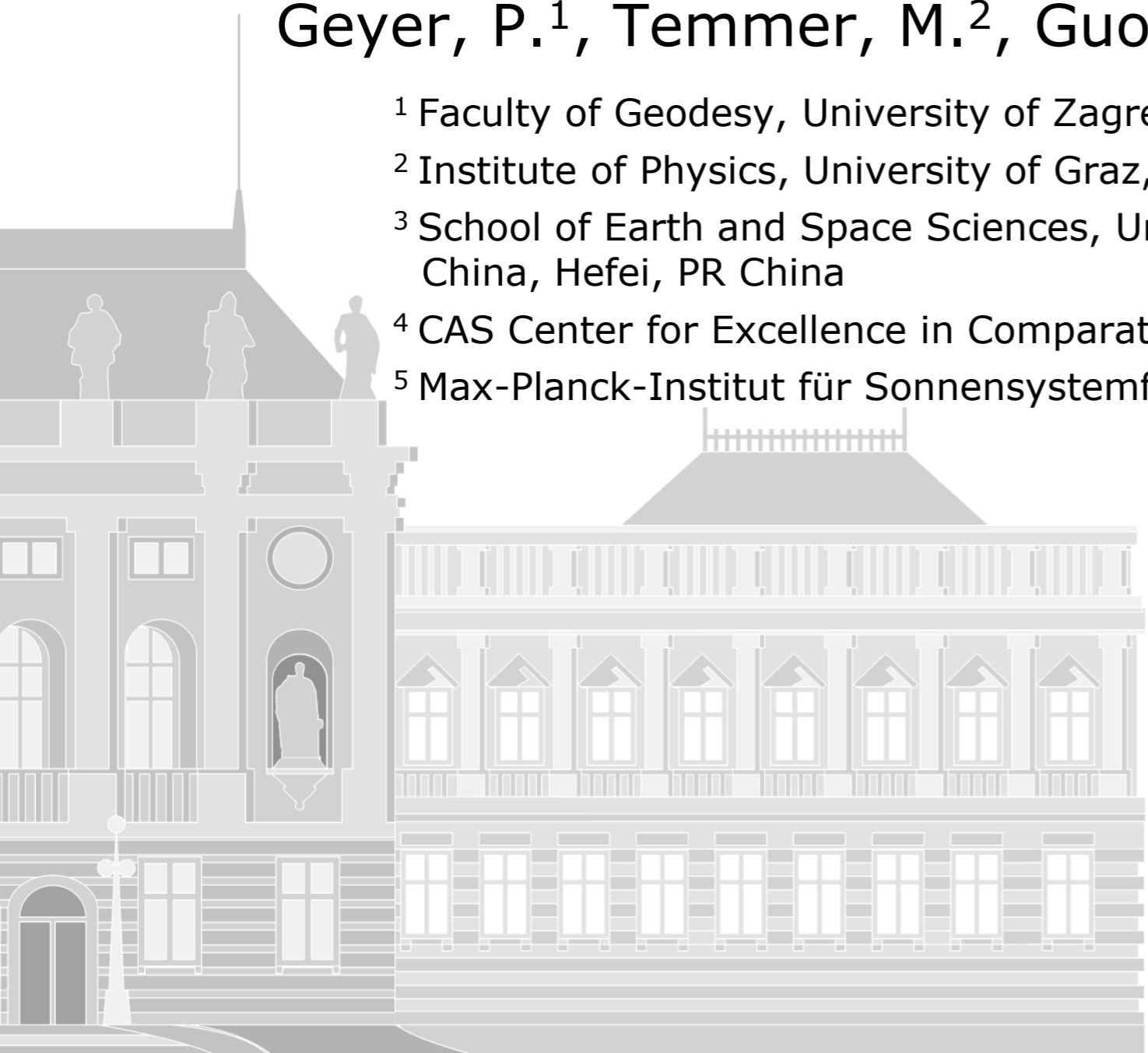
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FWF

Der Wissenschaftsfonds.

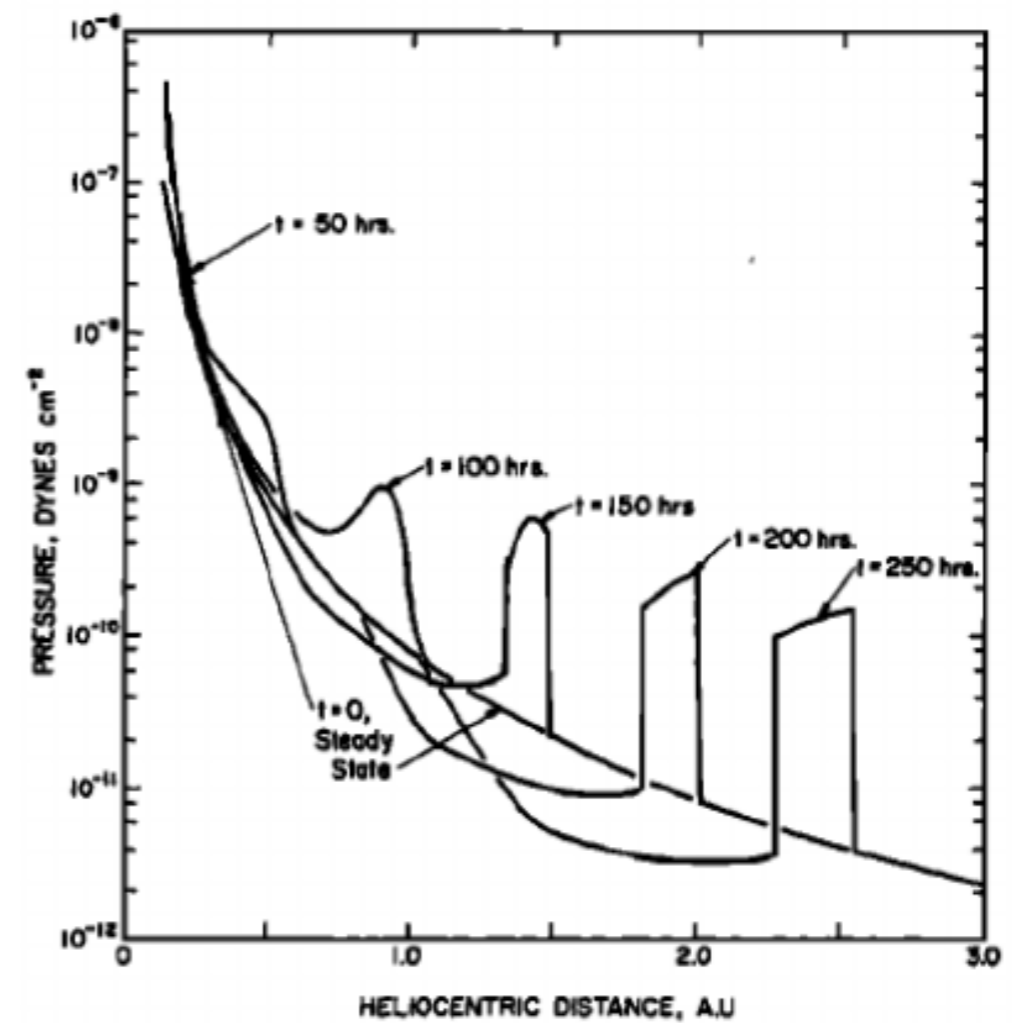
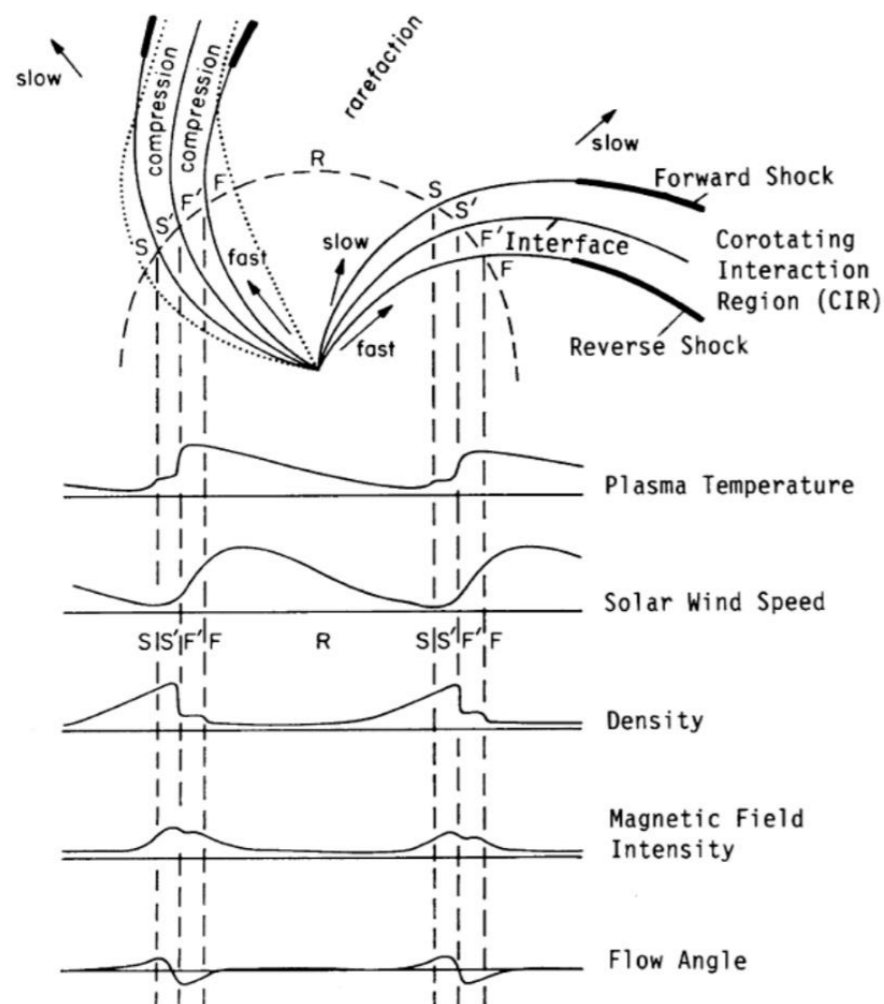


FFG



State of the Art

- Stream expansion of SIRs at the interface and shock formation predicted by models:
 - Hundhausen 1973
 - Gosling and Pizzo 1999
 - Cranmer+ 2007

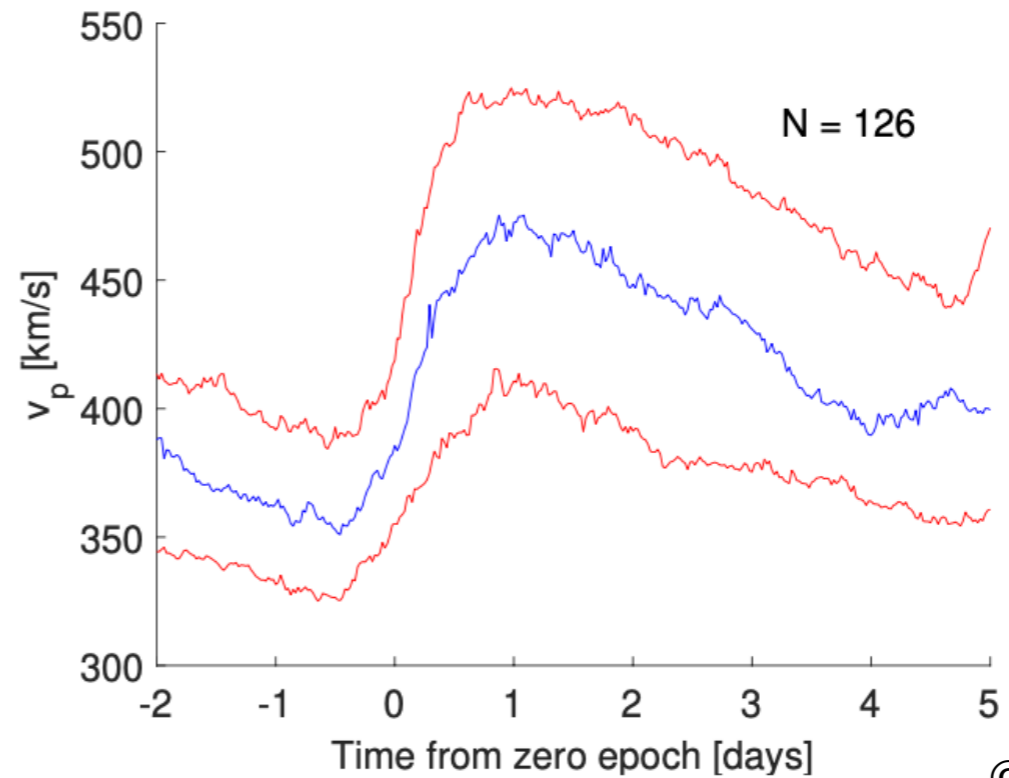
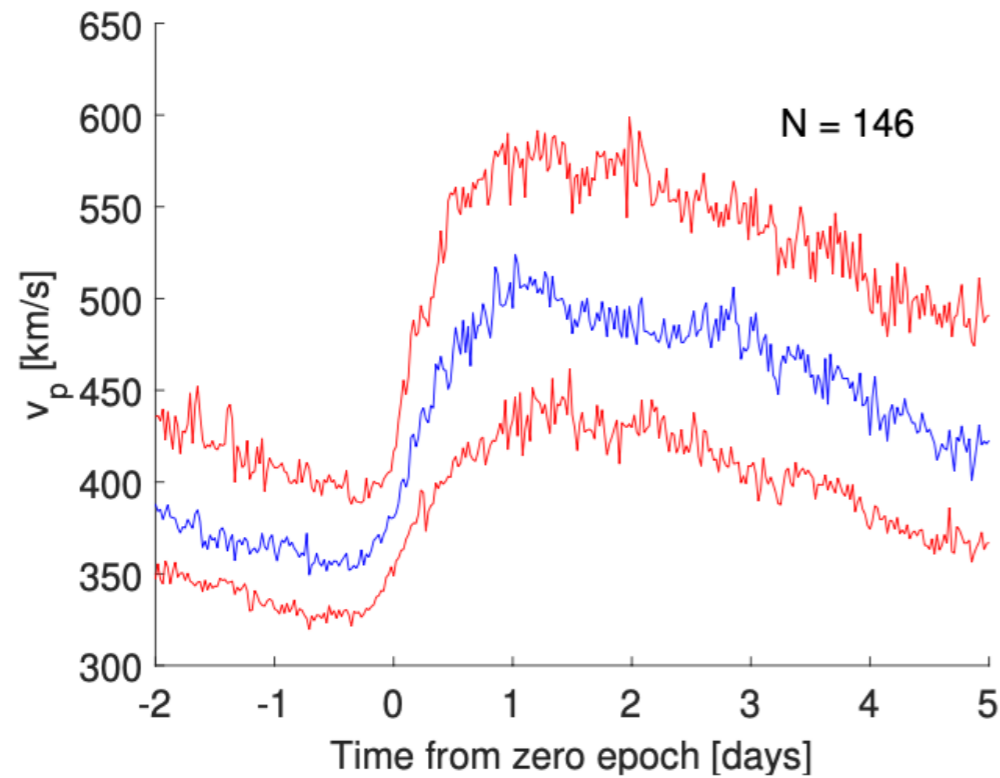


Data & Methods I

- **OMNI** and **MAVEN** in-situ plasma and magnetic field data, **SDO** EUV images
- Catalogs of **Grandin+ 2019** and **Huang+ 2019**
- **Superposed epoch analysis:** Extraction of statistical parameters of events shifted to reference (“epoch”) time
- **Aligned events analysis:** Correlation of **CHs** extent with SIR plasma properties

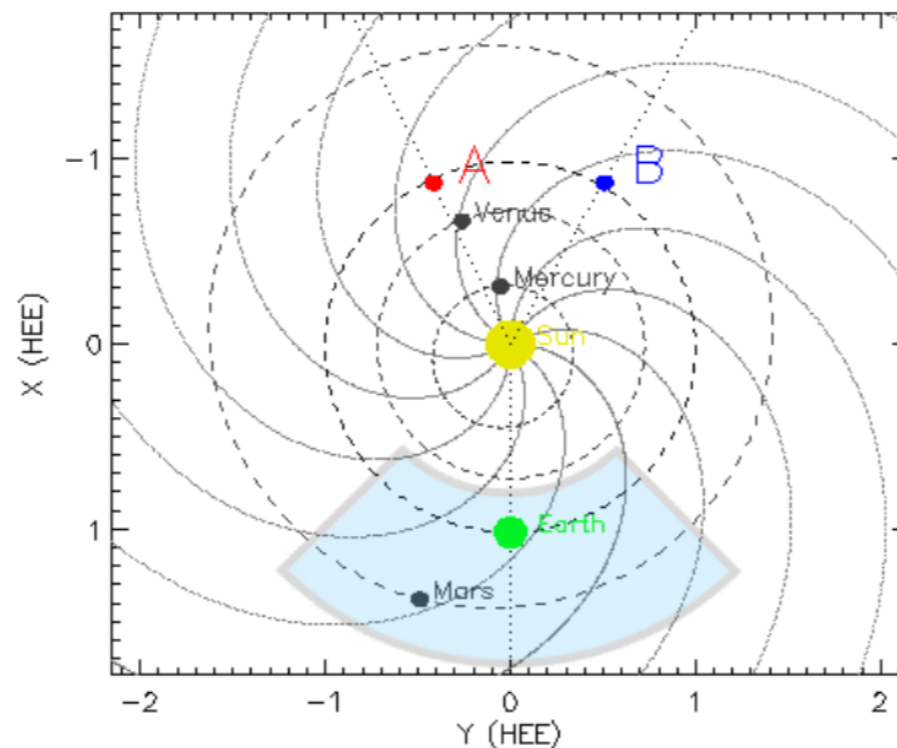
Data & Methods II

SEA example, SW bulk speed (Earth: left; Mars: right)



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Aligned events constrain:



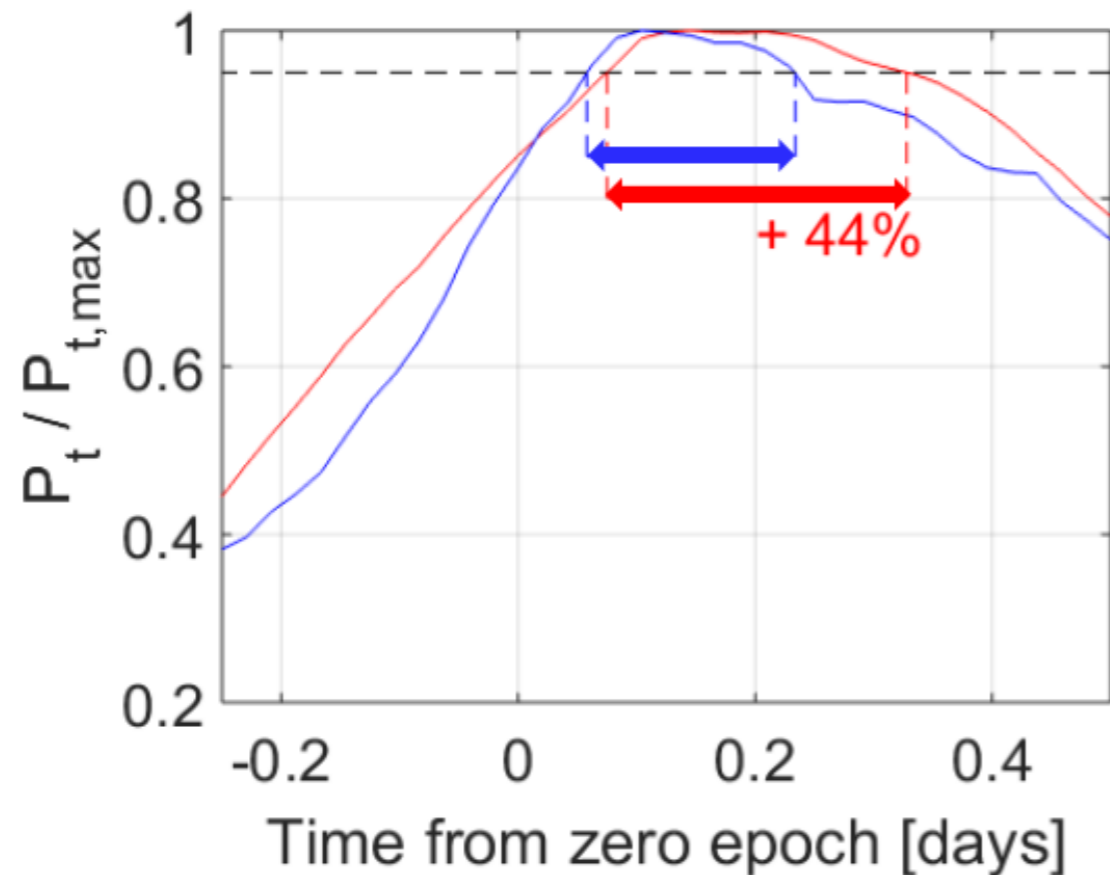
$$\Delta t = t_M - t_E = \frac{\phi_M - \phi_E}{\omega_{\text{sun}}} + \frac{r_M - r_E}{v_{\text{SW}}}$$

Opitz+ 2009

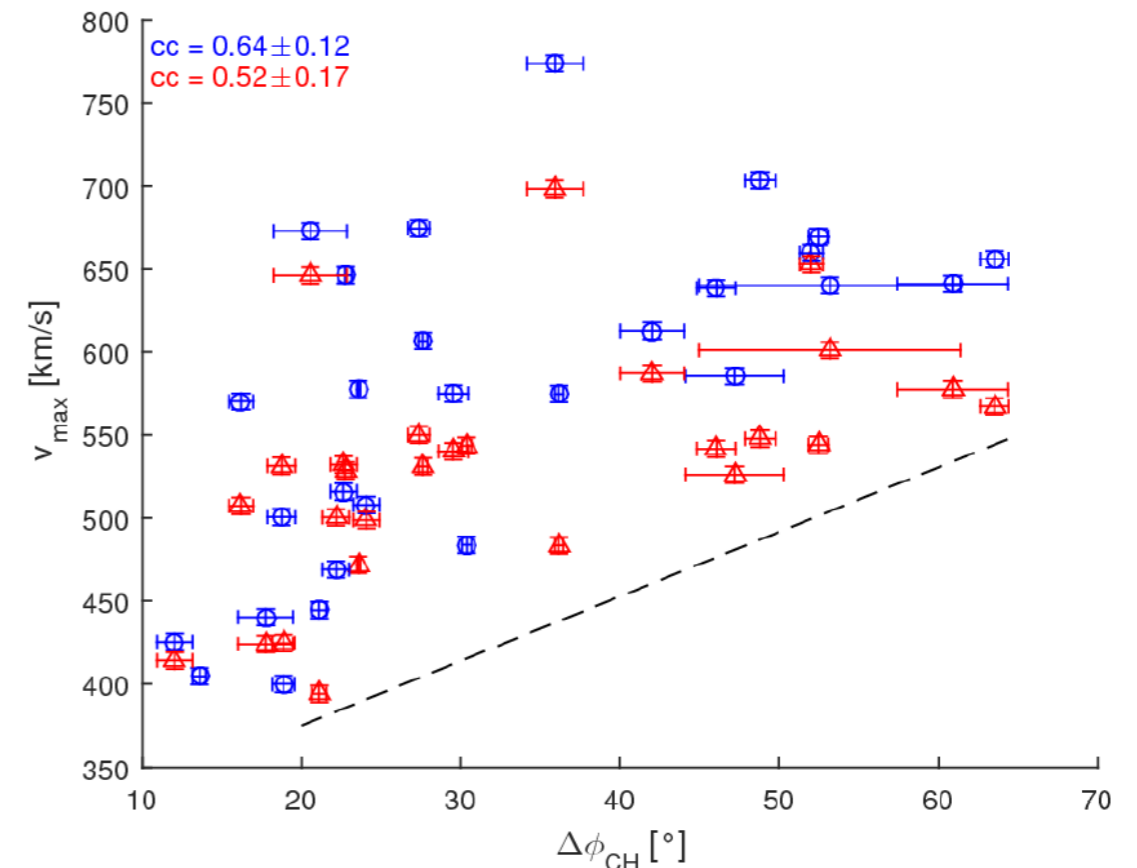
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Results

- Expansion of wave crest from 1 – 1.5 AU:



- Good correlation of maximum SW speed with CH $\Delta\phi$:



- Shock occurrence rate:

Shock type	Earth	Mars
FF only	6.7% (3)	20.0% (9)
FR only	6.7% (3)	6.7% (3)
FF and FR	0% (0)	8.9% (4)
FF and/or FR	13.3% (6)	35.6% (16)

Discussion & Conclusion

- Most evolution happens on front side of SIRs from 1 – 1.5 AU, as shock occurrence rates imply
- Expansion of streams linked to higher shock occurrence rate (Gosling and Pizzo 1999, Hundhausen 1973)
- Steepening of waves also consistent with amplitude analysis of Richter and Luttrell (1986, 0.35 – 0.95 AU)
- High correlation of HSS maximum speed with latitudinal extent → only valid for small to medium sized CHs? (c.f. Garton+ 2018)

References

Cranmer, Steven R., Adriaan A. van Ballegoijen, and Richard J. Edgar (Aug. 2007). „Self-consistent Coronal Heating and Solar Wind Acceleration from Anisotropic Magnetohydrodynamic Turbulence“. In: *The Astrophysical Journal Supplement* 171.2, pp. 520–551. doi: 10.1086/518001.

Garton, Tadhg M., Sophie A. Murray, and Peter T. Gallagher (2018). „Expansion of High-speed Solar Wind Streams from Coronal Holes through the Inner Heliosphere“. In: *The Astrophysical Journal* 869.1, p. L12. doi: 10.3847/2041-8213/aaf39a.

Geyer, P., Temmer, M., Guo, J. and Heinemann, S. (May 2021). “Properties of stream interaction regions at Earth and Mars during the declining phase of SC 24”. In: *Astronomy & Astrophysics* 649, id.A80, 20 pp. doi: 10.1051/0004-6361/202040162.

Gosling, J. T. and V. J. Pizzo (July 1999). „Formation and Evolution of Corotating Interaction Regions and their Three Dimensional Structure“. In: *Space Science Reviews* 89, pp. 21–52. doi: 10.1023/A:1005291711900.

Grandin, Maxime, Anita T. Aikio, and Alexander Kozlovsky (2019). „Properties and Geoeffectiveness of Solar Wind High-Speed Streams and Stream Interaction Regions During Solar Cycles 23 and 24“. In: *Journal of Geophysical Research: Space Physics* 124.6, pp. 3871–3892. doi: 10.1029/2018JA026396.

Huang, Hui, Jianpeng Guo, Zehao Wang, et al. (2019). „Properties of Stream Interactions and Their Associated Shocks near 1.52 au: MAVEN Observations“. In: *The Astrophysical Journal* 879.2, p. 118. doi: 10.3847/1538-4357/ab25e9.

Hundhausen, A. J. (Jan. 1973). „Nonlinear model of high-speed solar wind streams“. In: *Journal of Geophysical Research* 78.10, p. 1528. doi: 10.1029/JA078i010p01528.

Opitz, A., R. Karrer, P. Wurz, et al. (May 2009). „Temporal Evolution of the Solar Wind Bulk Velocity at Solar Minimum by Correlating the STEREO A and B PLASTIC Measurements“. In: *Solar Physics* 256.1-2, pp. 365–377. doi: 10.1007/s11207-008-9304-7.

Richardson, I. (Apr. 2004). „Energetic Particles and Corotating Interaction Regions in the Solar Wind“. In: *Space Science Reviews* 111, pp. 267–376. doi: 10.1023/B:SPAC.0000032689.