



On developing a physics-based Heliospheric modeling system: Background Solar Wind

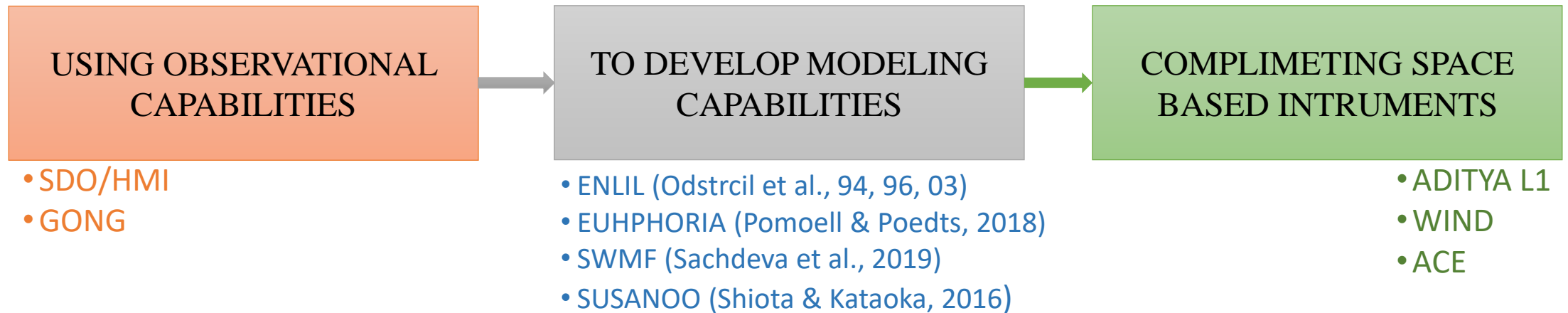
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Current Scenario

Requirement: *to quantify the physical parameters of the corona to get insight into solar wind outflow.*

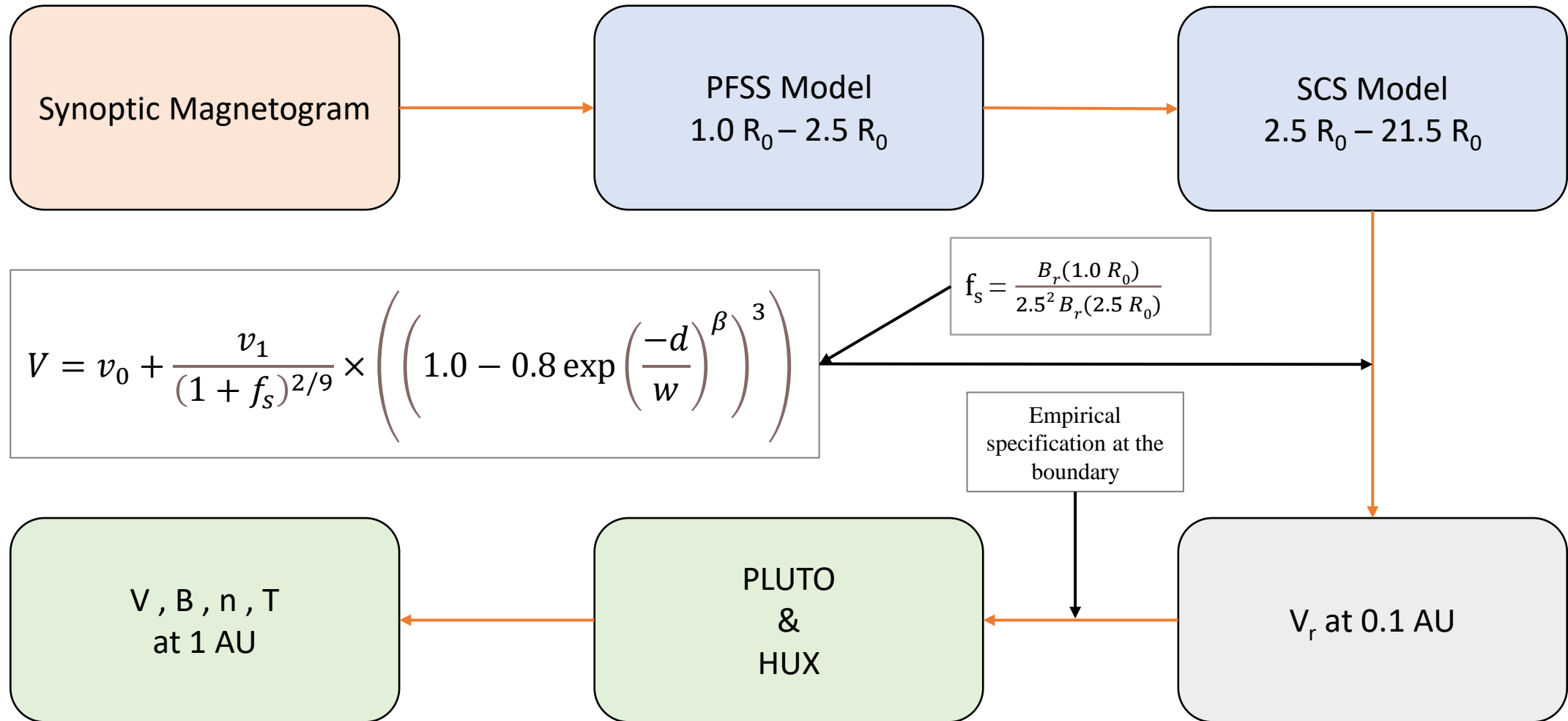
Problem: *measurement of coronal magnetic field still poses many challenges and at present there's no reliable approach to measure coronal field.*



FROM THE VIEWPOINT OF HELIOSPHERIC WORK

- Data-driven models for the global Solar surface-coronal field
- Data-driven ensemble models of the Solar Wind including magnetic field
- Develop the ability to incorporate the Coronal Mass Ejection in the model and measure its arrival time

Model Flowchart



WSA Model

$$V = v_0 + \frac{v_1}{(1 + f_s)^{2/9}} \times \left(\left(1.0 - 0.8 \exp\left(\frac{-d}{w}\right)^\beta \right)^3 \right)$$

- McGregor et. al. 2011 :

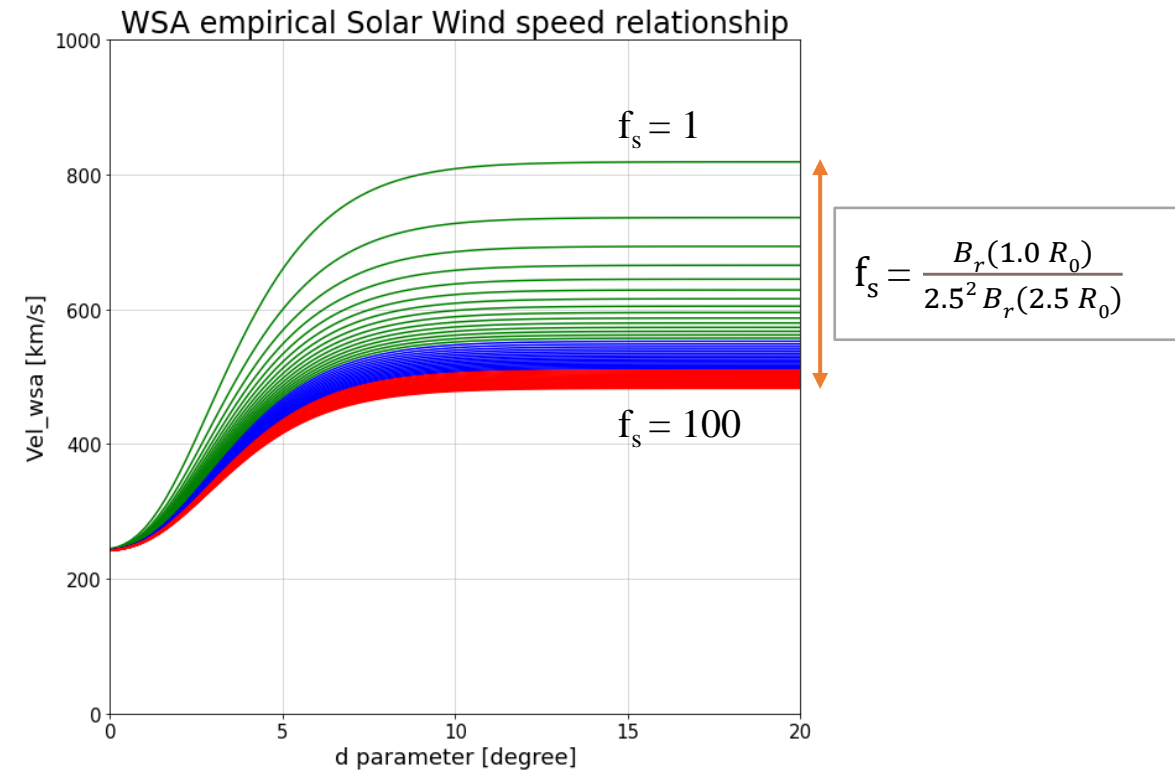
$$v_0 = 240 \text{ km/s}$$

$$v_1 = 675 \text{ km/s}$$

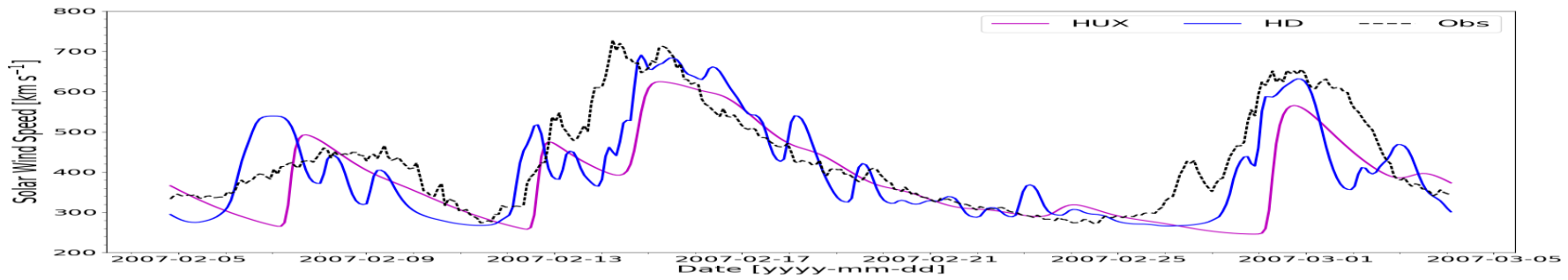
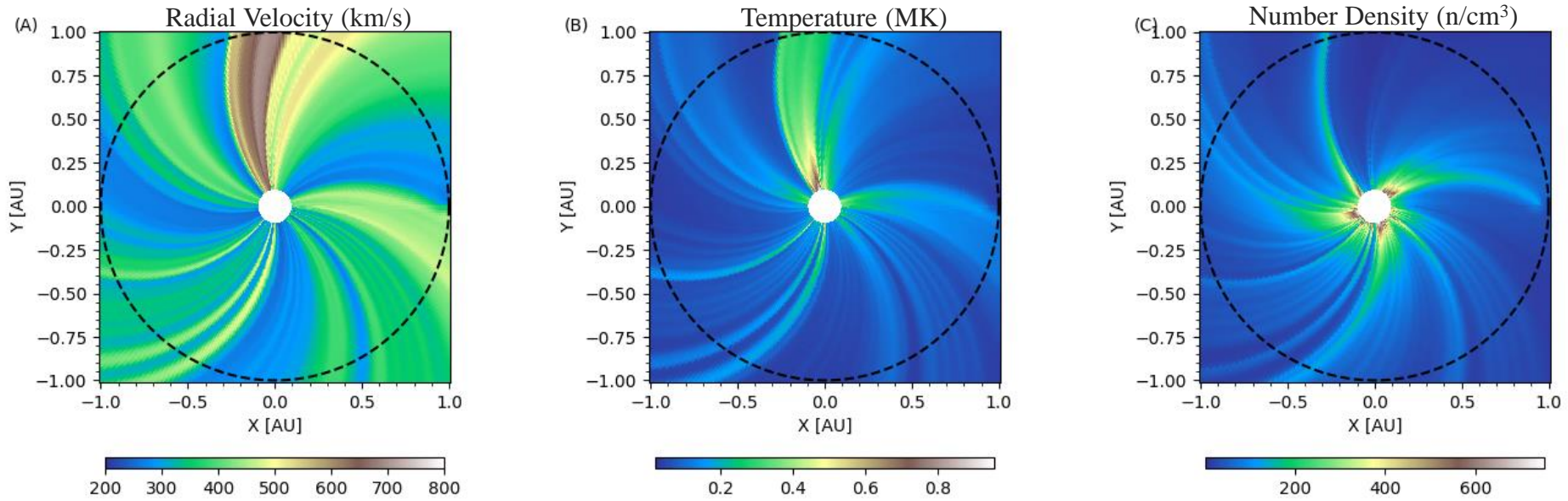
$$\beta = 1.25$$

$$w = 2.8 \text{ degree}$$

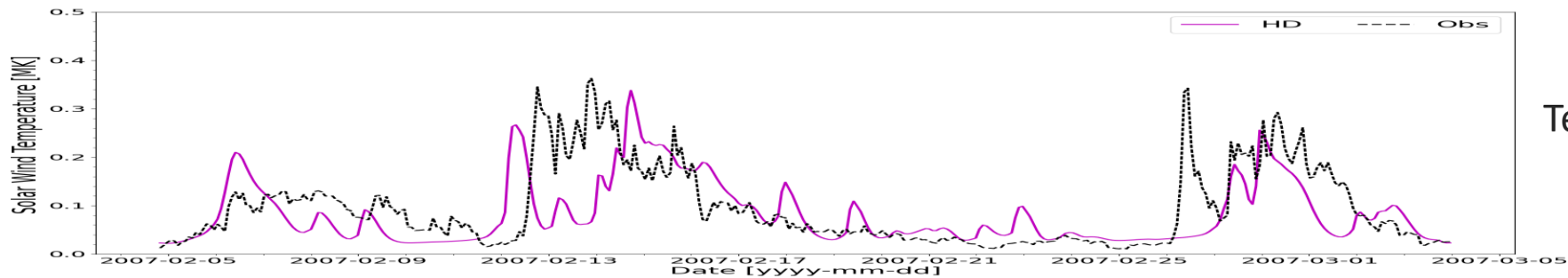
- We modified the parameter 'w' and replaced it with the *median of parameter 'd'*



Model Outputs



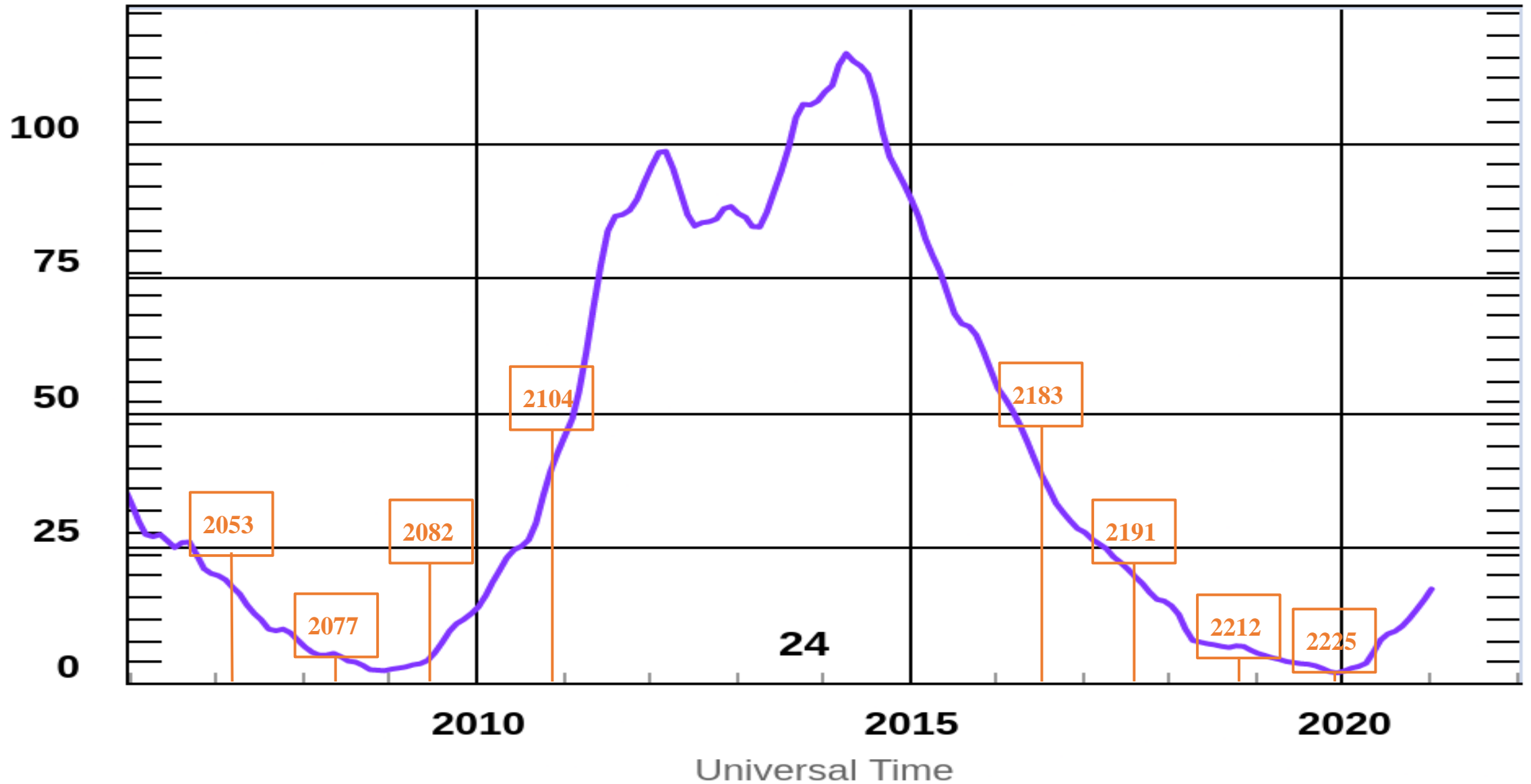
Speed Profile



Temperature Profile

CR 2053

Sample CRs

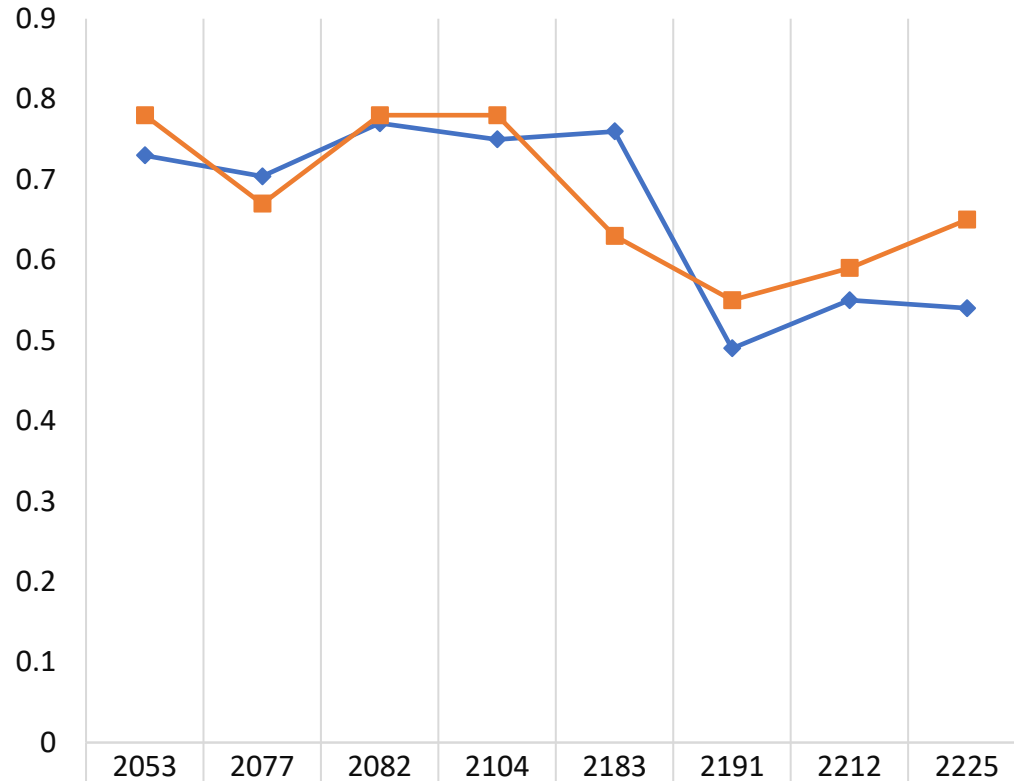


Statistics

Solar Wind Speed (km/s)

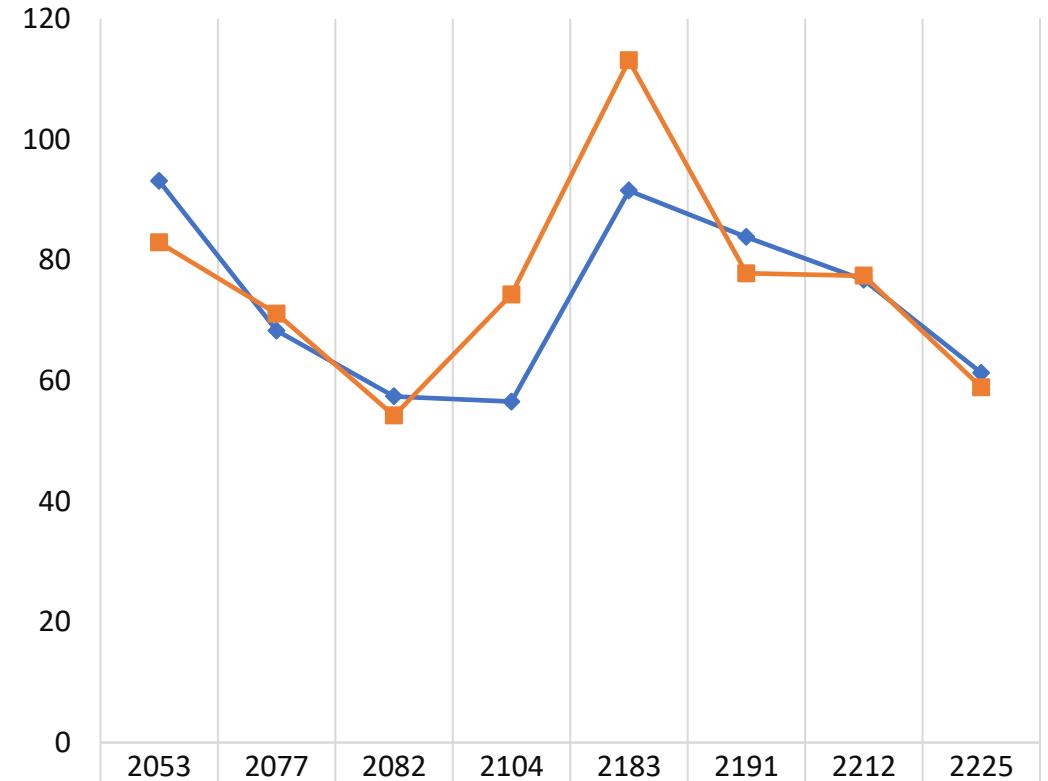
PEARSON CC

—◆— HUX —■— PLUTO



RMSE

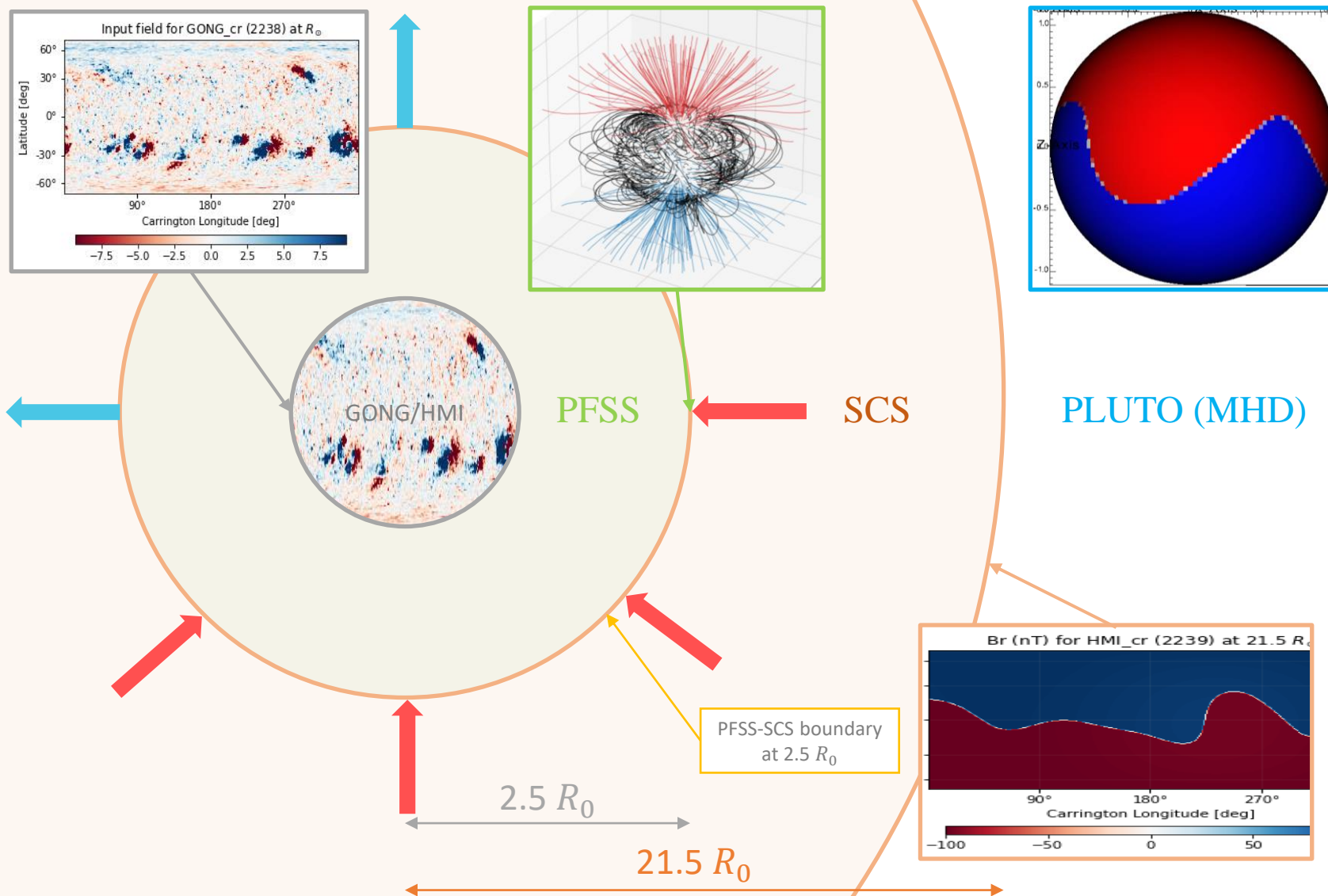
—◆— HUX —■— PLUTO



—◆— HUX	0.73	0.704	0.77	0.75	0.76	0.49	0.55	0.54
—■— PLUTO	0.78	0.67	0.78	0.78	0.63	0.55	0.59	0.65

—◆— HUX	93.1	68.3	57.4	56.5	91.5	83.8	76.7	61.3
—■— PLUTO	82.9	71.1	54.2	74.3	113.1	77.8	77.4	58.9

Summary



SOLAR WIND at L1

- Velocity (ASPEX)
- Magnetic Field (MAG)
- Temperature (ASPEX)
- Density (PAPA)

