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From highly-collisional to collisionless fluid models

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The atmosphere of the sun represents a complex physical environment where two traditional classes of fluid theories come to interplay. While the solar chromosphere is typically modeled with fluid models derived under the highly-collisional assumption, whereas the collisional frequencies in the solar corona can become so small, that collisionless fluid models seem to be more appropriate. We briefly overview these two distinct classes of fluid theories and discuss the major differences between them. We also discuss some preliminary results from a 3rd unifying class of arbitrary-collisional models that are suitable for any regime of collisionality.

Student poster?

Primary authors: Dr HUNANA, Peter (Instituto de Astrofísica de Canarias (IAC)); KHOMENKO, Elena (Instituto de Astrofísica de Canarias); COLLADOS, Manuel (Instituto de Astrofísica de Canarias)

Presenter: Dr HUNANA, Peter (Instituto de Astrofísica de Canarias (IAC))

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