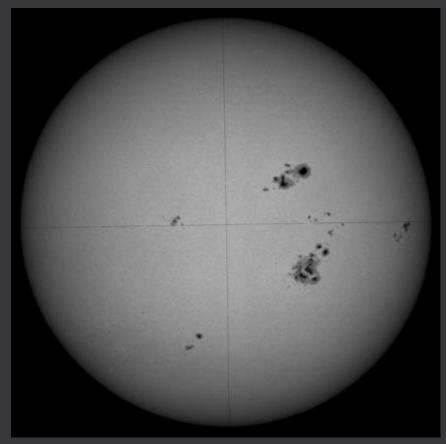


Skolkovo Institute of Science and Technology









Kanzelhöhe Observatory, Austria

Prediction of 11-year solar cycle strength with Hemispheric Sunspot Numbers

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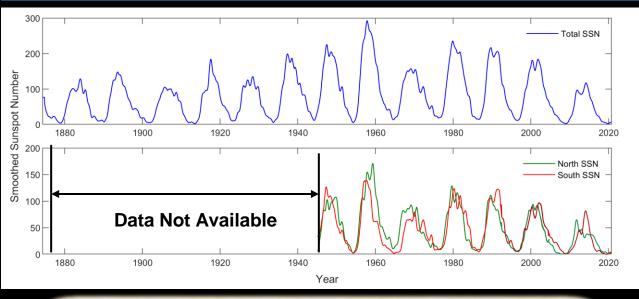
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(4) Hvar Observatory, University of Zagreb, Croatia

(5) World Data Center SILSO, Royal Observatory of Belgium

EPSM-16, 6-10 September 2021

Existing Data of Hemispheric Sunspot Number



We use the Hemispheric Sunspot Areas (available from 1874 onwards) to reconstruct Hemispheric Sunspot Number

How?

Take Relative Fraction of Sunspot Area and recalibrate with Total Sunspot Number (version 2.0)

$$HSN_{north} = \frac{N_{area}}{N_{area} + S_{area}} * TSN_{2.0}$$

$$HSN_{south} = \frac{N_{area}}{N_{area} + S_{area}} * TSN_{2.0}$$

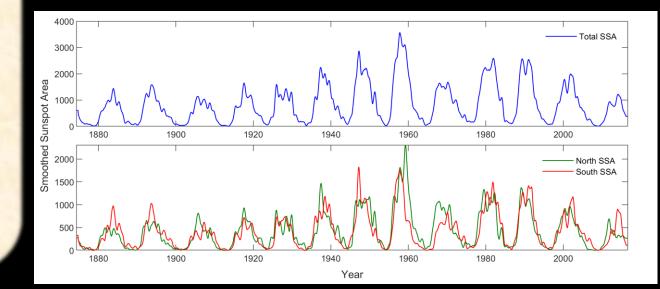
 $HSN_{north} + HSN_{south} = TSN_{2.0}$

Sunspot Number

Hemispheric Sunspot Numbers are available only from 1945 onwards !!

- 1945-2004: <u>Temmer et. al.(2006)</u> -1992-2020: SILSO World Data Centre

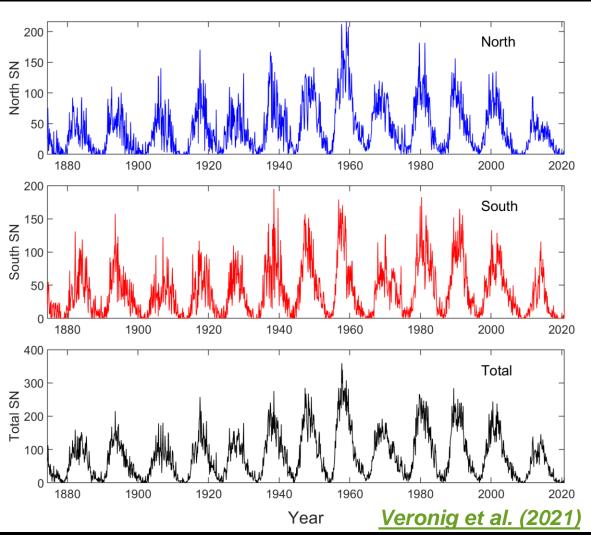
Sunspot Area



Hemispheric Sunspot Number 1874-2020

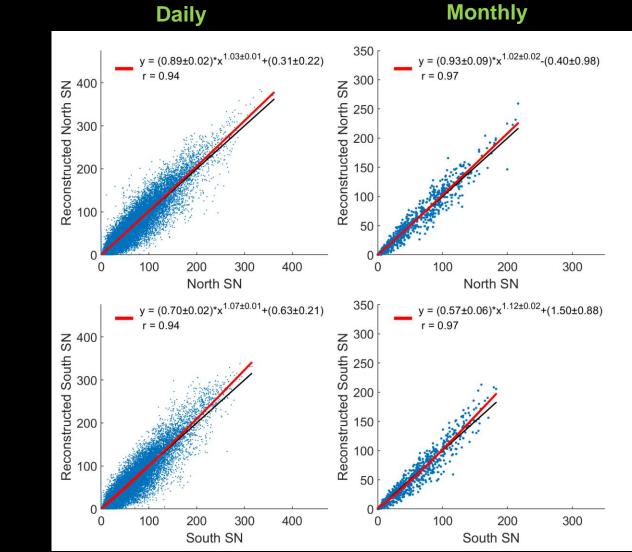
1874-1944 : Reconstructed from Sunspot Areas **1945-1991:** Derived from <u>Temmer et al (2006)</u> **1992-2020:** Taken from World Data Centre SILSO

Monthly Series (1874-2020)



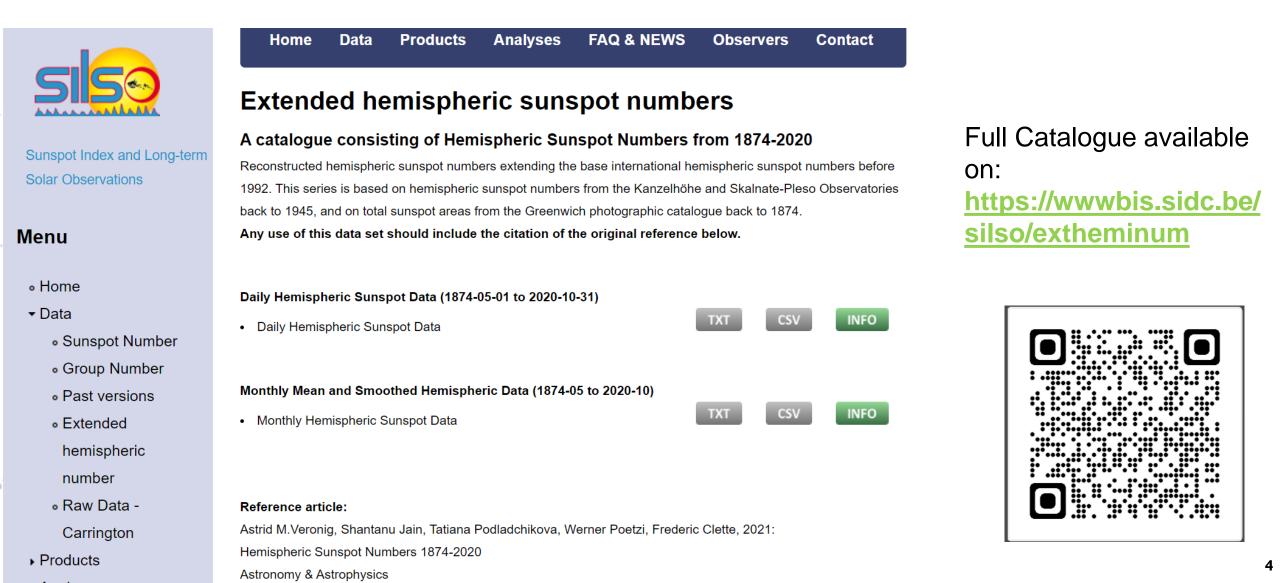
Validation

Reconstructed HSN to the existing HSN for the common time period (1945-2016)



Online Catalogue

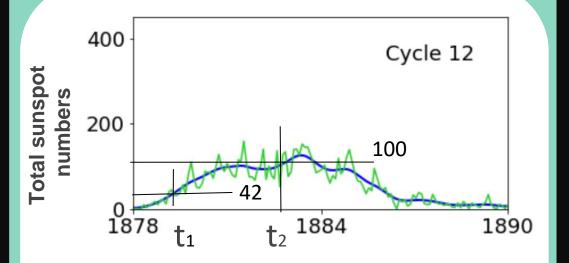
The Catalogue of Hemispheric Sunspot Numbers is already **embedded in SILSO**, which is the World Data Center for the production, preservation, and dissemination of the international sunspot number



Growth Rate Vs Cycle Amplitude

Existing approach:

Relation between <u>average</u> growth rate and cycle peak

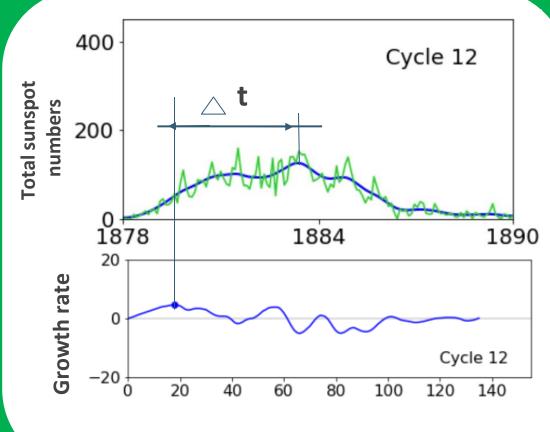


The average growth rate = $(100-42)/t_2 - t_1$

Cameron and Schüssler, 2008

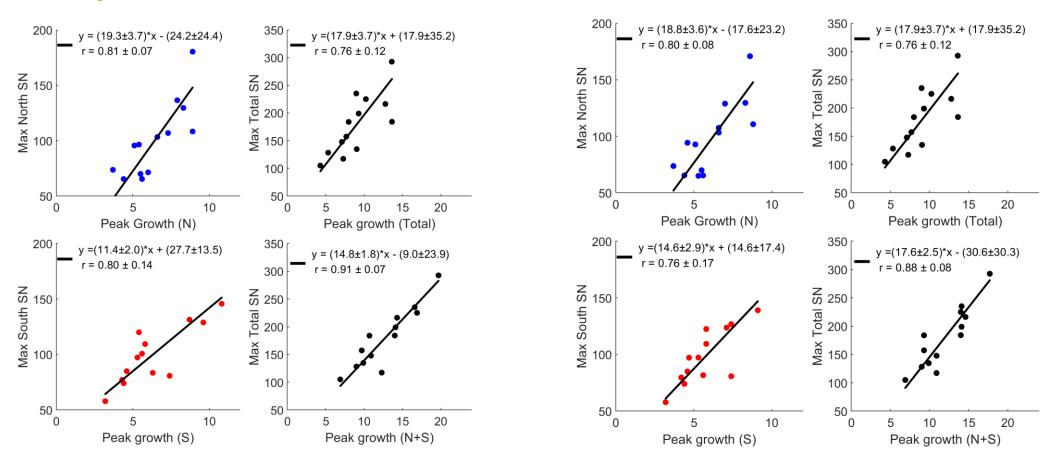
Our approach:

Analyze relation between <u>maximal</u> growth rate and cycle peak



 \triangle t≈ 21 months (on average)

Peak growth rate and relation to solar cycle amplitudes



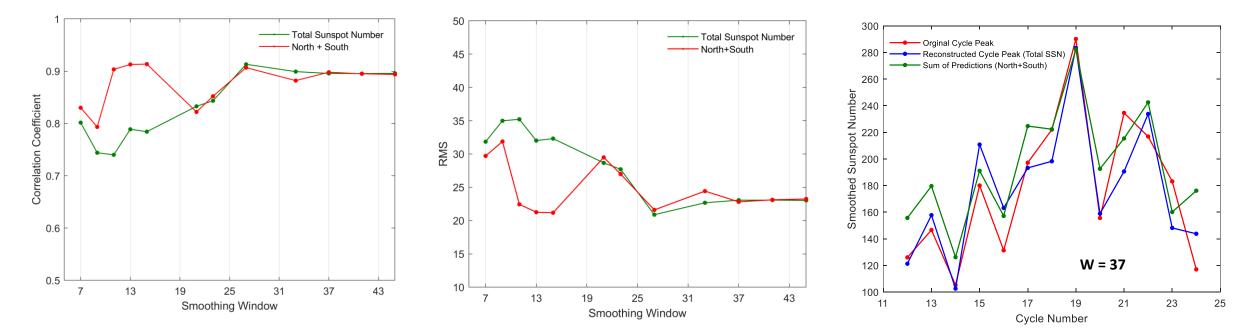
Purely Reconstructed from Area

Merged Dataset

- Peak Growth of hemispheric activity (N+S) provides higher correlation with cycle peak than than the peak growth derived from the total sunspot numbers characterizing the full Sun
- Supporting the importance of the regular monitoring, record, and analysis of solar activity separately for the two hemispheres for space weather science and predictions

PREDICTION OF CYCLE STRENGTH USING TOTAL AND HEMISPHERIC SUNSPOT NUMBERS(MERGED DATA)

RED: RMS/Correlation coefficient between true and predicted cycle peaks using Hemispheric Sunspot Number GREEN: RMS/Correlation coefficient between true and predicted cycle peaks using Total Sunspot Number

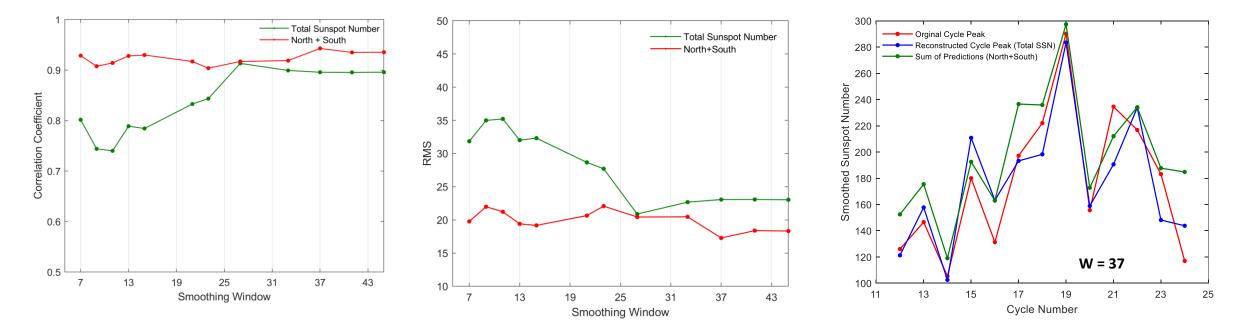


Smoothing Windows : 7,9,11,13,15,21,27,33,37,41,45



PREDICTION OF CYCLE STRENGTH USING TOTAL AND HEMISPHERIC SUNSPOT NUMBERS (PURELY RECONSTRUCTED FROM SUNSPOT AREAS)

RED: RMS/Correlation coefficient between true and predicted cycle peaks using **Hemispheric Sunspot Number** GREEN: RMS/Correlation coefficient between true and predicted cycle peaks using **Total Sunspot Number**



Smoothing Windows : 7,9,11,13,15,21,27,33,37,41,45



Conclusions and Outcomes

1. A new data product – Catalogue of Hemispheric Sunspot Numbers is developed. It is embedded in to World Data Center SILSO

Link : https://wwwbis.sidc.be/silso/extheminum

- 2. We demonstrated that the peak of growth rate has a higher correlation to cycle peak (>0.80) than mean growth rate (0.60-0.65) for all considered smoothing windows.
- The peak of growth rate from HSN purely reconstructed from Sunspot Area (1874-2016) has a higher correlation (0.90 0.95) than combined HSN (~0.80 0.90) and Total SN (~0.75 0.89) for all considered smoothing windows.
- 4. We showed that the solar cycle strength predictions can be done more accurately with the newly developed catalogue (r = 0.90 to 0.95) with respect to total sunspot numbers characterizing the full Sun (r = 0.75 to 0.89) depending on the smoothing window. The average improvement in RMS, Correlation Coefficient is 9%, 5% for merged HSN data set and 27%, 11% for HSN purely reconstructed from Areas compared to Total Sunspot Numbers. A new and important result for space weather predictions and necessity of regular monitoring and record of hemispheric sunspot numbers

