Rosseland Centre for Solar Physics

A Neural Network Approach to 3D NLTE Radiative Transfer

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Why 3D non-LTE Radiative Transfer? Mg II k3



3D

CS

R



Leenaarts et al. (2012, ApJ 749)

Leenaarts et al. (2013, ApJ 772)







Bjørgen & Leenaarts (2017, A&A 599)

Full 3D NLTE calculations very expensive. Many iterations, hundreds of thousands of CPU-hours.

3D effects can be important in the cores of strong lines, 1.5D (column-by-column) radiative transfer is not enough.



We consider not just 1 to 1 pixel mapping, but input is 3D, so we surround the pixel of interest by neighbouring columns, to account for inclined rays. Window size can be adjusted.



LTE atmosphere

"True" result is taken from a run of Multi3D

Input is LTE populations, output NLTE populations

CS R

SunnyNet

Goal is to learn final spectra from simulations, but SunnyNet learns mapping from LTE to NLTE populations of a given model atom. This allows later computation of any spectral line in the model, for any viewing angle (in CRD).









R C S

Learn "true" result from Multi3D calculations

LTE pops —> NLTE pops

Addition

From NLTE populations we then calculate the synthetic spectra.





Prediction

"True" result









Enhanced Network simulation

Flaring simulation

Prediction



"True" result

$H\alpha$ core



Flaring simulation



R

Training and testing from different snapshots of **same** simulation

- Excellent results based on simple network structure
- GPU-based, PyTorch, speeds up 3D NLTE by 10⁵

R

CS

Training and testing using **different** simulations ("out of sample")

• Open source, available at <u>https://github.com/bruce-chappell/SunnyNet</u>

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