Debriefing from IPMU Rebriefing in Rome (Alvio Renzini)

They will all come on line after (several) years of JWST operations will have transformed the field

- Their advantages over JWST
- Higher spatial resolution in imaging and slit & IFU spectroscopy
- Access to shorter optical wavelengths
- Higher spectral resolution
- Wider FoV (and multiplex?)
- Lifetime
- Did I forget something?

Their advantages over JWST

- Higher spatial resolution in imaging and slit & IFU spectroscopy
- Access to shorter optical wavelengths
- Higher spectral resolution
- Wider FoV (and multiplex?)
- Lifetime
- Did I forget something?

- The problems we wish they will solve (if JWST did not):
- First galaxies and reionization
- Galaxy-SMBH coevolution
- Baryon cycle (in and out of galaxies)
- Relative weights to the various SF quenching channels
- Giant colorful posters of high-z galaxies with ~30pc resolution (3D) ~1000 resolution elements per kpc²
- Did I forget something?

Three telescopes, each with limited instrumentation (2-3 each). Competition with a bit of collaboration? Best/exclusive capabilities of each of them, e.g.:

- Widest FoV → GMT
- Northern Hemisphere → TMT
- Highest spatial resolution → ELT
- UV response → TMT
- Largest collecting area → ELT
- Grassroots intercontinental collaborations to orchestrate their usage for problem-solving observations.
- Did I forget something? Yes

Three telescopes, each with limited instrumentation (2-3 each). Competition with a bit of collaboration? Best/exclusive capabilities of each of them, e.g.:

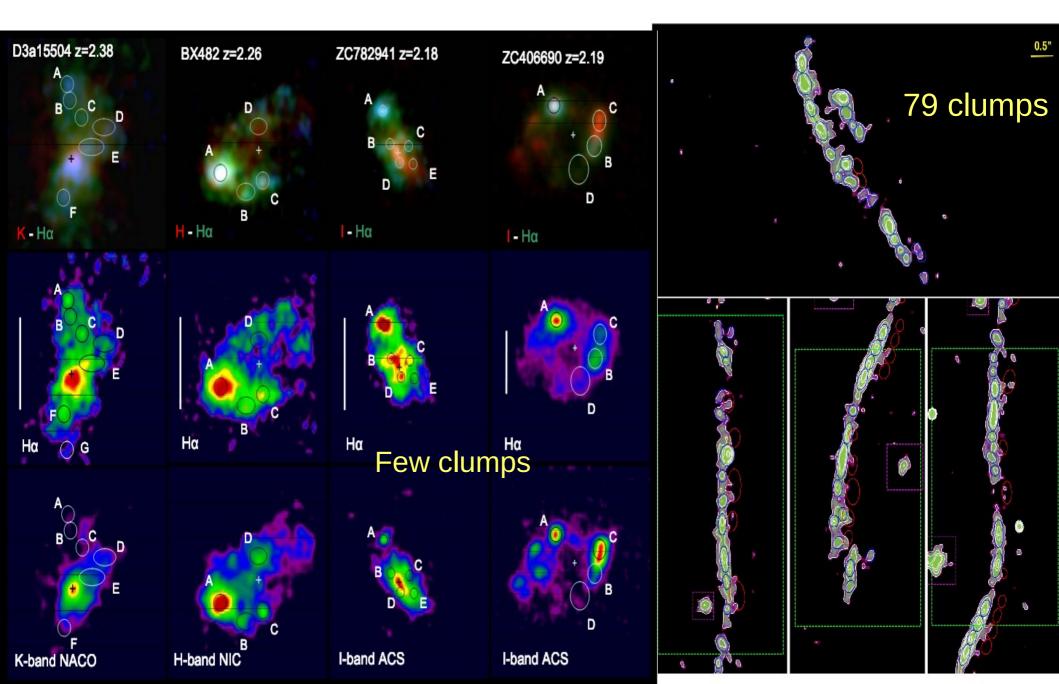
- Widest FoV → GMT
- Northern Hemisphere → TMT
- Highest spatial resolution → ELT
- UV response → TMT
- Largest collecting area→ ELT
- Grassroots intercontinental collaborations to orchestrate their usage for problem-solving observations.
- Did I forget something? Yes to some extent these telescopes will transform us! i.e., our way of carrying on Large Programs.

Science requirements:

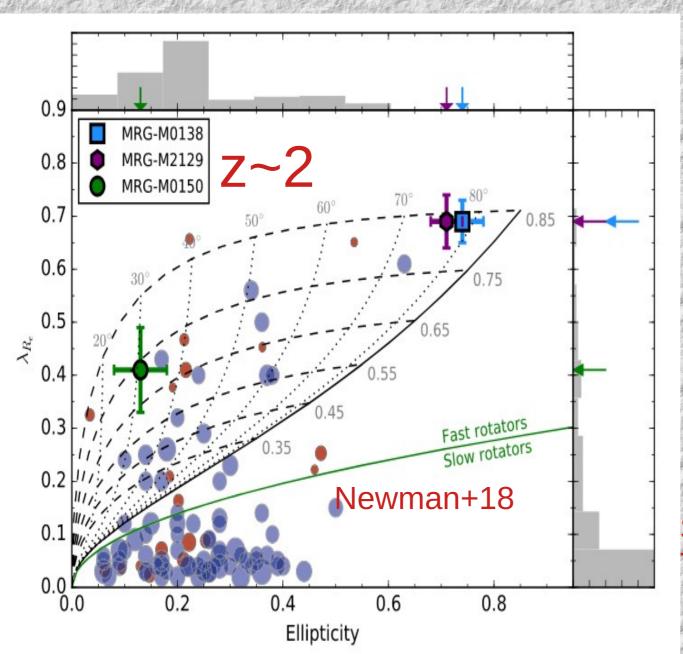
- Needs for high spatial & spectroscopic resolution + IFU on single targets seem finely met by first generation instruments
- Needs for high multiplex spectroscopy not met quite as well ...
- "Instruments need to have high multiplex in order to exploit the higher target densities" (KG Lee)
- "Difficult to have high multiplex and multi-IFUs in the same instrument" (M Akiyama?)
- Did I forget something?

What is now possible only thanks to magnification and stretching by Gravitational Lensing will become possible for unlensed high redshift galaxies: two examples

~1kpc, Genzel+11Spatial Resolution~30pc, Cava+18



No Morpho Change at Quenching (!)



Space-resolved 1D and 3D spectroscopy of freshly-quenched, high redshift galaxies

See also Toft+17

