GLaD
Gravitational Lensing and Dynamics
combined analysis to unveil properties of
high-redshift galaxies.

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MPA Garching

Extremely Big Eyes on the Early Universe

Rome, September 10th, 2019
Early Type Galaxies (ETGs)

Early Type Galaxies can be broadly characterised by:

- old population
- red colours
- small amount of gas and dust
- lack of spiral arms

Fast Rotators

- less massive
- outer disk
- distributed like spiral galaxies
- evolution is gas-accretion driven

Slow Rotators

- dominate at high mass end
- weakly triaxial, no disk
- in high density environments
- evolution dry-merger driven

Emsellem et al. 2007, 2011

How did they evolve to be the most massive slow rotators galaxies today?
Early Type Galaxies (ETGs)

Evolution of mass-size relation early-type galaxies (red) through cosmic time shows rapid size evolution. Star formation cannot account for this growth in the past 10 Gyr.
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Kinematical measurements are a key tool to study the dynamical properties of galaxies and provide strong constraints on galaxy formation and evolution.

Van Dokkum et al. 2010
Van der Wel et al. 2014
low signal-to-noise ratios and low spatial resolution

observations are limited to bright central regions (bulge).


How can this be overcome?
Gravitational lensing

Distorts and **MAGNIFIES** light from background source
GLaD Method

- Lensing and dynamics analysis are performed \textit{simultaneously}.

- We compare the dynamical model of the source kinematics on the image plane.

- Use \textit{parametrised profiles} for both source and lens (do not rely on any source reconstruction).

- Trace gravitational potential using stellar kinematics (no gas).
COMPARE

to

DATA
RX J1131−1231

Used for studies on:
- quasars and region around BH
- time delay cosmography
- dark matter substructures
- planet searches with microlensing

$z_d = 0.295$
$z_s = 0.654$

HST/ACS F814W filter
**Mock Lensing Data: MOCK 1**

**LENS:**
- MASS: power law + external shear
- LIGHT: Sersic profile

**SOURCE:**
- MASS: isothermal profile
- LIGHT: Sersic profile

Adjust amplitudes and effective radii to mimic HST observations of the SB distribution and the S/N.
Mock Kinematic Data: MOCK 1

Source Plane

Lens Plane

unlensed source kinematics

lensed source kinematics

Chirivì et al. in prep
Mock Kinematic Data: MOCK 1

19 bins

This is a very optimistic case, we will probably not be able to measure anything on the right hand side in a realistic scenario.
Models

We perform the following tests:

- consistency test
- lensing analysis vs. lensing + dynamics
- systematics for lens and source mass profiles
Mass-Follows-Light Profile: model prediction

Test **systematic errors** on **SOURCE MASS**, by modelling it with a mass-follows-light (MFL) profile.

Assuming a **constant** mass-to-light ratio.
Mass-Follows-Light Profile: model prediction

INPUT ISOTHERMAL PROFILE

\( \chi^2_{\text{red}} = 42. \)
Mass-Follows-Light Profile: model prediction

$\chi^2_{\text{red}} = 42$. 

PRELIMINARY

Chirivì et al. in prep

25
Mock Lensing Data: COMPOSITE MOCK

LENS:

MASS: NFW+Chameleon x M/L+external shear
LIGHT: Sersic profile

SOURCE:

MASS: isothermal profile
LIGHT: Sersic profile

Adjusted amplitudes and effective radii to mimic HST observations of the SB distribution and the S/N
Mock Kinematic Data: COMPOSITE MOCK

22 bins

Chirivì et al. in prep
We find that the **lensing analysis gives a good fit, and the dynamics analysis doesn’t**. This means that the addition of the dynamics analysis gives a huge contribution to distinguishing between models.
Results:

- The combined lensing and dynamics analysis:
  - significantly **tightens the constraints** on the source and lens mass parameters (by up to a factor 3-20).
  - allows us to **constrain the source mass parameters**, which are otherwise unconstrained, and reconstruct the **source total mass profile**.
  - allows us to **break degeneracies between lens mass parameters and source kinematic properties**.
  - helps to **better discern between models**.

- Thanks to **instruments** on future telescopes that would provide improved data quality **we would be able to study the source galaxies in even more detail**, and simultaneously tighten the constraints on the lens galaxy mass.