

The Host Galaxies of Fast Radio Bursts with ASKAP and the VLT



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***On behalf of the CRAFT
collaboration***

Overview



-
- What is a Fast Radio Burst (FRB)?
 - Circumstantial localisation: FRB 171020.
 - The first ASKAP localisation: FRB 180924.
 - Probing a galaxy halo with FRB 181112.
 - Do FRBs have SN-like optical counterparts?
 - Recap

What is an FRB?



-
- *Fast*: typically < 10 milliseconds.
 - *Radio*: detected at 0.5–10 GHz, no counterparts at other wavelengths.
 - *Bursts*: (almost) never repeat, flux densities 10^{10} x greater than pulsars.
 - Rate of $10^3 - 10^4 \text{ sky}^{-1} \text{ day}^{-1}$.
 - Occur at cosmological distances.

The “Repeater”



-
- FRB 121102 (Spitler et al. 2014) displays sporadic repeat bursts with same DM \Rightarrow not cataclysmic.
 - Enabled localisation to 0.01" with JVLA & EVN.
 - Host is a low-metallicity, star-forming dwarf $r \sim 25$ mag galaxy at $z=0.2$ (Tendulkar et al. 2017).
 - FRB associated with a persistent, compact radio source and star-forming knot \Rightarrow magnetar?

“Flys eye” FRB survey



Shannon et al. (2018, Nature) found 23 FRBs in 14 months during ASKAP commissioning.



FRB 171020



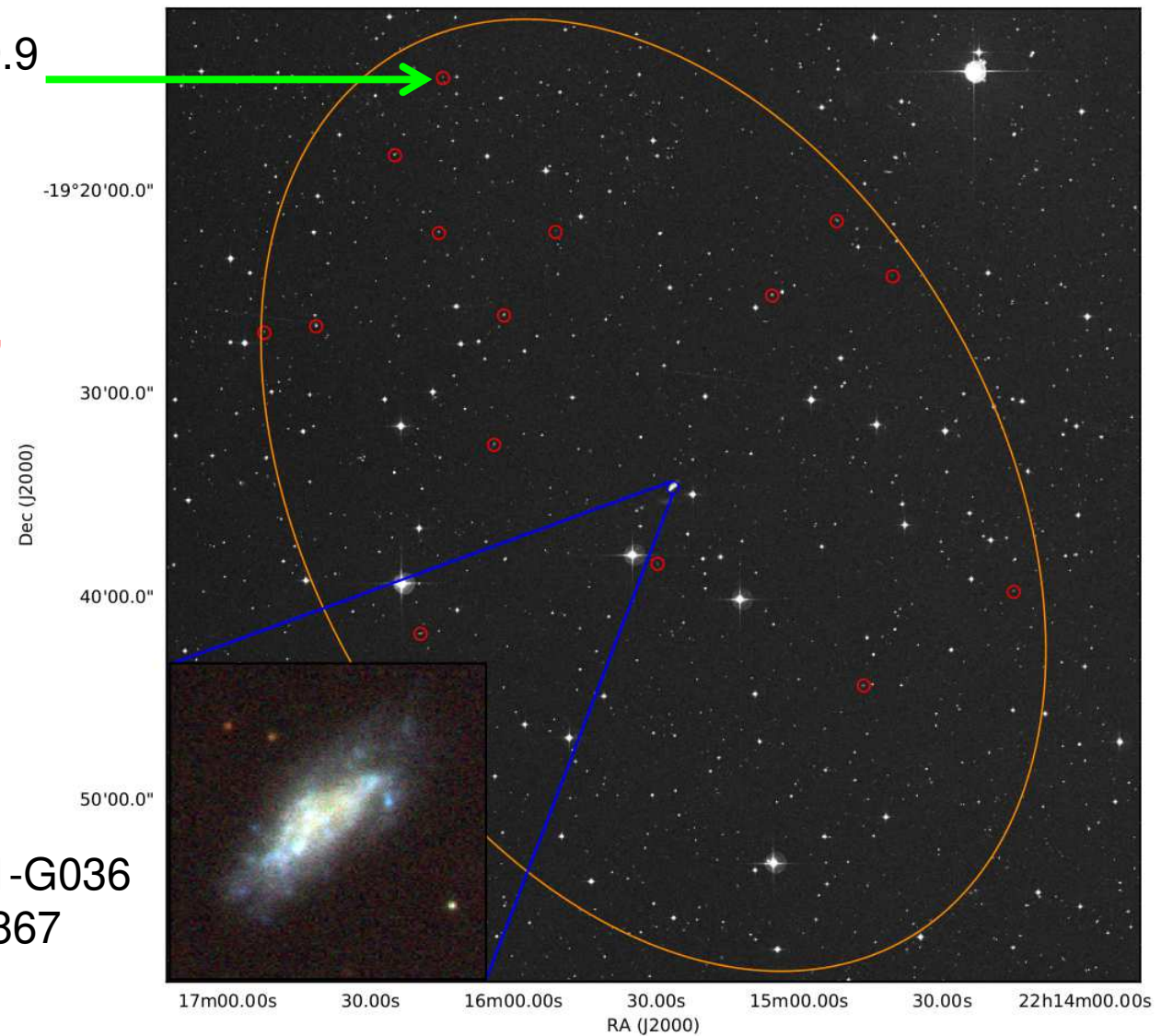
-
- FRB 171020 had $DM = 114 \text{ pc cm}^{-3}$, so expect $z < 0.08$.
 - Spectroscopy with VLT/X-shooter all but eliminated WISE/SuperCOSMOS sources as host.
 - Most likely host is ESO 601-G036, an Sc galaxy at $z=0.009$, with signs of tidal interaction.
 - ~ 1 mag brighter in R than FRB 121102 host + no continuum radio emission comparable to the persistent source.

FRB 171020

J221621.59-191829.9
 $z = 0.024$

Mahony et al. 2018,
ApJL, 867:L10

ESO 601-G036
 $z = 0.00867$



The search begins

Two weeks of searching in interferometric mode in Sep 2018 turned up nothing, until...

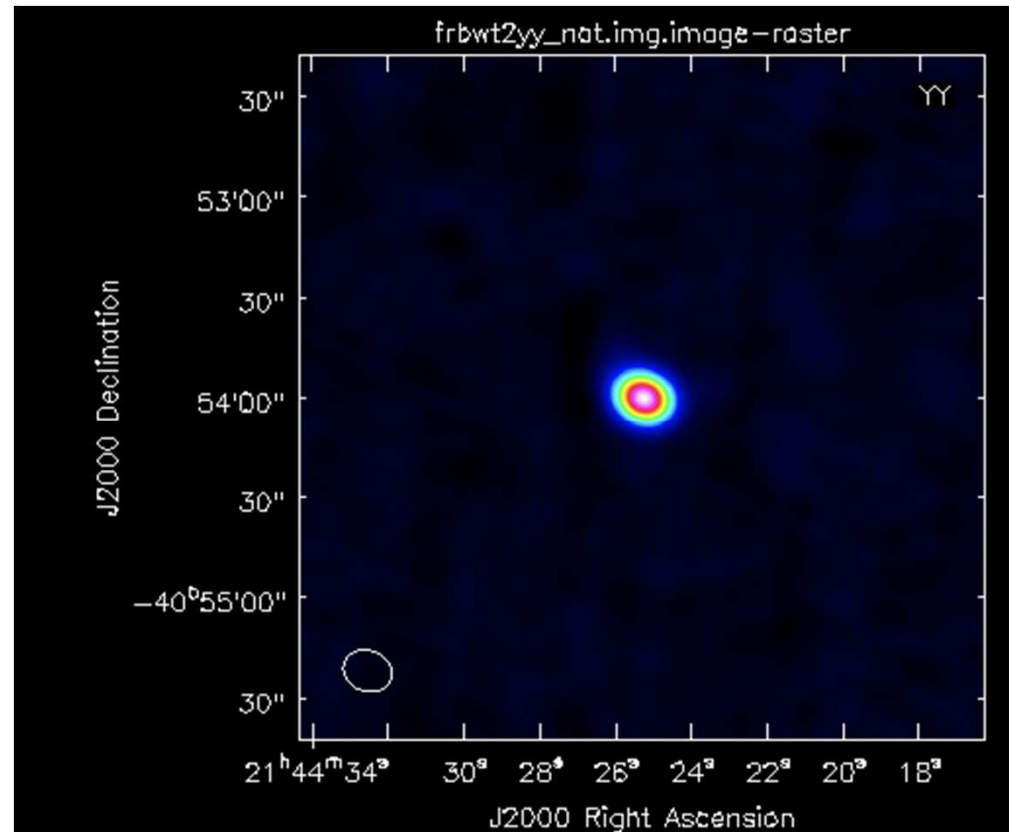


FRB 180924

RA = $21^{\text{h}} 44^{\text{m}} 25.255^{\text{s}} \pm 0.008^{\text{s}}$

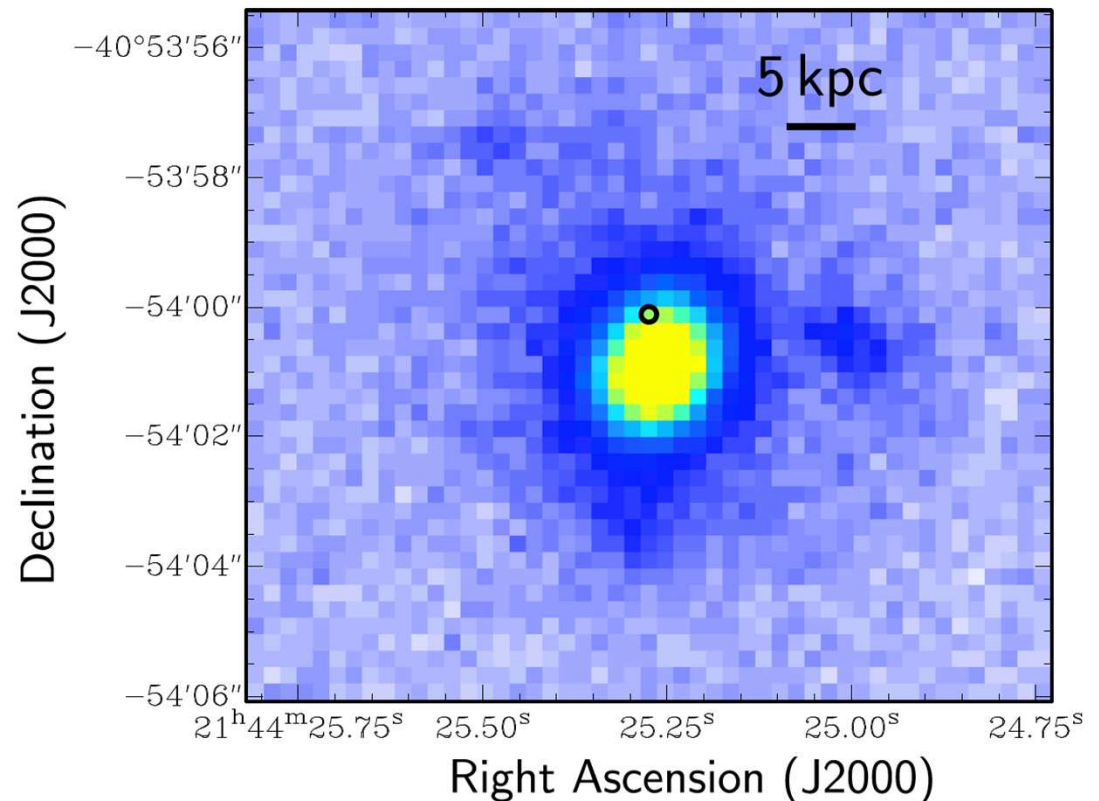
Dec = $-40^{\circ} 54' 00''.1 \pm 0''.1$

No repeat burst, or
persistent radio
counterpart
observed.



The host galaxy of FRB 180924

- VLT FORS2 *g*-band 2500s exposure.
- DES J214425.25–405400.81.
- $r = 20.54 \pm 0.02$.
- $z = 0.3214$ (KCWI)
 $\Rightarrow D \sim 1300$ Mpc.
- $M_* \sim 2.2 \times 10^{10} M_{\odot}$.
- $\text{SFR} < 2 M_{\odot} \text{ yr}^{-1}$.



Cite as: K. W. Bannister *et al.*, *Science*
10.1126/science.aaw5903 (2019).

A single fast radio burst localized to a massive galaxy at cosmological distance

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Alien signal? Cosmic mystery of 'fast radio bursts' from space baffles astronomers

ASTRONOMERS have pinpointed the location of unexplained radio signals from a distant galaxy – but could this be proof of alien life?

By TOM FISH

PUBLISHED: 18:29, Thu, Jul 4, 2019 | UPDATED: 19:35, Thu, Jul 4, 2019

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International journal of science

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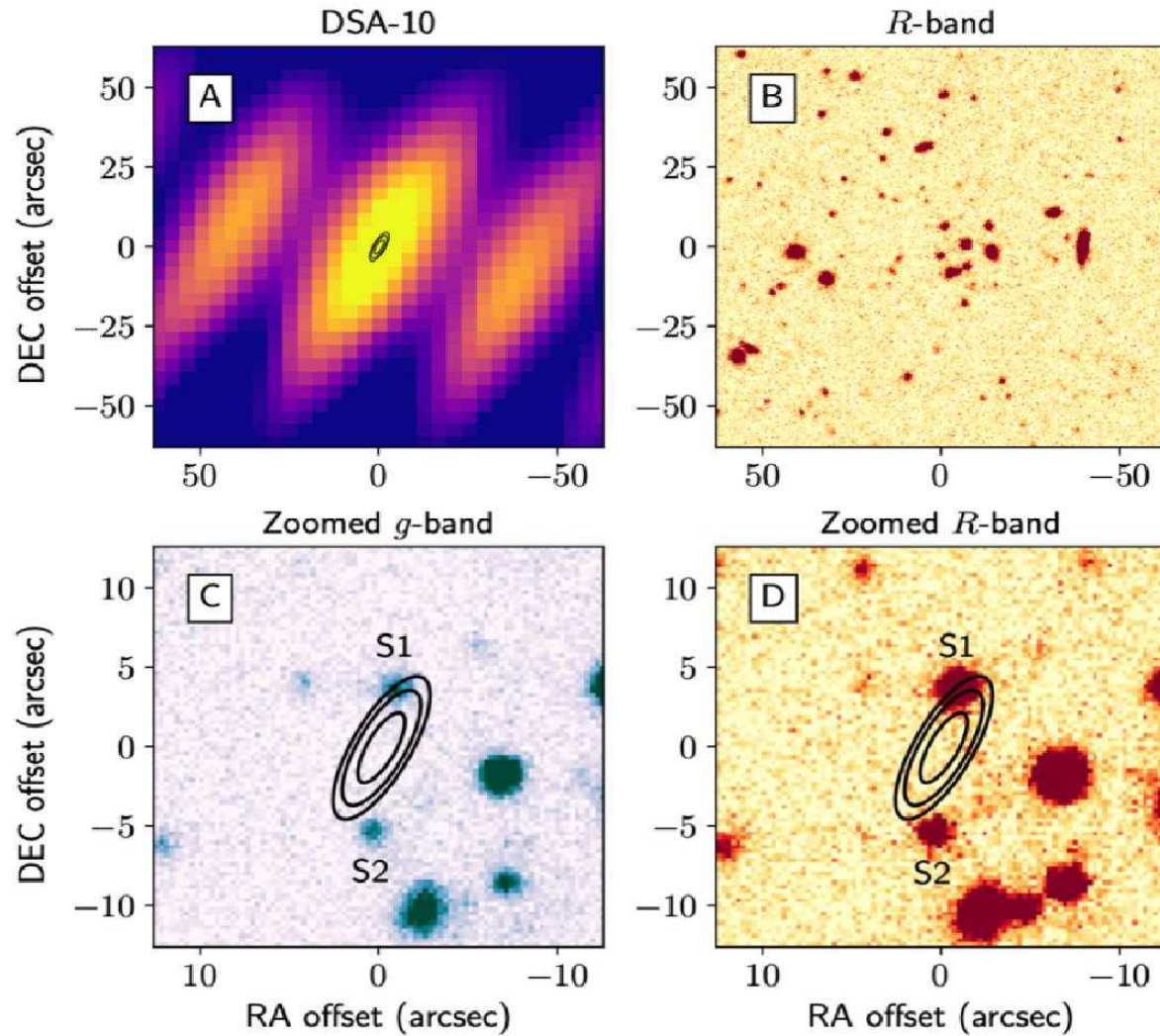
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A fast radio burst localized to a massive galaxy

V. Ravi , M. Catha, L. D Addario, S. G. Djorgovski, G. Hallinan, R. Hobbs, J. Kocz, S. R. Kulkarni, J. Shi, H. K. Vedantham, S. Weinreb & D. P. Woody

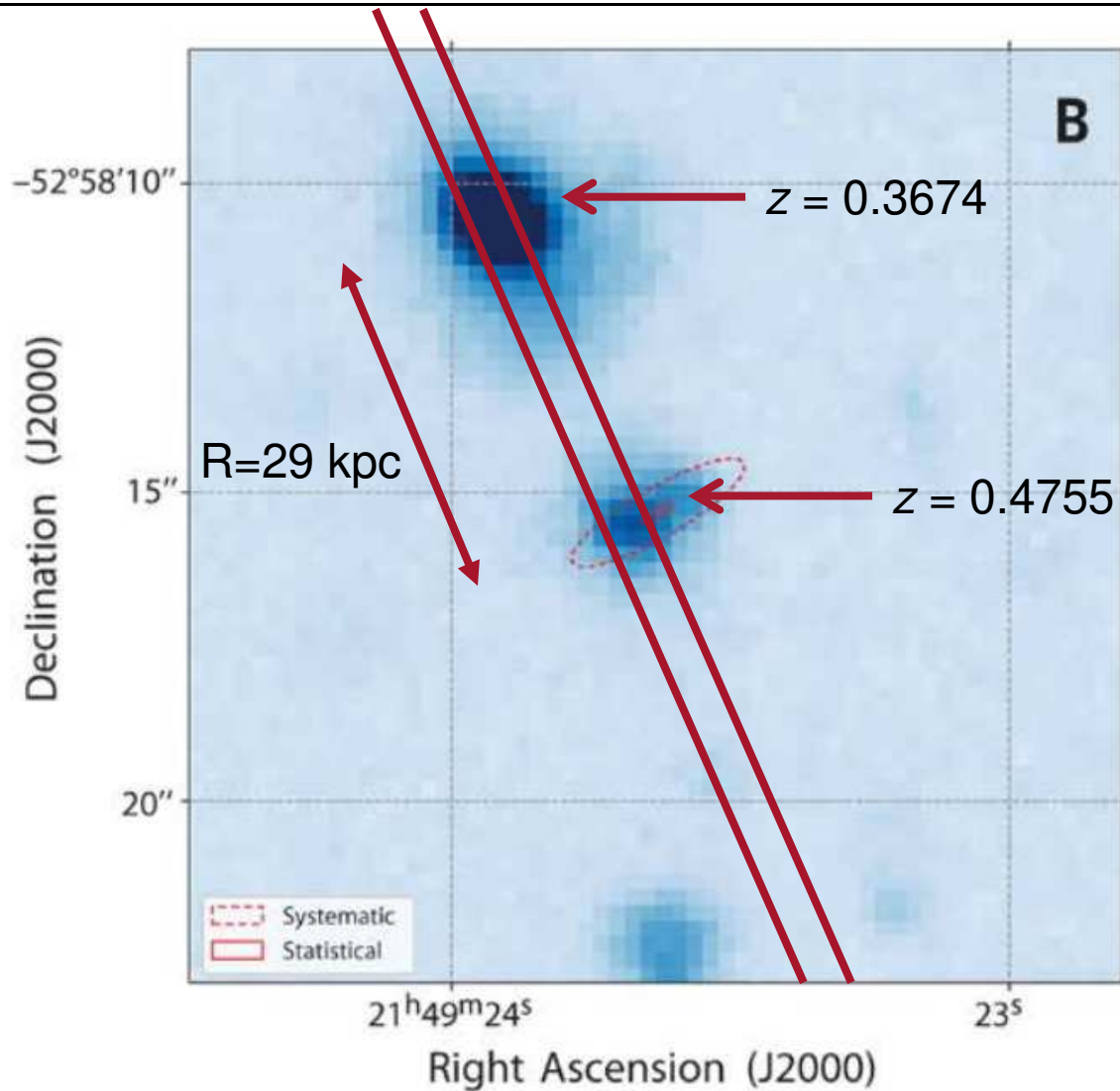
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FRB 190523

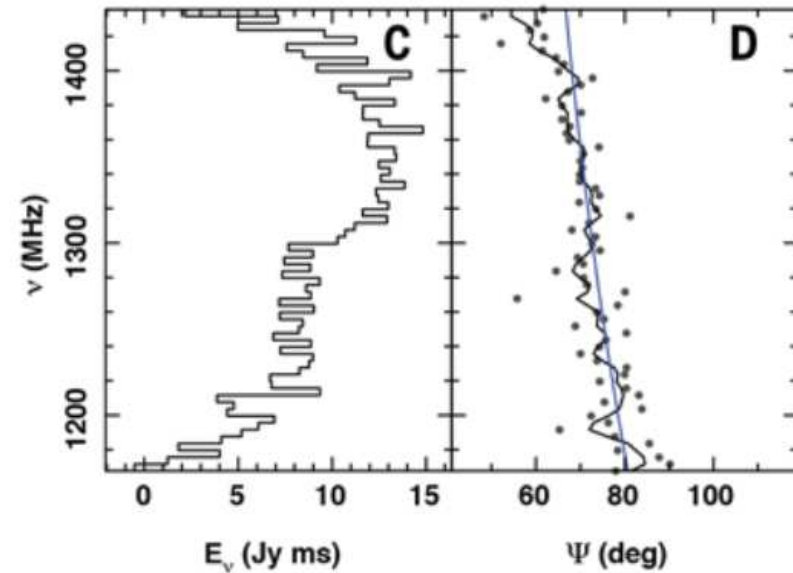
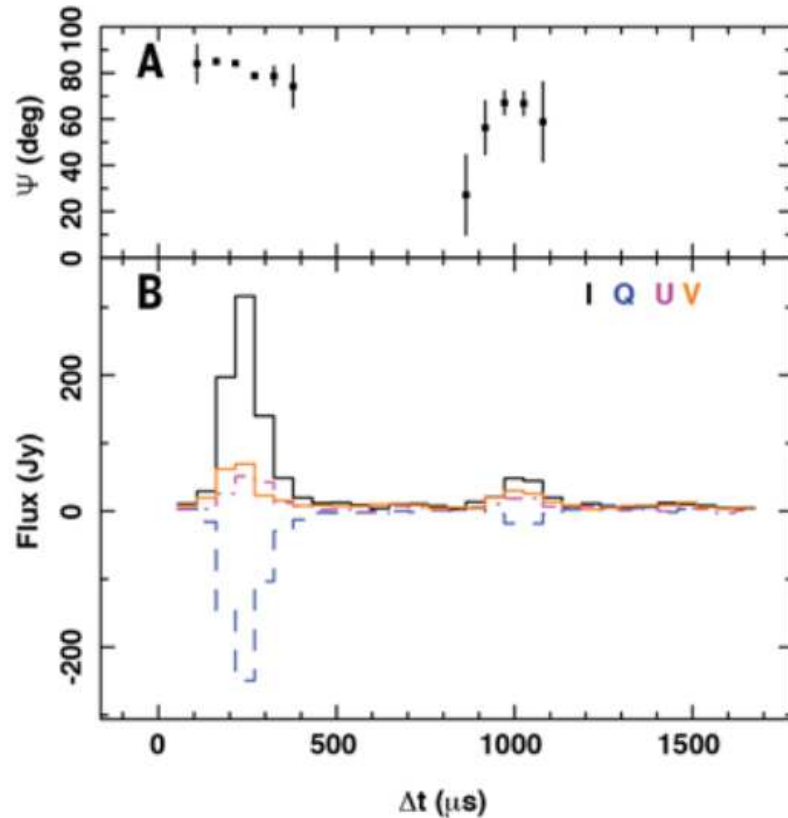


FRB 181112

- Foreground galaxy has $M^* = 10^{10.7} M_{\odot}$
- Seyfert nuclear spectrum \Rightarrow expect temporal smearing of FRB pulse due to turbulence in halo.



Halo? What halo?



$\tau_{\text{scatt}} < 40 \mu\text{sec}$ constrains $\langle n_e \rangle < \sim 10^{-3} \text{ cm}^{-3}$
(cf. $\sim 0.1 \text{ cm}^{-3}$ for pressure equilibrium with $T \sim 3 \times 10^6 \text{ K}$ halo gas) and low turbulence.

$\text{RM} = 10.9 \text{ rad m}^{-2} \Rightarrow$
 $B_{\parallel} (\text{max}) < 0.8 \mu\text{G}$
 (cf. $\sim 10 \mu\text{G}$ from RM of Mg II
 absorbers in halos of normal
 galaxies probed by QSOs –
 Bernet+08).

Cite as: J. X. Prochaska *et al.*, *Science*
10.1126/science.aay0073 (2019).

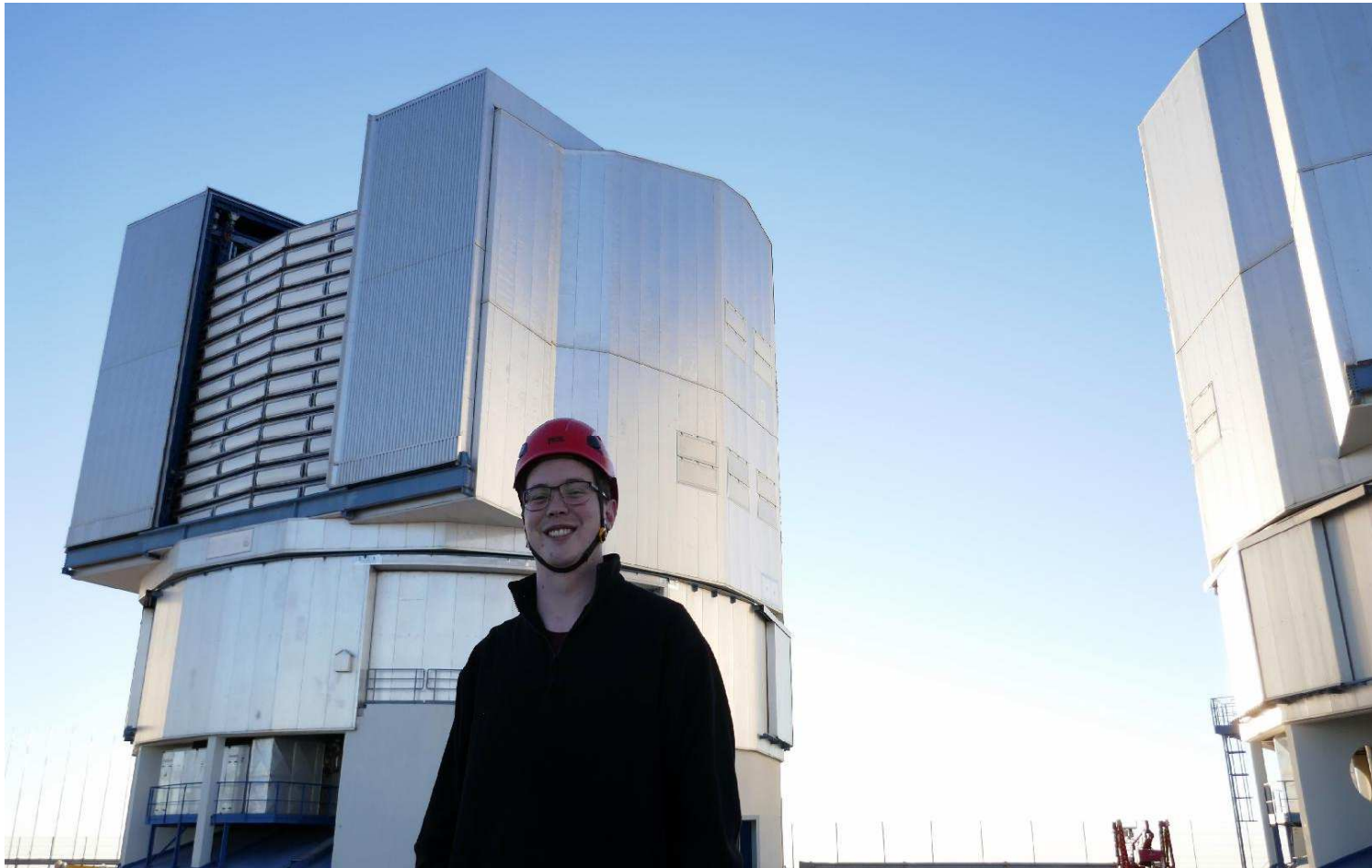
The low density and magnetization of a massive galaxy halo exposed by a fast radio burst

J. Xavier Prochaska^{1,2*}, Jean-Pierre Macquart³, Matthew McQuinn⁴, Sunil Simha¹, Ryan M. Shannon⁵, Cherie K. Day^{5,6}, Lachlan Marnoch^{6,7}, Stuart Ryder⁷, Adam Deller⁵, Keith W. Bannister⁶, Shivani Bhandari⁶, Rongmon Bordoloi⁸, John Bunton⁶, Hyerin Cho⁹, Chris Flynn⁵, Elizabeth K. Mahony⁶, Chris Phillips⁶, Hao Qiu¹⁰, Nicolas Tejos¹¹

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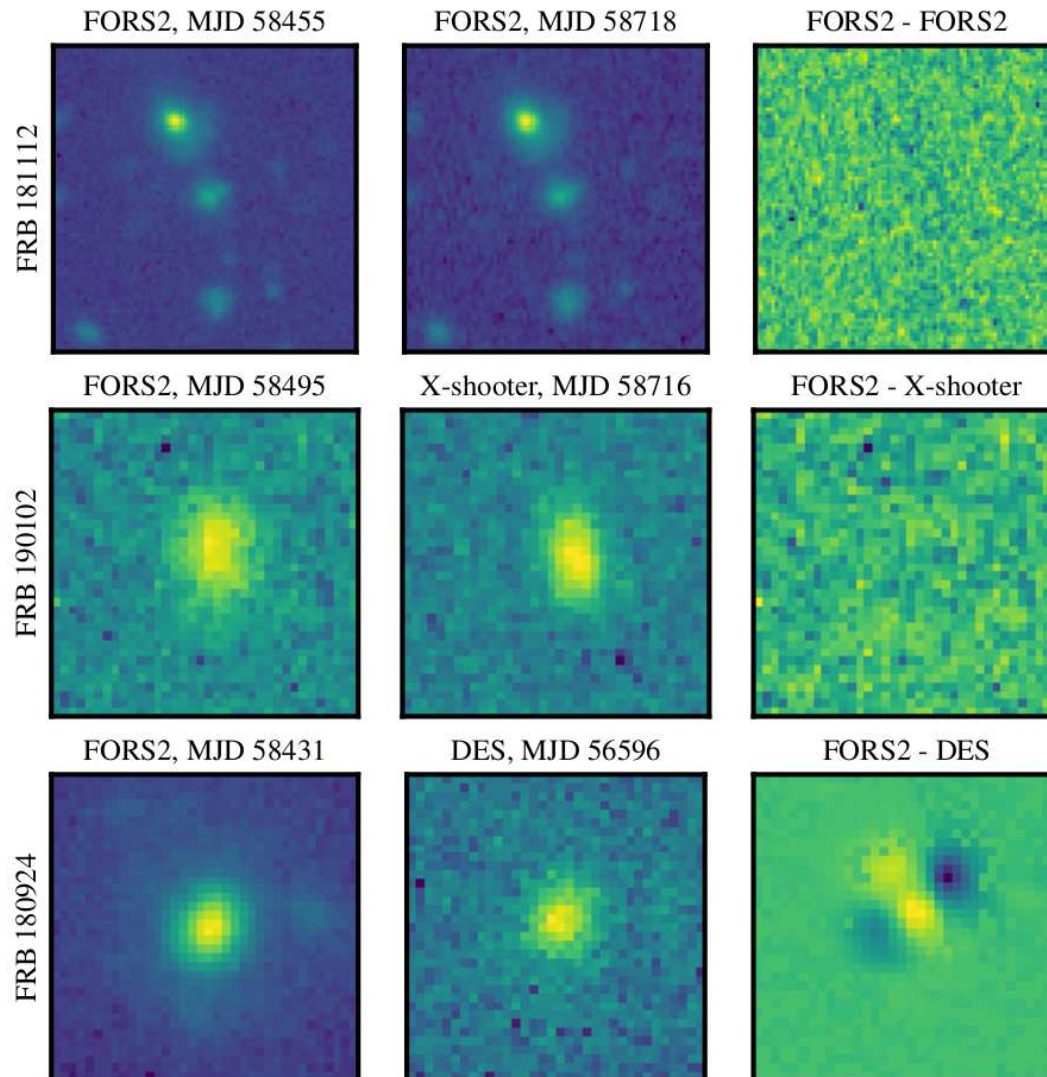
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Are FRBs associated with SN-like optical transients?

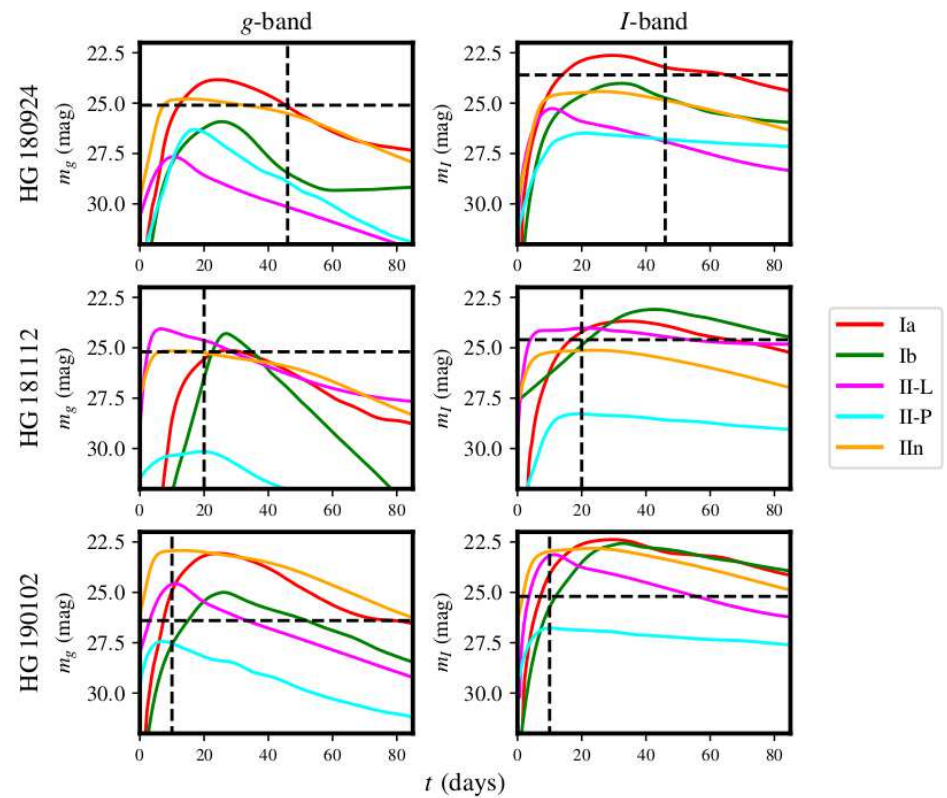
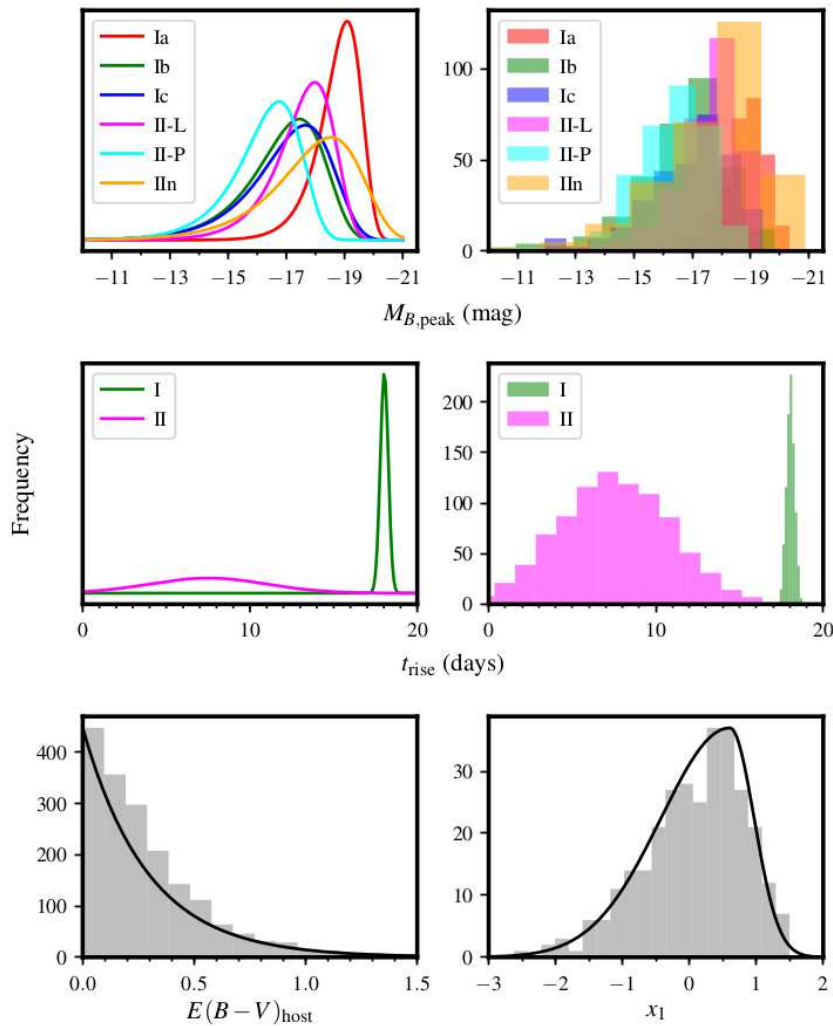


Lachlan Marnoch, MRes 2019, Macquarie University

Are FRBs associated with SN-like optical transients?



Are FRBs associated with SN-like optical transients?



Are FRBs associated with SN-like optical transients?



FRB	180924		181112		190102		Combined		190608	
Band	<i>g</i>	<i>I</i>	<i>g</i>	<i>I</i>	<i>g</i>	<i>I</i>	<i>g</i>	<i>I</i>	<i>g</i>	<i>I</i>
Limit at burst position (mag)										
FORS2 template	25.1	23.6	25.2	24.6	26.4	25.2	-	-	-	-
X-shooter template	24.7	22.7	25.9	24.0	24.3	-	-	-	22.6	22.5
Probability of non-detection					T + 10d					
Type Ia	34%	12%	44%	34%	78%	88%	12%	4%	95%	69%
Type Ib	94%	54%	91%	91%	99%	99%	85%	49%	98%	77%
Type Ic	94%	49%	88%	80%	98%	98%	82%	39%	97%	81%
Type II _n	40%	41%	34%	47%	39%	65%	6%	13%	82%	70%
Type II-L	69%	48%	78%	79%	63%	76%	34%	29%	90%	79%
Type II-P	89%	73%	75%	89%	74%	90%	50%	59%	96%	86%

Are FRBs associated with SN-like optical transients?



- Type Ia / II in all 3 hosts? Almost certainly not.
- Type Ib/c, IIL, IIP? Possible, but unlikely.
- Superluminous SNe? No.
- Kilonovae? Can't say.
- TDEs/AGN flares? No, radial offsets too large.

In summary

- Localising FRBs to/within host galaxies is shedding new light on the nature of progenitors, and the IGM.
- Galaxy halos may not be as turbulent, or magnetically supported as we assumed.
- We can already all but rule out some progenitors (Type Ia & II SNe, AGN) to FRBs.
- ASKAP + ESO/VLT are ideally suited to exploiting FRBs as cosmological tools.