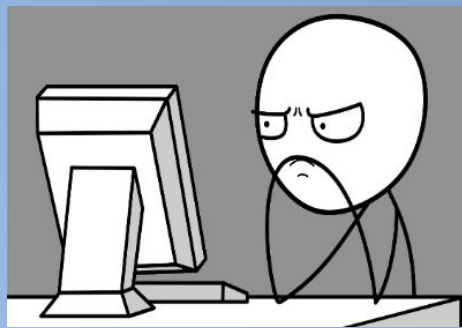


2th ASTRI Science

Tools F2F meeting

Milano, 19-20 Dicembre 2018



Punto della situazione

M. Cardillo (IAPS-INAF)

Cosa abbiamo a disposizione

- **REDMINE:** <http://astriweb.oa-roma.inaf.it/redmine/projects/astriscience tools?jump=welcome/>

Open issues in “**Astri science tools**” (<http://astriweb.oa-roma.inaf.it/redmine/projects/astriscience tools/issues>):

- 389: Analysis ON-OFF in Gammapy
- 391: Reconstructed IRF with Energy Dispersion (High Priority)
- 404: Analysis of HESS public data with CTOOLS
- 405: Issue about ASTRI Prototype tests → work in progress
- 411: Analysis and benchmarks for parallel computing with CTOOLS/Gammapy

Open Issues in “**Astri-irfs-and-event-lists**” (<http://astriweb.oa-roma.inaf.it/redmine/projects/astri-irfs-and-event-lists/issues>):

- 390: Interface between ASTRIPIPE and tools developed for IRF reconstruction
- 401: DL2b data sample for RTA tests
- 403: DL2a data sample for RTA tests

- **REPOSITORY:**

<http://astri03.oa-roma.inaf.it/gitlab/>

- **Audioconf and F2F meeting**

<https://indico.ict.inaf.it/category/104/>

TOOLS:

❖ **CTOOLS** → version 1.5.2

❖ **Gammapy**

❖ **ASTRI:**

ASTRISIM

Simulation of EVT list

ATSPEC

Generation of source spectrum

Prototype IRF

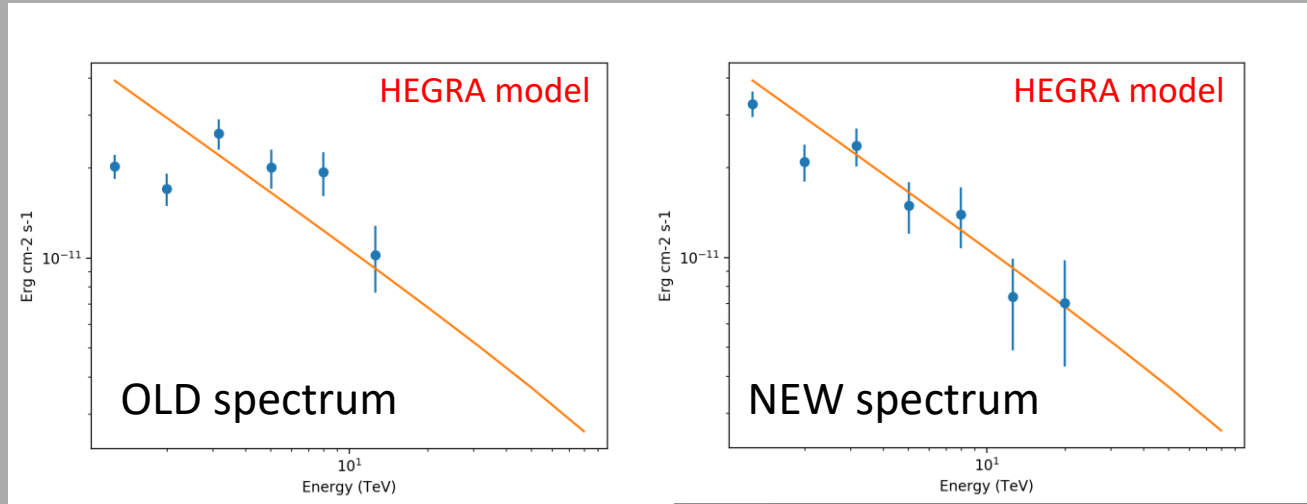
Index	Extension	Type	Dimension	View				
0	Primary	Image	0	Header	Image	Table	7 energy bins [0,6309;398] TeV 25 angular bins	
1	EFFECTIVE AREA	Binary	6 cols X 1 rows	Header	Hist	Plot	All	Select
2	POINT SPREAD FUNCTION	Binary	10 cols X 1 rows	Header	Hist	Plot	All	Select
3	ENERGY DISPERSION	Binary	7 cols X 1 rows	Header	Hist	Plot	All	Select
4	BACKGROUND	Binary	7 cols X 1 rows	Header	Hist	Plot	All	Select

- CTA compliant
- Direction cut;
- We need to fill OFF-AXIS bin ;
- Modelization of the PSF through a gaussian fit;
- Normalization of the effective area with reconstruction efficiency



Implementation in **ASTRIPIPE**

ASTRIIRF
(IRF2)
ASTRIANA
(IRF2 → IRF3)



CRAB
simulation

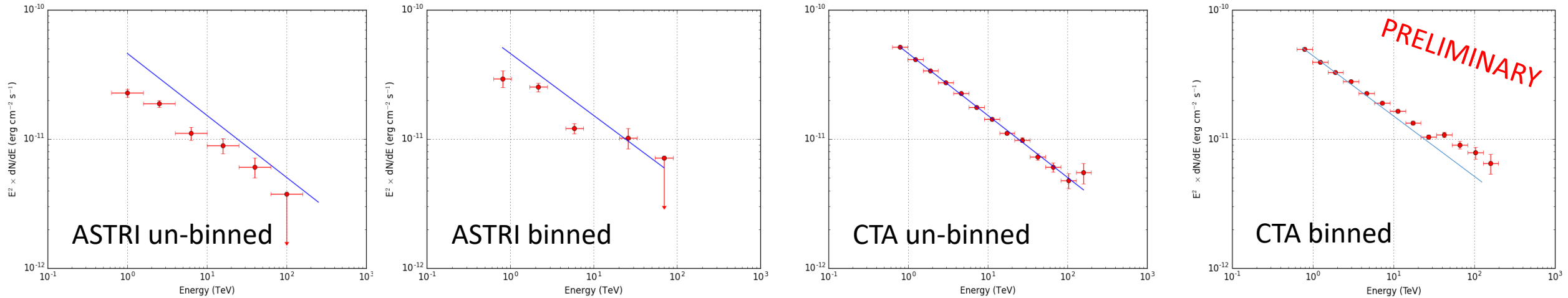
- 50h of exposure,
- E Range 0.631-398 TeV
- ROI=5.0 deg
- (no Energy dispersion)

Prototype: IRF and energy dispersion

Issue #391 REDMINE in “ASTRI Science Tools”

(<http://astriweb.oa-roma.inaf.it/redmine/issues/391>)

Using the same IRF with the same parameters
but taking into account the Edisp



- Underestimation of the flux and of the spectrum
- Test Statistic value lower than the one obtained without energy dispersion
- Likelihood binned a few better but there are however spectral problems
- In the binned analysis, we have a problem with the spectrum also without energy dispersion
- In the binned analysis, there is a problem at the highest energies also with CTA IRFs

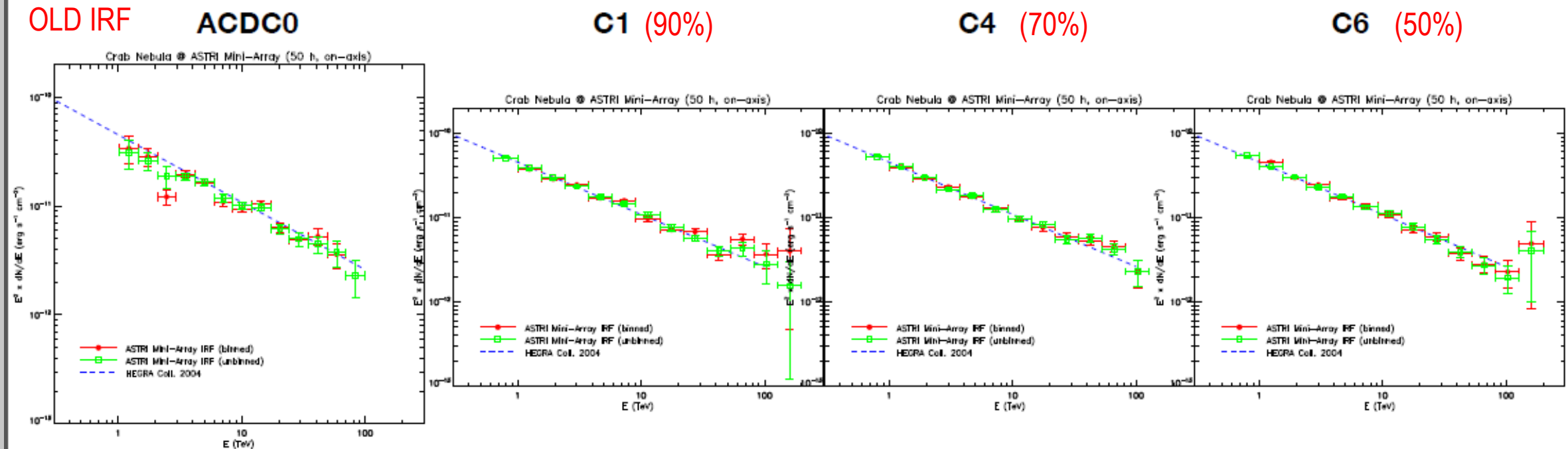
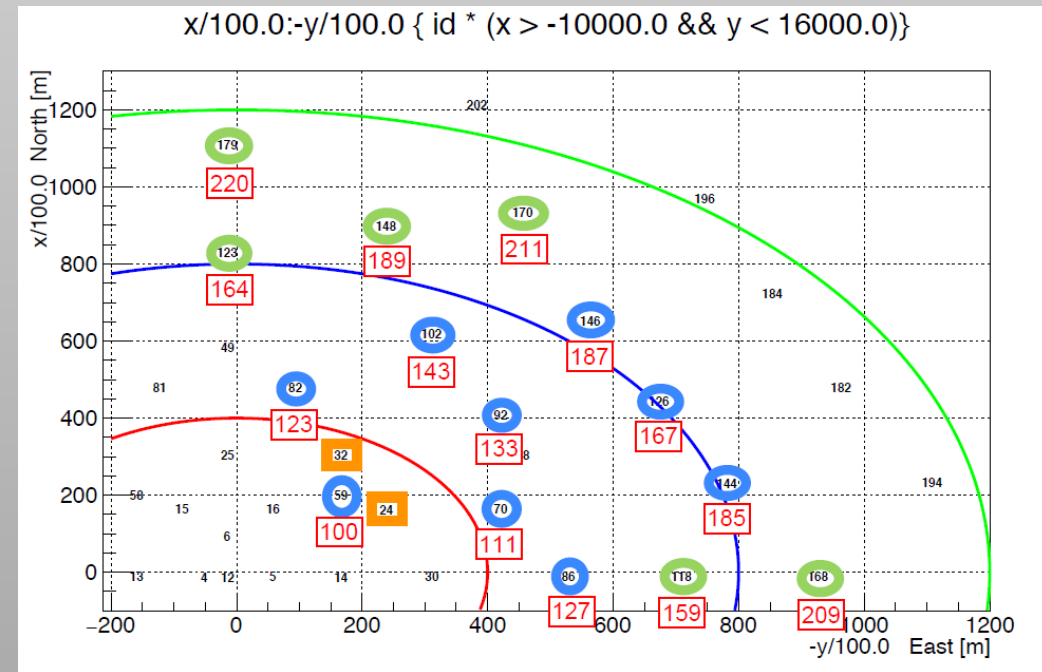
Mini-Array IRF (ACDC1)

Issue #394 REDMINE in “WP-IRFS”

(<http://astriweb.oa-roma.inaf.it/redmine/issues/394>)

- ❑ Monte-Carlo data reduced from DL0 to the IRF with ASTRIPIPE
- ❑ Realistic Layout;
- ❑ 21 energy bins: [0,0126;199,5] TeV
- ❑ 10 angular bins (OFF-AXIS present)
- ❑ 70% in gamma/hadrons separation efficiency
- ❑ Circular smearing of the background
- ❑ Extensively checked off-axis

50h - ON AXIS



Mini-Array: IRF and energy dispersion (ADCD1)

- No spectral problems taking into account the energy dispersion
 - 'Pile-up' problem due to the first point
- No problems using the CTA IRFs
- There are problems with the maximum likelihood (log-parabola):

Prefactor IN: $3.23e-17$

Prefactor OUT: $(3.888 \pm 0.037)e-17$

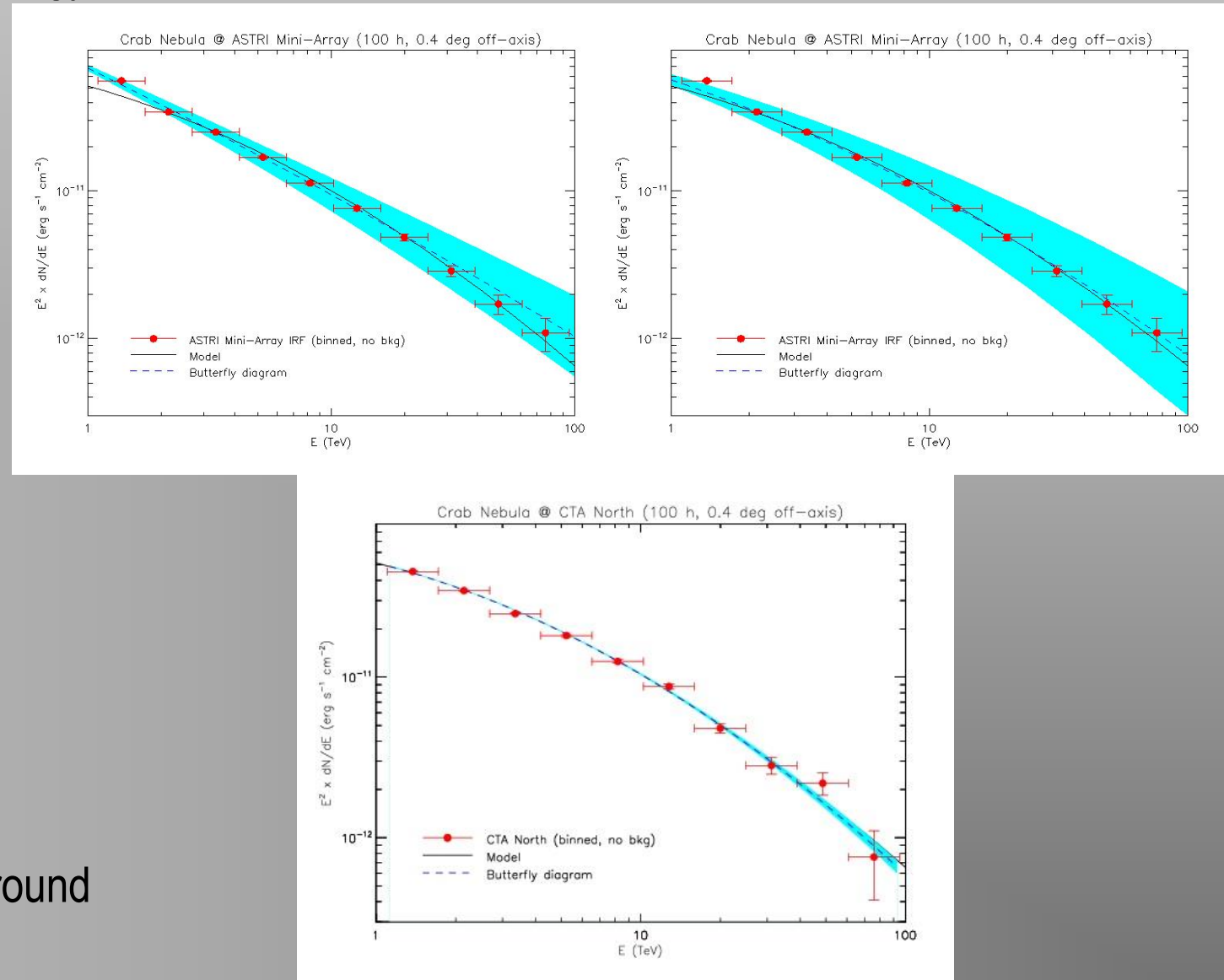
Index IN: -2.47

Index OUT: -2.789 ± 0.020

Curvature IN: -0.104

Curvature OUT: -0.0181 ± 0.0078

- The problem persists also excluding the background



Prototype Data Challenge

Issue #393 REDMINE in “ASTRI science tools” (<http://astriweb.oa-roma.inaf.it/redmine/issues/393>) → closed

FEATURES

- File EVT directly from MC simulations
- 5.8 h CRAB ON
- 80% gamma/hadrons separation efficiency

UNBINNED

EVT:

TS=70

Flux = $(6 \pm 1)e-16$

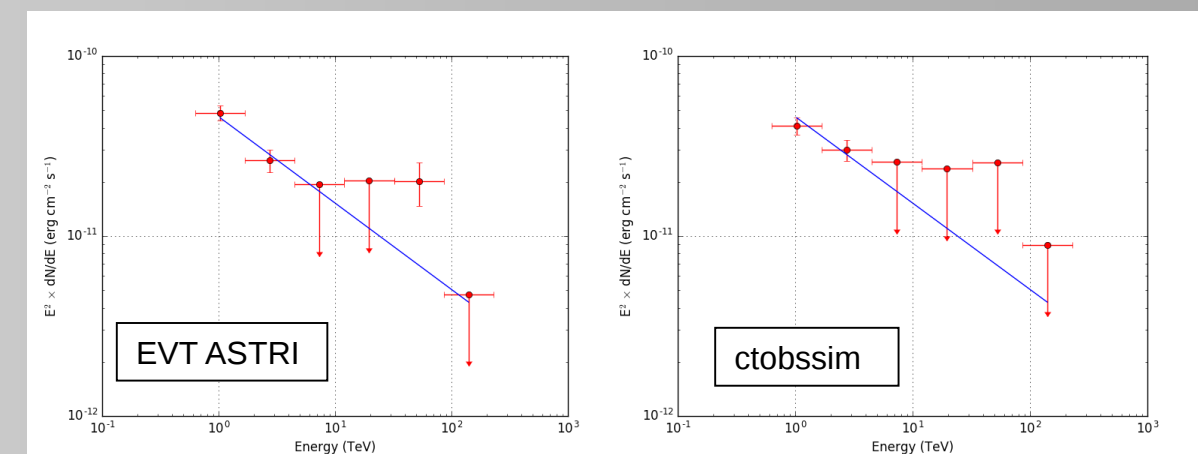
Index = (2.6 ± 0.1)

CTOBSSIM:

TS = 211.307

Flux = $(5.4 \pm 0.9)e-16$

Index = (2.50 ± 0.08)



CTOOLS ISSUES

- Feedback to CTOOLS → overwriting temporal labels (correct on “.dev” version 1.6.0)
- XML model create with CTBKGCUBE has CTA as instrument by default

BINNED (13 bin)

EVT

TS=203

Flux = $(5.3 \pm 0.9)e-16$

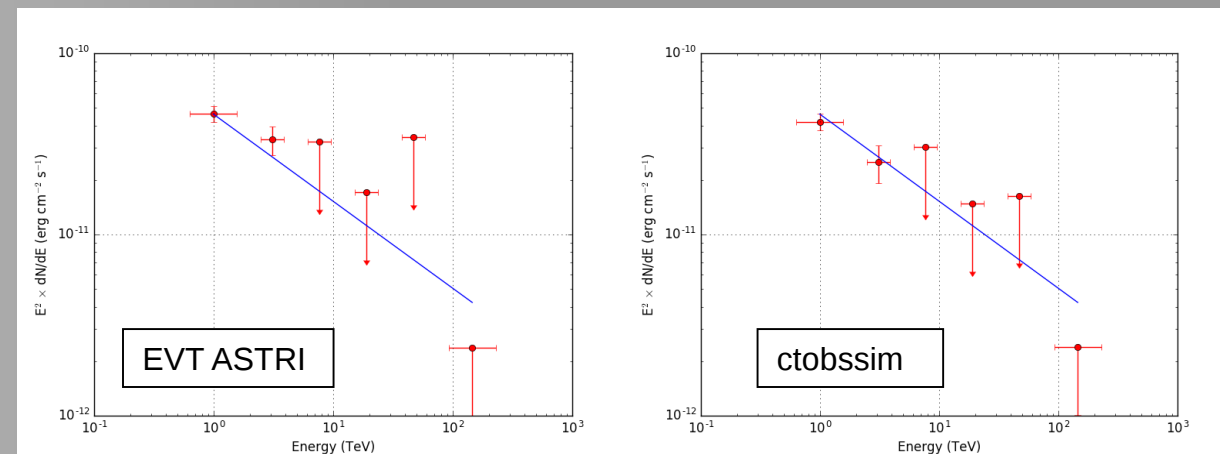
Index = (2.48 ± 0.08)

CTOBSSIM:

TS=179

Flux = $(6 \pm 1)e-16$

index = (2.56 ± 0.09)



Cross-check con i dati HESS

Issue #404 REDMINE in “ASTRI science tools”

(<http://astriweb.oa-roma.inaf.it/redmine/issues/404>)

HESS data:

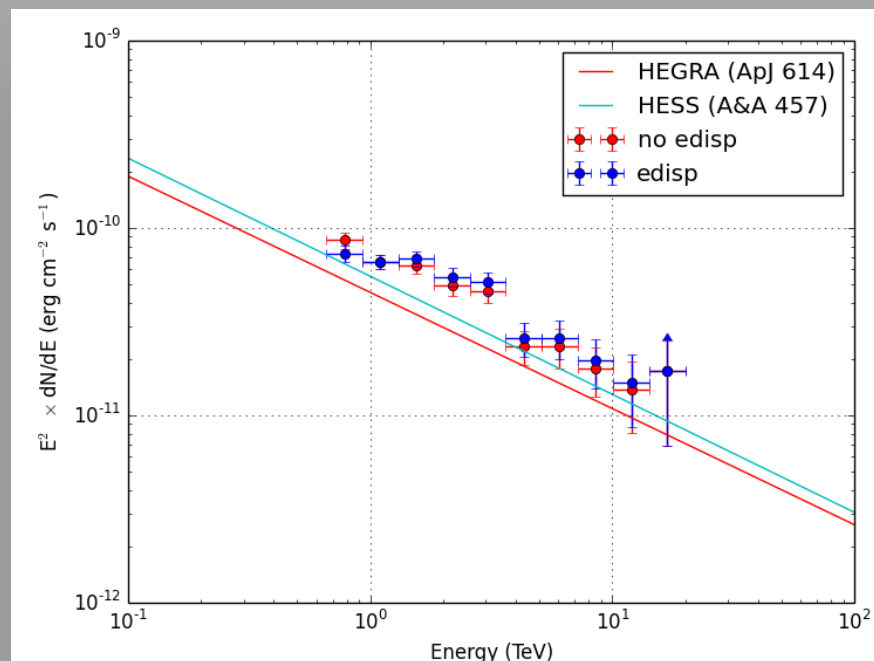
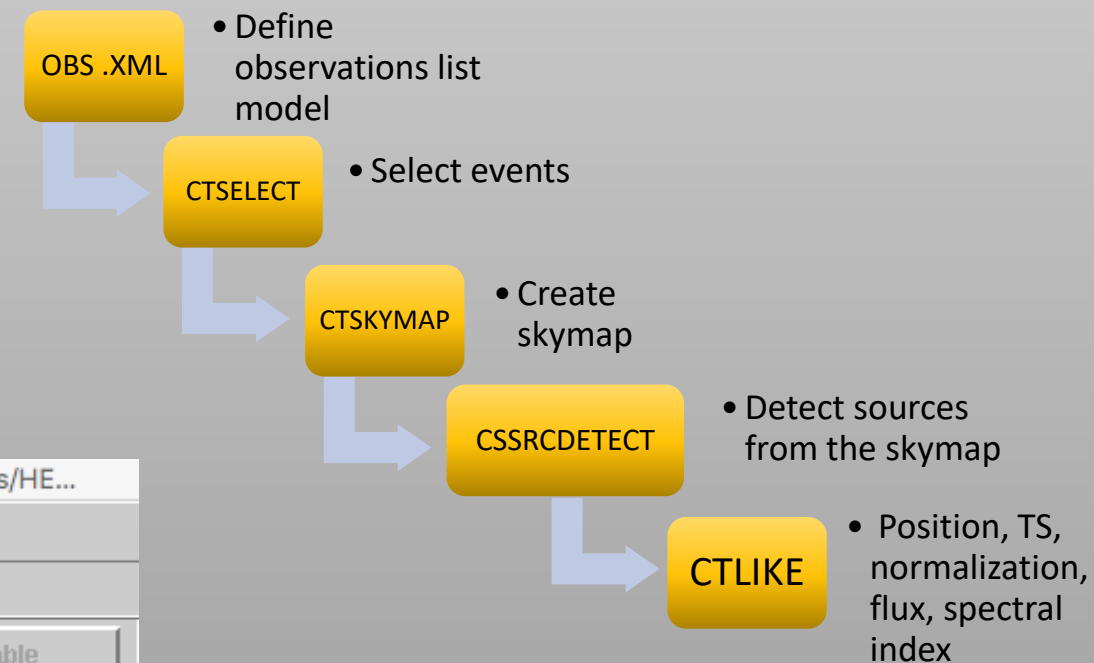
<https://www.mpi-hd.mpg.de/hfm/HESS/pages/dl3-dr1/>

Data sets (IRF, EVT) on: **Crab**, PKS 2155-304, MSH 15-52,
RX J1713.7-3946, Off data

fv: Summary of hess_dl3_dr1_obs_id_023526.fits.gz in /Users/flucarelli/Analisi/ASTRI/DL3_analysis/HE...

File	Edit	Tools	Help					
Index	Extension	Type	Dimension	View				
0	Primary	Image	0	Header	Image		Table	
1	EVENTS	Binary	5 cols X 7581 rows	Header	Hist	Plot	All	Select
2	GTI	Binary	2 cols X 1 rows	Header	Hist	Plot	All	Select
3	AEFF	Binary	5 cols X 1 rows	Header	Hist	Plot	All	Select
4	EDISP	Binary	7 cols X 1 rows	Header	Hist	Plot	All	Select
5	PSF	Binary	7 cols X 1 rows	Header	Hist	Plot	All	Select

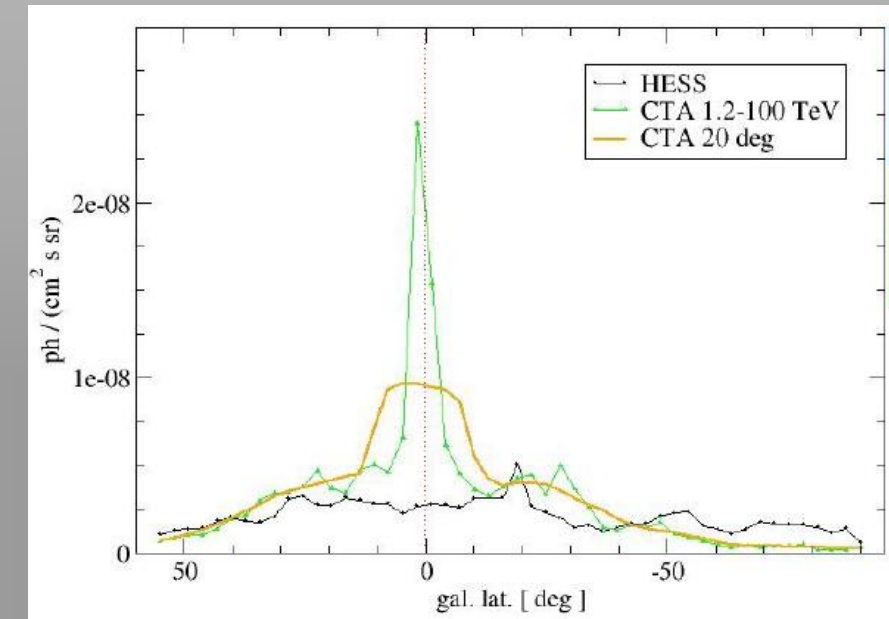
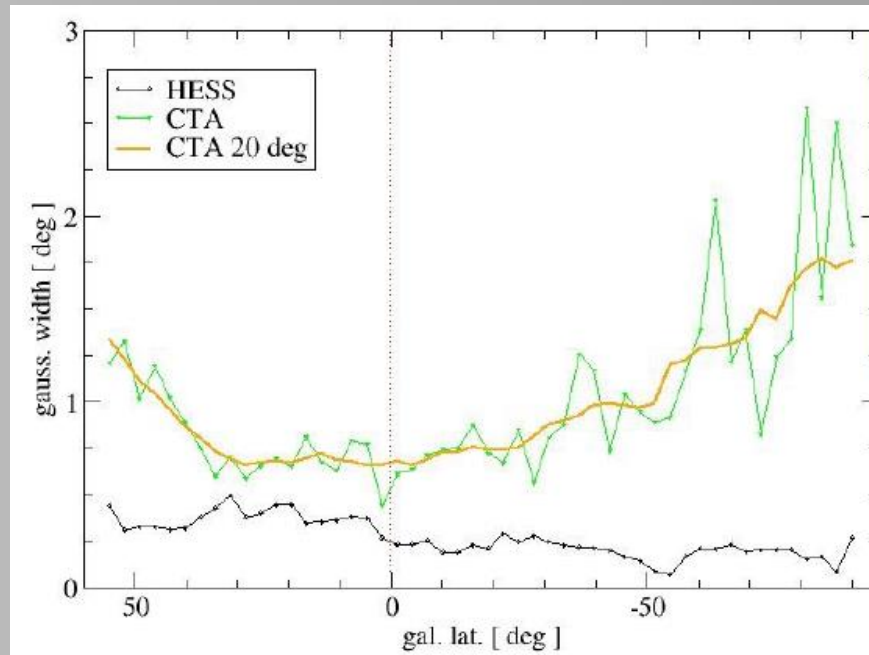
- Some keywords lacking in the ASTRI IRF/EVT;
- Only one fits file with both EVT and IRF (no BKG template provided) → consolidated?
- BKG from different sources (Aeff, IRF, ON/OFF)
- There are some problems with the generation of the spectrum



Diffuse model: CTA vs HESS

Bisogna aprire una Issue nel redmine

- Comparison between HESS (first estimation of the galactic TeV) and CTA diffuse model (theoric, used in DC) with an integration of CTA maps in the energy band of HESS [0,2-100] TeV
- Expected photons from CTA are many more than the ones seen by HESS (even using a more realistic energy band for CTA)
- Gaussian amplitude with the latitude are anti-correlated
→ maybe due to the fact that HESS considers also the instrument PSF



Wednesday 19 December 2018

- **Punto della Situazione (10:30-11:00)**
CARDILLO, Martina
- **Simulazioni, IRF e dati reali (11:00-11:30)**
LOMBARDI, Saverio
- **Ricapitolazione sul lavoro fatto e "in progress" su Prototipo e MA (11:30-12:00)**
GIULIANI, Andrea; PINTORE, Fabio
- **Risultati da ultimi test su Prototipo e MA (12:00-12:30)**
CARDILLO, Martina
- **Pranzo (12:30-14:00)**
- **Gammapy: problematiche e piano di lavoro (14:00-14:30)**
D'Ai, Antonino; PINTORE, Fabio
- **Dati HESS e MAGIC (14:30-15:00)**
LUCARELLI, Fabrizio
- **Modello del diffuso: confronto CTA-HESS (15:00-15:30)**
MARCHILI, Nicola
- **Repository: accesso e utilizzo (15:30-16:00)**
LOMBARDI, Saverio
- **Tutorial su ambiente ASTRI e utilizzo ASTRIPIPE modificata (16:00-16:30)**
PINTORE, Fabio
- **Discussione e completamento issues precedenti (16:30-17:30)**
TUTTI

AGENDA

Thursday 20 December 2018

- **Riassunto punti salienti del giorno precedente (09:30-10:00)**
CARDILLO, Martina
- **Hando to hand e discussione (10:00-13:00)**
TUTTI
- **Pranzo (13:00-14:00)**
- **Azioni future e gruppi di lavoro (14:00-17:00)**
TUTTI

AGENDA

Obiettivi

- Individuazione dei singoli task da portare avanti e di quelli da sospendere
- Organizzazione in sottogruppi di lavoro per procedere nei vari tasks
- Capacità nell'utilizzare il repository
- Capacità nel maneggiare ASTRIPIPE e la sua implementazione
- Maggiore compattezza del team e maggiore comunicazione al suo interno (con fondamentale utilizzo dei repository)
- Collacazione del nostro lavoro all'interno di CTA
- Piano di lavoro sui dati reali