

Astrometric Gravitation Probe (AGP)

Riva, A., on the behalf of AGP team

Rome, Nov 20th 2018

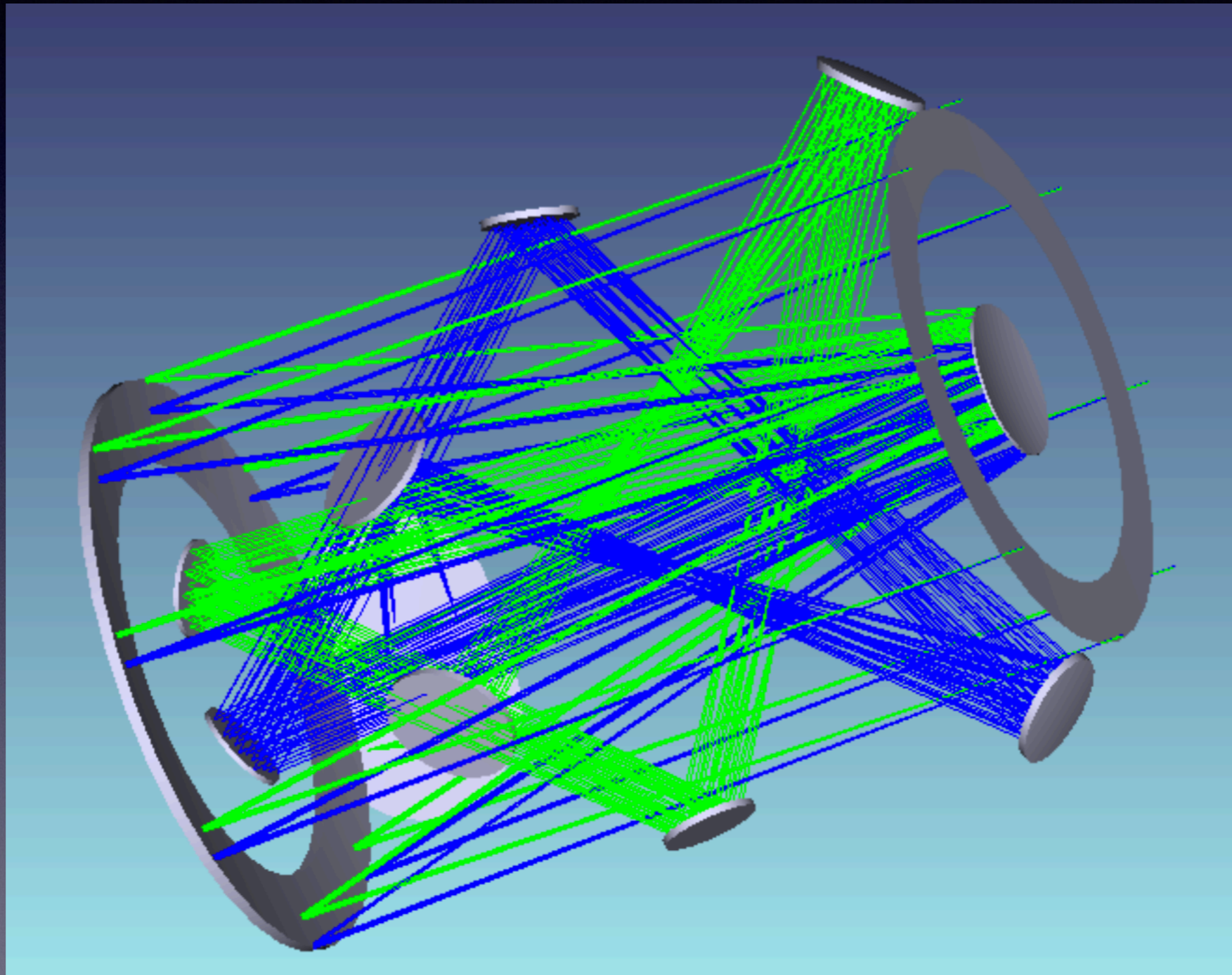
Final message

- Let's enhance the “sparse pupil” technique research activities

Overview

- AGP Mission
- Science goals
- Scientific requirements
- Measurement principles
- Instrument design
- Performances

What's AGP?



AGP: Astrometric Gravitation Probe

Astrometry = measurement of apparent star position

Gravitation = Tests in weak field (solar system)

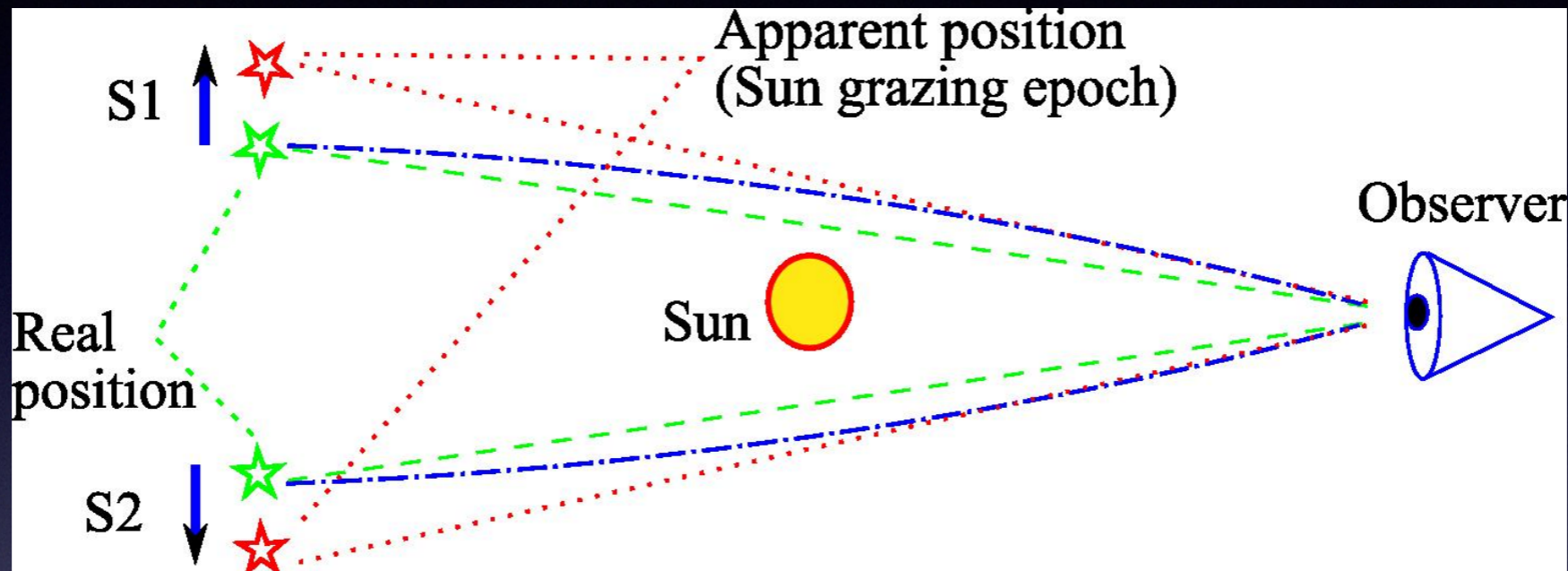
- 1) Light deflection close to the Sun
- 2) High precision dynamics in Solar System

Probe = Medium size space mission

**Mission design driver:
Light bending around the Sun @ sub- μ as level**

Proposed in ESA M4 call and 2015 Call for Ideas

AGP concept: Dyson-Eddington-Davidson experiment (1919)



A space mission in the visible range to achieve

- long permanent artificial eclipses
- no atmospheric disturbances, low noise

Differential measurement for
systematic error control

AGP Mission Profile

Baseline launcher:	Vega
Sun-Synchronous orbit (SSO, $i=99.48^\circ$), elevation:	1000 (1500) km
Useful mass (satellite + payload):	1140 (1000) kg
Spacecraft dry mass (incl. payload and propulsion systems)	~ 800 kg
Payload mass	~ 300 kg
In-orbit operations	3-5 years
fields of view	16' x 16'
Effective Focal Length	~ 25 m
Telescope primary diameter:	1.15 m
Effective Fizeau pupil diameter:	1.10 m
	lightweight mirror zerodur or SiC
Payload envelope:	2.1 m diameter x 1.5 m height
Detector:	CCD mosaic @ -20 C
Main science focal plane:	8 x 4 CCDs - 2kx4k
Auxiliary (pointing) focal plane:	1 x 2 CCDs - 2kx4k

AGP Science goals

Characterisation of weak field gravity in the Solar System

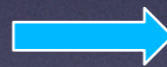
- Deflection of light in the solar system
- Non-linearity of gravity, preferred frame
- Relativistic effects of oblate and moving giant planets
- Solar system dynamics [High precision ephemerides]

AGP Science goals

Characterisation of weak field gravity in the Solar System

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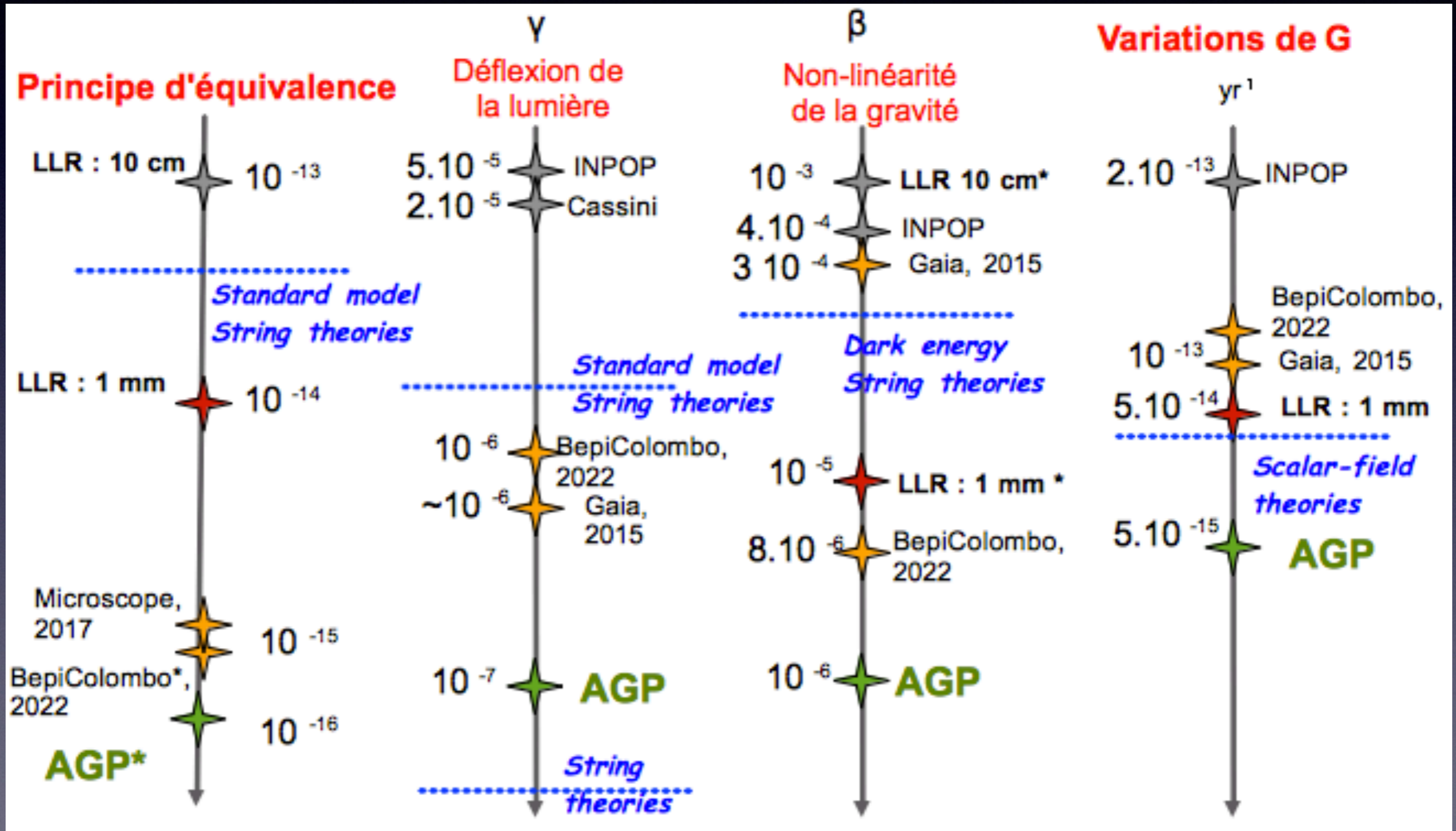
- Detection of dilaton
- Limits on Lorentz invariance
- Preferred frame detection
- Anisotropy of light deflection
- Test of Equivalence Principle



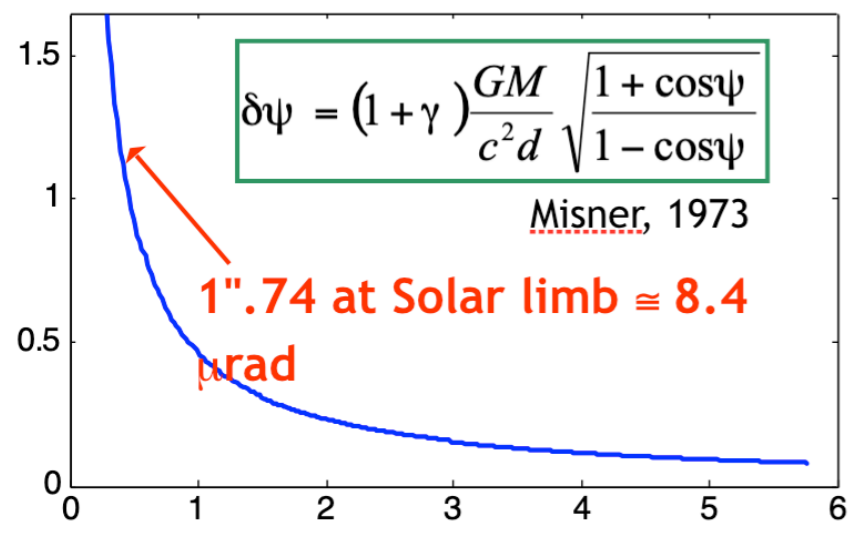
Exclusion / validation of alternative theories of gravity

Crosta and Mignard, 2006
Colladay, Kostelecký, 1998
Kostelecký, 2004
Kostelecký, Russell, 2014
A. Hees et al., 2015

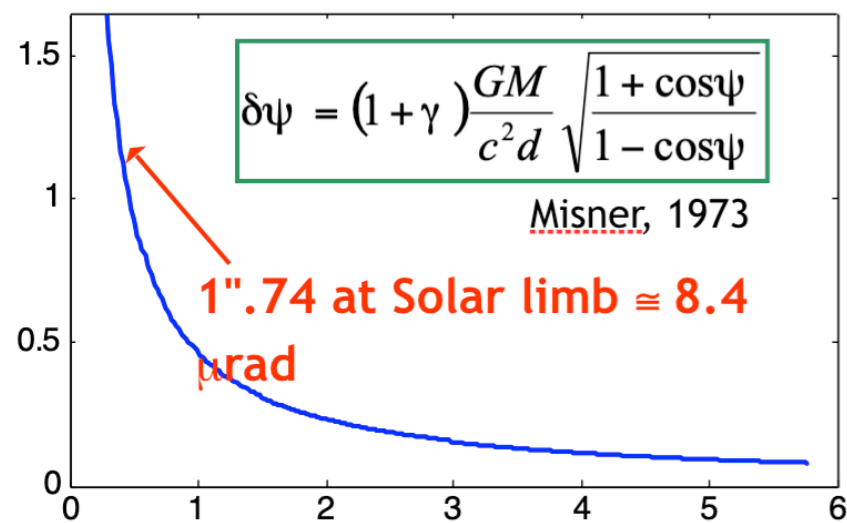
Science goals summary



AGP scientific requirements driven by light deflection case



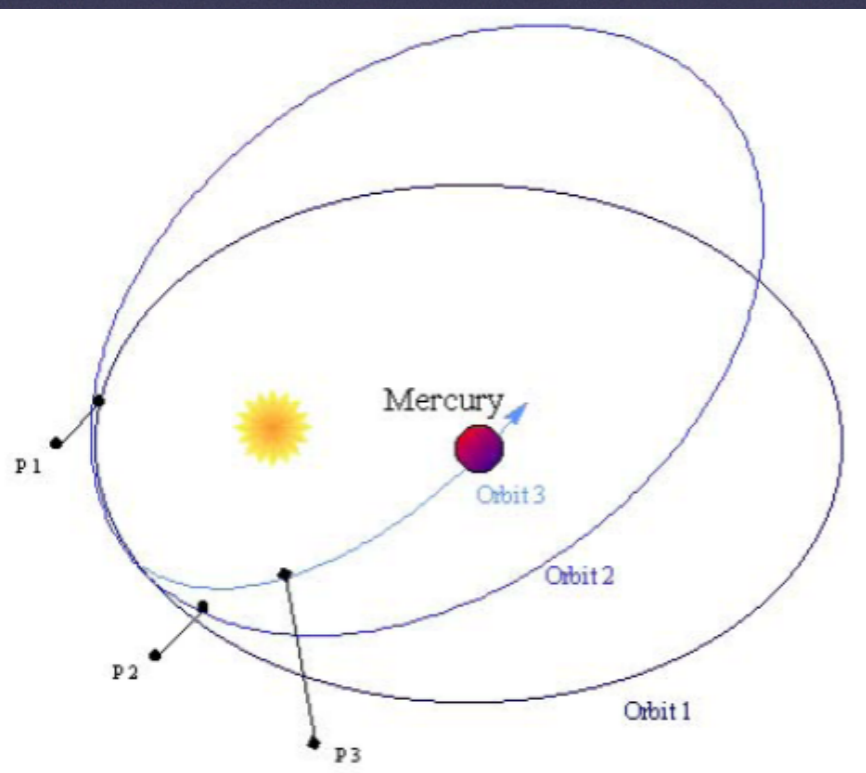
AGP scientific requirements driven by light deflection case



Measurement precision required:

- Final collective accuracy (α, δ) : 0.1 to 0.01 μ as
- 1 μ as (α, δ) for Mercury
- Individual precision: $\sigma_{\text{star}} \sim$ 100 to 10 μ as

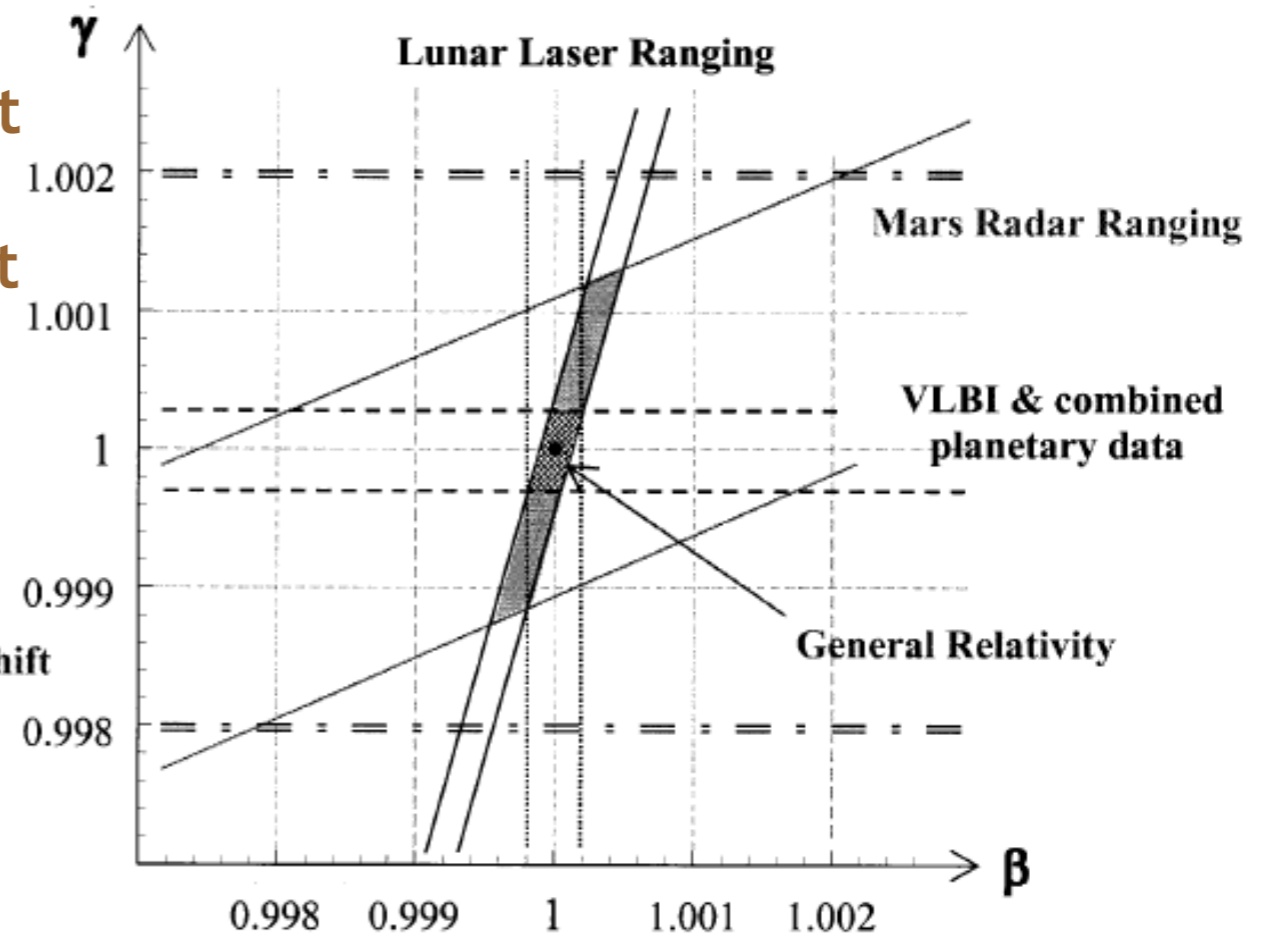
Complementary GR tests:



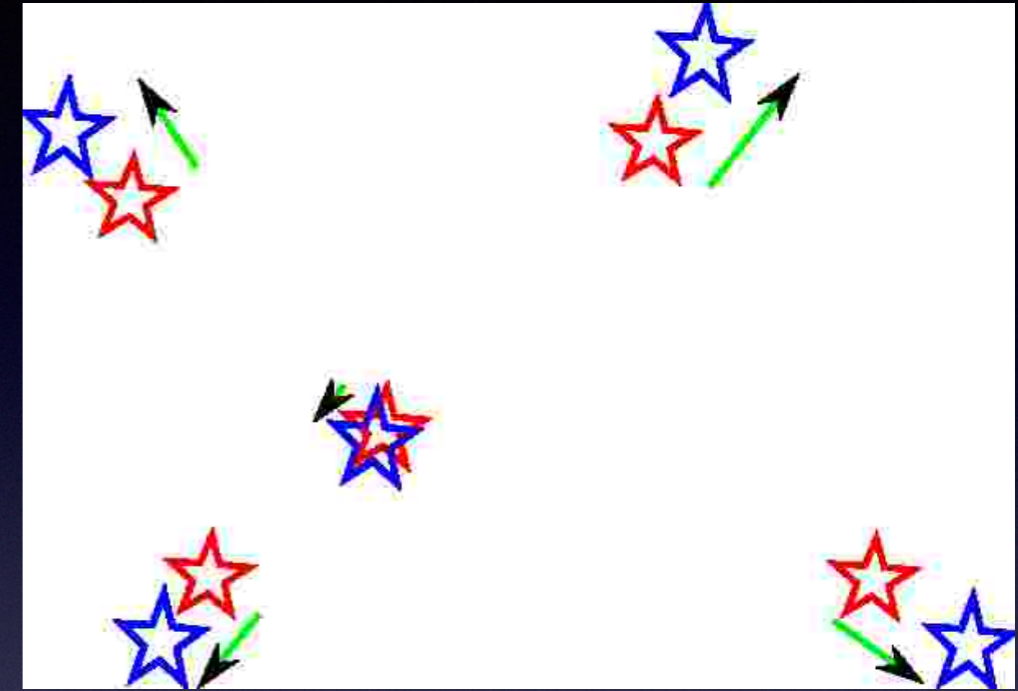
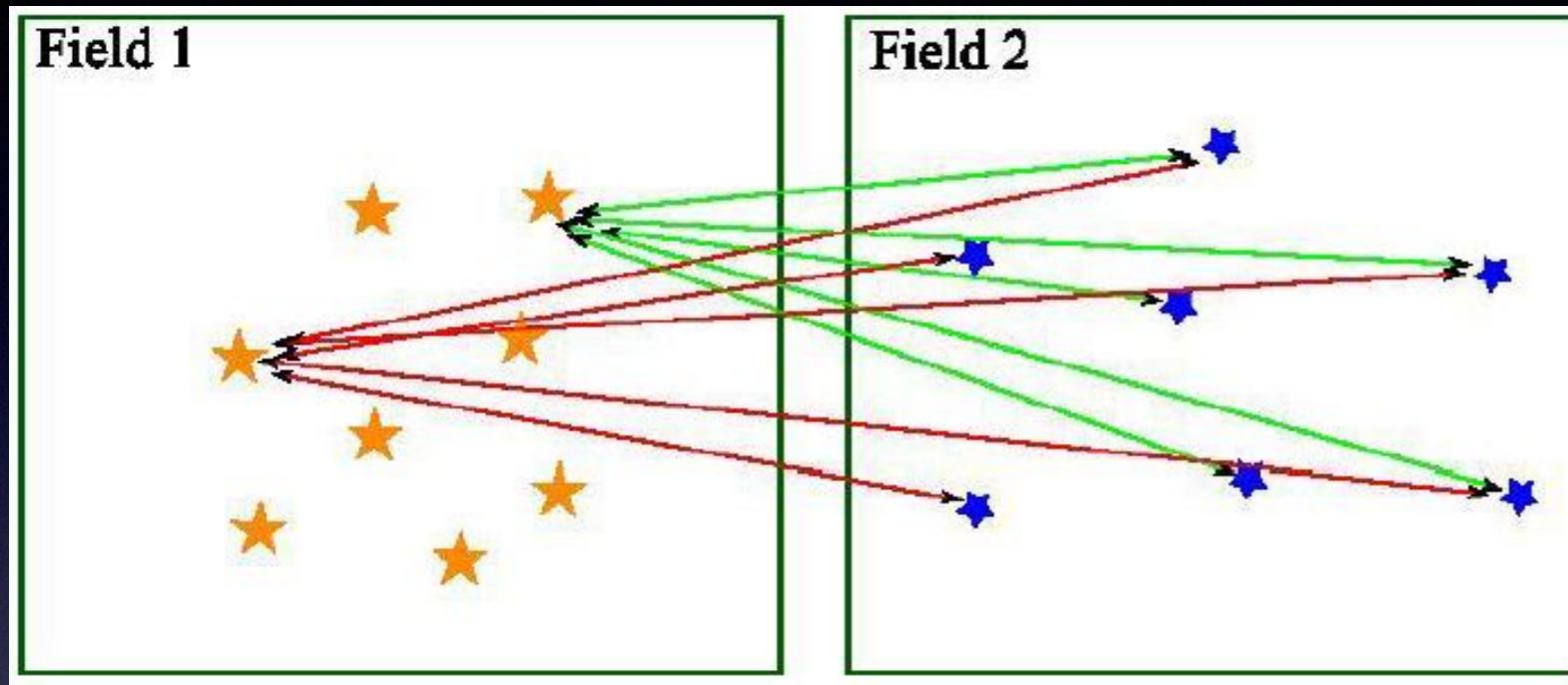
measurement of β from Mercury orbit

$$2 + 2\gamma - \beta$$

Mercury Perihelion Shift

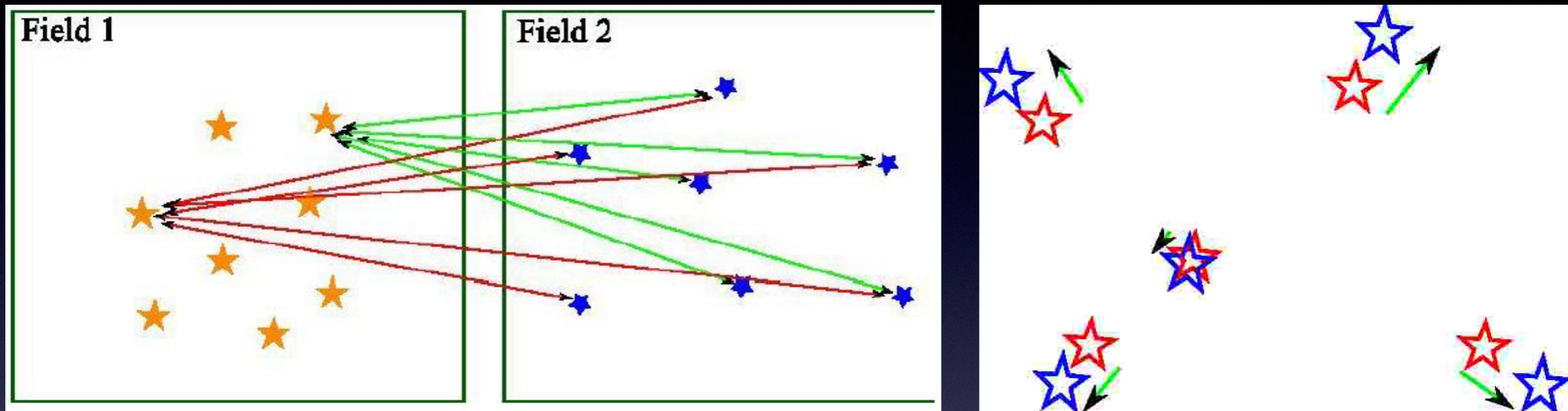


Fully differential measurement



Measurement of angular separation of stars between fields

Fully differential measurement

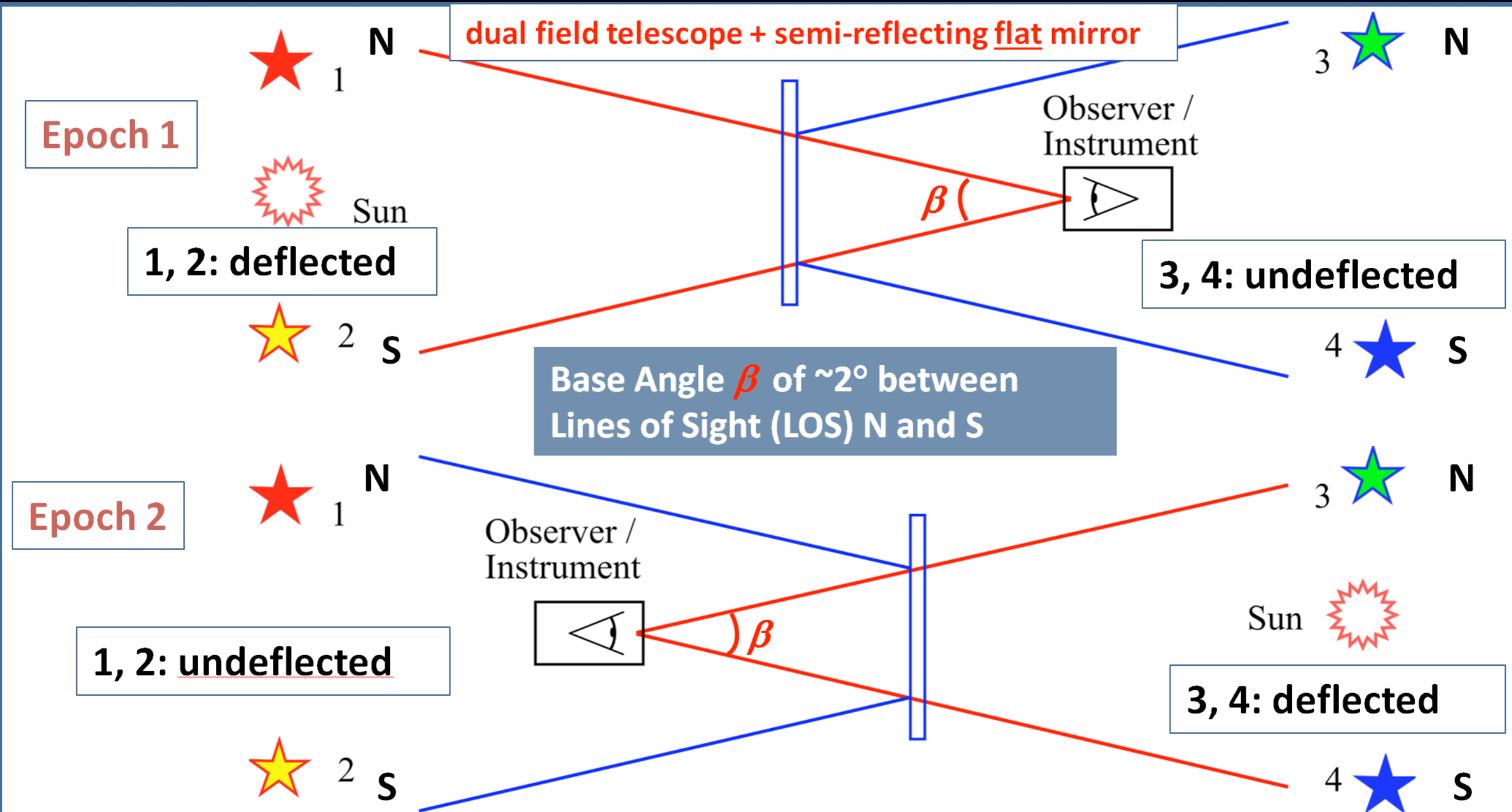


Measurement of angular separation of stars between fields
Different collective effects on field images from

- instrument evolution (focal length, distortion)
- deflection (field displacement)

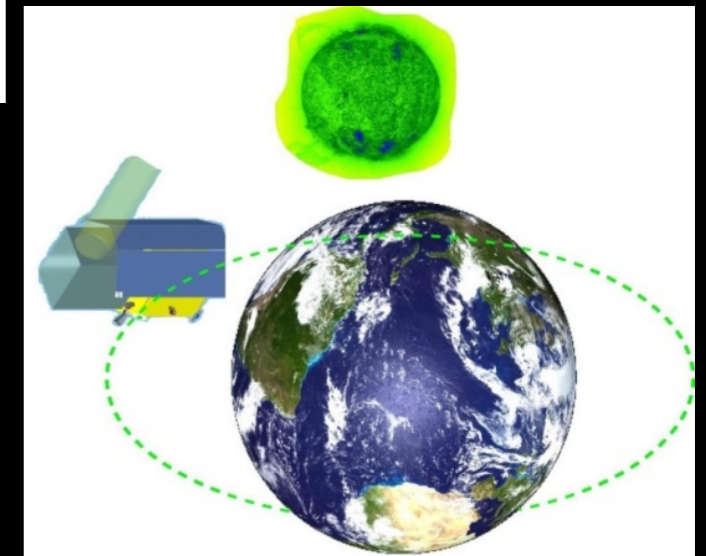
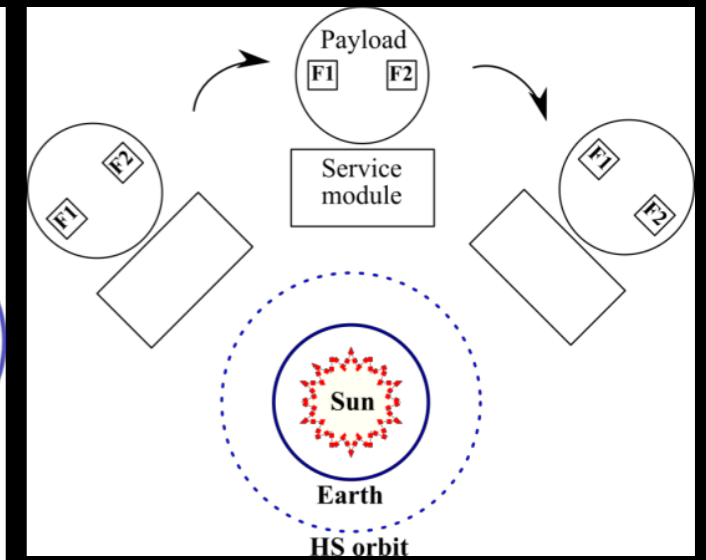
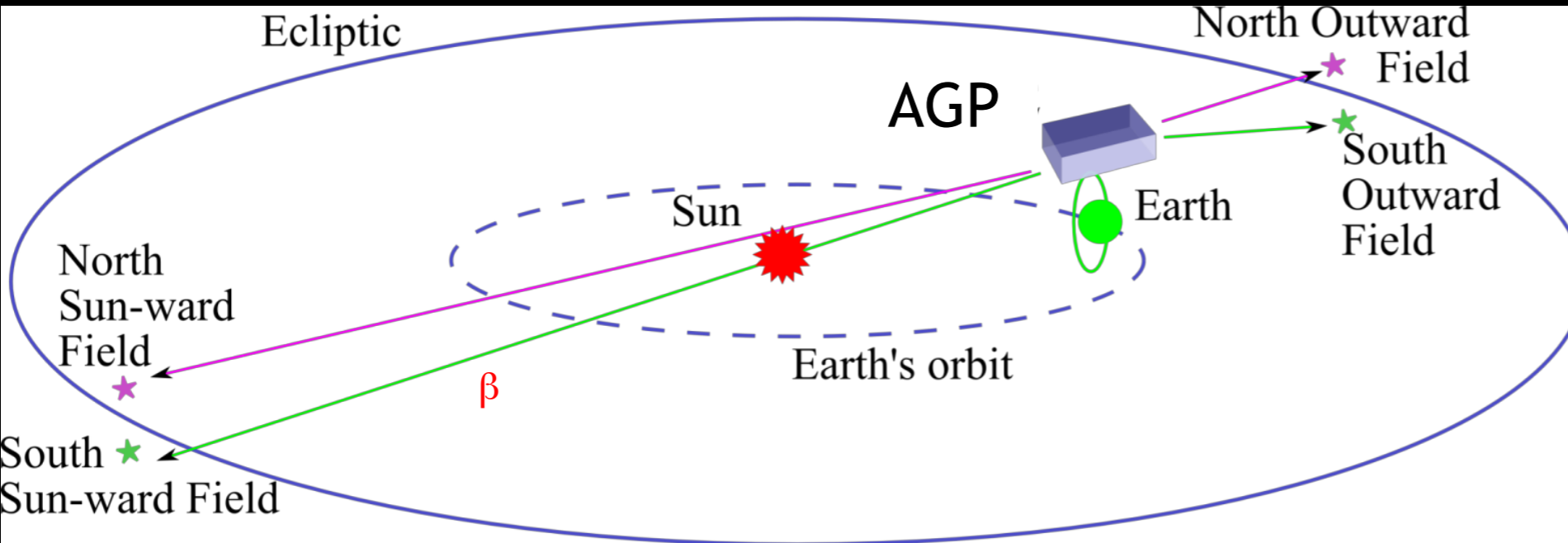
⇒ simple calibration of **MULTIPLICATIVE** terms

Principle



AGP Observation Strategy

Multiple field telescope on Helio-synchronous orbit @ h=1000 km

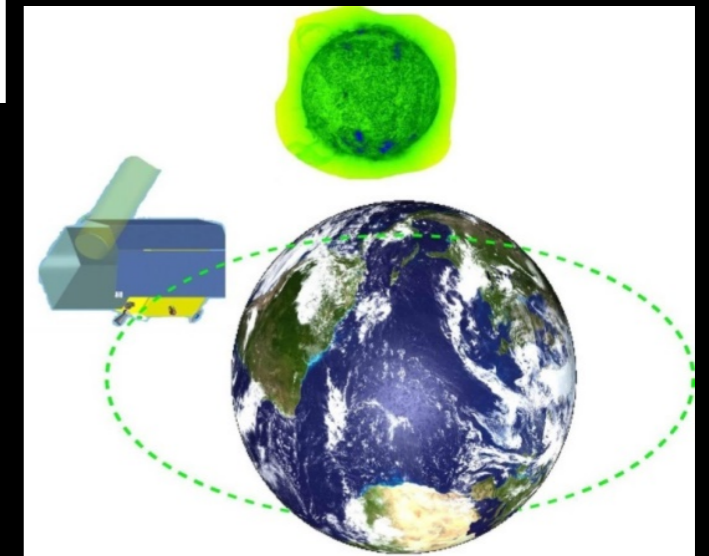
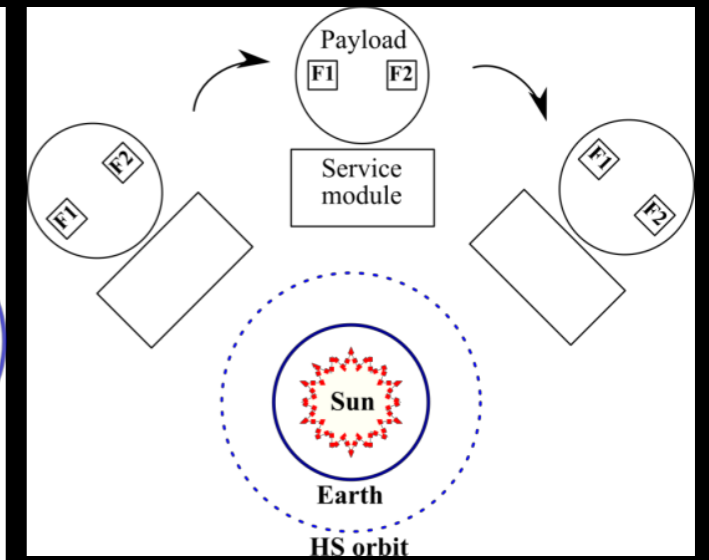
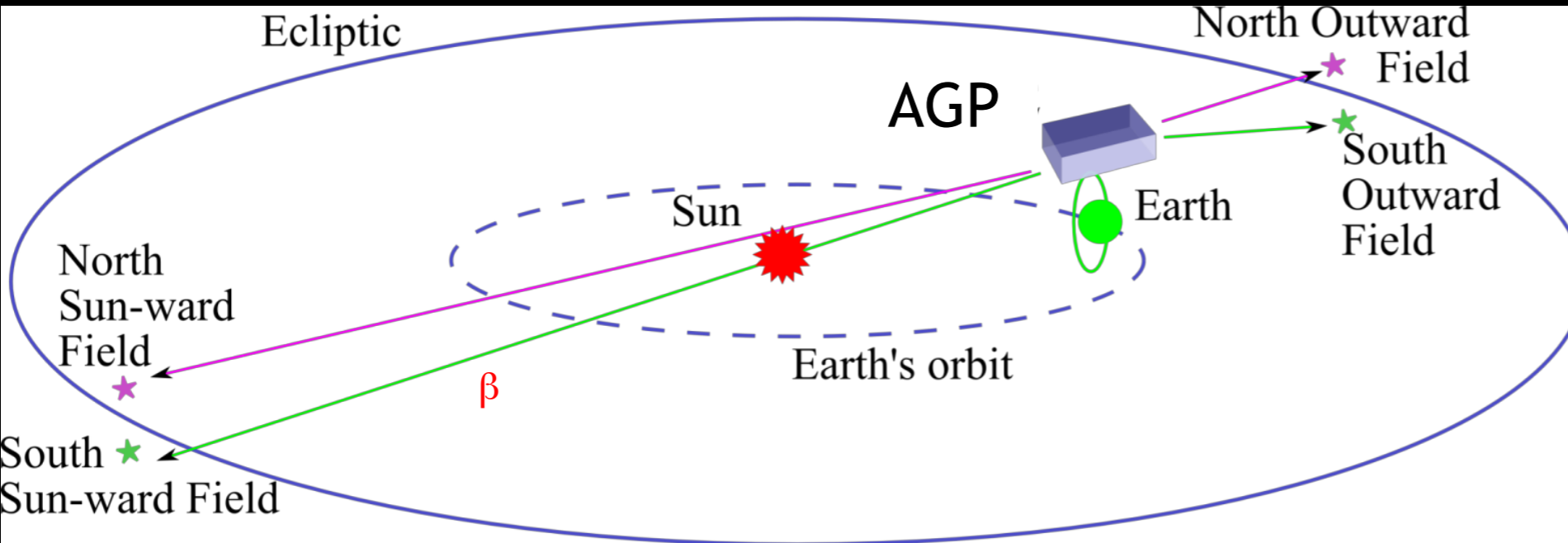


- 105 minute period
 - Satellite spin with orbit period
- 100% observing time

Stable energy supply + *thermal environment*

AGP Observation Strategy

Multiple field telescope on Helio-synchronous orbit @ h=1000 km



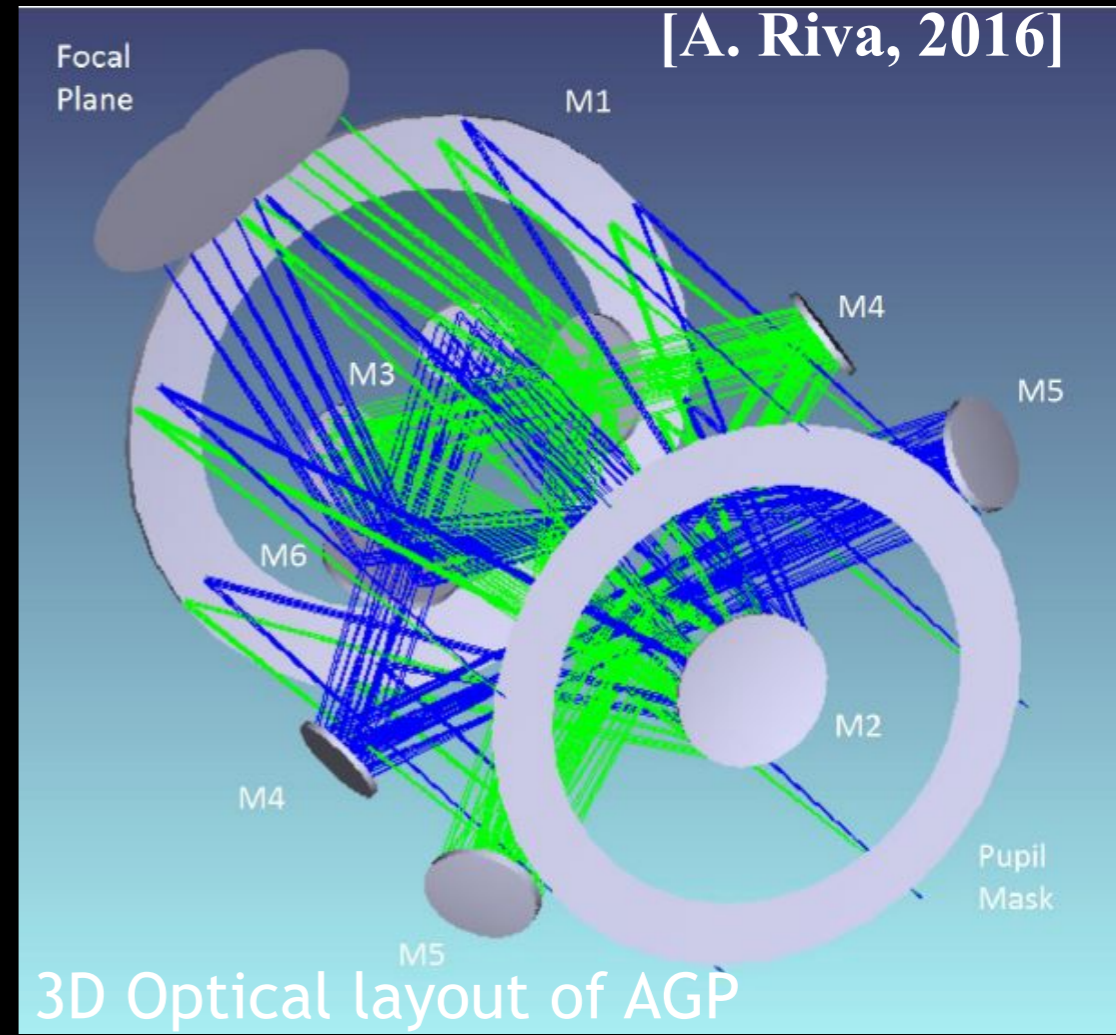
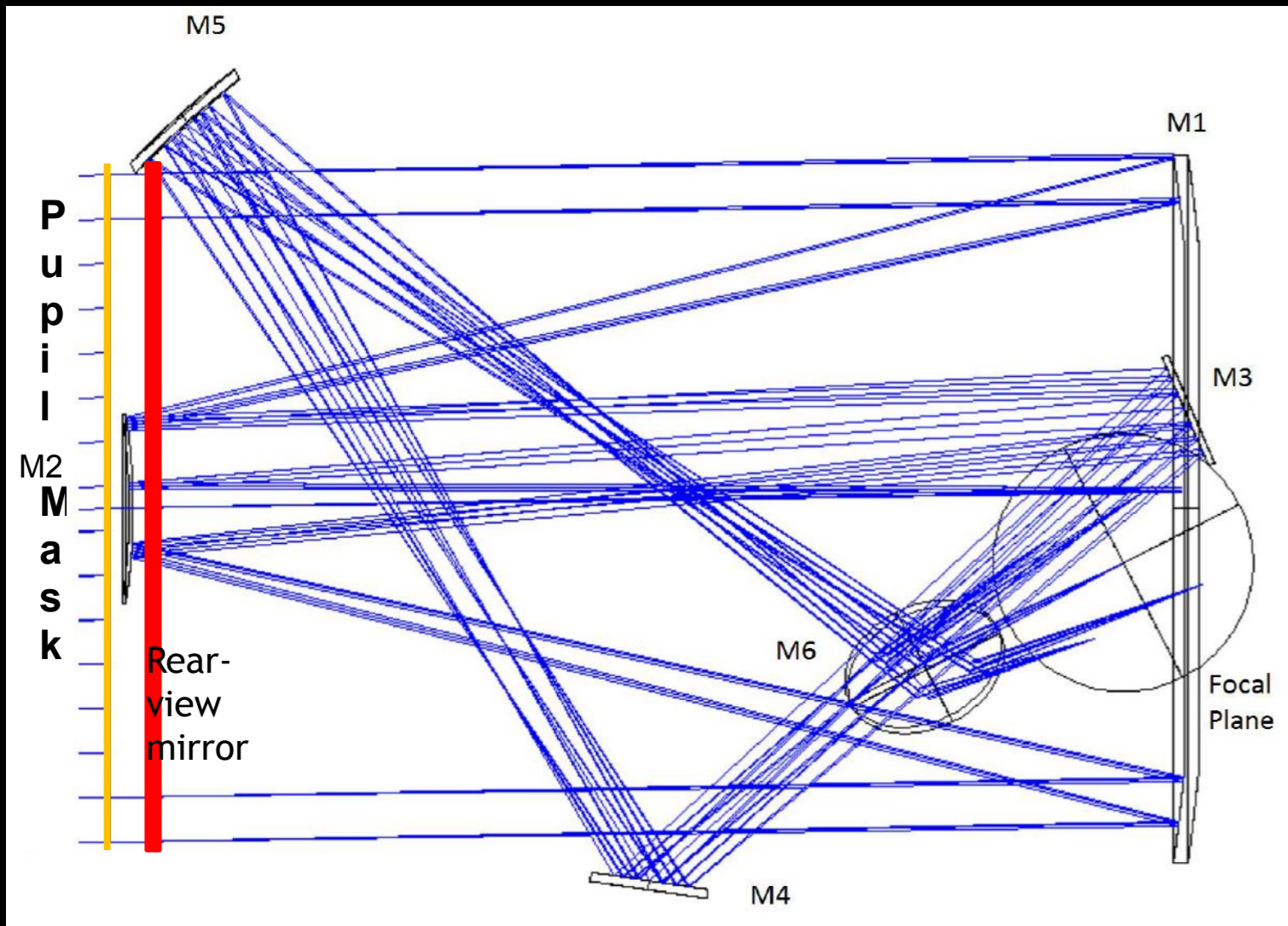
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Stable energy supply + *thermal environment*

System rationale: preserve satellite orientation vs. Earth (nearly stable thermal environment, filtered by service module)

Science rationale: average instrument response among channels (strengthen calibration) by field rotation

AGP Instrument design

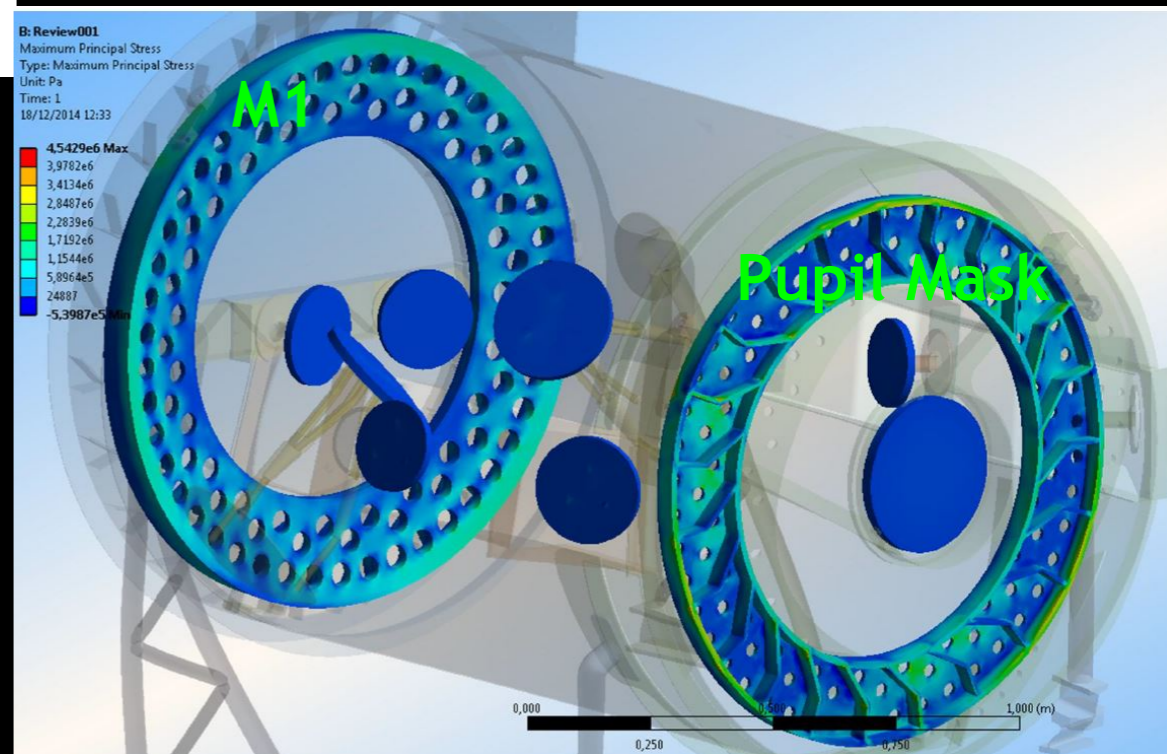


[A. Riva, 2016]

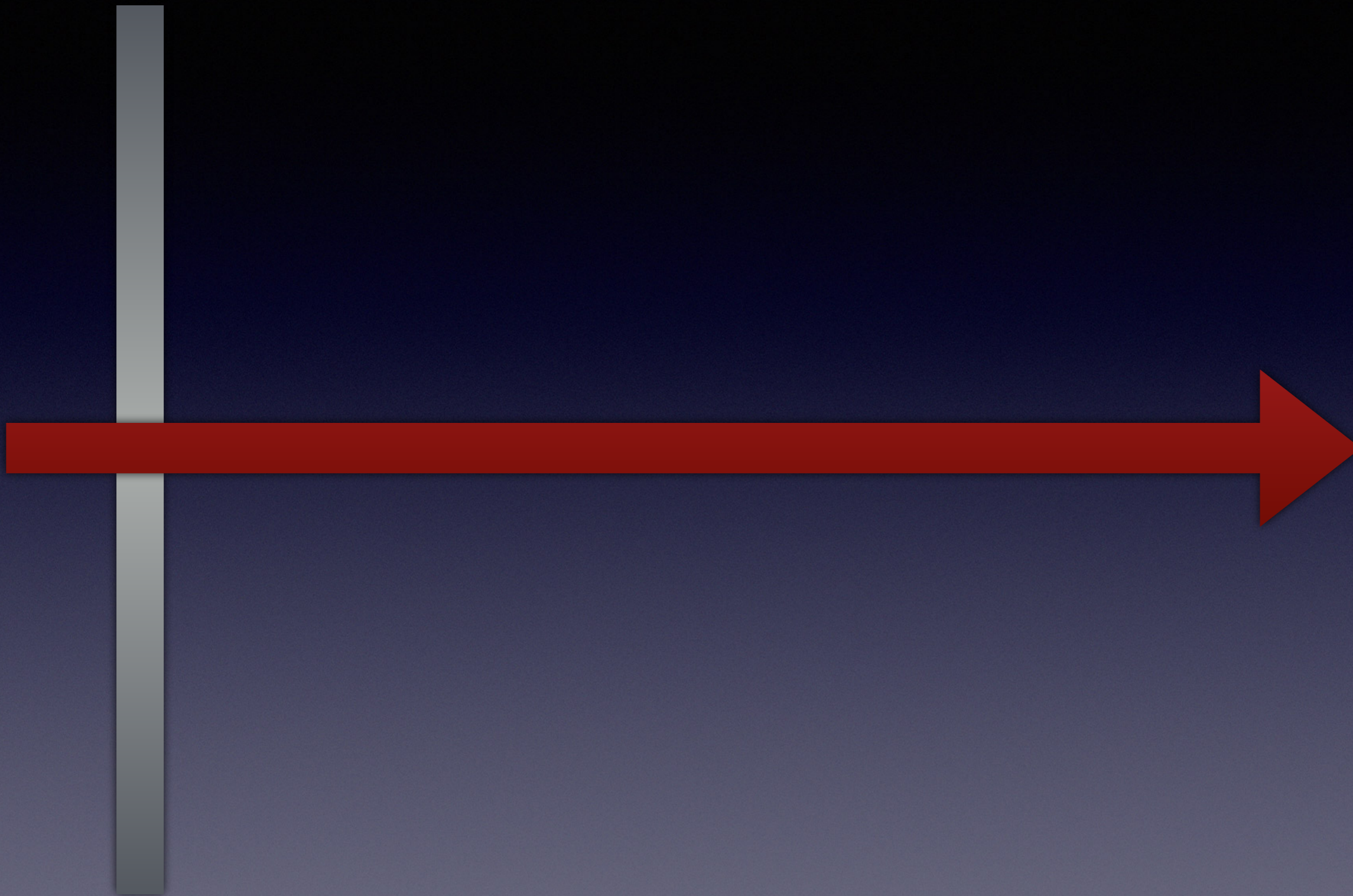
3D Optical layout of AGP

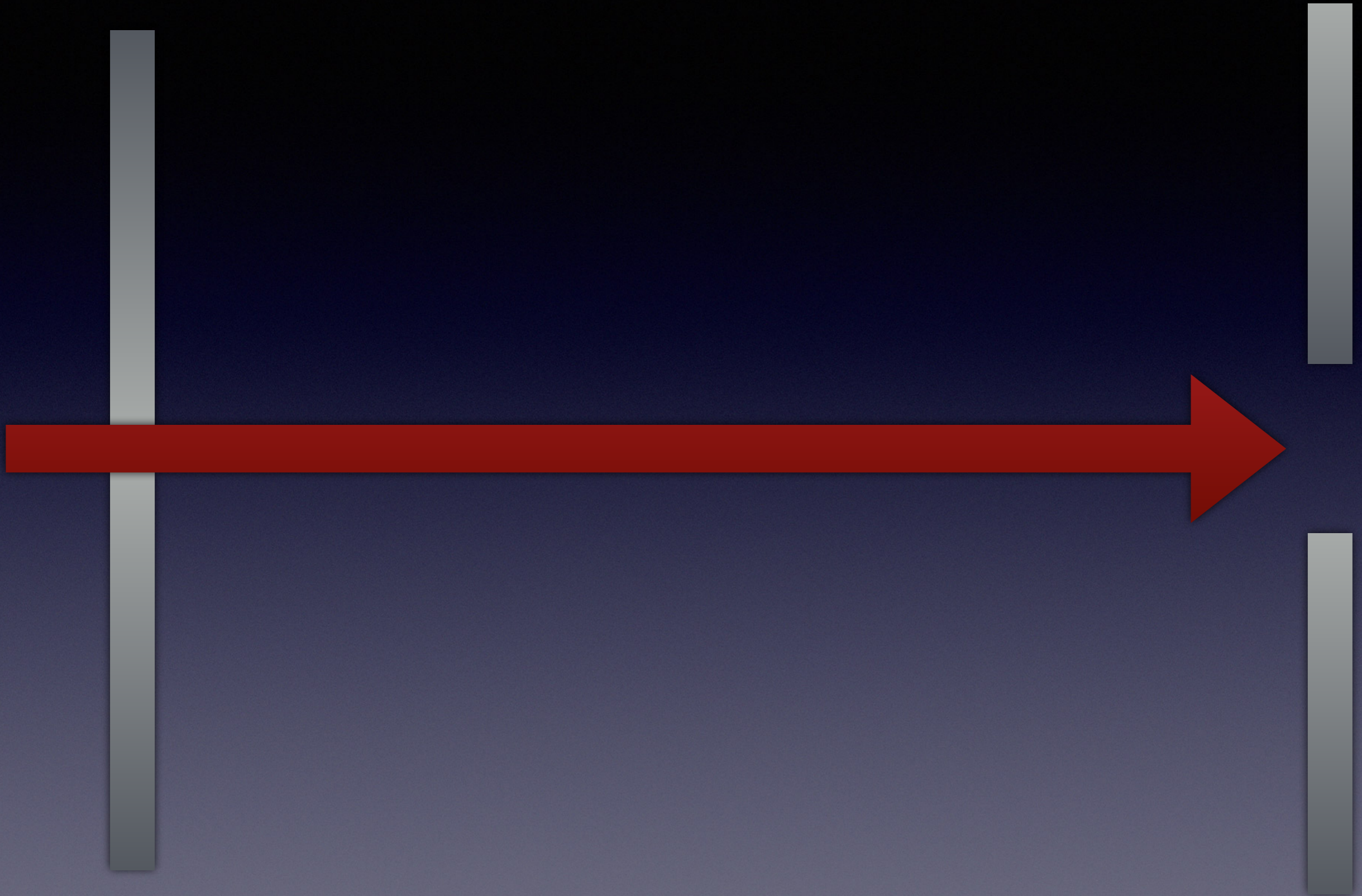
Optical layout of AGP for the single channel

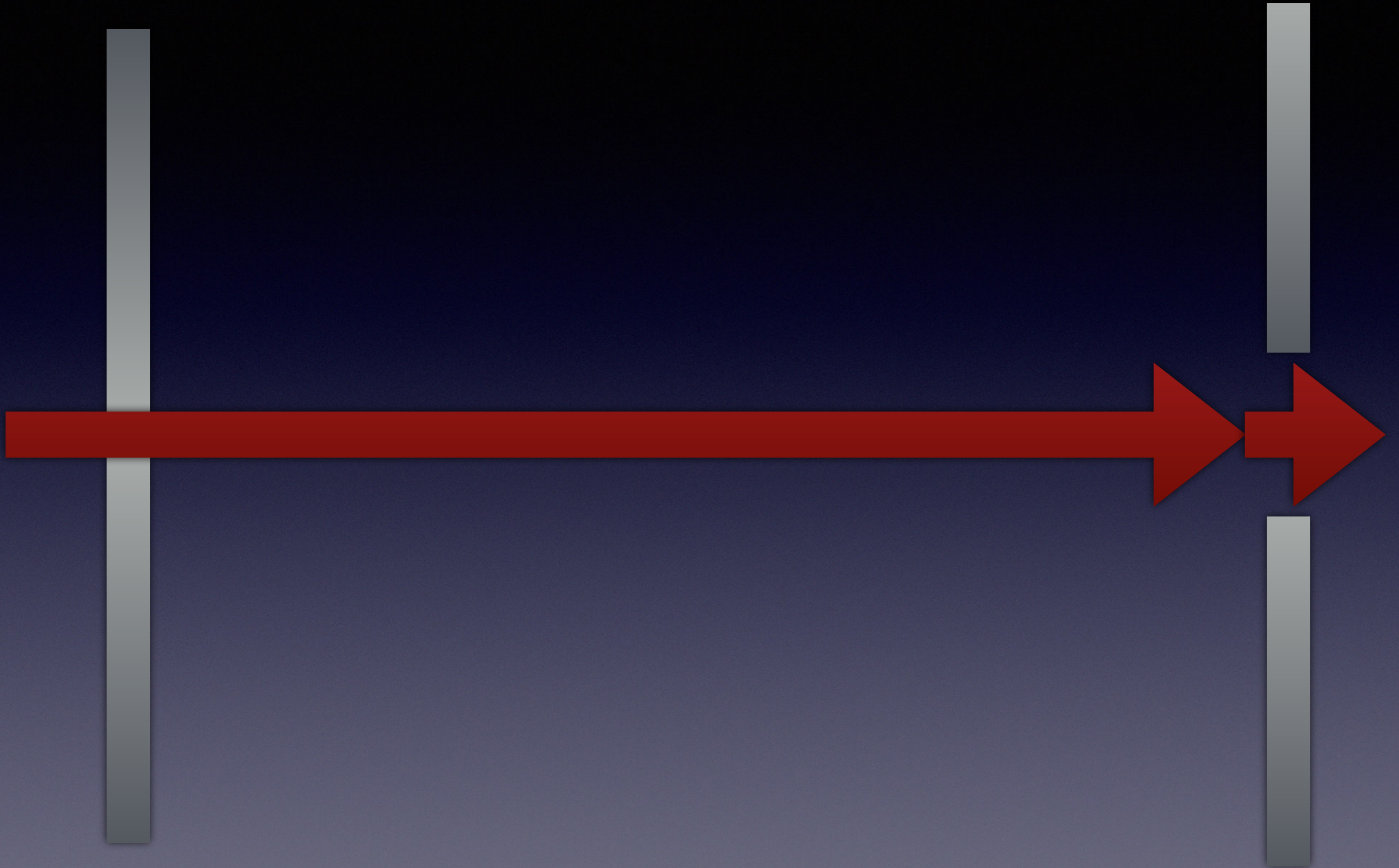
Element	Diameter	Radius	Conic	Distance to next element
Mask	1080.00	flat	0	1750.00
M1	1148.26	-4367.56181	-0.99265207	-1730.00
M2	309.08	-1045.77804	-1.65031622	1760.00
M3	200.00	flat	0	-1000.00
M4	180.00	flat	0	1650.00
M5	220.00	-1340.96793	-0.71077101	-1280.00
M6	240.00	flat	0	860.00

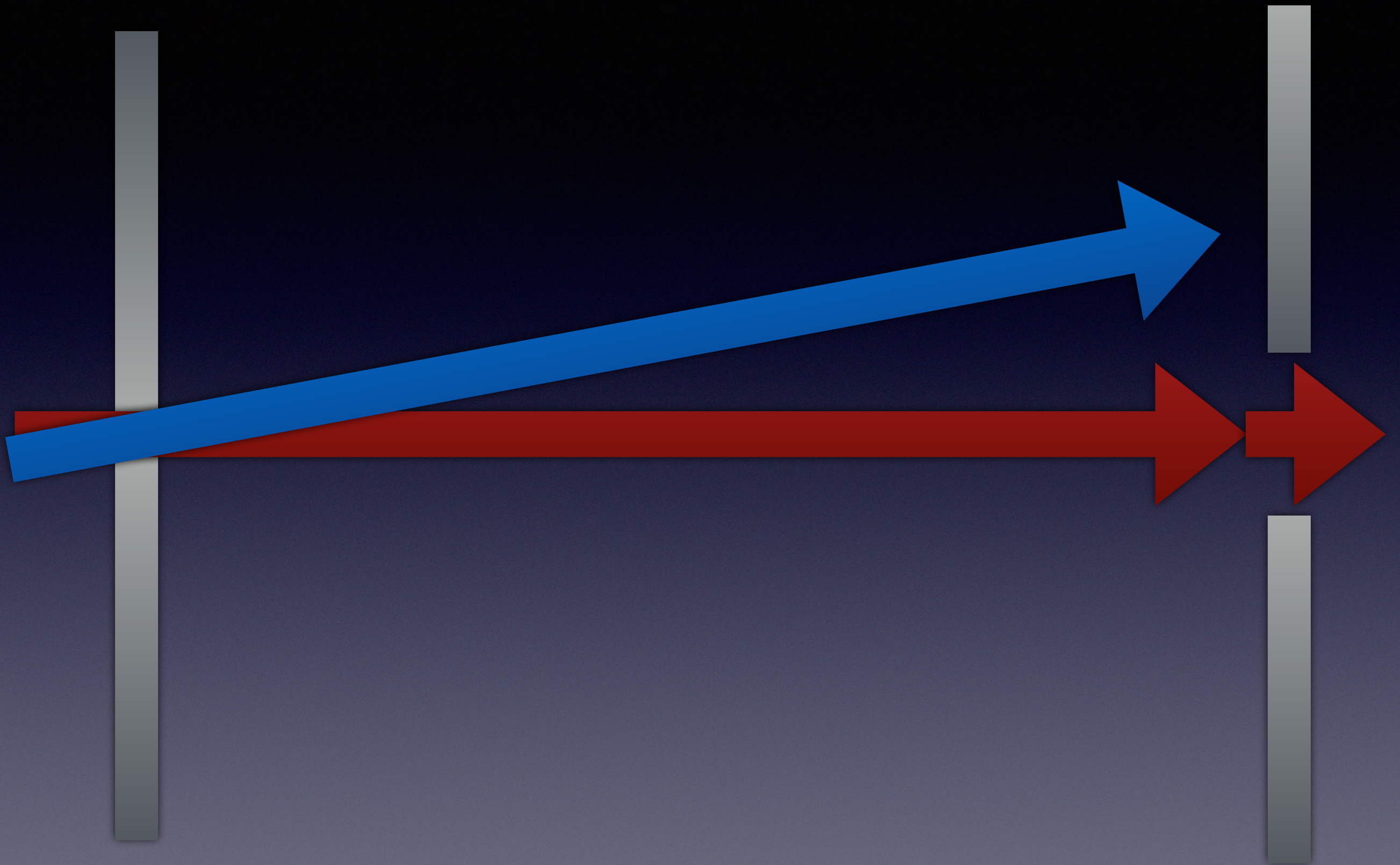


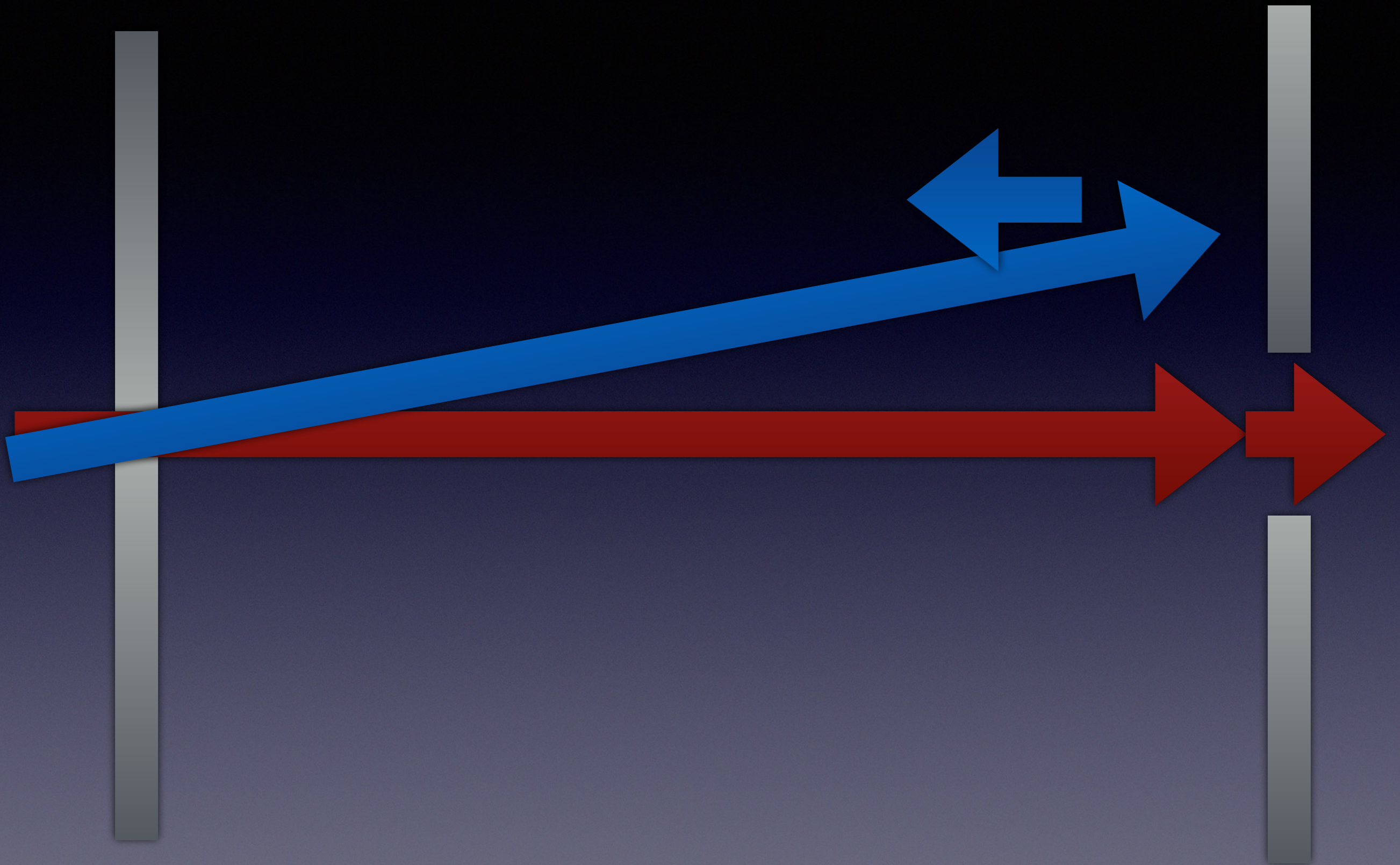


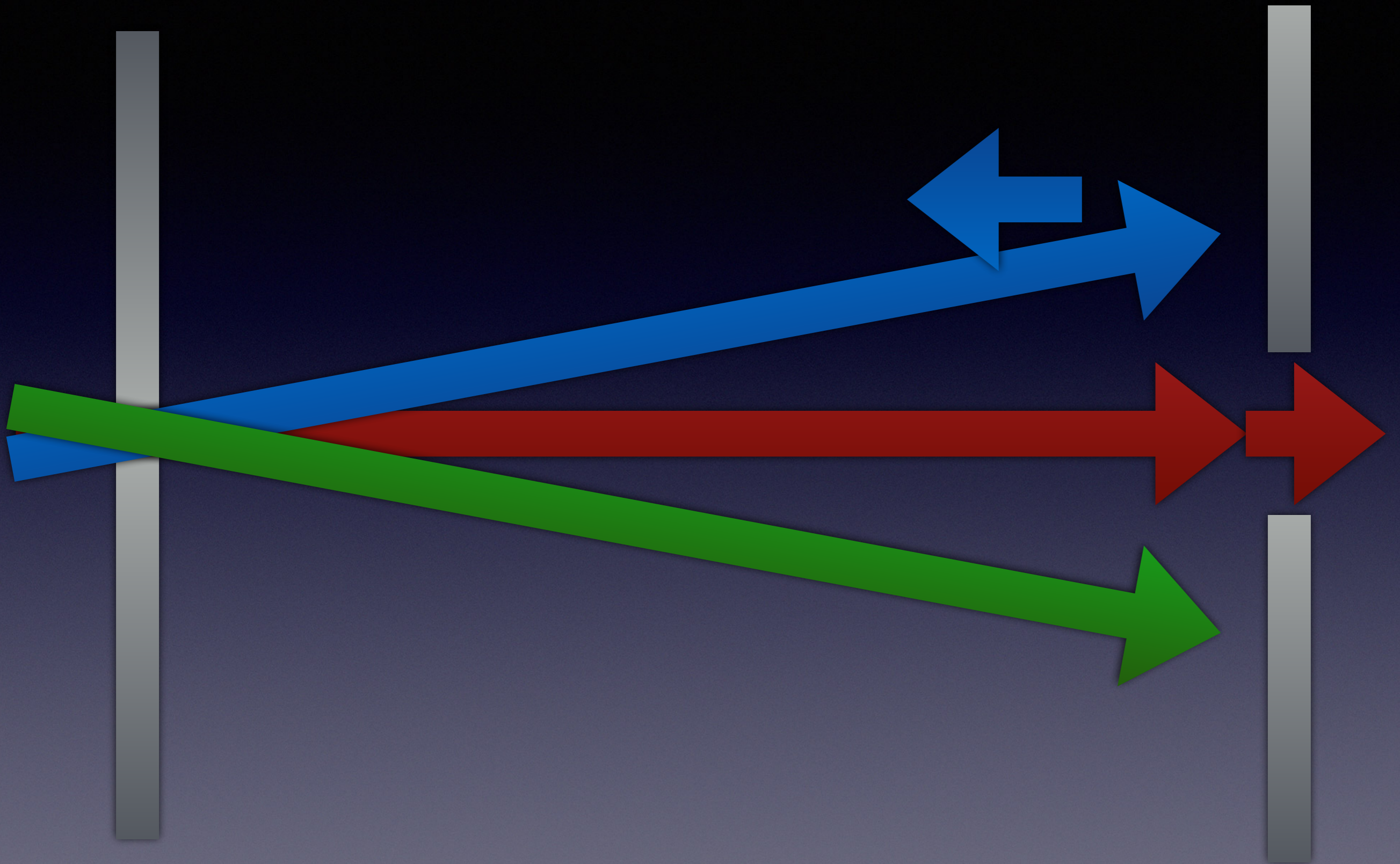


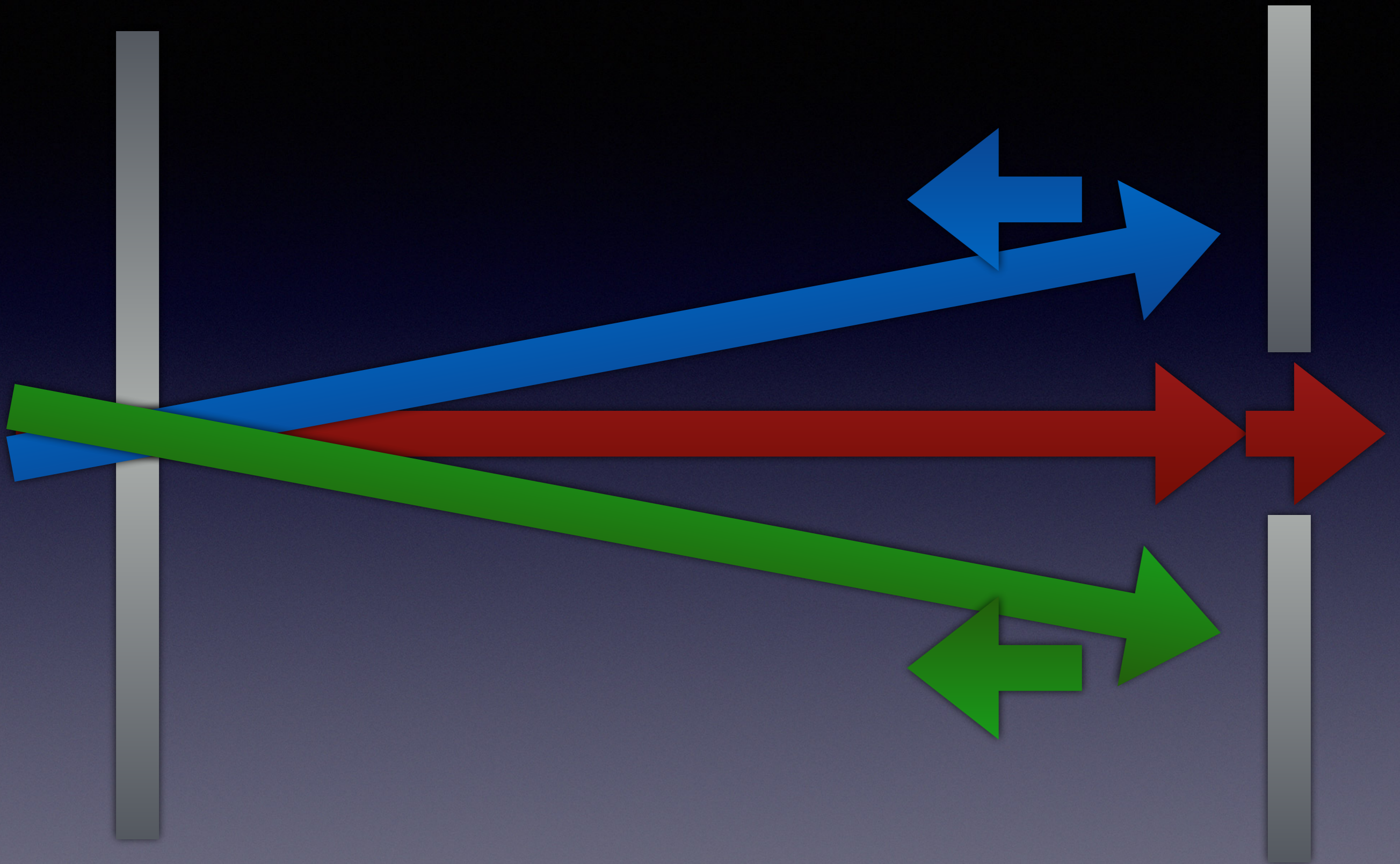


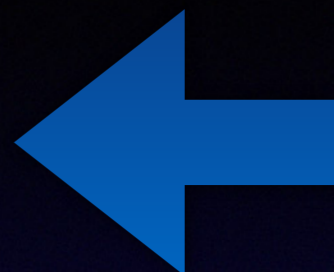
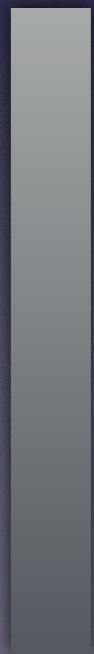
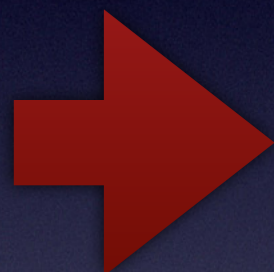


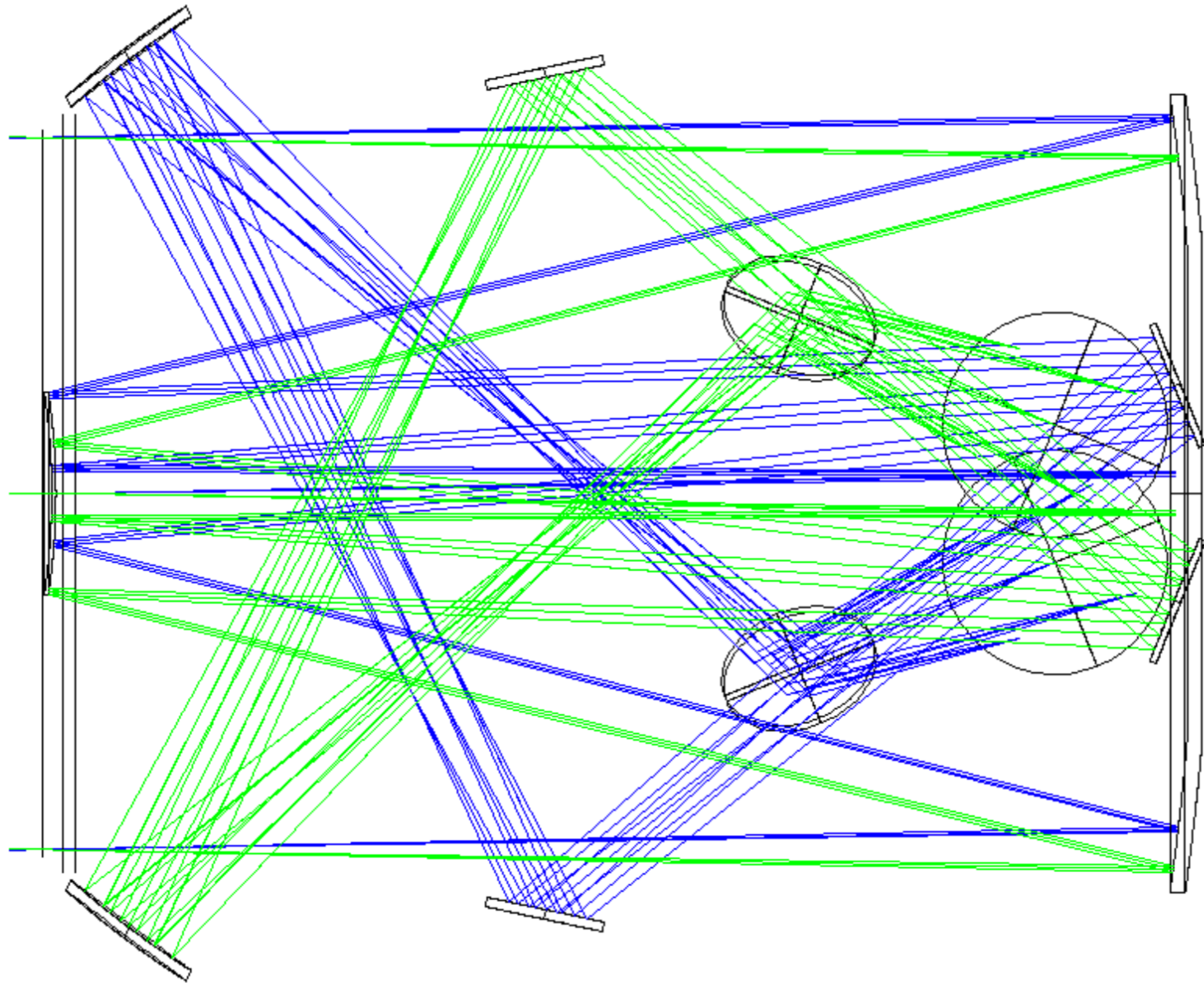


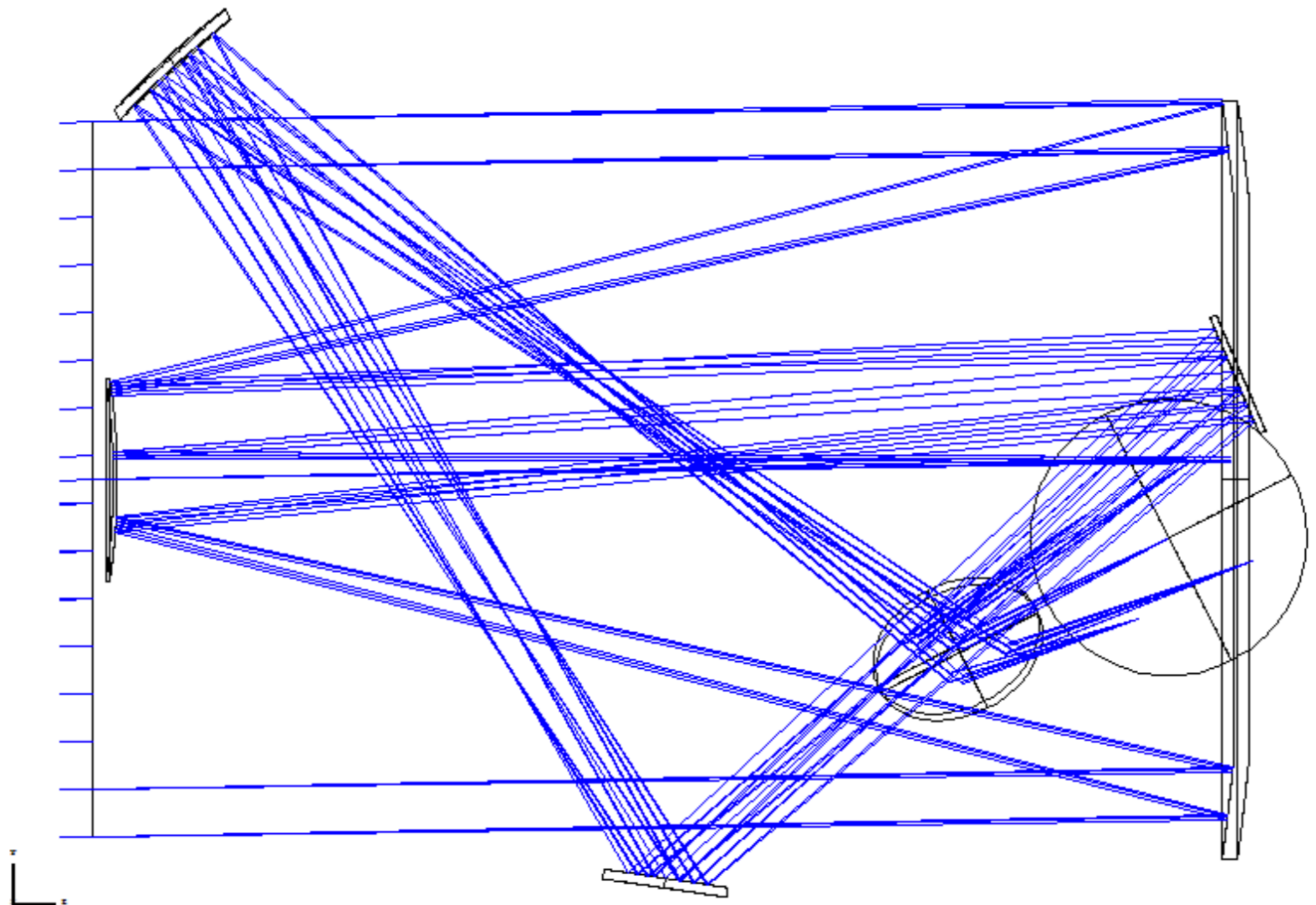




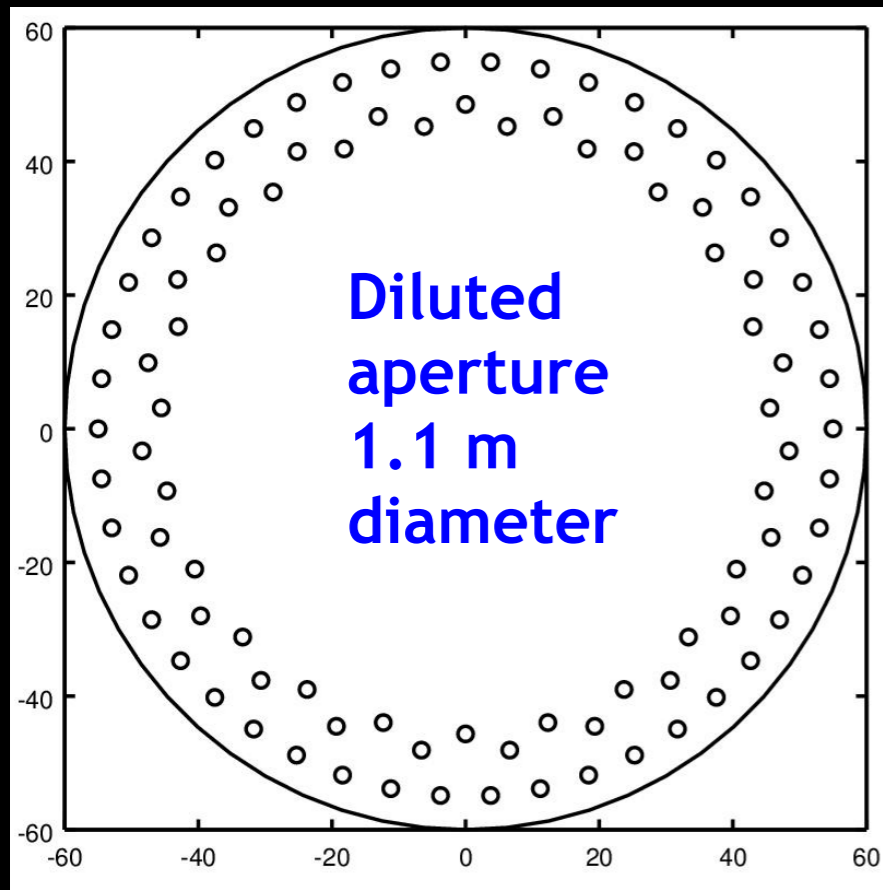








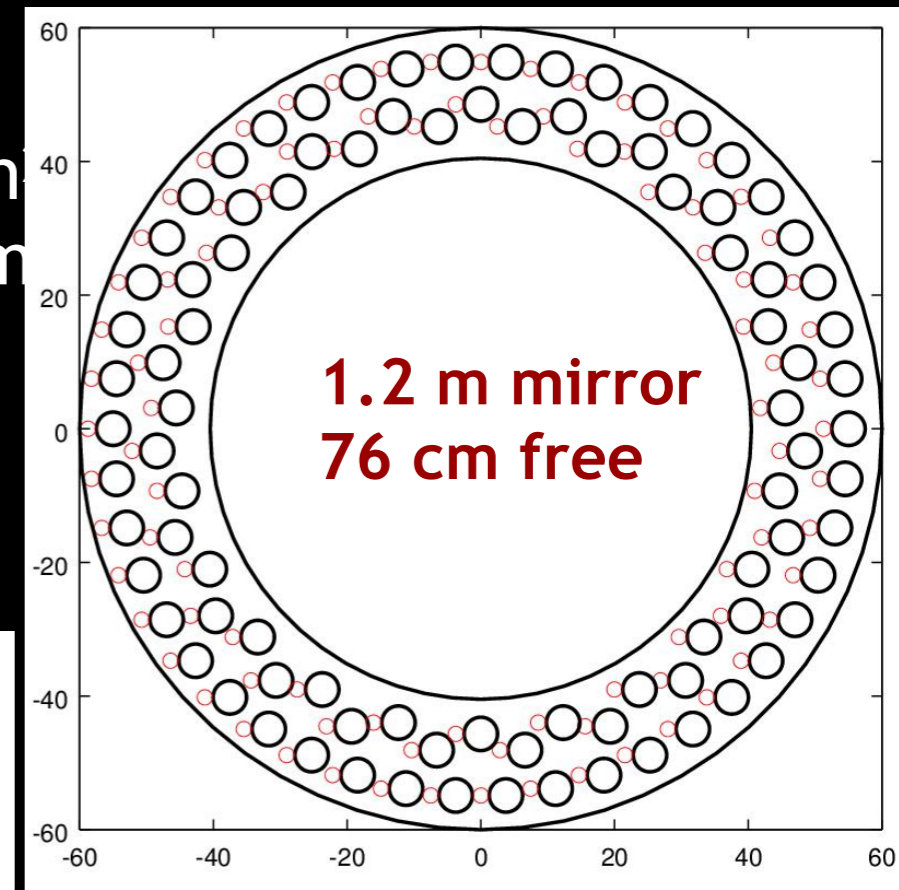
Coronagraphic system - multiple inverted occulter



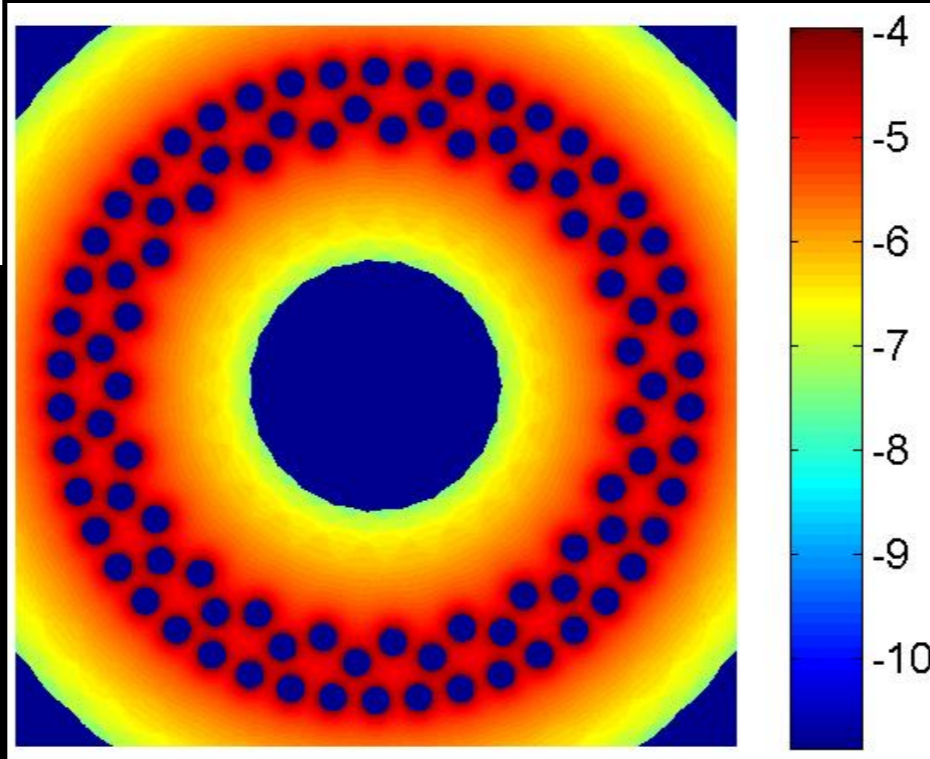
Collecting area

Sun direction: 375 cm

Anti-Sun direction: 1500 cm



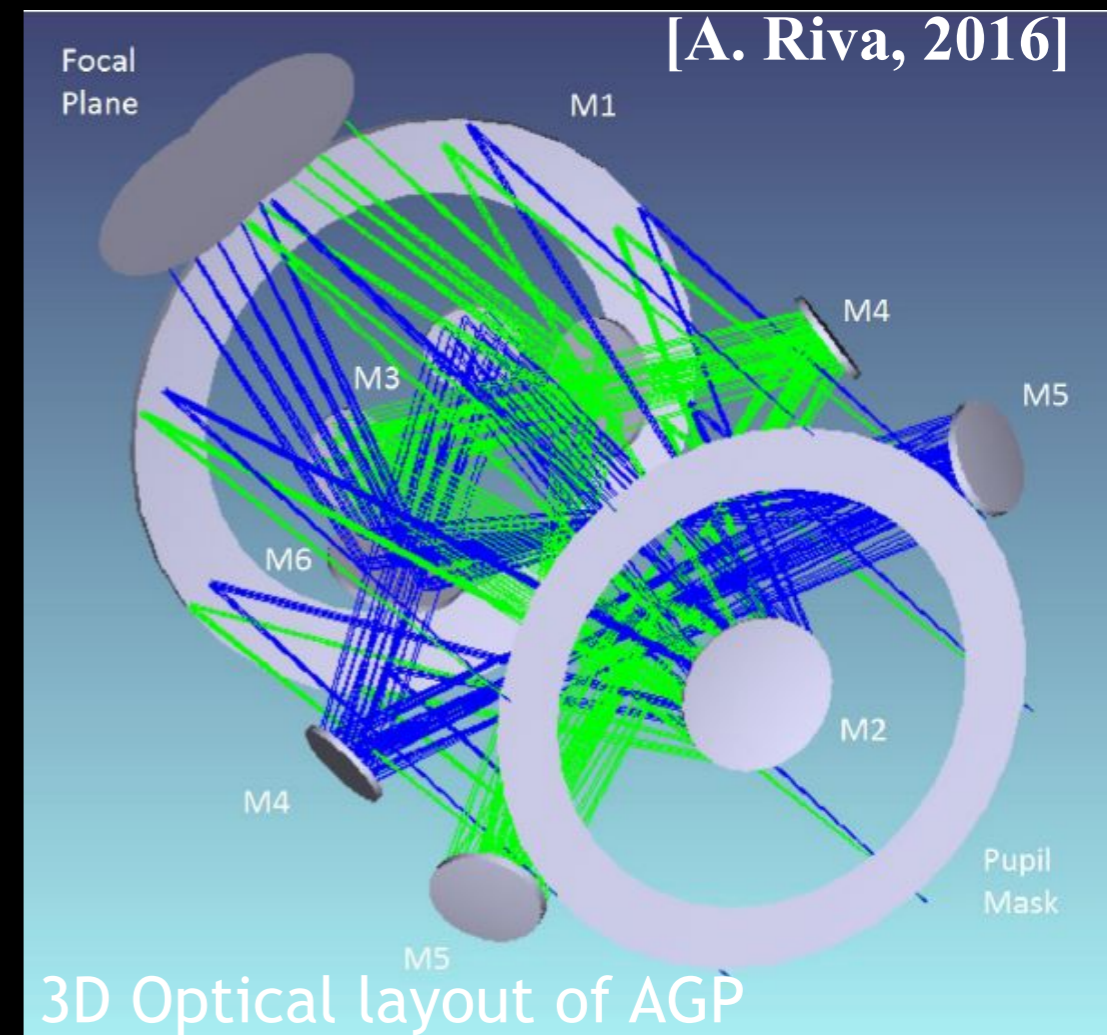
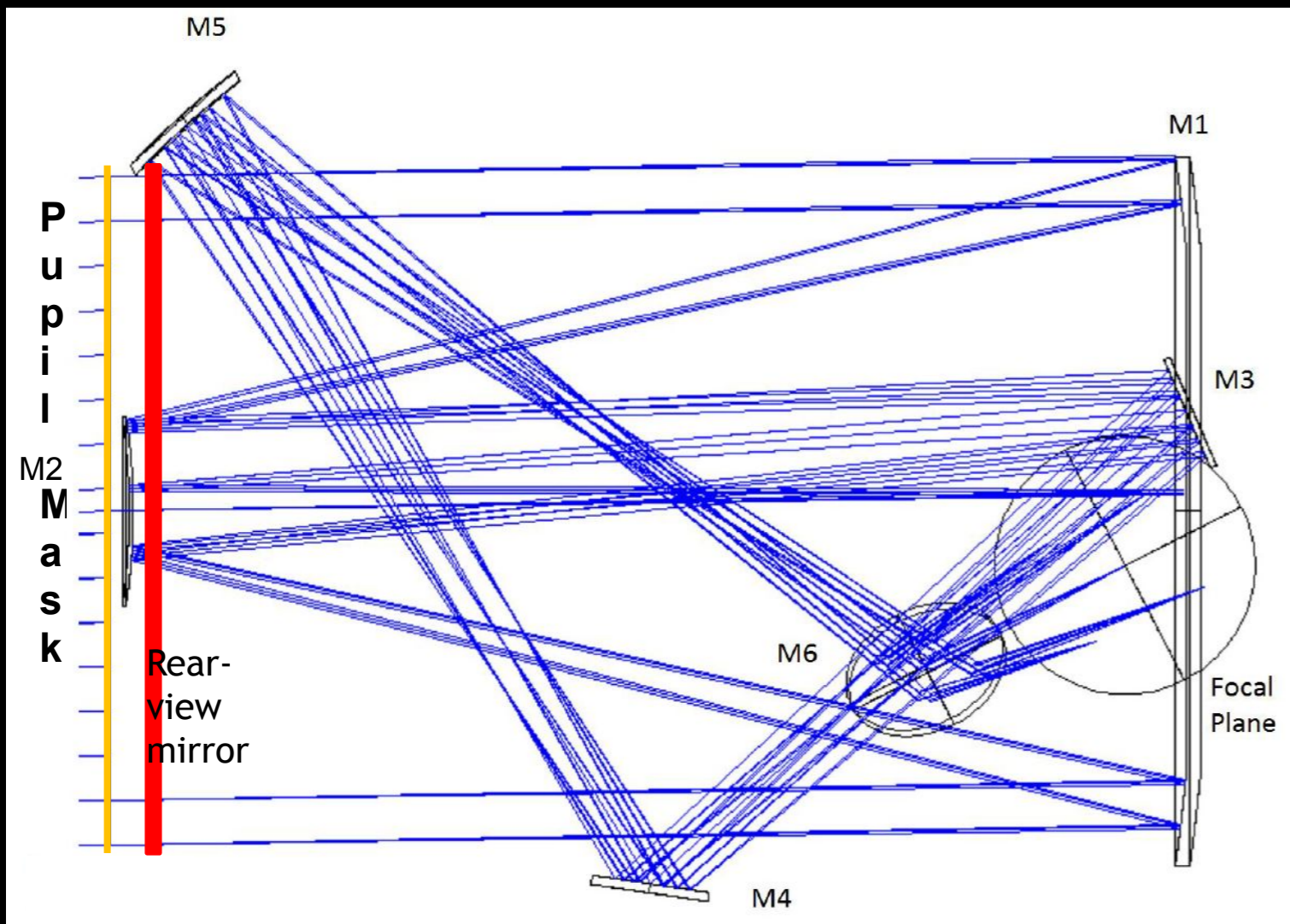
Diffraction on primary mirror [log units]



Background on focal plane: $\sim 10^{-7} B_{\text{Sun}}$

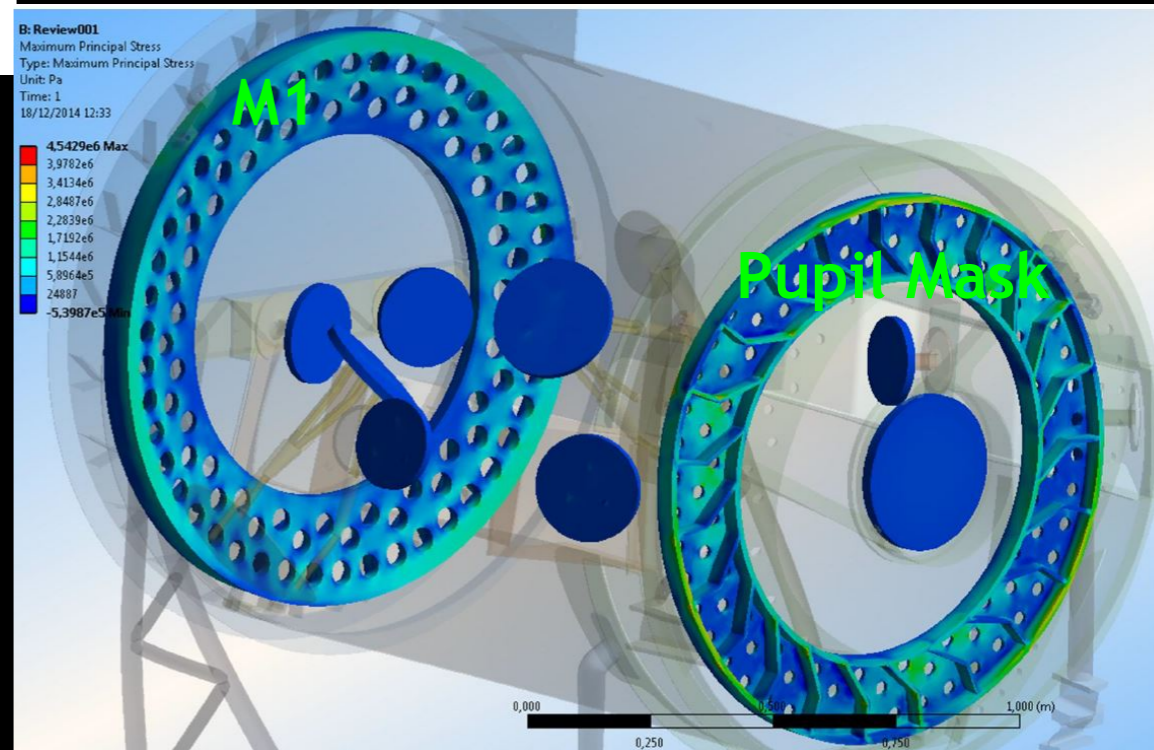
[F. Landini, 2016]

AGP Instrument design



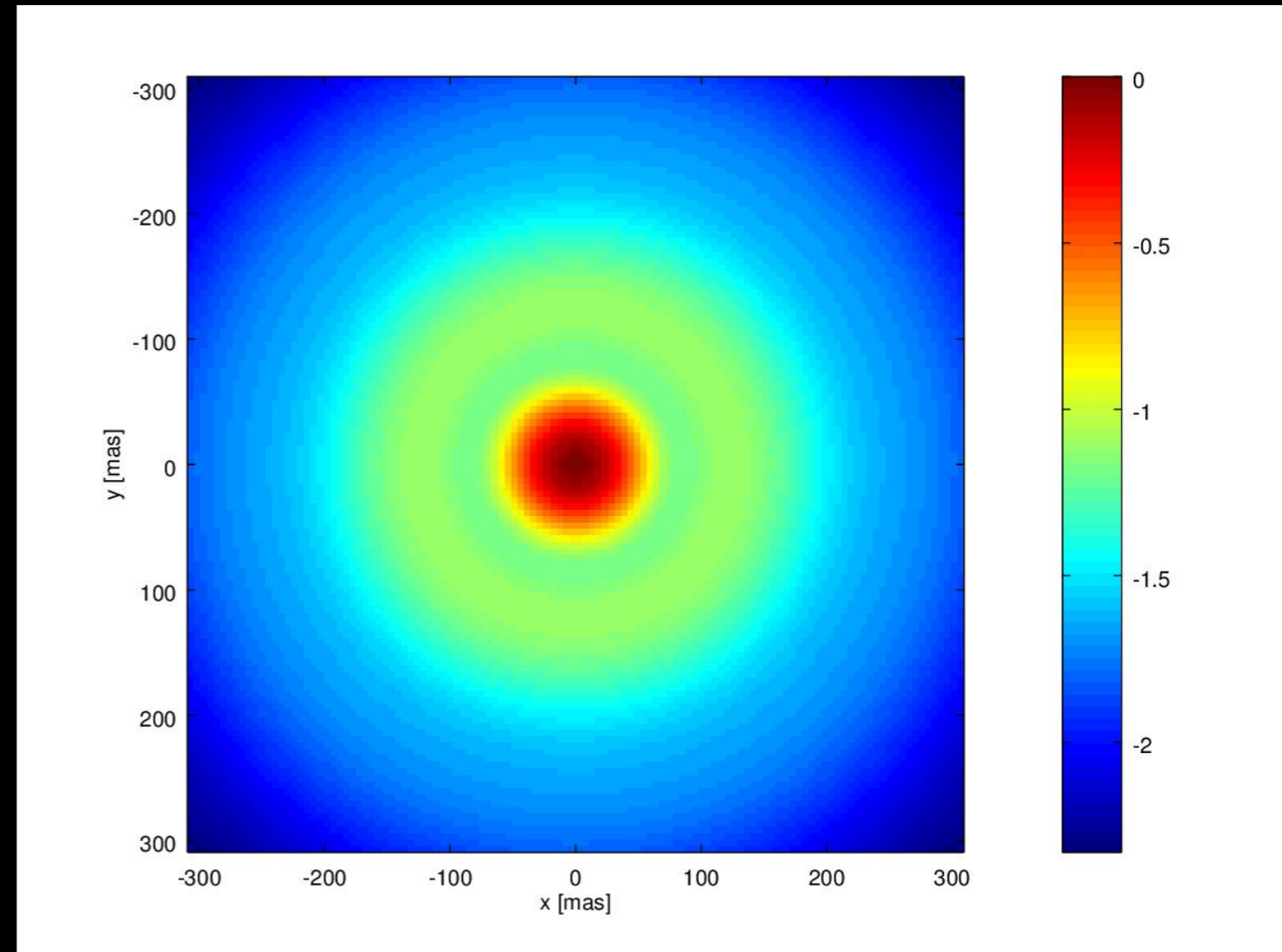
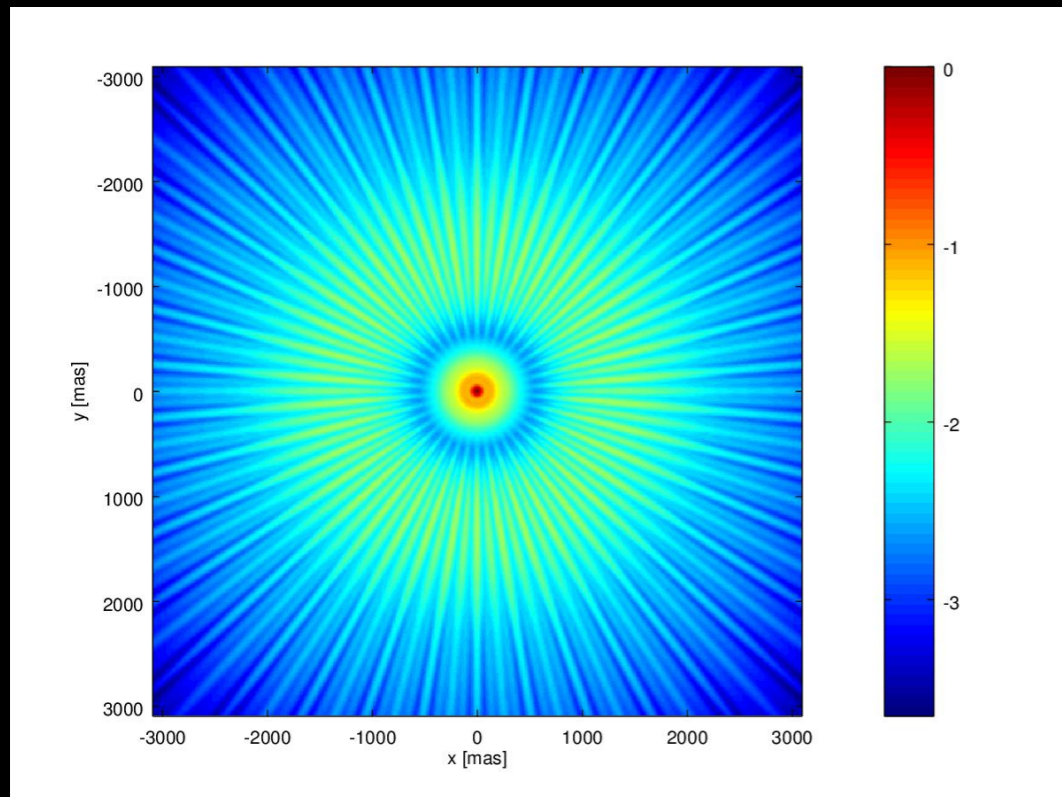
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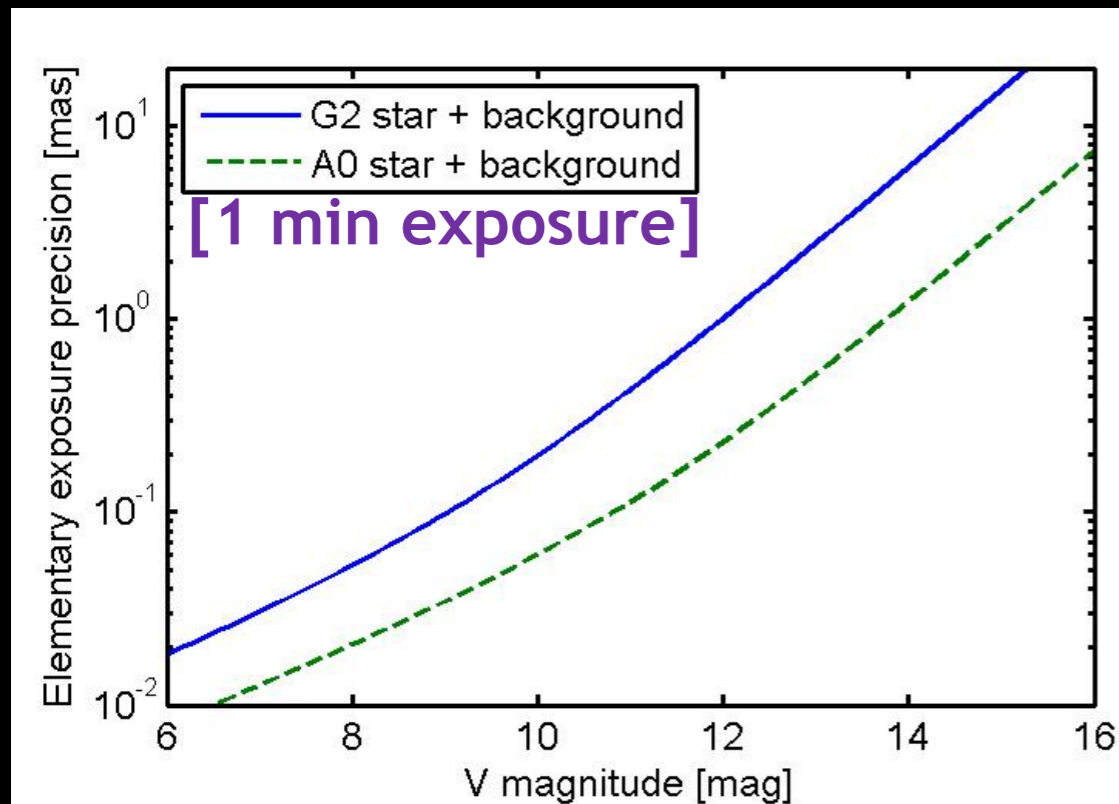
AGP Instrument design

AGP Point Spread Function for a solar-type star



Zoomed in the central region

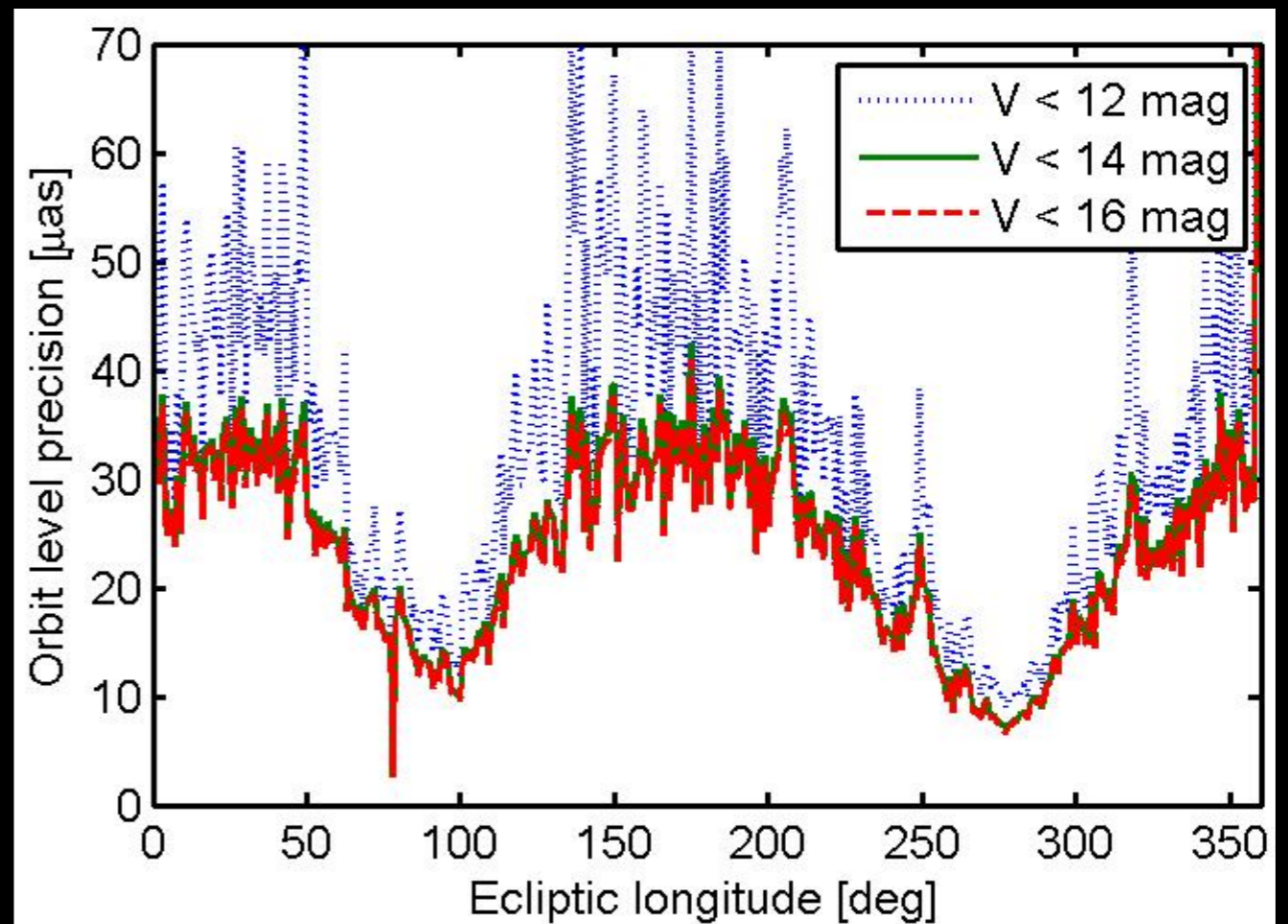
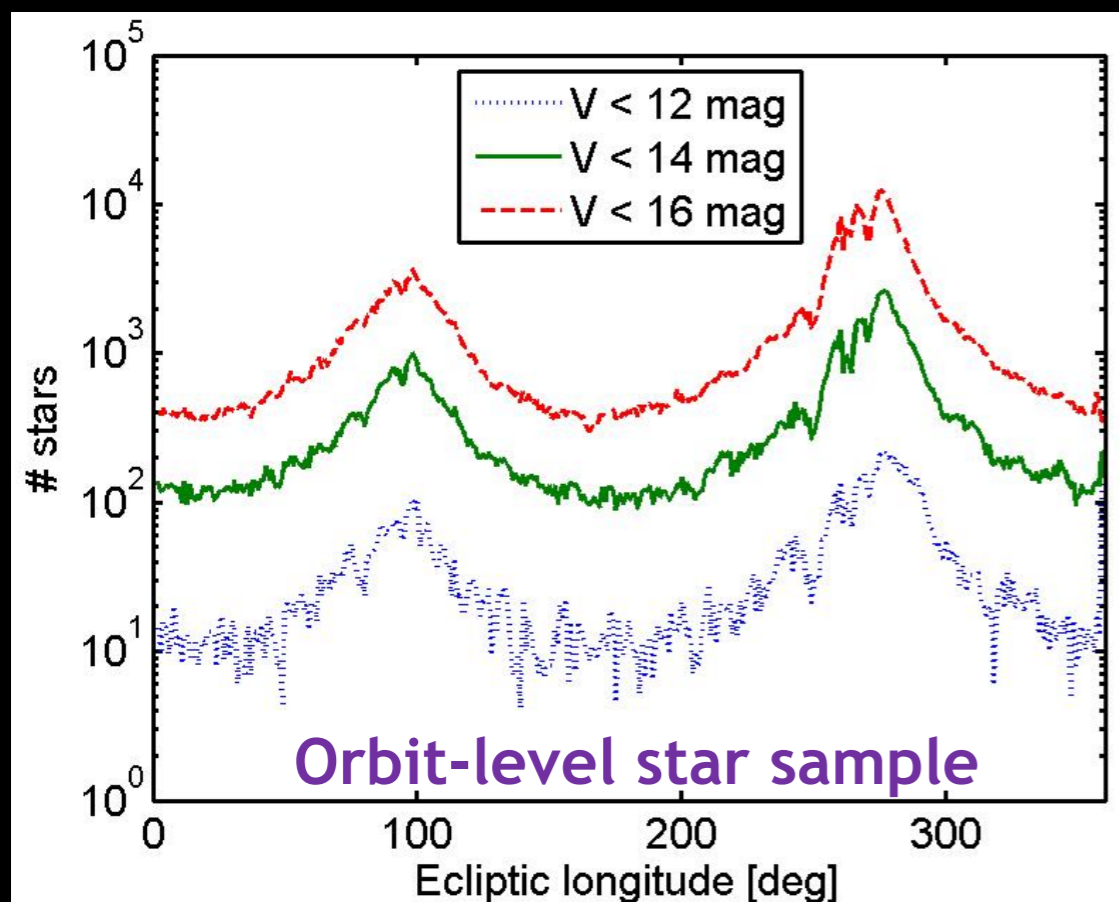
Individual location error and orbit-level performance



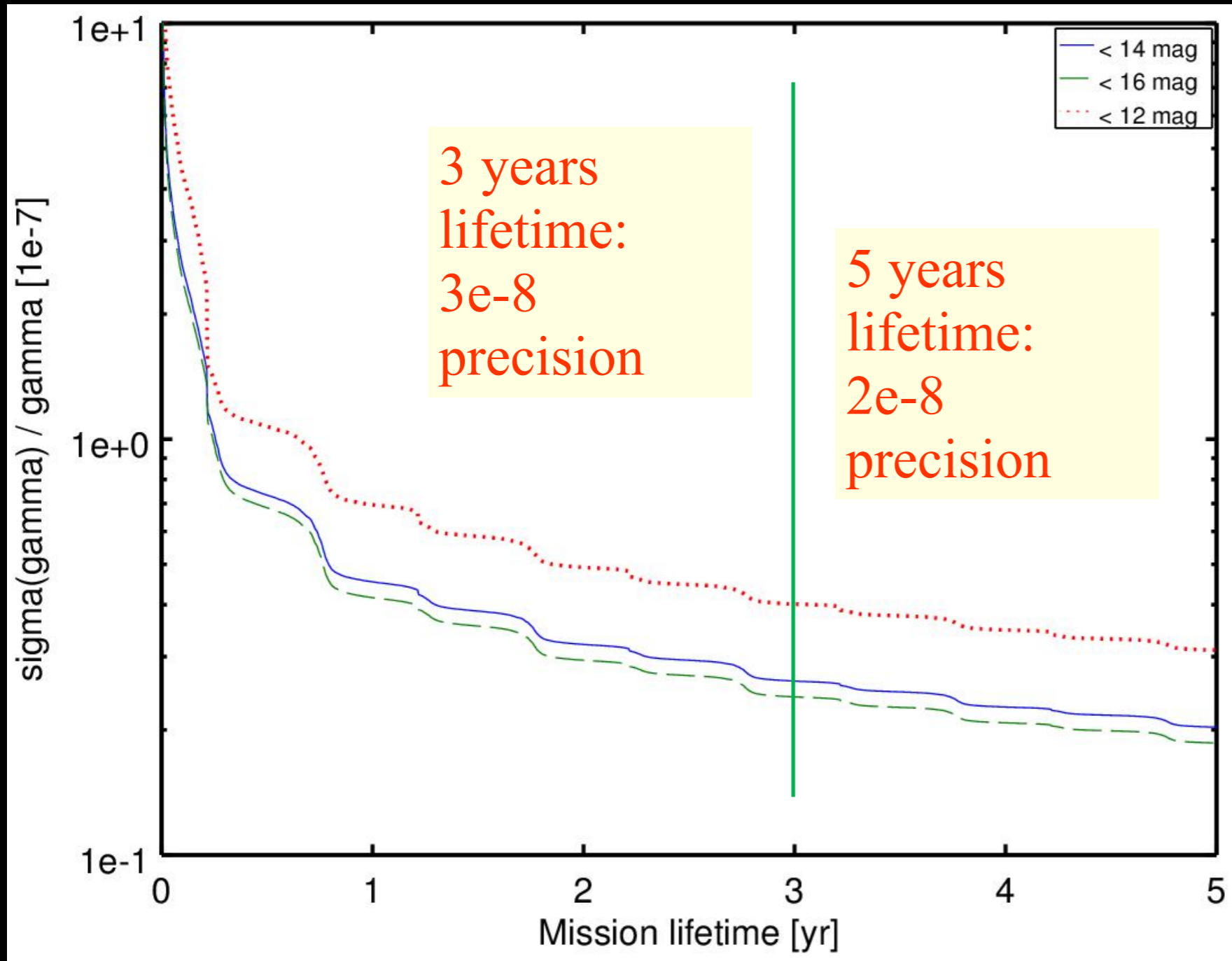
Observation close to Sun \rightarrow trade-off PSF accuracy / background / deflection

Measurement focused on bright stars $V < 16$

10^3 - 10^4 stars with cumulated precision $\sim 30 \mu\text{as}$ / orbit (+metrology/calibration)



Mission performance on light deflection



Deflection measurements averaged over ~ 5000 orbits/year

Simultaneous estimate of star and planet astrometric parameters

Years of mission	$obs.n. \times 10^{-7}$	$\sigma_{\beta} / \beta \times 10^6$
1	9.5	5.5
2	19.8	3.9
3	28.4	3.2
4	38.8	2.8
5	47.8	2.5

Tab. 1 – Accuracy for β and total number of observations throughout the mission lifetime.

Performance factors: $\sim D^2$, $FoV^{3/2}$, $t^{1/2}$

...but performance on orbits scales as $time^{3/2}$: factor > 2

Final message

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