

## On the origin of steep emissivity profiles in AGN accretion discs

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## Objective

X-ray observations suggest high compactness of coronæ in active galactic nuclei as well as in X-ray binaries. The compactness of the source implies a strong radial dependence in the illumination of the accretion disc. This will, for any reasonable radial profile of the density, lead to a radial profile of the disc ionization. Svoboda et al. (2012) showed on a single example that assuming a radially structured ionization profile of the disc can cause an artificial increase of the radial emissivity parameter. We further investigate how the X-ray spectra are modified and quantify this effect for a wide range of parameters.





The radial ionization profile  $\xi(r)$ , for various lamp-post heights. We assume the same inner ionization parameter  $(\log \xi_{in} = 2.5)$  for all the cases.

The innermost regions are highly ionized. The outer regions are neutral  $\Rightarrow$  The resulting spectrum is the sum of the contributions from each radius and might affect the radial emissivity parameter.

## Method



## Results



- For low and large values of  $\xi_{in}$ , the parts of the disc contributing the most to the observed spectrum would be either neutral or highly ionized  $\Rightarrow$  a small gradient of ionization would be expected in these parts of the disc  $\Rightarrow$  constant-ionization approximation holds  $\Rightarrow$  lower values of q.
- For the intermediate values of  $\xi_{in}$ , the gradient of ionization is more important. The innermost regions will be more ionized and will have softer reflection spectra. The outer regions of the disc are less ionized  $\Rightarrow$  model assuming a single ionization parameter of the disc will underestimate the ionization from the innermost regions by assuming an average ionization.  $\Rightarrow$  This effect will then be compensated by requiring a steep emissivity profile.