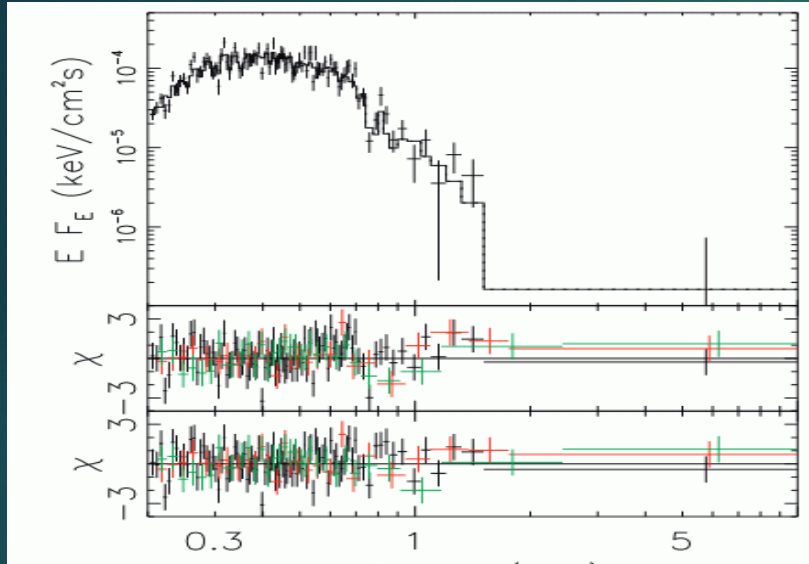




Comptonisation in TDE

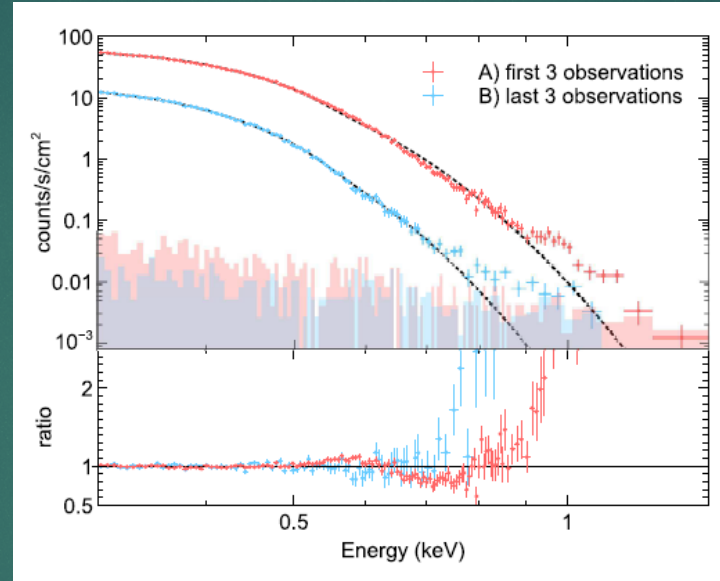
R. SAXTON, S. KOMOSSA, K. ALEXANDER, A.M. READ, P. LIRA, P. JONKER, S.
SMARTT, F. FUERST, G. MINIUTTI, M. GIUSTINI, T. WEVERS

Early-phase X-ray spectra



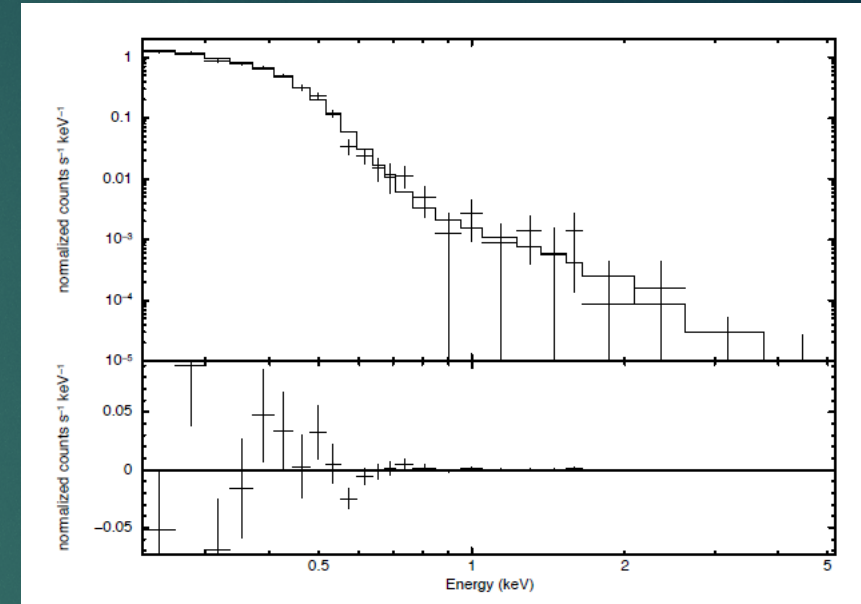
3XMM J 1521+07 - Lin et al. 2015

$kT=99\pm 3$ eV



ASASSN-14li – Kara et al. 2018

$kT=57\pm 1$ eV

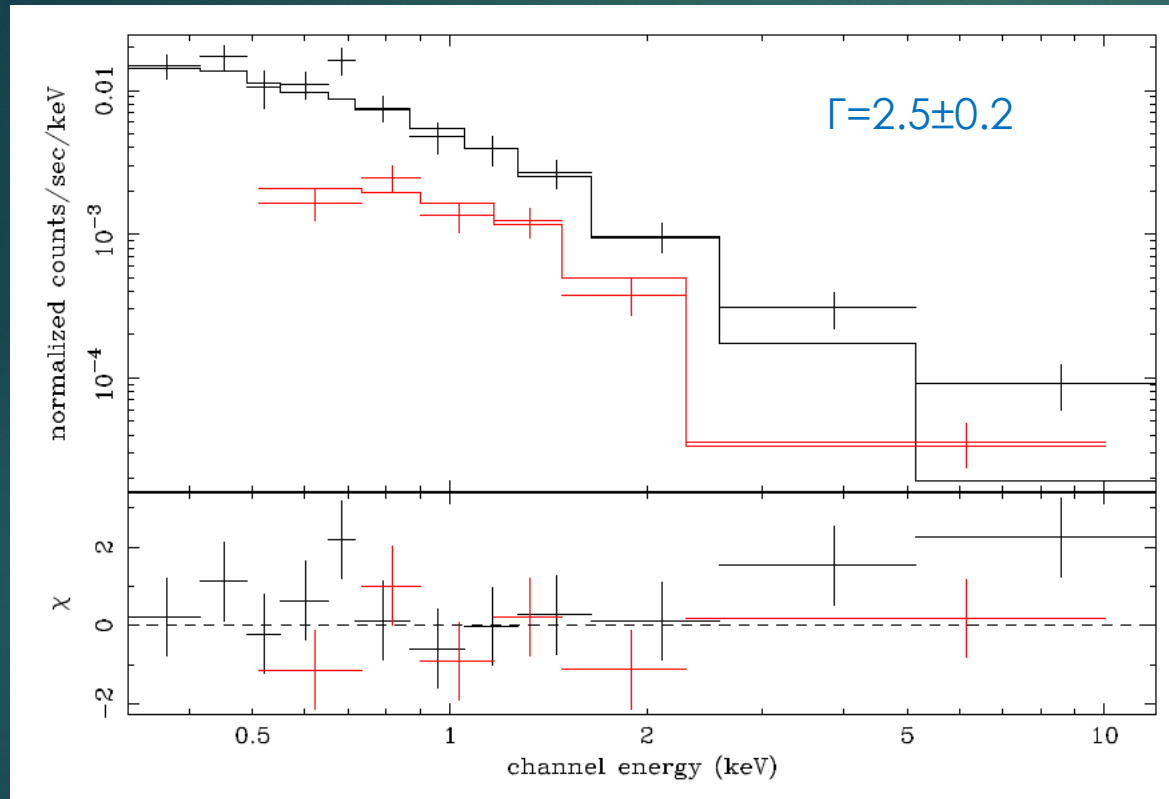


ASASSN-15oi – Gezari et al. 17

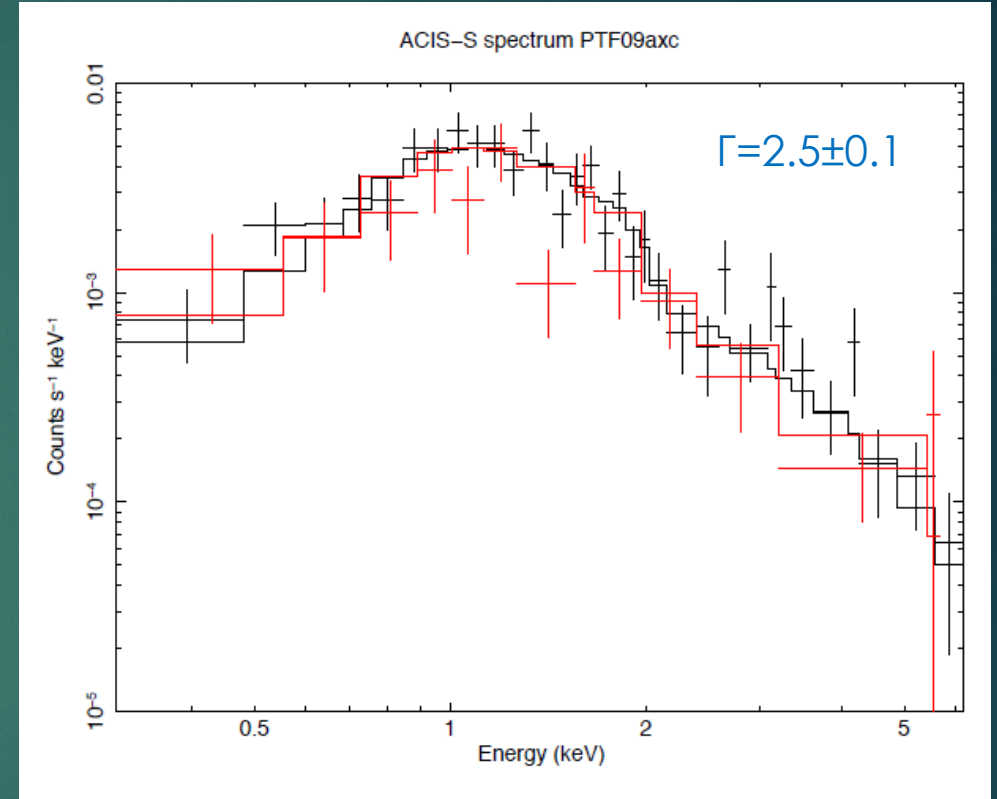
$kT=42\pm 1$ eV

Initially TDE X-ray spectra tend to be dominated by thermal emission with **$kT\sim 40\text{-}100$ eV**.

Late-phase X-ray spectra



RXJ~1242.6-1119 XMM-Newton spectrum taken 10 years after discovery – Komossa et al. 2004

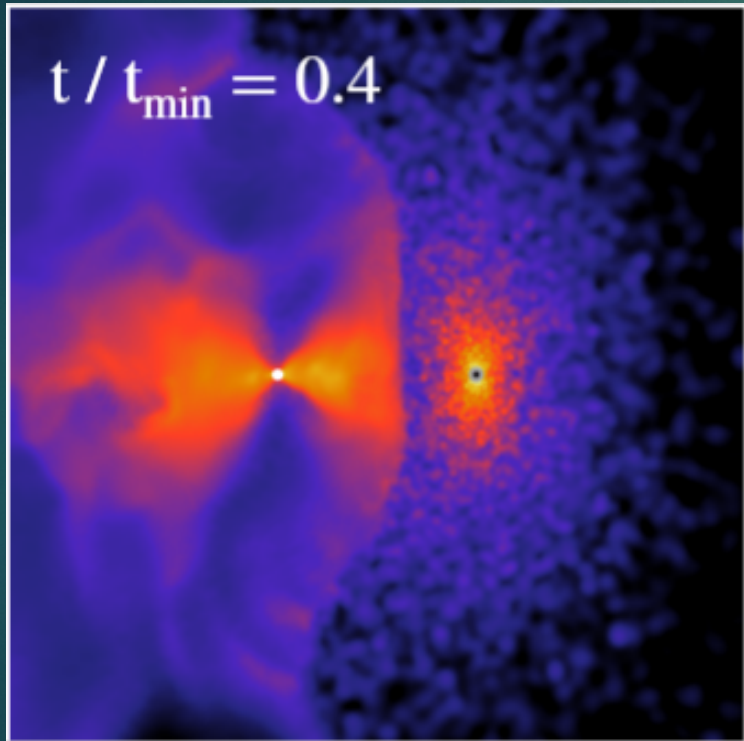


PTF09axc, Chandra spectrum, 8 years After discovery – Jonker et al. 2019

After a few years the spectrum is dominated by a steepish power-law component $\Gamma \sim 2.5-3.5$

What causes the transition?

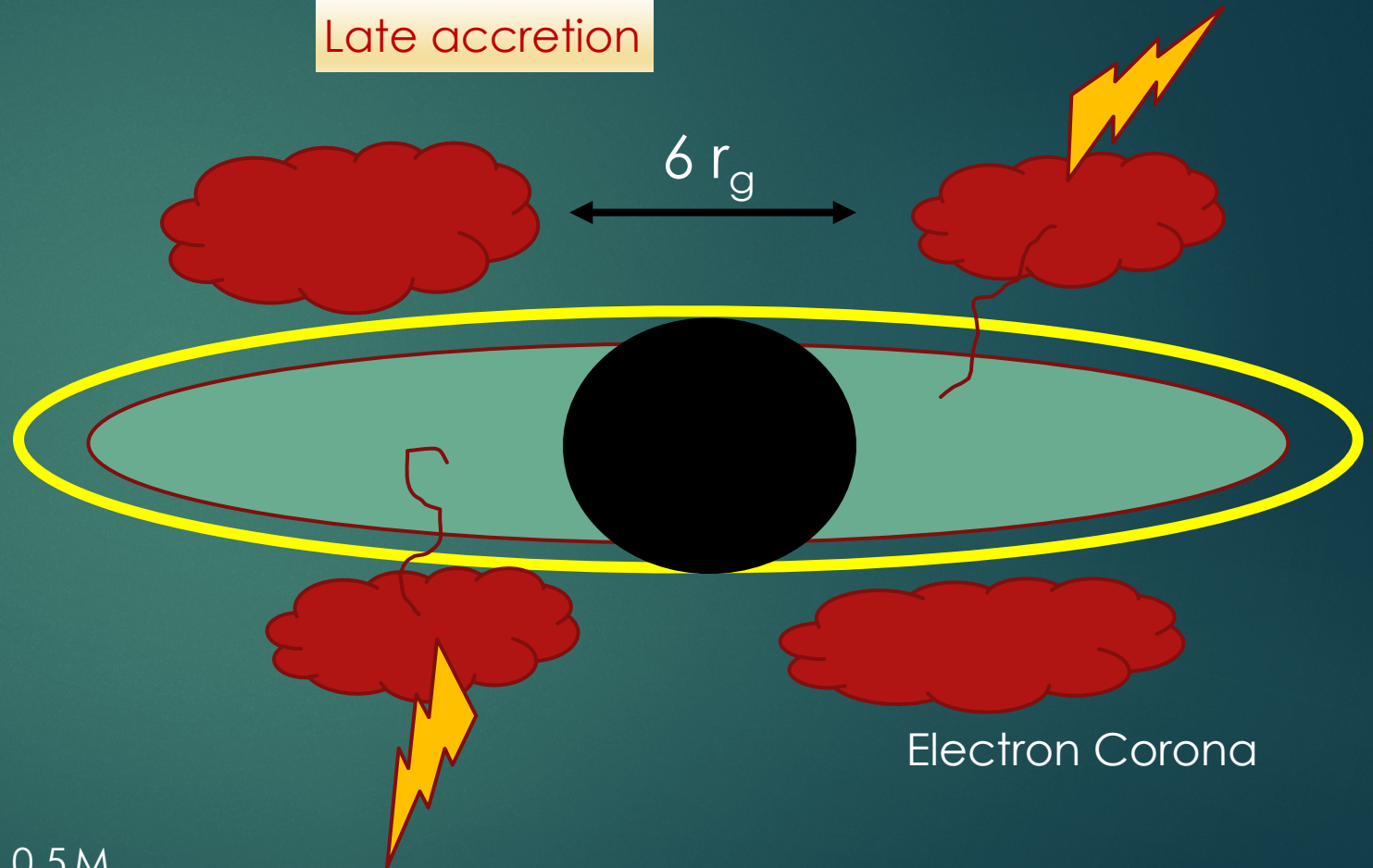
Early accretion



Bonnerot & Lu 2019

$M_{\text{BH}} = 2.5 \times 10^6$, $M_* = 0.5 M$,
16 days after
stream interaction

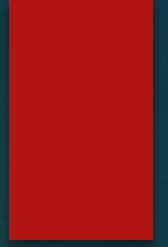
Late accretion



Electron Corona

Cannizzo, Lee & Goodman 1990

How and when does transition occur ?



? Is Comptonising corona always there but becomes increasingly important as thermal diminishes ?

? If corona develop later does it build up slowly or quickly ?

New source: XMMSL2 J1404-25

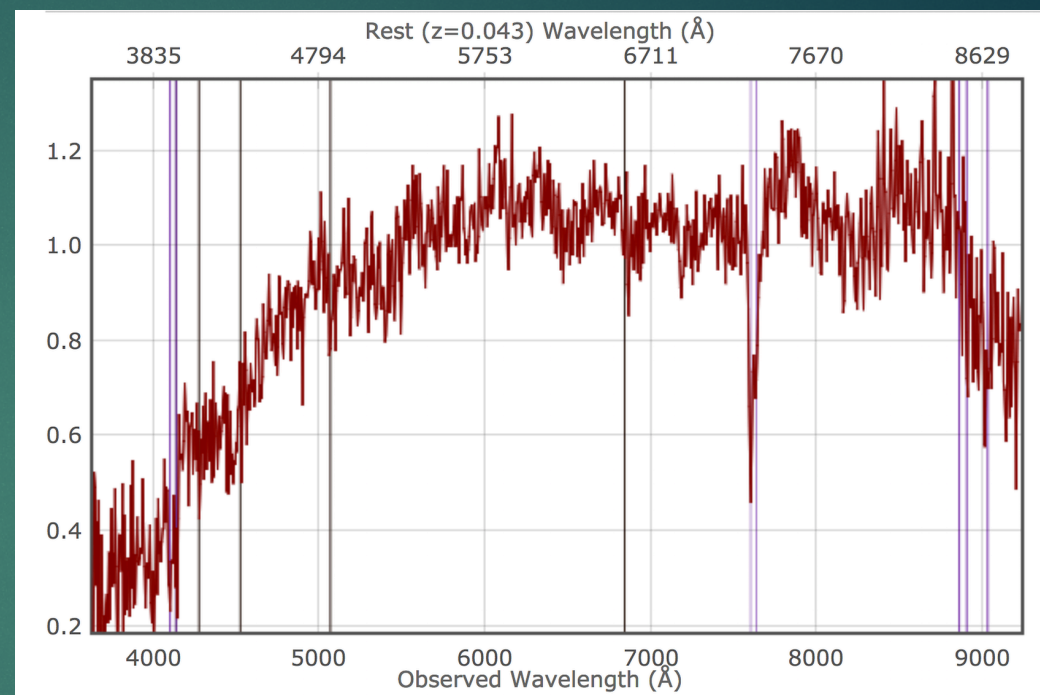


Discovery image – 15 Feb 2018

Atel #11394

3.4 s, 13 photons

>20 brighter than ROSAT in 1990



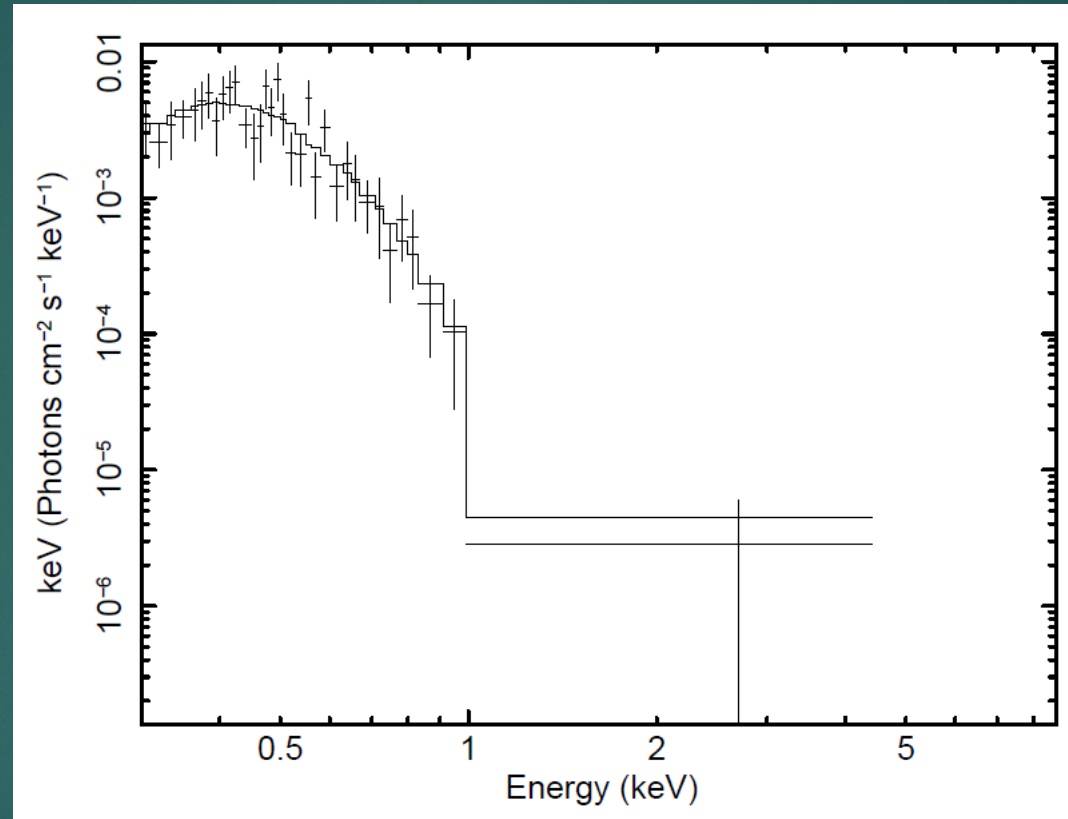
ePESSTO NTT, EFOSC2 spectrum (8 Mar 2018)

Atel #11395

2MASX 1404671-1511433

$z=0.043$

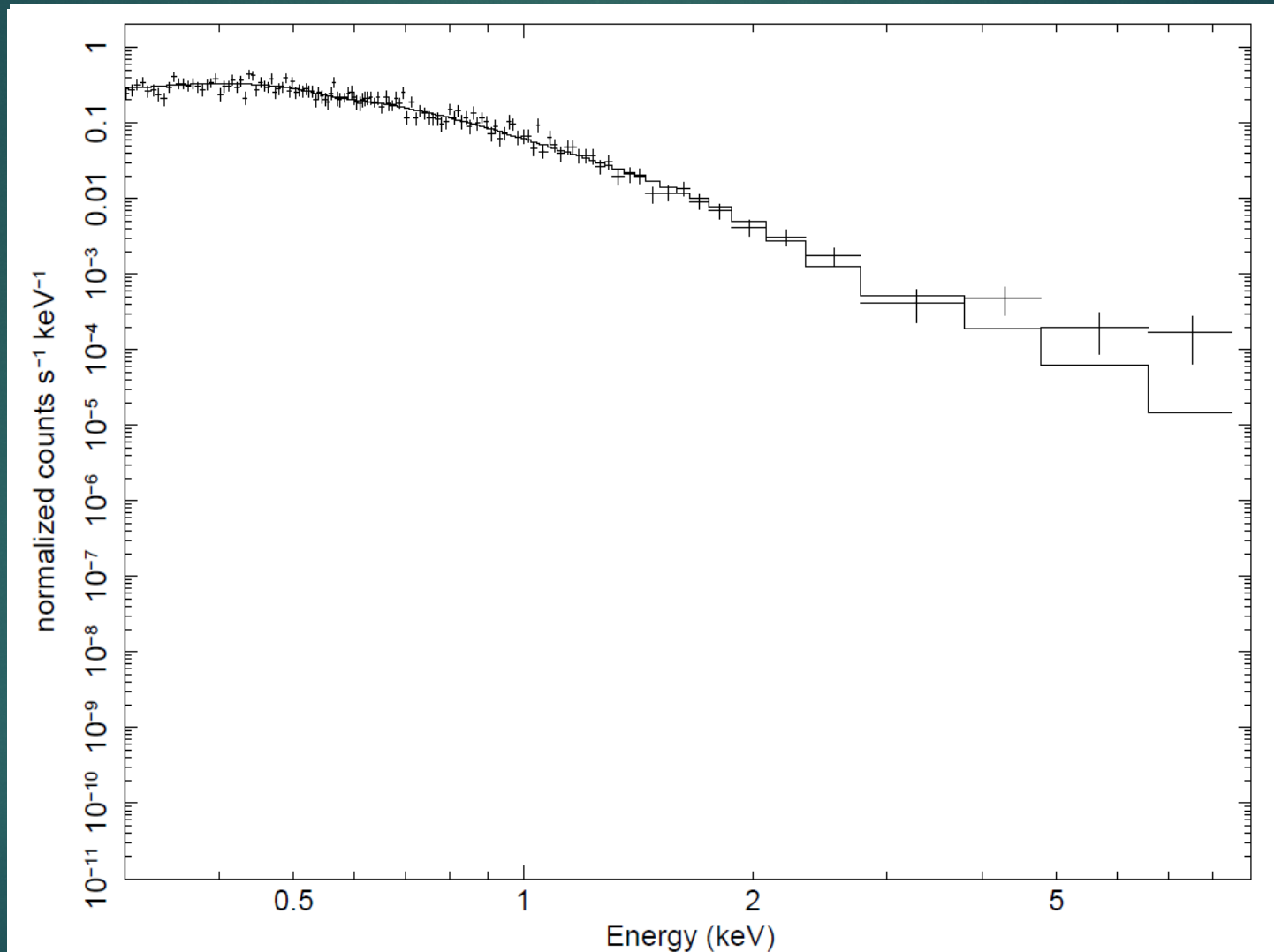
New source: initial spectrum



Black-body
 $kT=77\pm 2$ eV

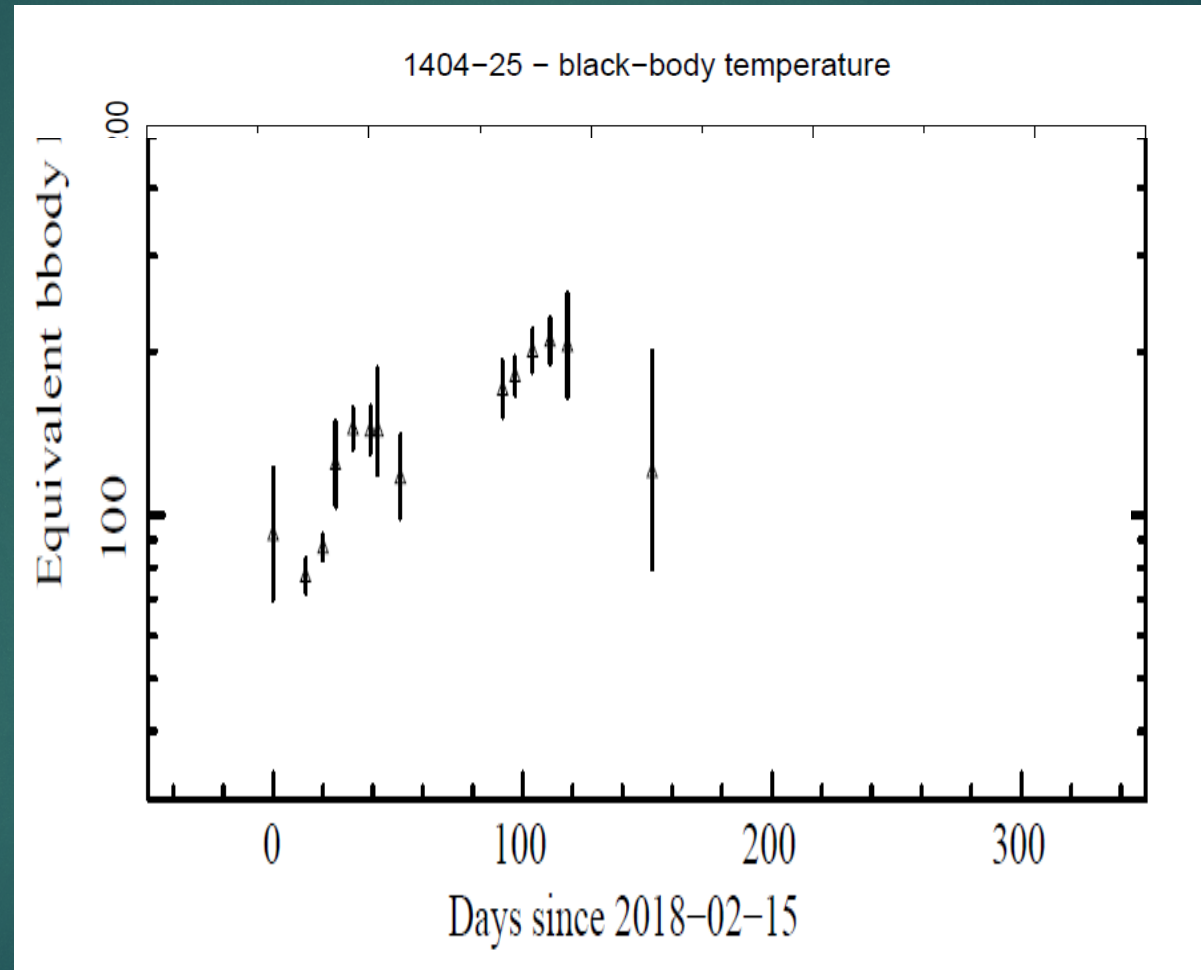
Swift-XRT, 28 Feb 2019

Changing spectrum



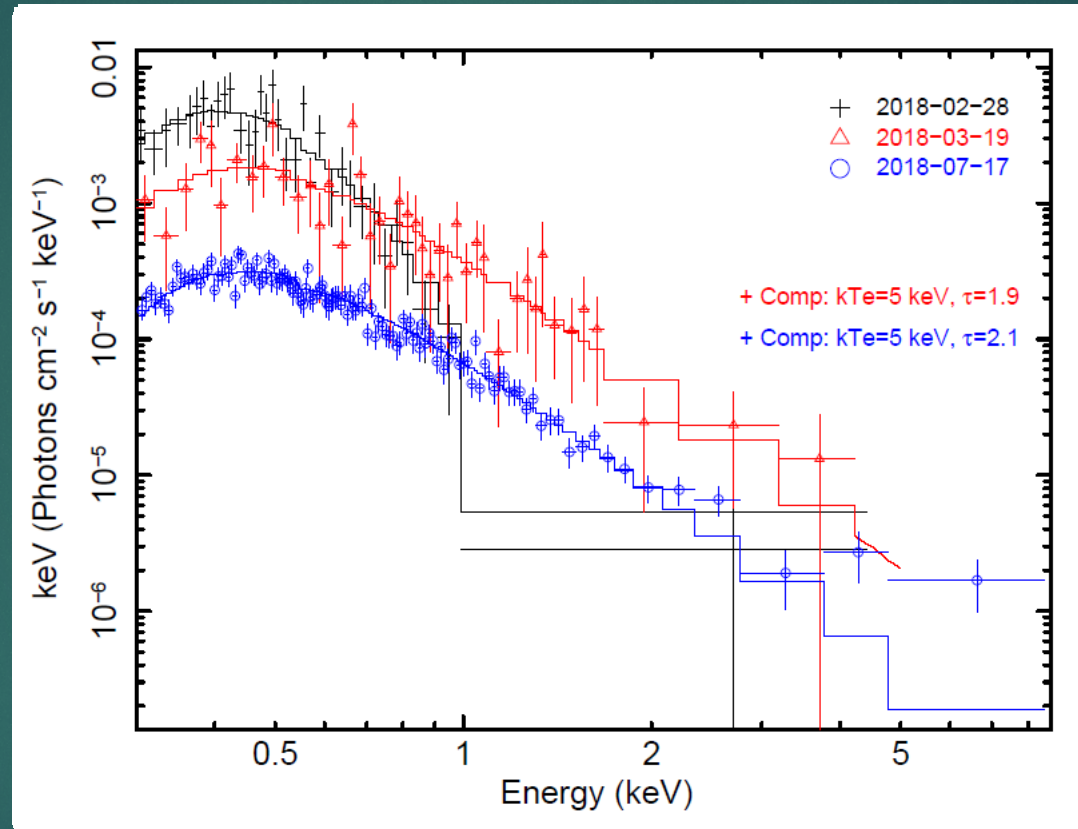
(fit with TBABS*COMPBB)

New source: hardness



Evolution of equiv temperature

New source: initial spectrum



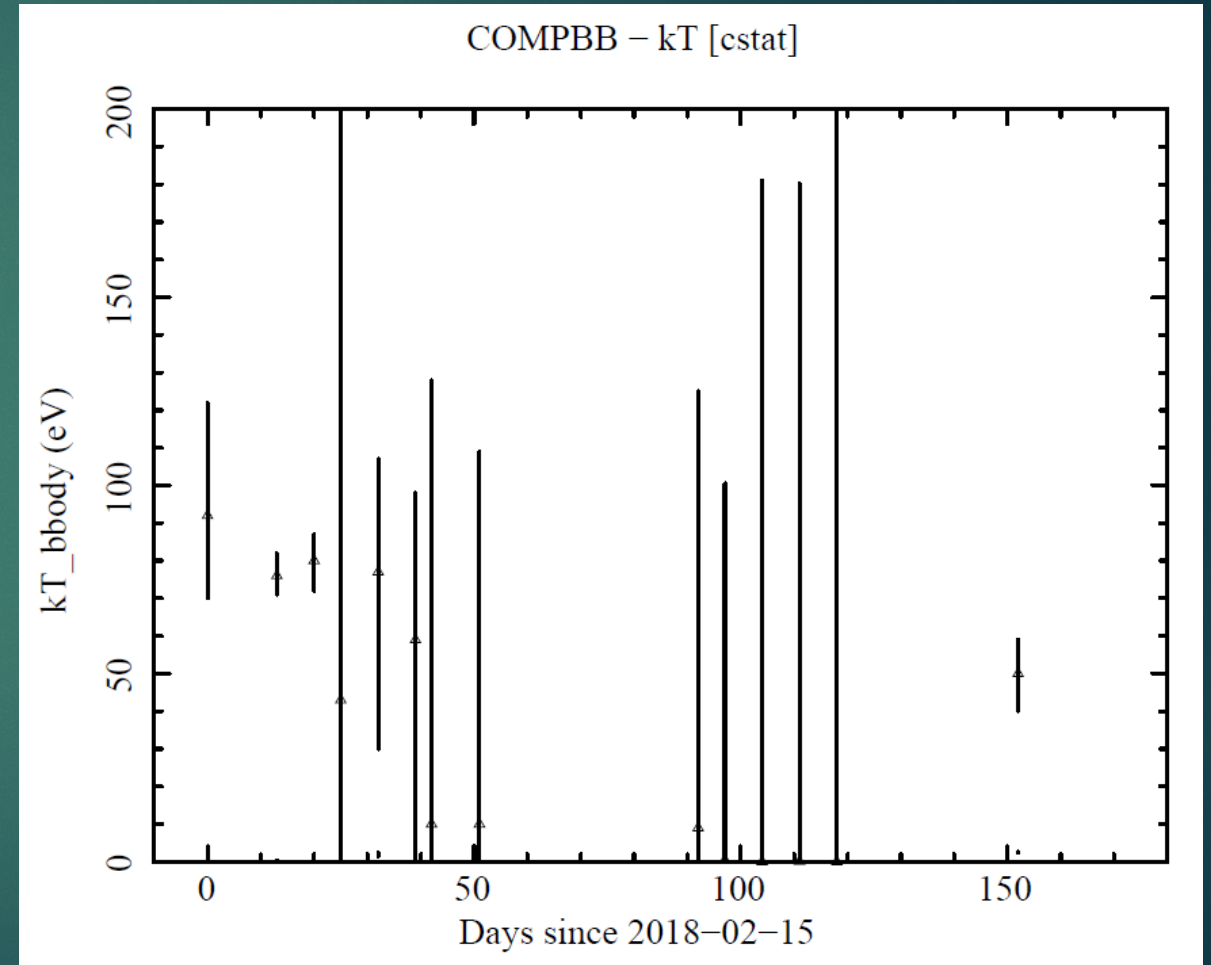
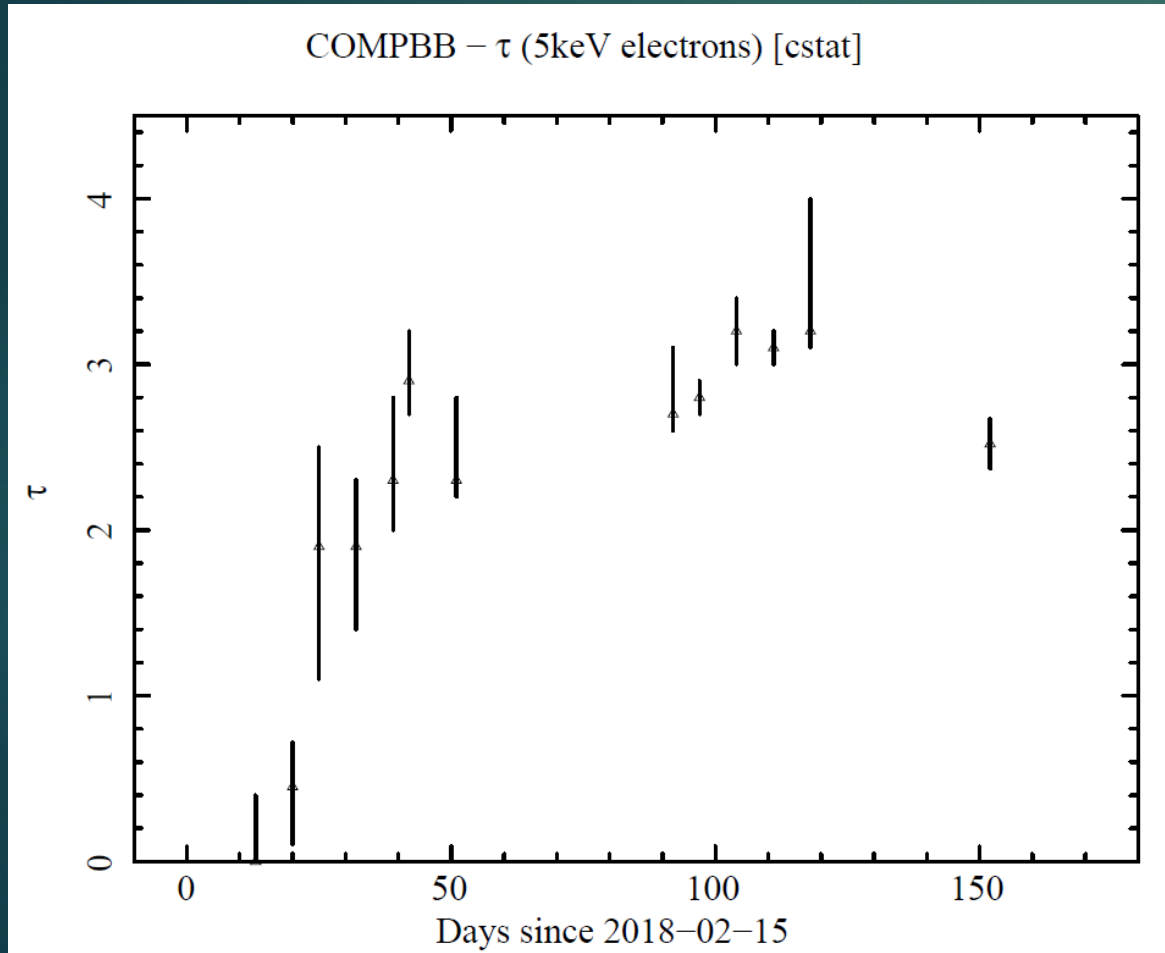
Black-body
 $kT=77\pm 2$ eV

Comptonisation:

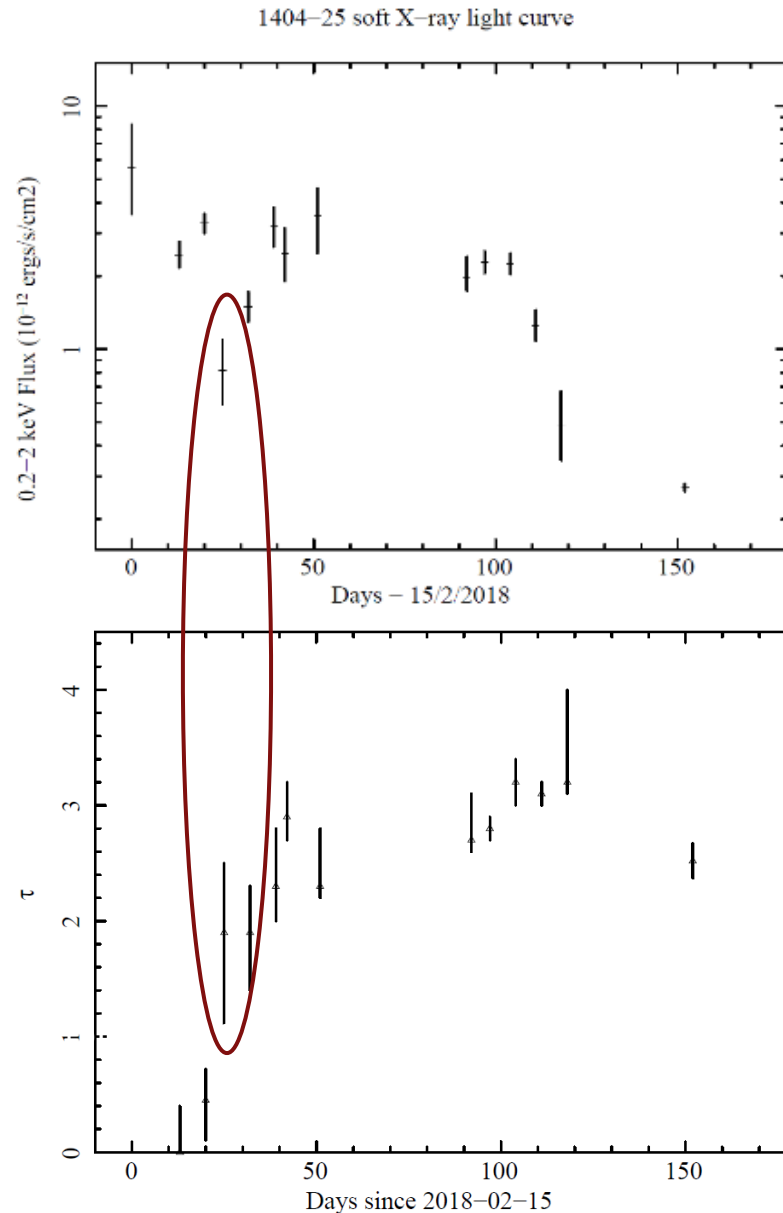
$\tau=1.9$
 $kTe=5$ keV

Developed within
7 days.

New source: Compton evolution



Sudden Spectral Change ?



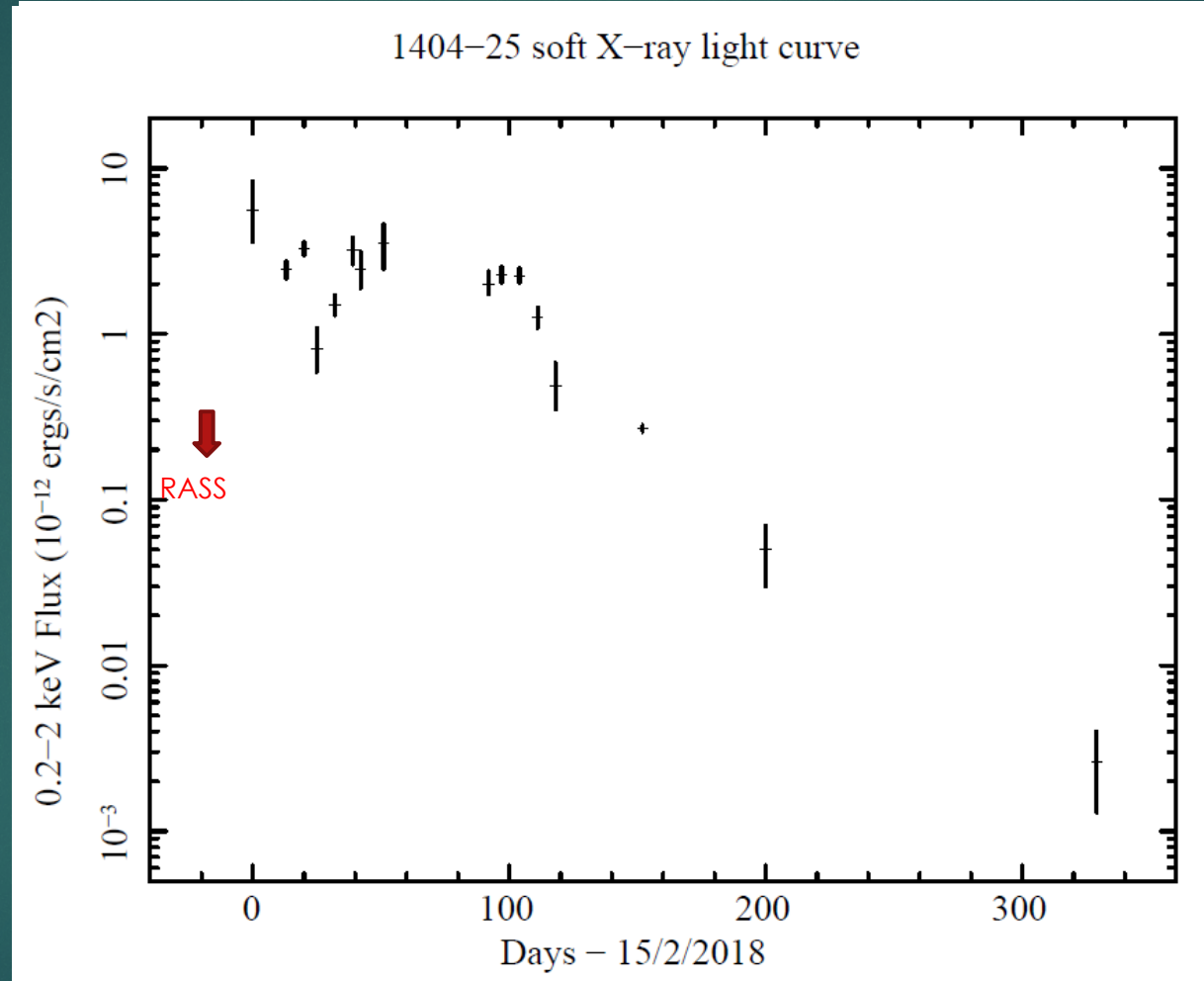
Electron optical depth

Sudden change between
Swift obs 2 and 3 ?

Coincides with factor 4
Dip in flux ?

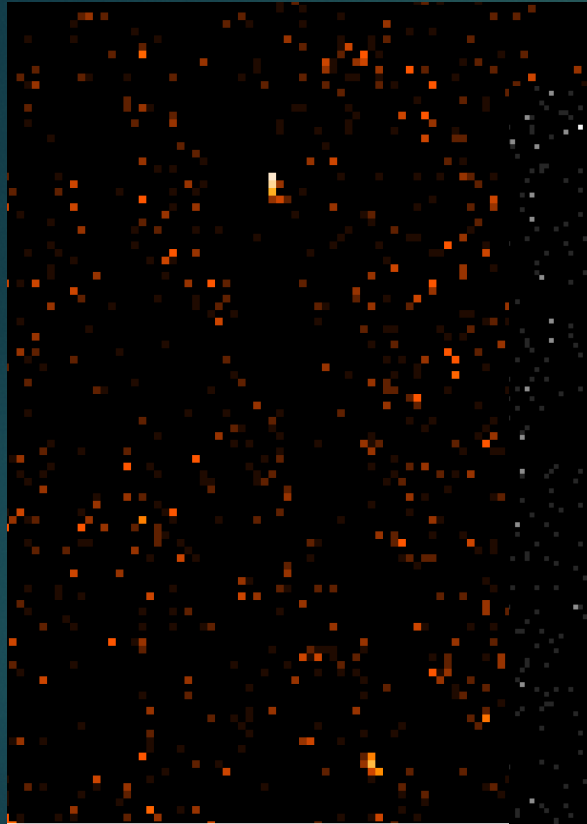
No obvious increase in
Obscuration (NH)

New source: evolution of light curve

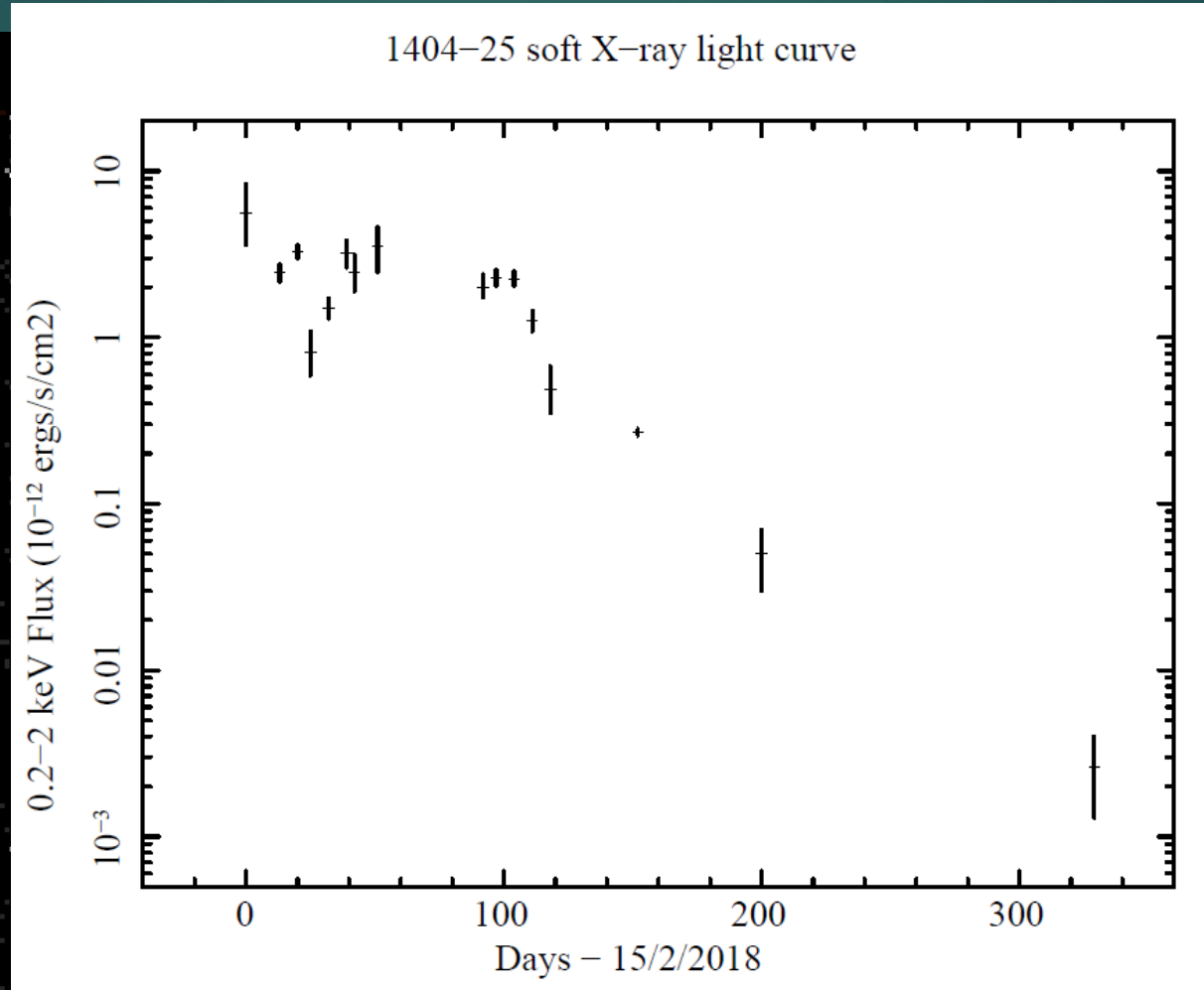


0.2-2 keV observed light curve

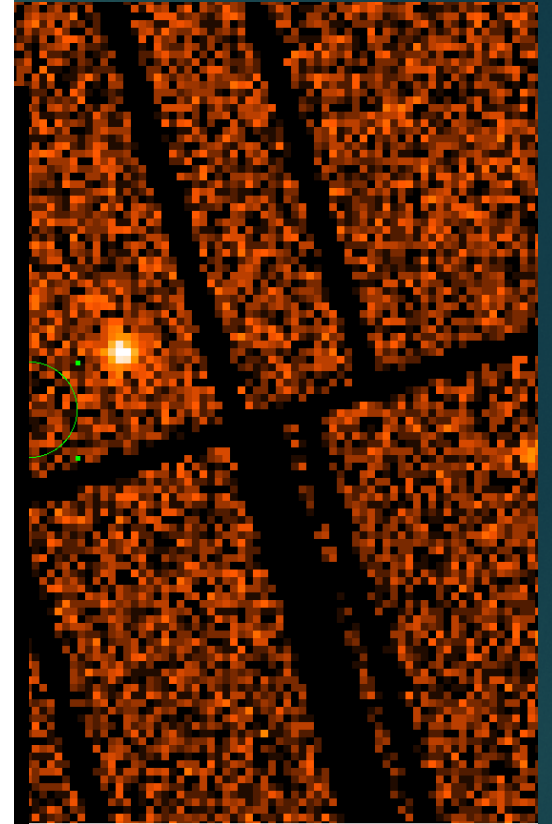
New source: 320d after discovery



EPIC-pn, 0.2-2 keV

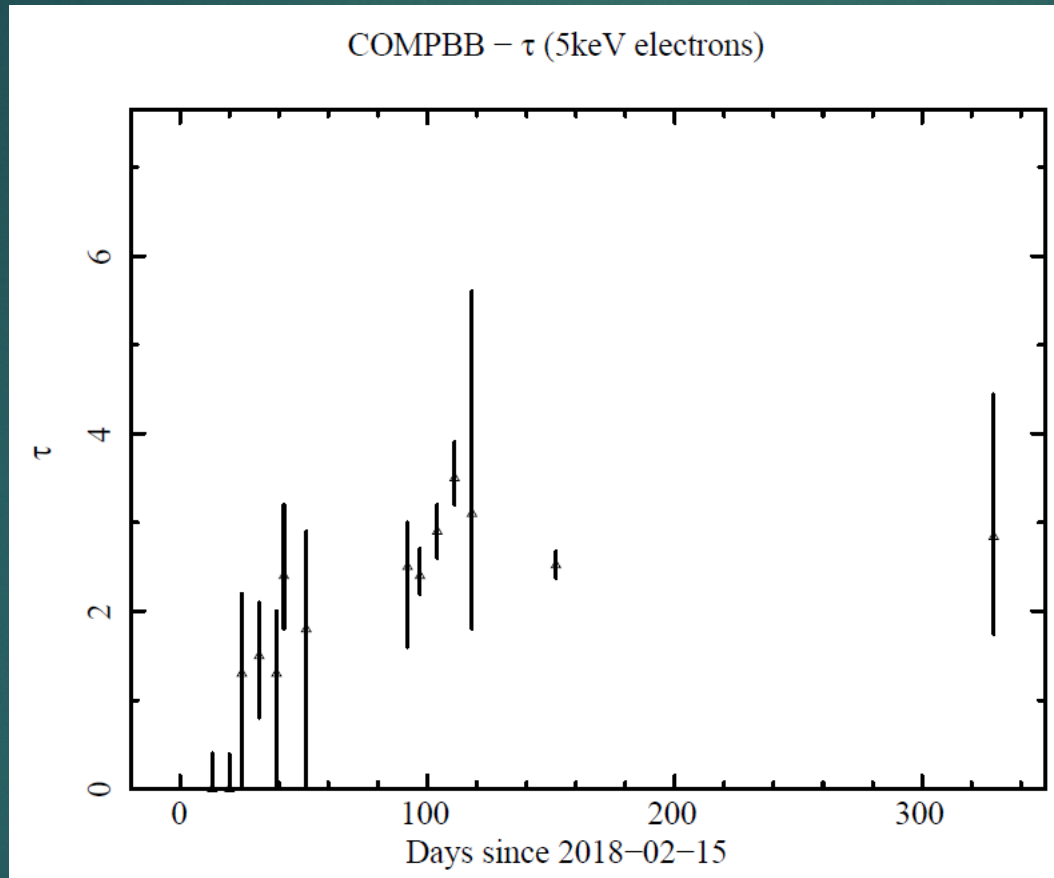


NuSTAR A, 3-100 keV



Shown to be intrinsic – thanks NuSTAR

Full evolution of electron optical depth



Build up to peak electron optical depth after 120d.

Maintained until day 320.

Black hole mass

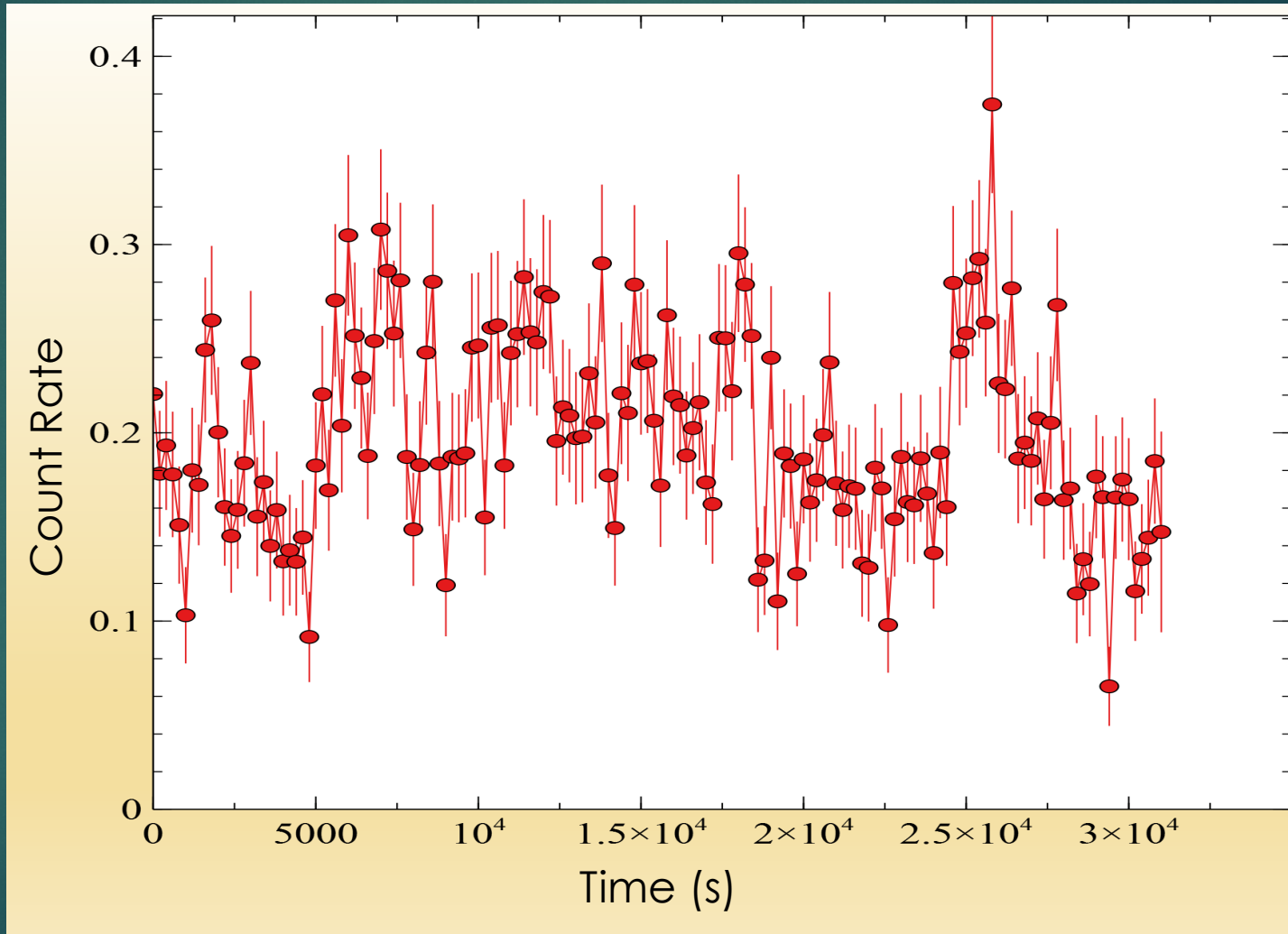
$$M_{\text{BH}} = 2.5 \pm 2 \times 10^6$$

From variability
(Ponti et al. 2012)

$$M_{\text{BH}} = 4 \pm 2 \times 10^6$$

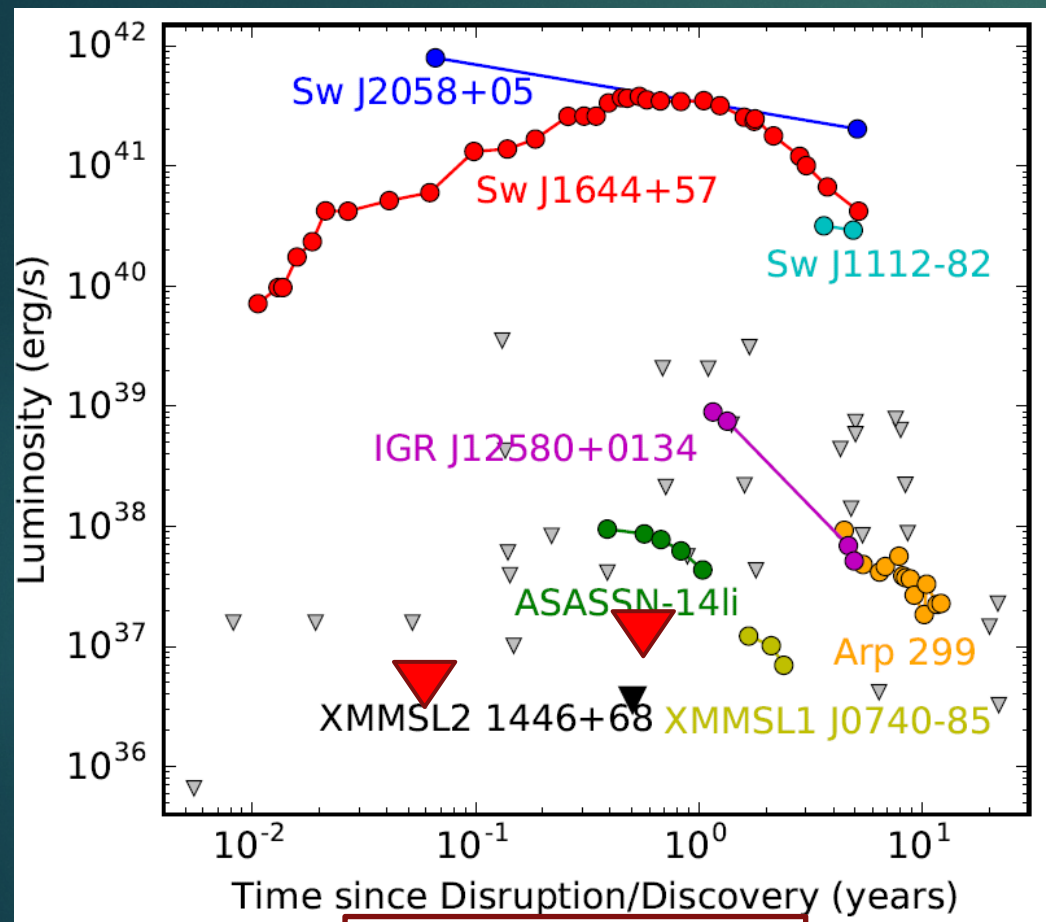
From $L_{\text{X}} = 10^{40} \text{ erg s}^{-1}$

$$m_{\text{edd}} \sim 0.5 \rightarrow 5 \times 10^{-4}$$

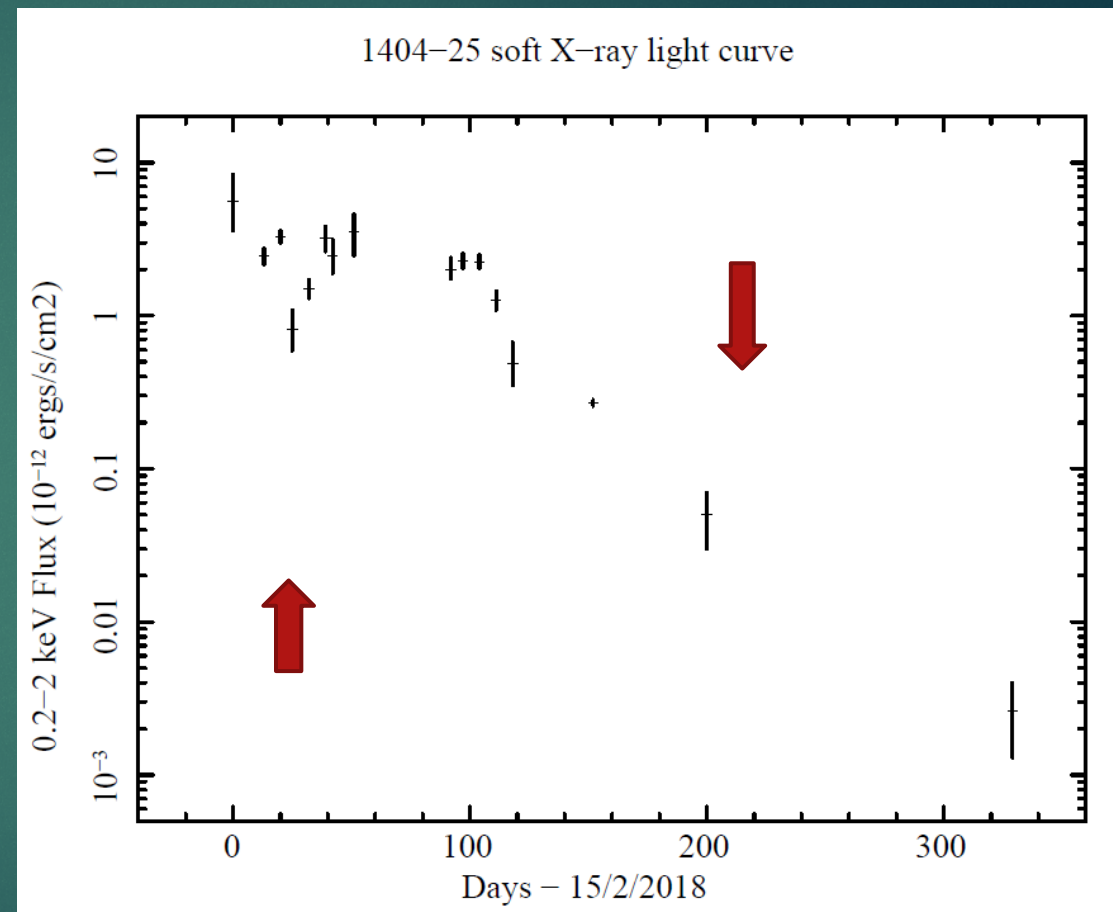


XMM-Newton, 0.2-2 keV – 17/7/2018

Radio (non) emission



$$L_{6\text{GHz}} < 2 \times 10^{37} \text{ ergs/s}$$



$$F_{6\text{GHz}} < 11, 35 \text{ uJy (eVLA)}$$

Summary

- ▶ X-ray TDE discovered with rapidly evolving X-ray spectrum in $M_{\text{BH}} \sim 2 \times 10^6 M_{\odot}$ galaxy
- ▶ Comptonisation region begins to form within 7 days and is “complete” after ~ 120 days
- ▶ Luminosity decays by factor ~ 1000 in 320 days
- ▶ After 320 days event has reached the typical late-time luminosity ($L_X = 2 \times 10^{41}$ ergs/s) and spectral shape ($\Gamma \sim 3$).