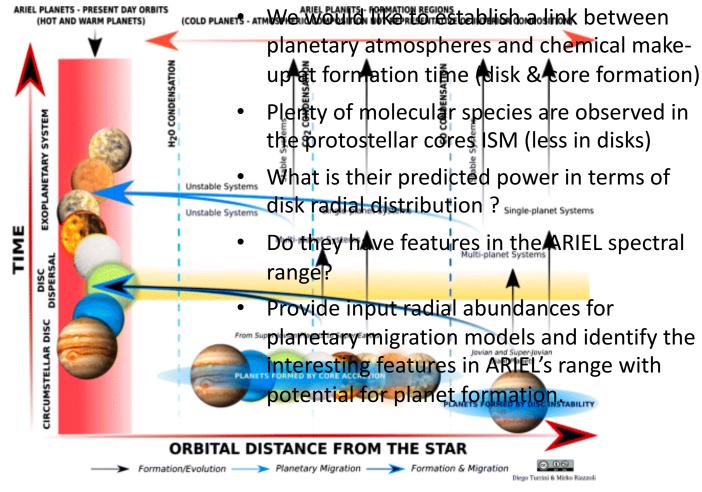
Hints to planet formation history from the chemistry of disks and natal clouds

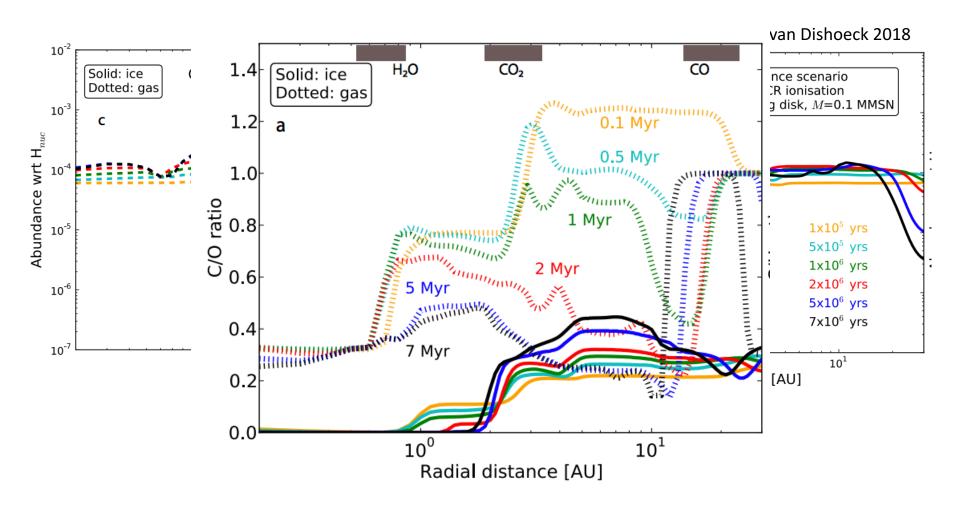
S. Molinari – INAF/IAPS

Context



Turrini+ 2018

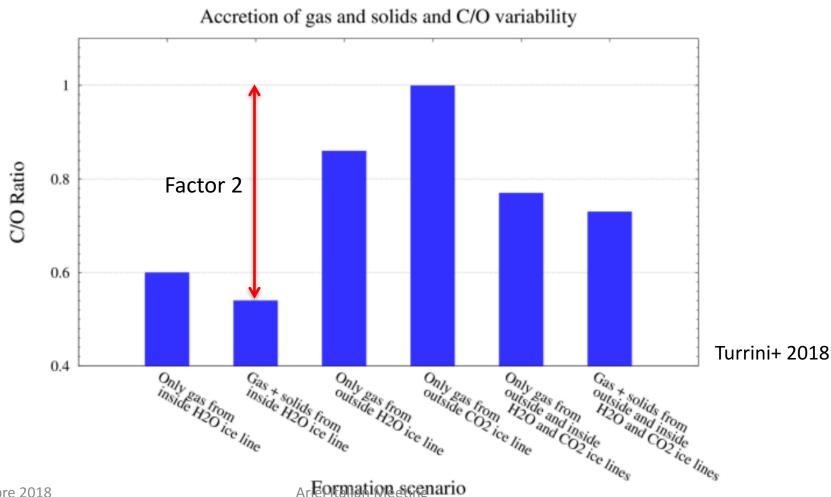
Carbon & Oxygen in disks



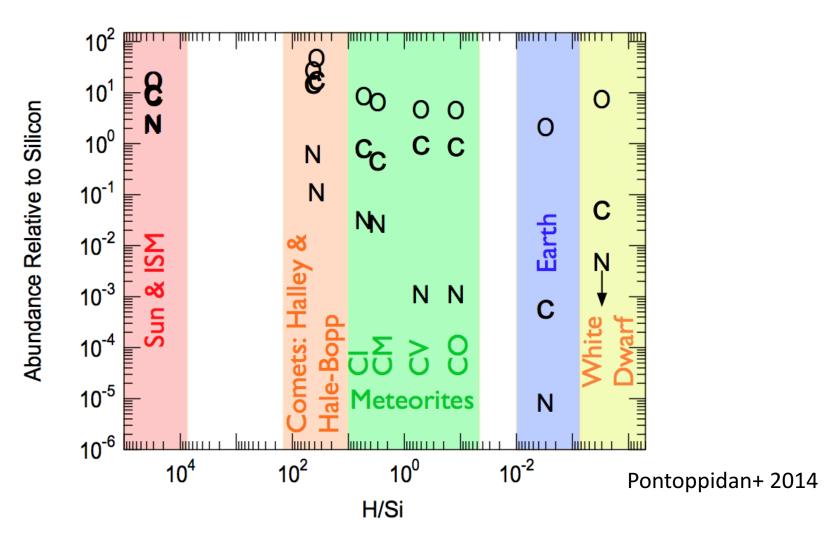
Factor of a few at most can be traced in gas (and much less in ices) below 10AU

Carbon & Oxygen in planets

Limited dynamical range in C/O ratio to distinguish between radically different scenarios?

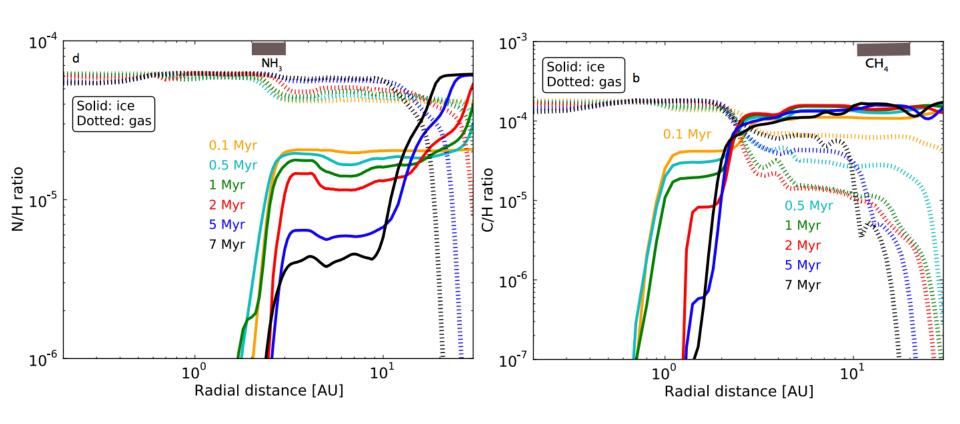


Alternatives



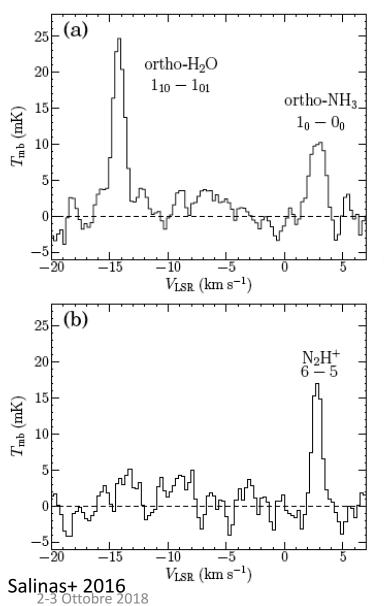
Nitrogen, Sulphur, ...?

Nitrogen



Large dynamical range in C/N ratio both in gas and ices can be traced to tens of AUs as a function of time

N-bearing tracers



NH₃, N₂H⁺, HCN, CN are routinely observed in protostellar cores



AND also in protostellar disks: TW Hydrae with Herschel/HIFI (Far-IR)

N-bearing species detection in the radio toward protoplanetary disks is a driving requirement for the

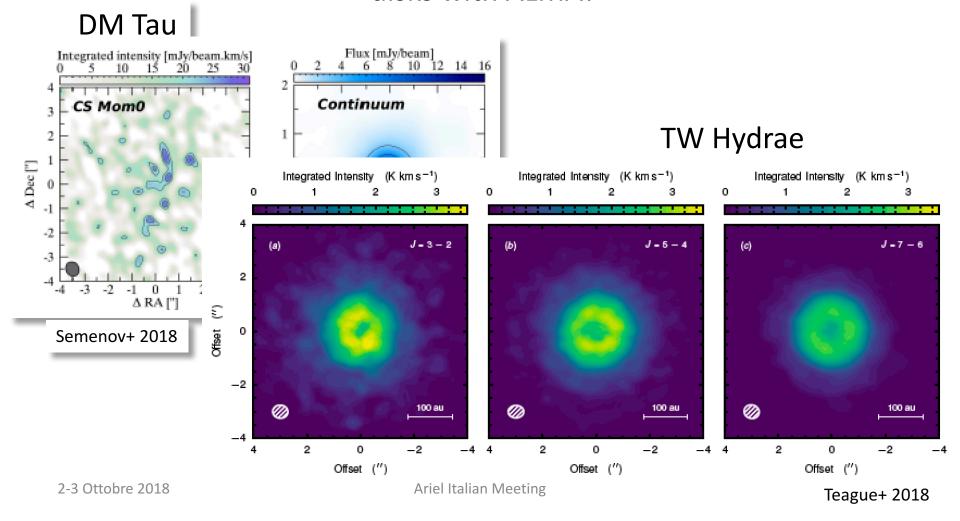
ngVLA

(early science 2028)

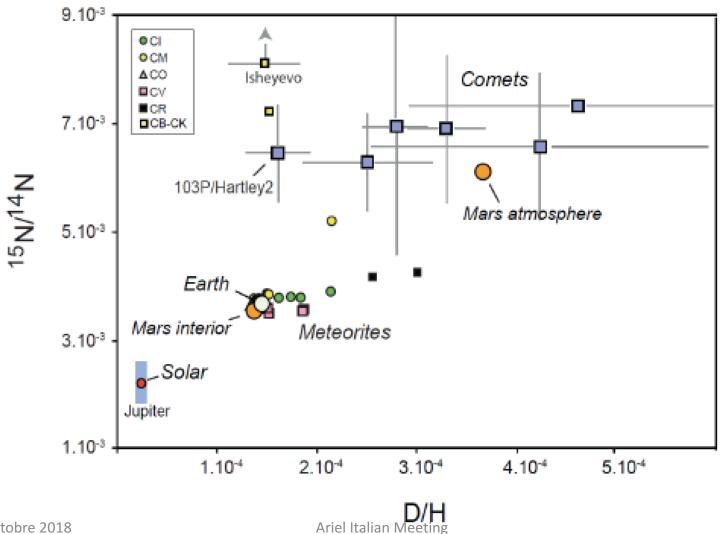
Ariel Italian Meeting

Any chance for S-bearing?

CS is routinely observed in protostellar cores, and now also in disks with ALMA!



Isotope ratios: a long shot?



Short term plan (phase B1)

- Inventory of species detected till now toward protoplanetary disks
- Inventory of chemistry models with evolving disks to identify promising species with predictive power X/H=f(r, t) → extension to N, S and others
- Search for corresponding features in ARIEL's range
- Define possible input conditions for planetary formation models and run migration models to predict what we could see with ARIEL
- Help define science requirements for the mission