

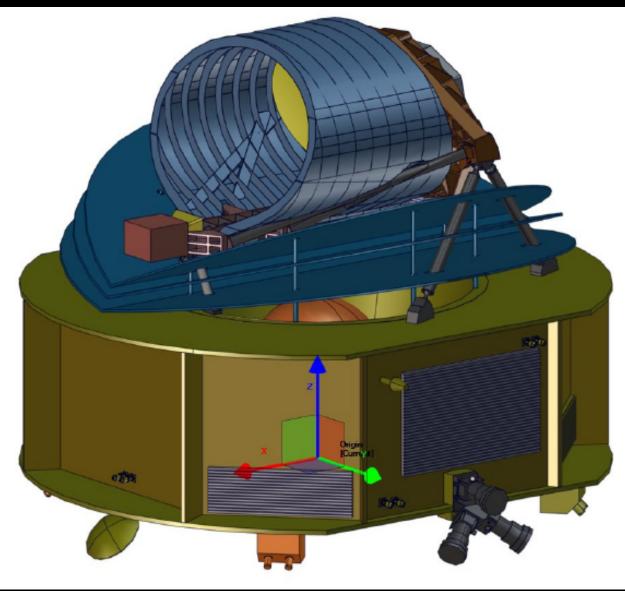
# The ARIEL payload

### **Emanuele Pace**

INAF, Università di Firenze

On the behalf of the ARIEL collaboration

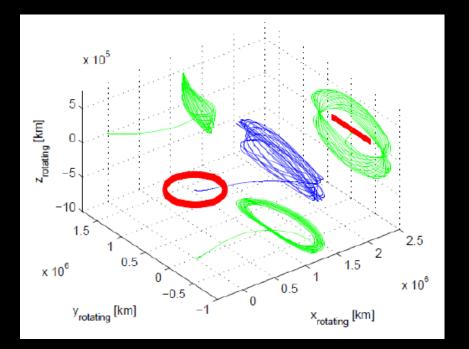






### Mission analysis





Mission sized for **3.5 years** dedicated to the nominal science operations phase, with additional 2 years goal extension

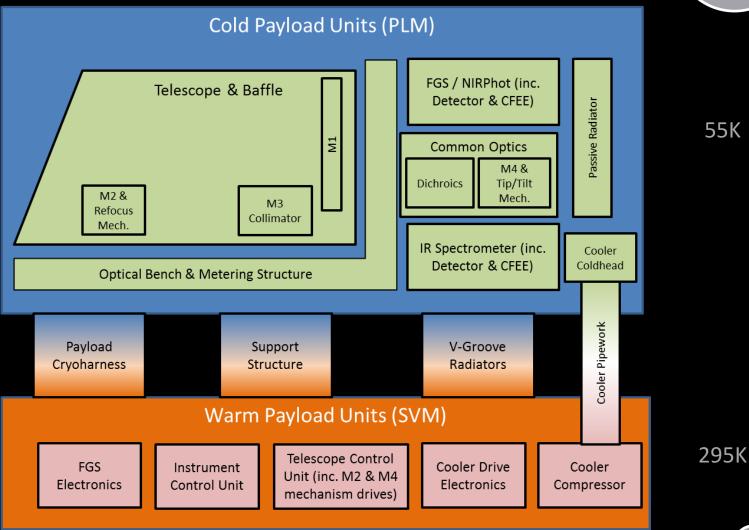
- orbit around Sun-Earth L2
  - ✓ very stable thermal environment
  - very large instantaneous field of regard
  - ✓ simple design of the comm's and power subsystems
  - benign radiation environment
- observation efficiency  $\geq 85\%$
- Solar Aspect Angle range to  $\pm 5^{\circ}$  around the S/C X-axis and to  $\pm 25^{\circ}$  around the Y-axis

## **ARIEL** Payload



- Off-axis Cassegrain telescope, 1.1m x 0.7m elliptical M1; diffraction limited at 3 µm. Mirrors, optical bench and telescope all manufactured from Aluminium alloy for isothermal design with minimal thermo-elastic deformation and Silver coated to improve efficiency in the VIS-NIR.
- The ARIEL InfraRed Spectrometer (AIRS) provides low/medium resolution (R = 30 - 200) spectroscopy between 1.95 and 7.8 μm.
- FGS module includes <u>3 photometric channels</u> (two used for guiding as well as science) between 0.5 and 1.2 μm and low resolution NIR spectrometer from 1.2 1.95 μm.
- Thermal: Warm SVM, cryogenic PLM cooled passively to ~55K with the thermal shield assembly. Active cooler (Neon JT) included to ensure AIRS detector operating temperature of ≤42K

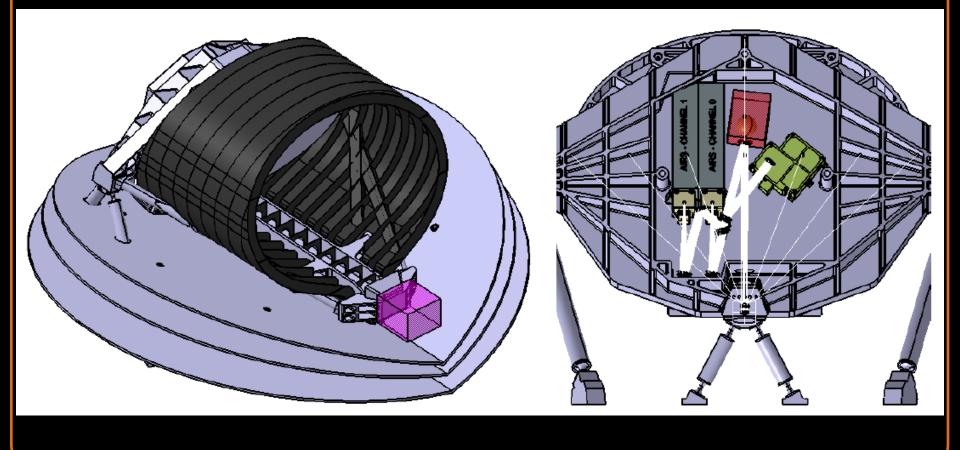
### Architecture



ARIEL

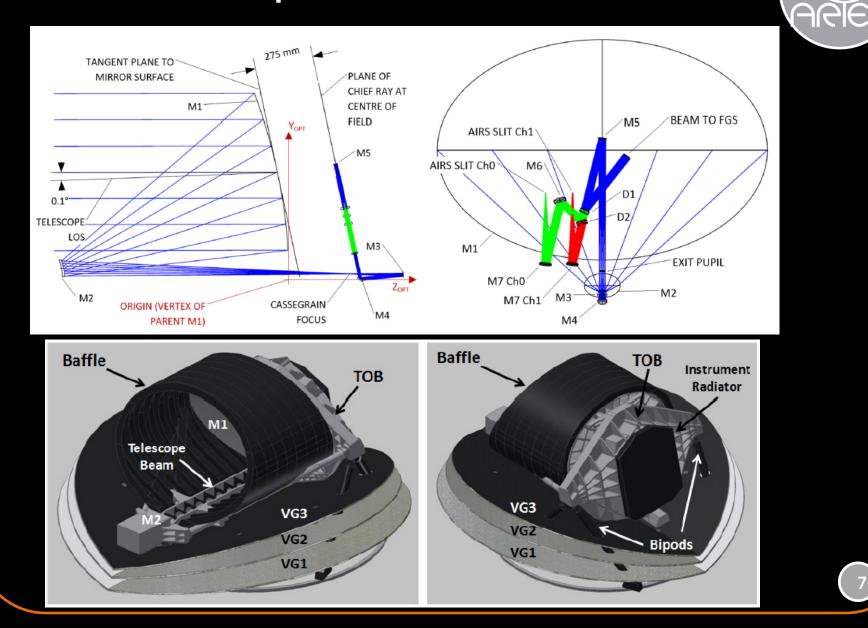
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# Mechanical design



ARIE

## Telescope characteristics



## Telescope parameters

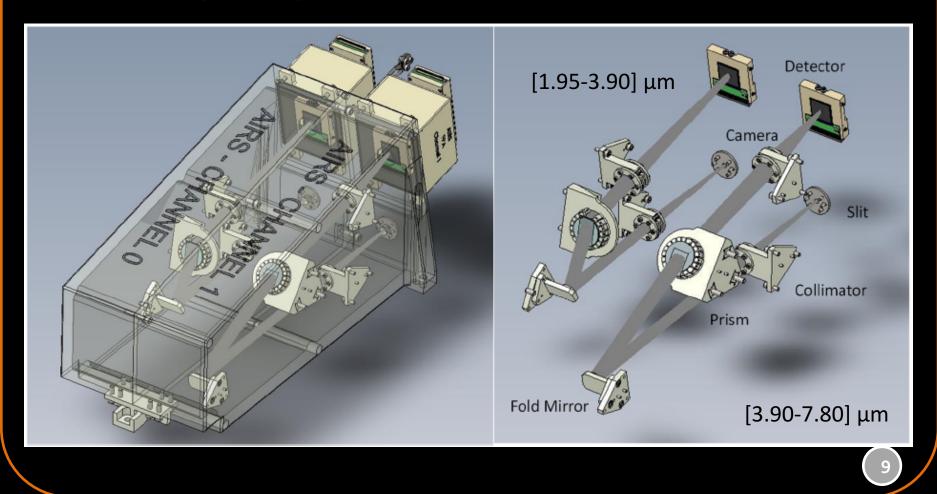


Parameter	Value
Size	1110 x 730 mm <sup>2</sup>
Collecting area	> 0.6 m <sup>2</sup>
FoV	<ul><li>30" with diffraction limited performance</li><li>41" with optical quality TBD allowing FGS centroiding</li><li>50" unvignetted</li></ul>
M1 WFE	Diffraction limited $\leq$ 3 $\mu m$ (~220 nm wv RMS)
M1 Roughness	< 100 nm
Focal ratio (@ AIRS entrance slit)	f/12 (pupil major axis) f/18 (pupil minor axis)

### Spectrometer AIRS

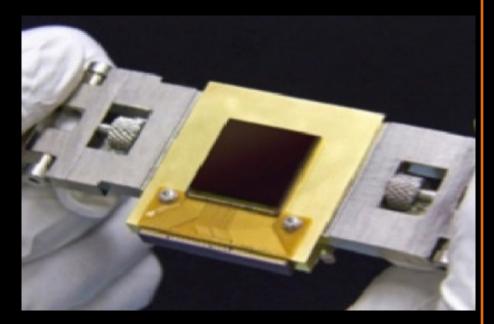
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### 2-channel prism spectrometers



### **Spectrometer Detector**

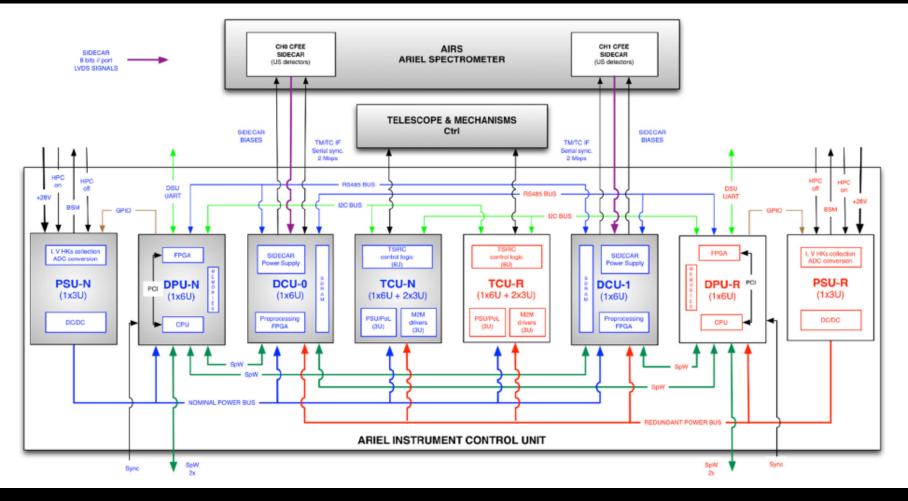
- Baseline NEOCam Teledyne MCT detector
- Work is continuing within consortium both on developing concepts to allow detectors to run warmer



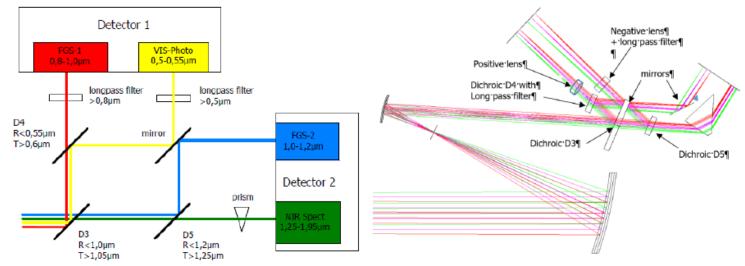




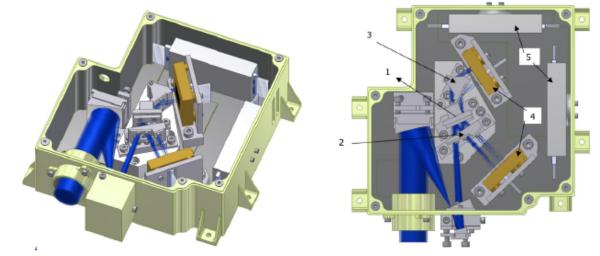
### Instrument Control Unit



### Fine Guidance System / NIRPhot



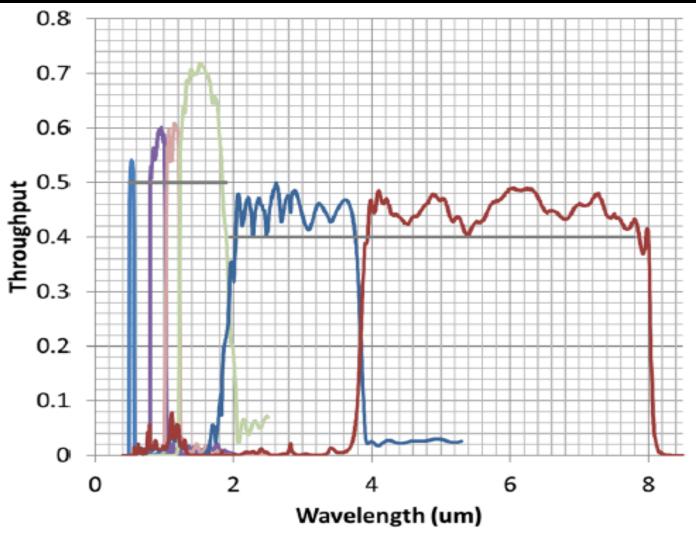
#### centroiding performance: 10 milli-arcsec at 10 Hz



ARIE

Throughput





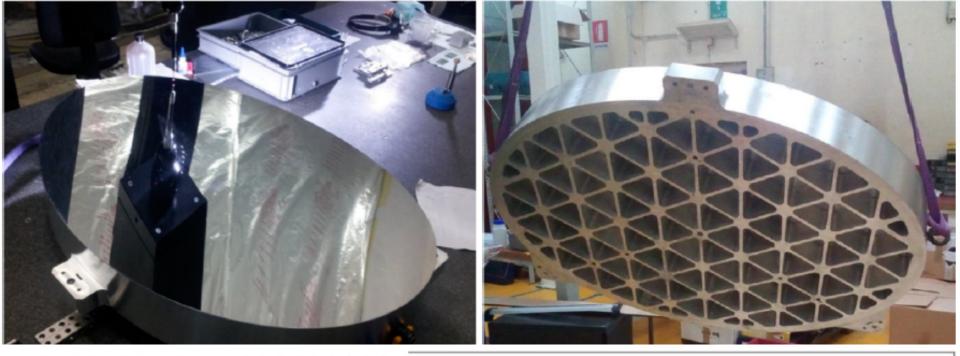
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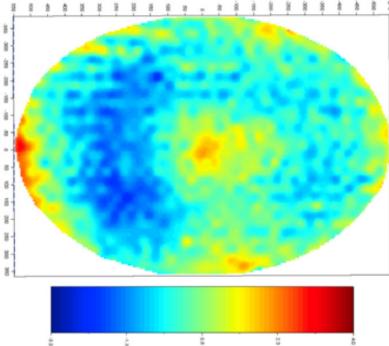
## Pathfinder Telescope Mirror

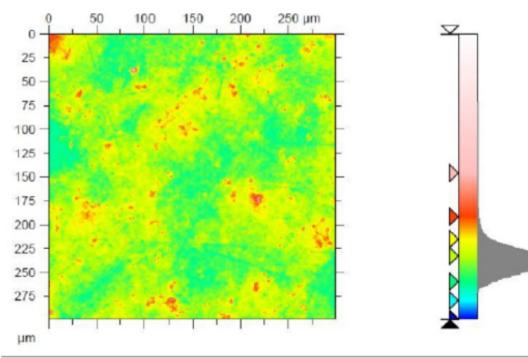


In order to assess the manufacturability of M1, a Pathfinder Telescope Mirror (PTM) has been produced with the following objective:

- manufacture a representative mirror with the same size and shape of the nominal design, with relaxed req's on shape accuracy and roughness and mild light-weighting;
- verify the structural performance of the mirror, in terms of figure error, under gravity;
- Identify a support strategy of the mirror during metrology and testing that minimizes the effect of gravity.







### Overall M<sub>4</sub> Timeline



#### Timeline targeted for M4 mission

Phase 0 (ESA internal CDF)	June to September 2015
Phase 0 completed	End September 2015
ITT for Phase A industrial studies	October 2015
Phase A kick-off	March 2016
Mission Selection Review completed (ARIEL, THOR, XIPE)	April 2017
Selection of M4 mission (SPC)	March 2018
Phase B1 kick-off of the selected M4 mission	March 2018
Phase B1 completed	March 2020
Mission Adoption Review	Q2 2020
Adoption of the M4 mission (SPC)	November 2020
Phase B2/C/D kick-off	TBD
Launch	2028