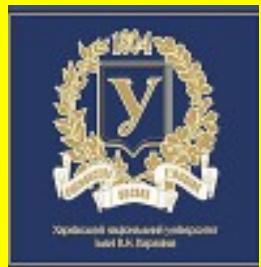




BILATERAL WORKSHOP ON ASTROPHYSICS
V.N. KARAZIN KHARKIV NATIONAL
UNIVERSITY - INAF



UKRAINIAN OBSERVING FACILITIES AND
TECHNOLOGIES

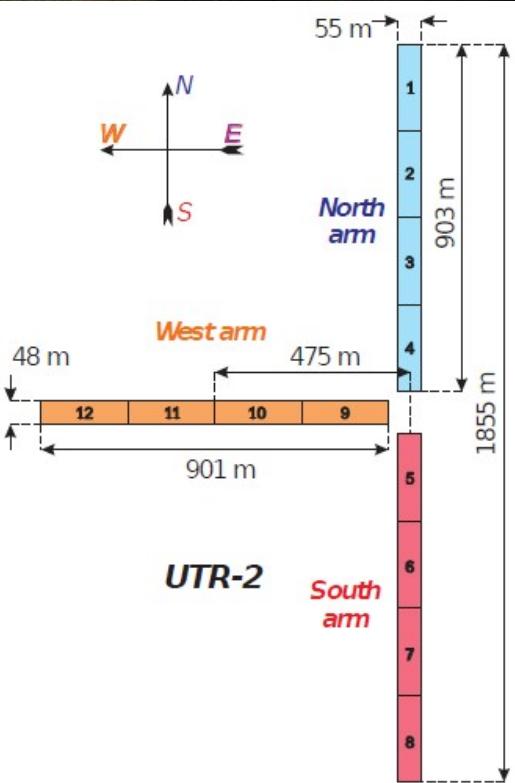
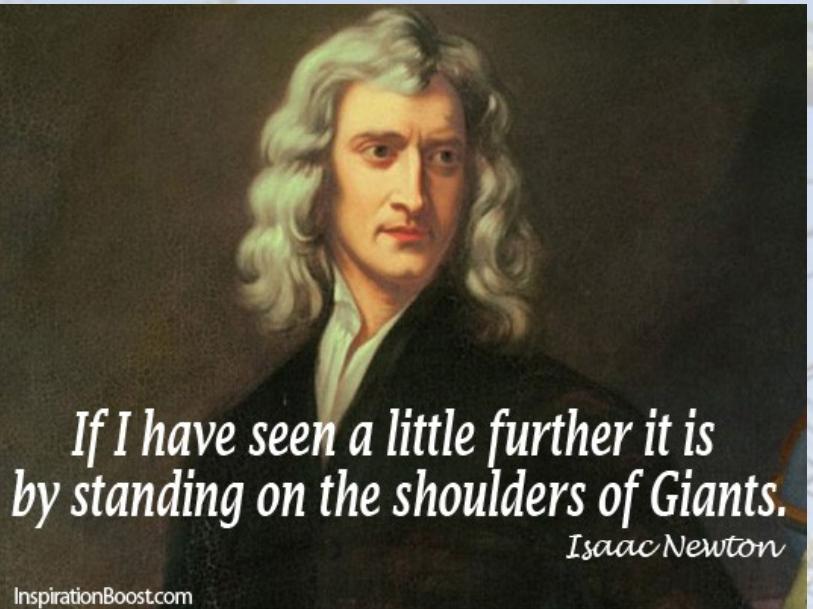
22-23, March 2018, Rome, Italy

O.M. Ulyanov and IRA NASU team

Institute of Radio Astronomy of NAS of Ukraine

oulyanov@rian.kharkov.ua





Scientific Structure of the Institute of Radio Astronomy NASU

- Division of Low-frequency Radio Astronomy
 - Decameter Radio Astronomy
 - Astrophysics
 - Radio-astronomical equipment and methods of observations
- Space Radio Physics
- Millimeter-Wave Radio Astronomy
- Microwave Spectroscopy
- Theoretical Radio Physics
- Microwave Electronics
- Geospace Radio Physics

<http://rian.kharkov.ua/index.php/en/structure-en>

Institute's structure - Institute of Radio Astronomy of NASU - Mozilla Firefox

rian.kharkov.ua/index.php/en/structure-en

Search

Directorate

Scientific departments

Division of decameter radio astronomy
Department of decameter radio astronomy (No. 11)
Department of astrophysics (No. 12)
S. Braude Radio Astronomical Observatory (No. 13)
Department of radio-astronomical equipment and methods of observation (No. 15)

Department of space radio physics (No. 16)
Department of millimeter-wavelength radio-astronomy (No. 17)
Department of microwave radiospectrometry (No. 18)

Department of theoretical radio physics (No. 20)
Experimental department of microwave high technologies (No. 14)
Microwave Electronics Department (No. 21)
Department of geospace radio physics (No. 22)
Low Frequency Observatory (No. 23)

Center of collective use of scientific equipment
"CRYOELECTRONICS"

Maintenance departments

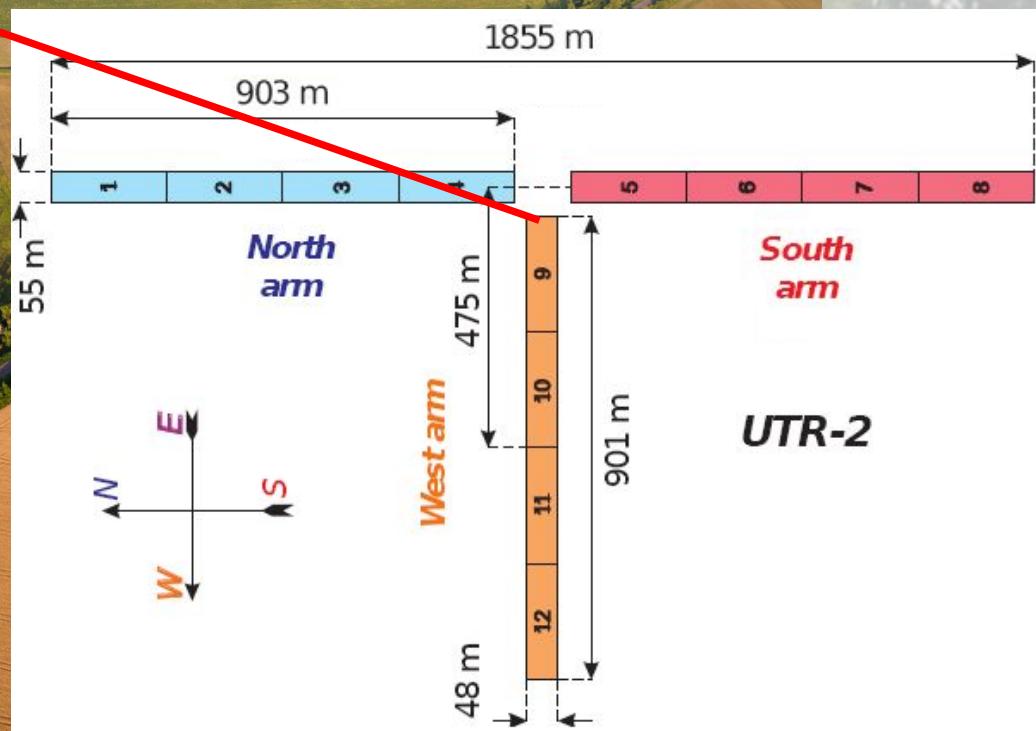
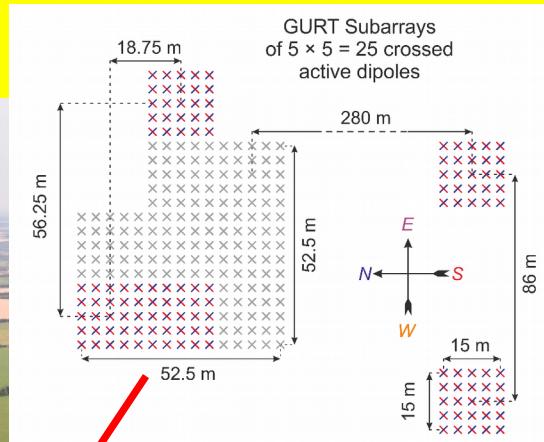
Research and organizational departments

The screenshot shows the 'Institute's structure' page of the Institute of Radio Astronomy of NASU. At the top, there are two main categories: 'Directorate' and 'Scientific departments'. Under 'Scientific departments', there is a box containing several sub-departments: Division of decameter radio astronomy, Department of decameter radio astronomy (No. 11), Department of astrophysics (No. 12), S. Braude Radio Astronomical Observatory (No. 13), and Department of radio-astronomical equipment and methods of observation (No. 15). Below this, there are three columns of boxes: the first column contains 'Department of space radio physics (No. 16)' and 'Department of theoretical radio physics (No. 20)'; the second column contains 'Department of millimeter-wavelength radio-astronomy (No. 17)' and 'Experimental department of microwave high technologies (No. 14)'; the third column contains 'Department of microwave radiospectrometry (No. 18)' and 'Department of geospace radio physics (No. 22)'. At the bottom, there is a box for 'Center of collective use of scientific equipment "CRYOELECTRONICS"'. At the very bottom, there are two more categories: 'Maintenance departments' and 'Research and organizational departments'. The page is displayed in Mozilla Firefox, with the address bar showing the URL rian.kharkov.ua/index.php/en/structure-en.



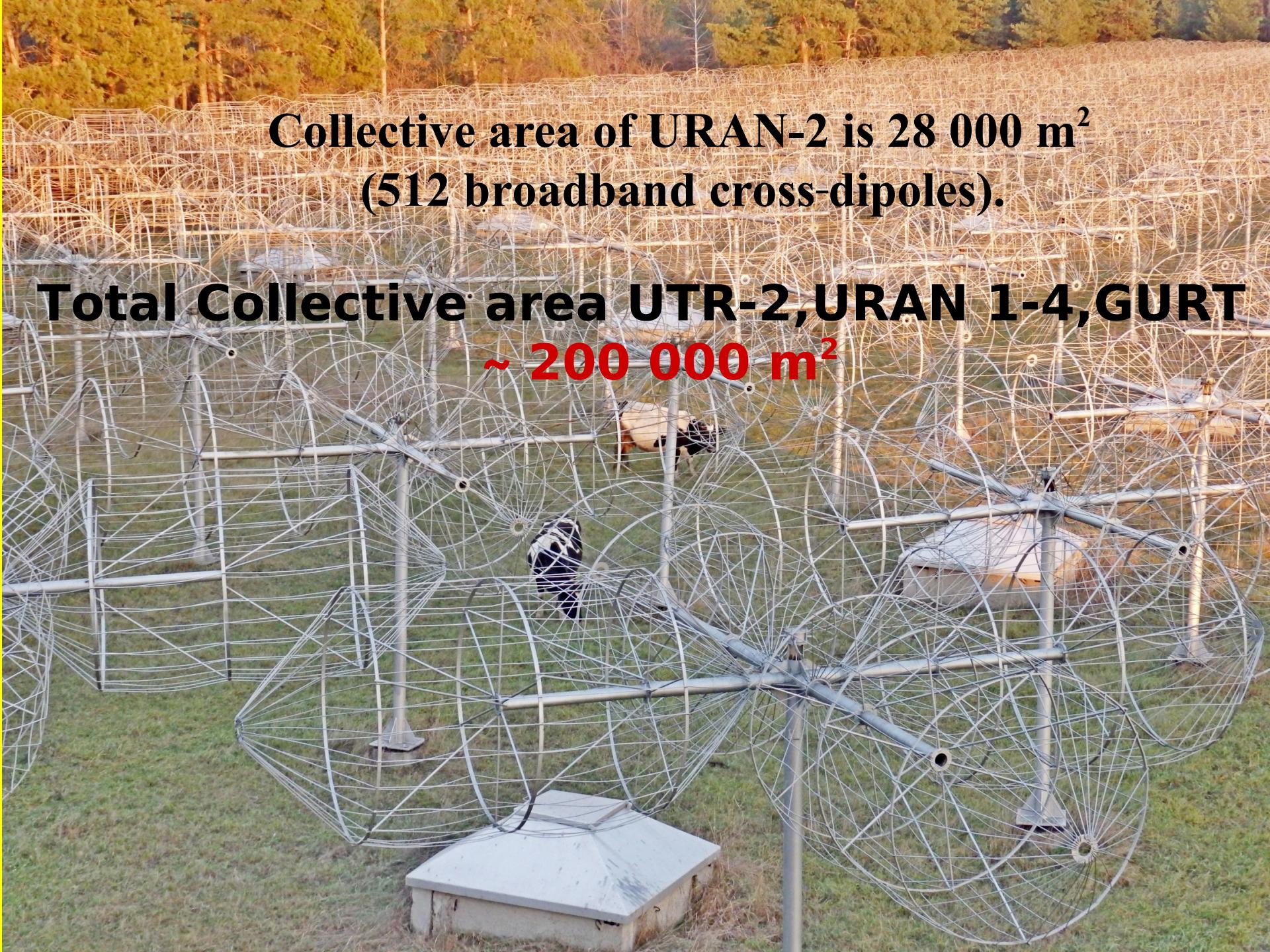
Radio telescopes UTR-2 and GURT

(September, 2017)





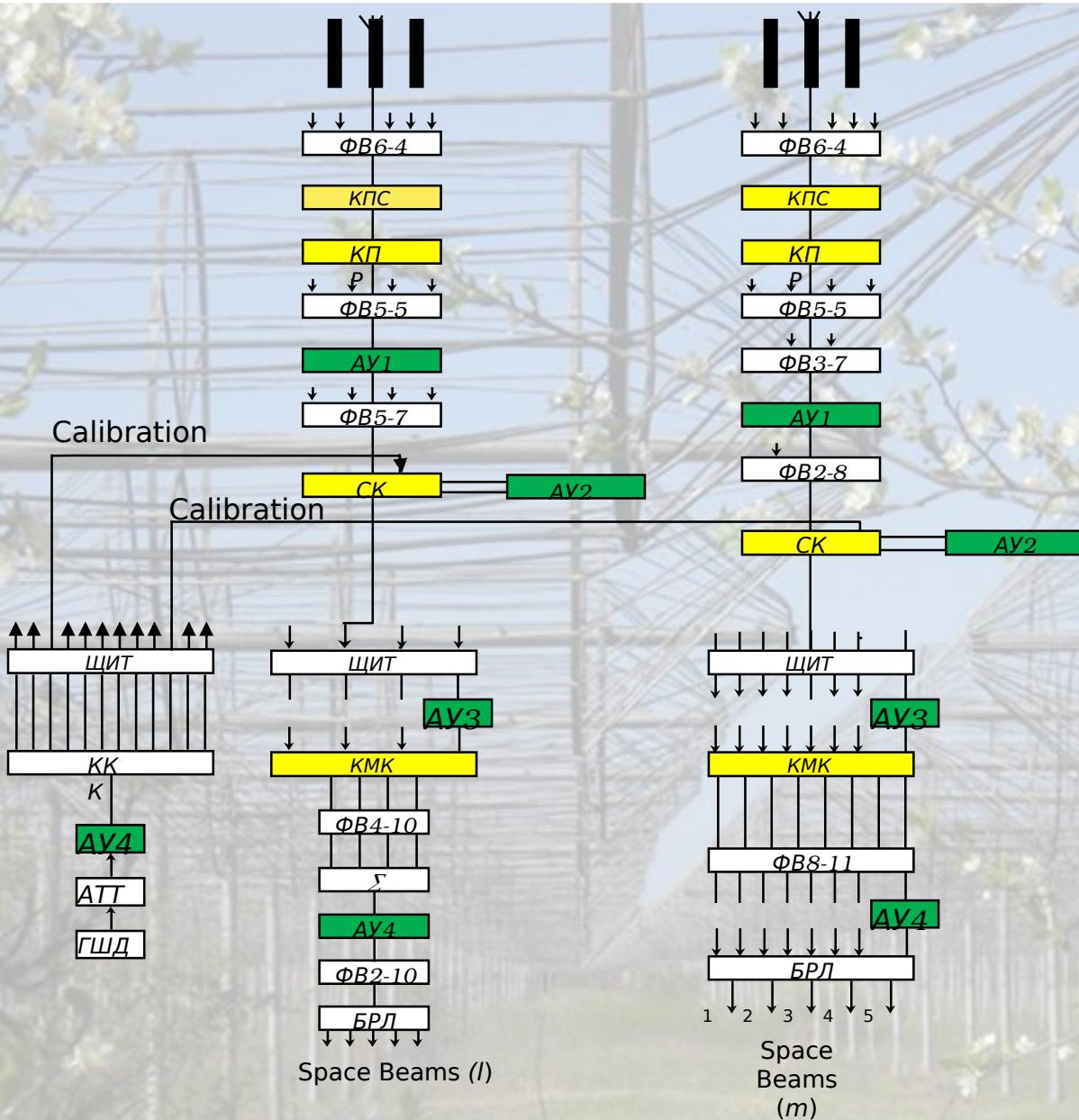
**The largest in the world radio telescope UTR-2
(1.8 km x 60 m) Frequency range 8...32 MHz; 2040 Dipoles;
Collective area ~ 150 000 m².**



**Collective area of URAN-2 is 28 000 m²
(512 broadband cross-dipoles).**

**Total Collective area UTR-2, URAN 1-4, GURT
~ 200 000 m²**

Modernization of Amplification System of UTR-2



1. Array Amplification System (1988-1994)

2. Commutators (1997-2004)

3. Digital Receivers (2000 - ...):

- 60-channal receiver,
- Pulsar Portable Receiver (WF 2 x 1.5 MHz; 2 bits),

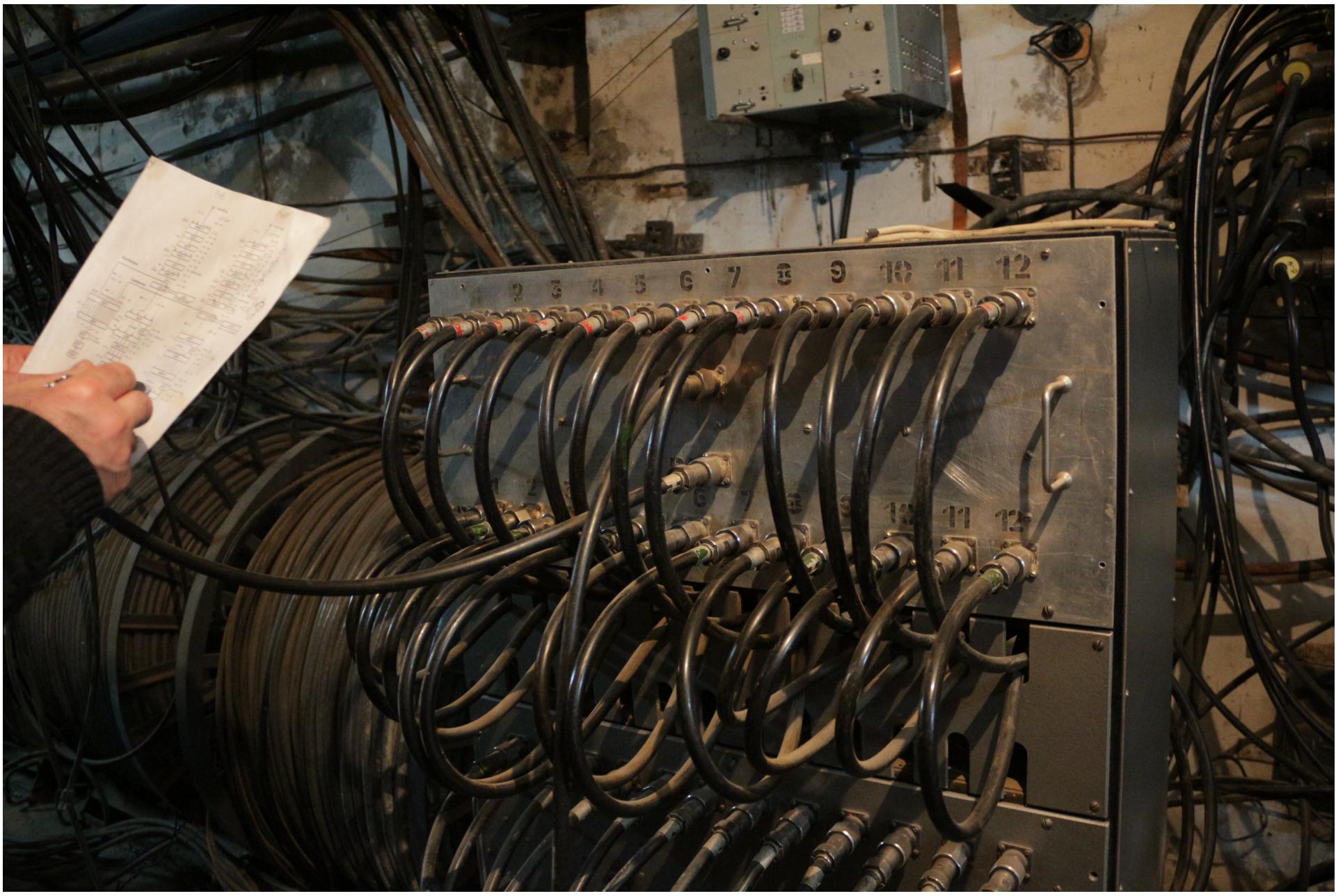
WF Receiver (WF 2 x 7.5 MHz; 8 bits),

4 DSPZ (Sp; Corr; WF 16 bits) (2x8192; 4x8192; 15ns)

Cable system of UTR-2

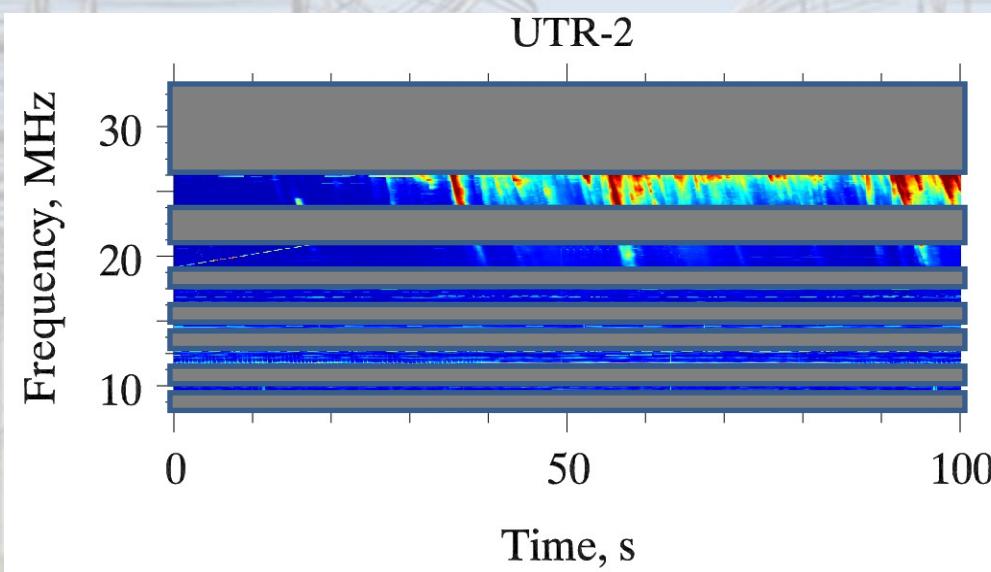
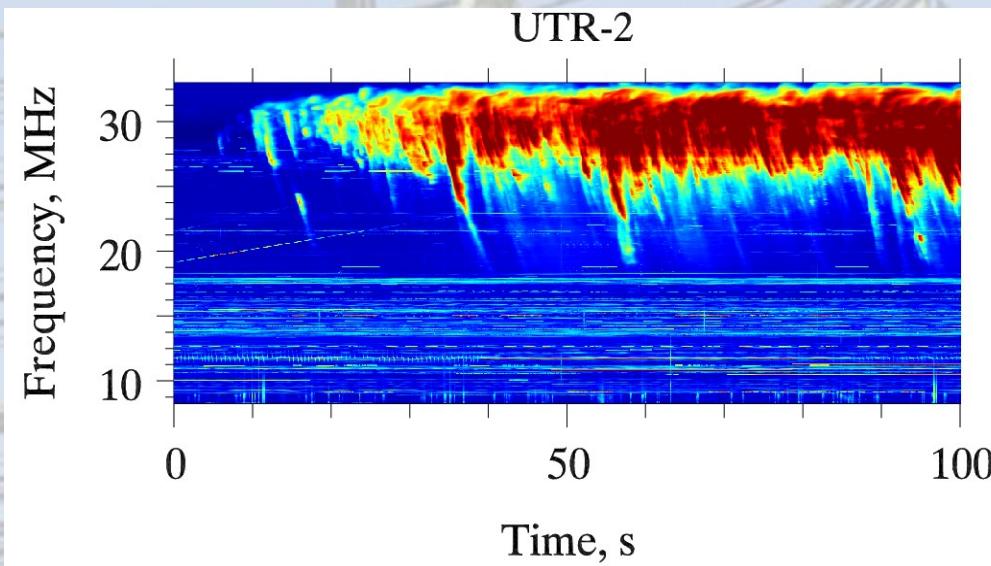


Huge Phase Shifters (left side)



Broad Bandwidth 8-32 MHz

Wide Frequency Band Instead of 6 Narrow Frequency Ranges

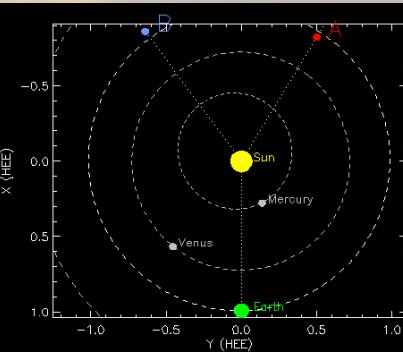
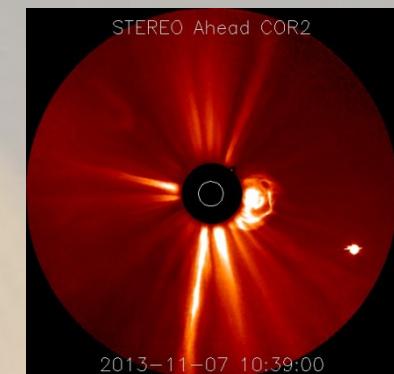
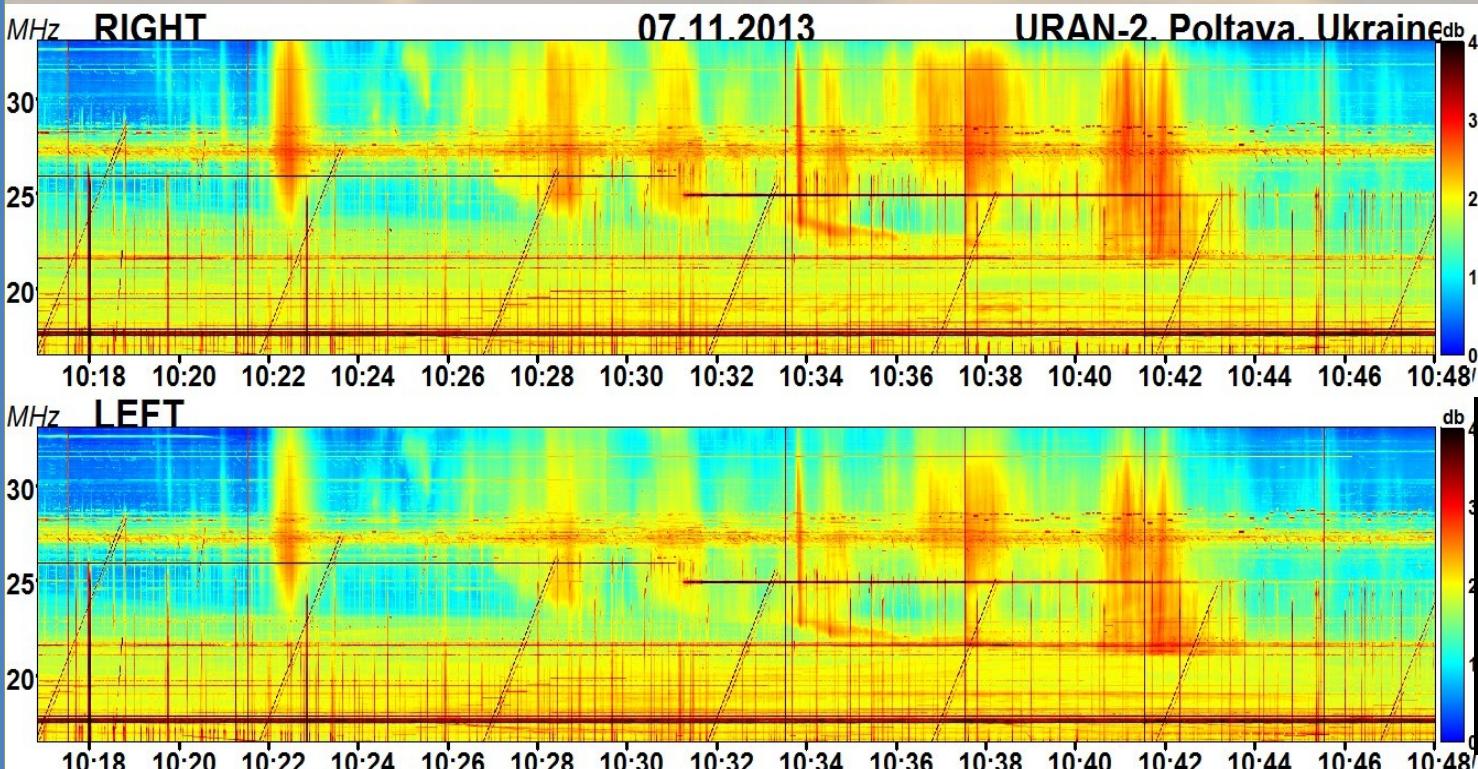




Solar Hurricane = Coronal Mass Ejections (CME) Bursts of the IV Type as indicator of CMEs

The Dynamic Spectrum of the IV Type Burst.
Observations on URAN-2 on November 7, 2013.

Coronal Mass Ejections (CME)
moving into south-west direction
(STEREO A)



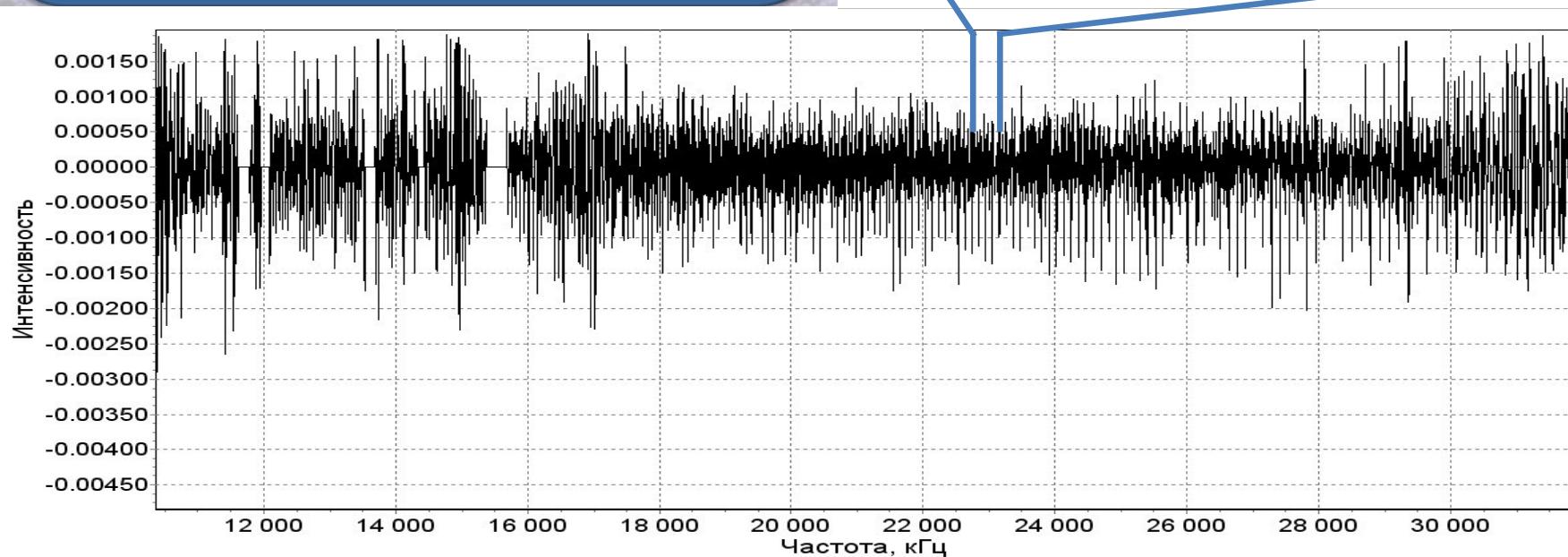
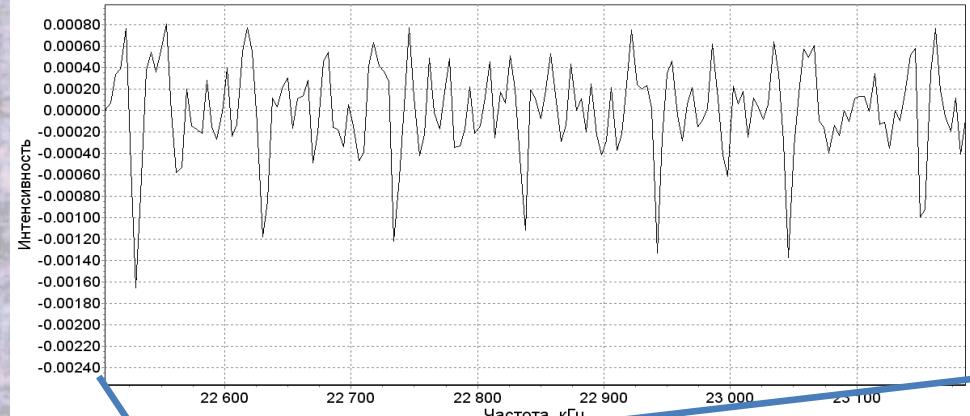
V.N. Melnik, A.I. Brazhenko, A.A. Konovalenko, V.V. Dorovskyy, H.O. Rucker, M. Panchenko, A.V. Frantsuzenko, M.V. Shevchuk "Decameter type IV burst associated with behind-limb CME observed on November 7, 2013", Solar Physics 2017 (submitted)



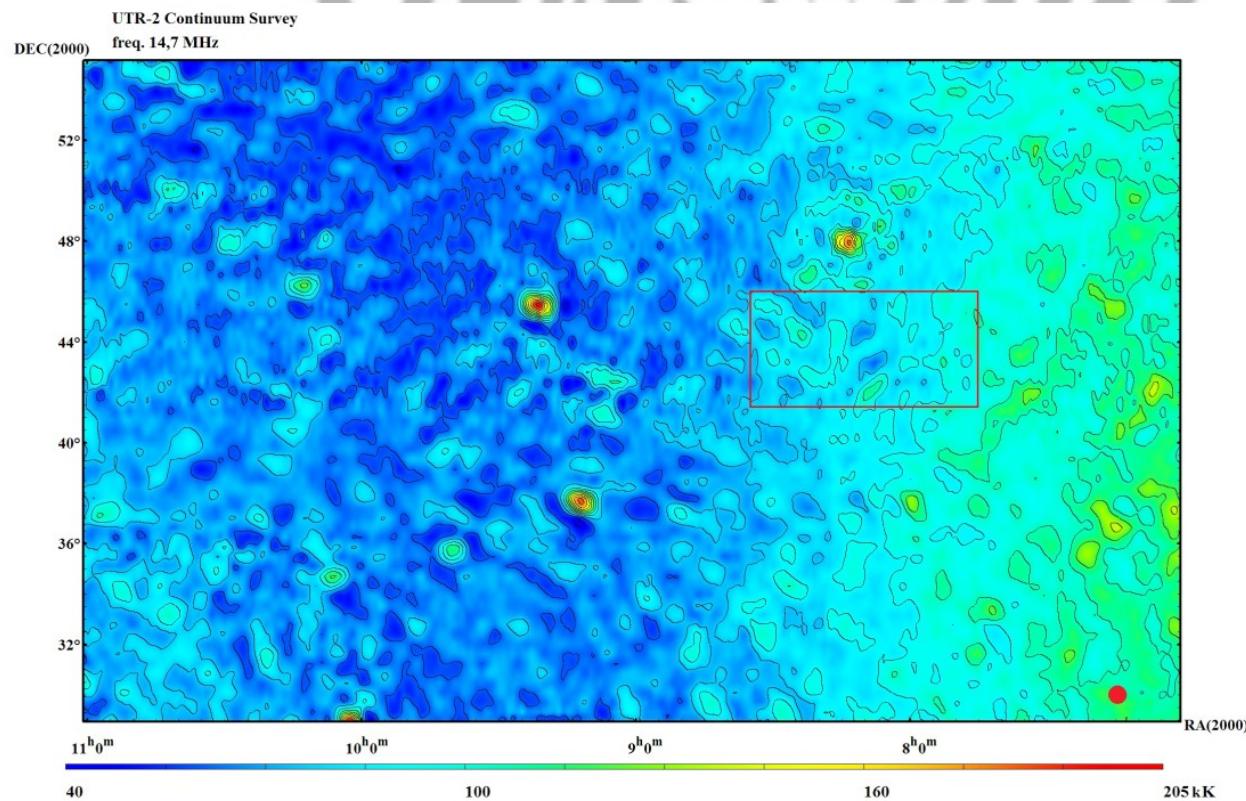
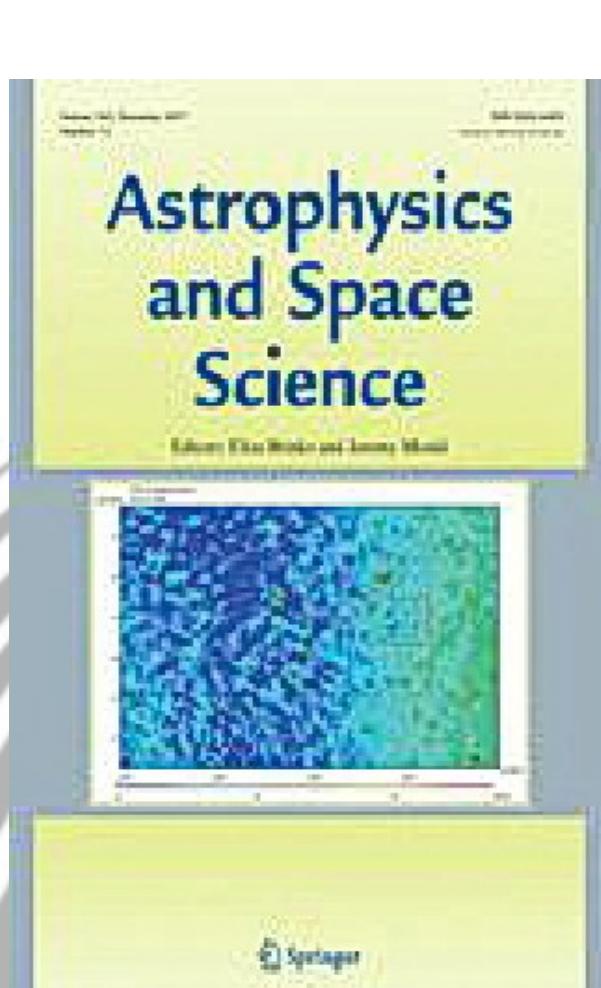
Investigations of the interstellar medium

Sequence of the Carbon Radio Recombination Lines (C854α ... C596α) in Cas-A direction.

Frequency range from 10,5 MHz up to 31 MHz. Cumulative time is 55 minutes and ~250 lines



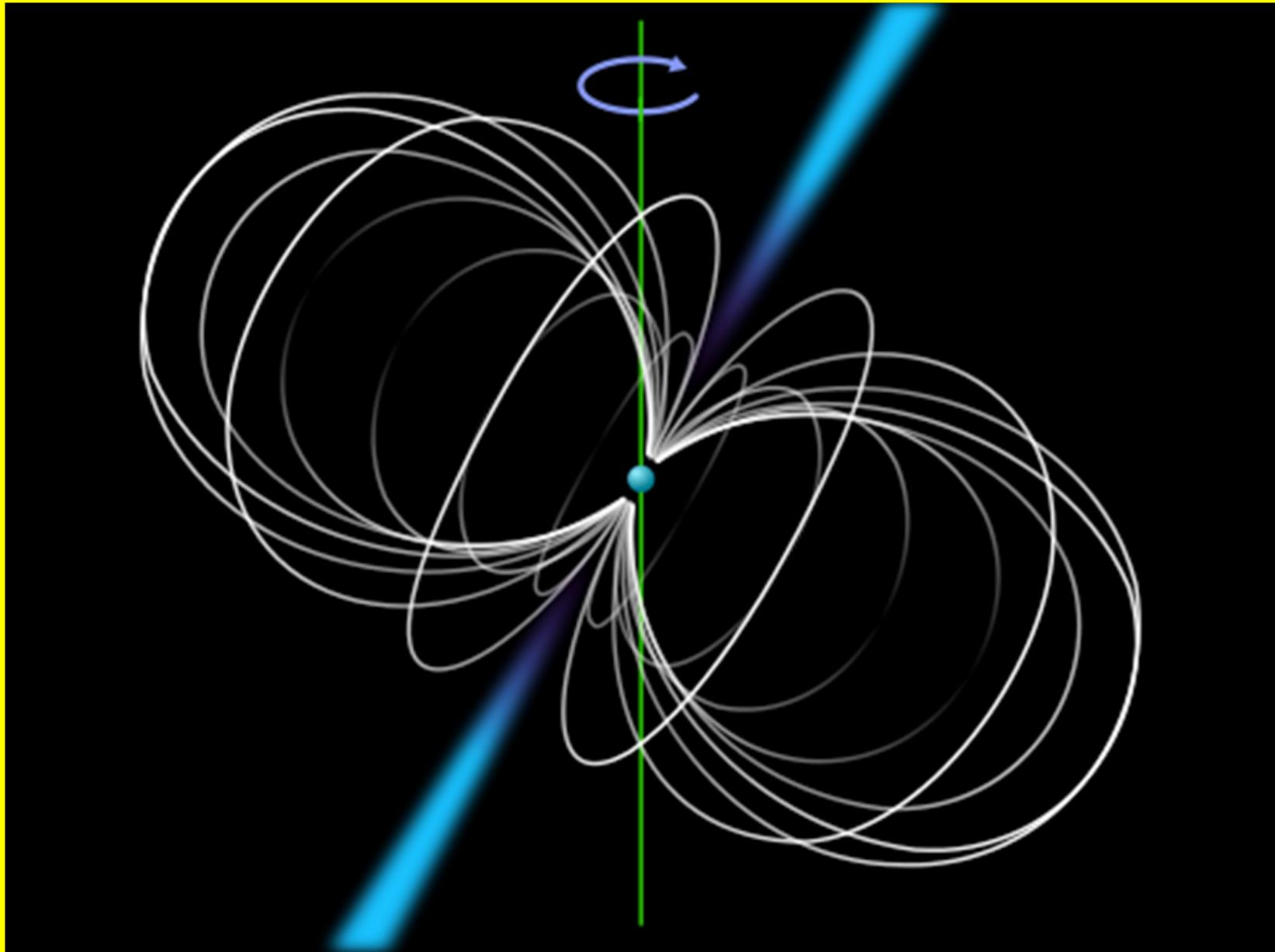
Separation of galactic and extragalactic radio emission at decameter wavelengths



Cold region of the Northern sky $F = 14,7$ MHz
(UTR-2). HPBW $\approx 47' \times 53'$, $15\,000 \text{ K} < T_b < 40\,000 \text{ K}$

N.M.Vasilenko, M.A.Sidorchuk. Separation of galactic and extragalactic radio emission at decameter wavelengths // Astrophys Space Sci (2017)- Vol.362-Is.12-P.221.

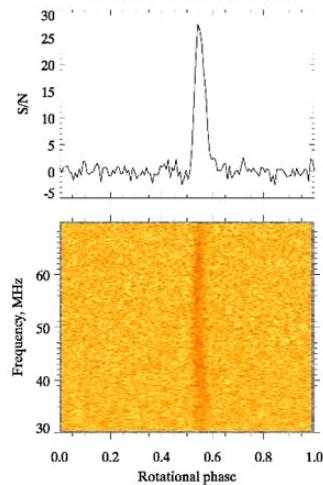
Model of Pulsar magnetosphere



Simultaneously PSRs observations at broad frequency range by UTR-2 and GURT

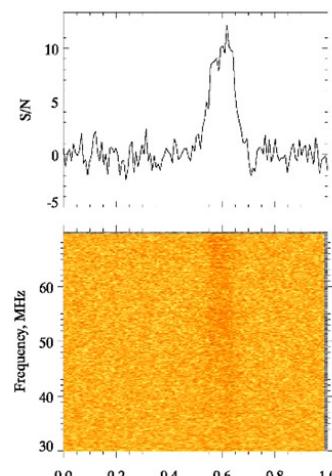
6

PSR B0834+06 P = 1.274 s
Observations 31.01.2017 at 19:00:00 UTC



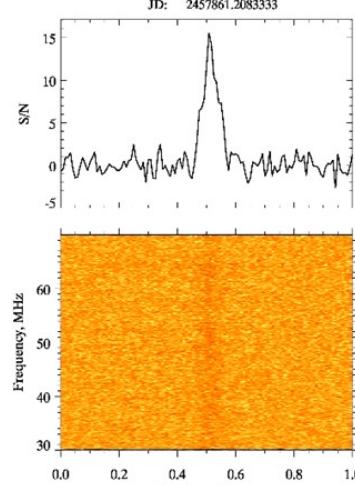
4

PSR B0809+74 P = 1.292 s
Observations 30.01.2017 at 17:59:42 UTC



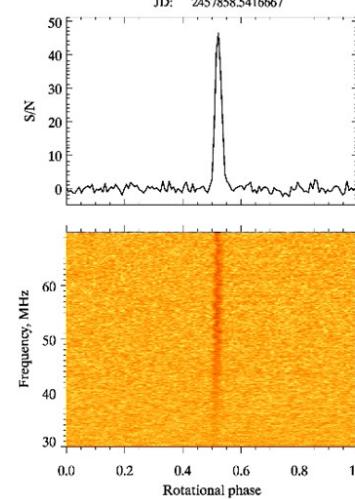
6

PSR B1133+16 Period = 1.188 s
Observations: 17.04.2017 at 17:00:00 UTC
JD: 2457861.2083333



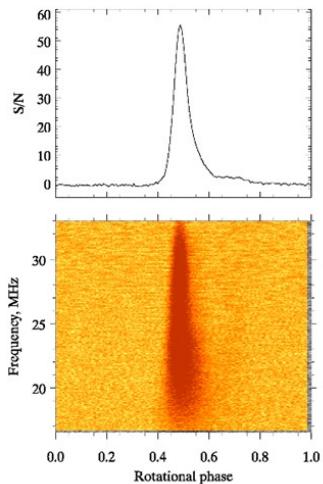
1

PSR B1919+21 Period = 1.337 s
Observations: 15.04.2017 at 01:00:00 UTC
JD: 2457858.5416667

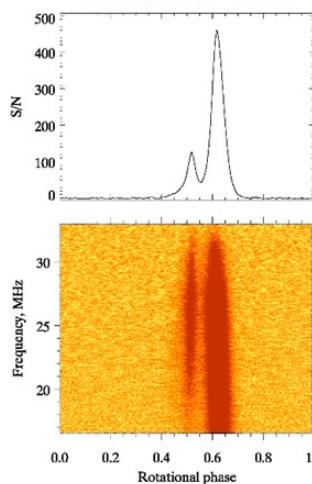


$f = 8 \dots 80 \text{ MHz}$

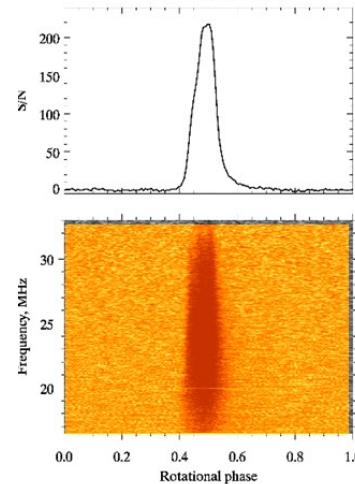
PSR B0834+06 P = 1.274 s



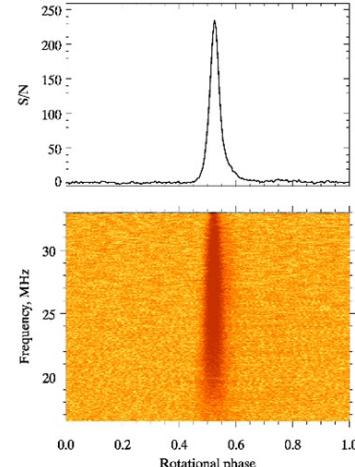
PSR B0809+74 P = 1.292 s



PSR B1133+16 Period = 1.188 s
Observations: 13.04.2017 at 16:59:59 UTC
JD: 2457857.2083218



PSR B1919+21 Period = 1.337 s
Observations: 13.04.2017 at 01:00:00 UTC
JD: 2457856.5416667



Σ



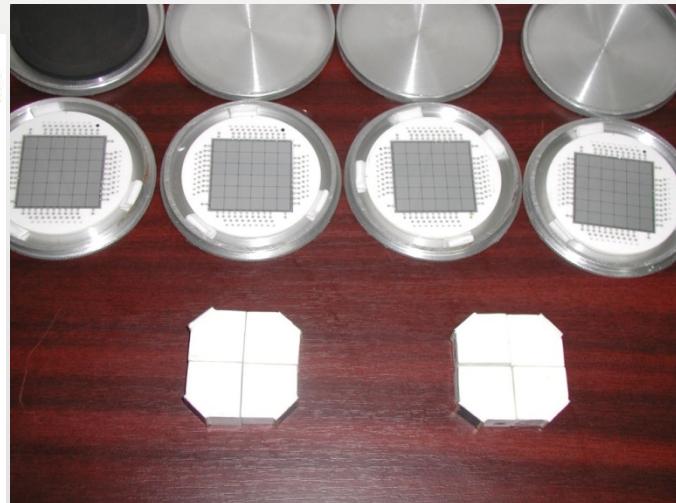
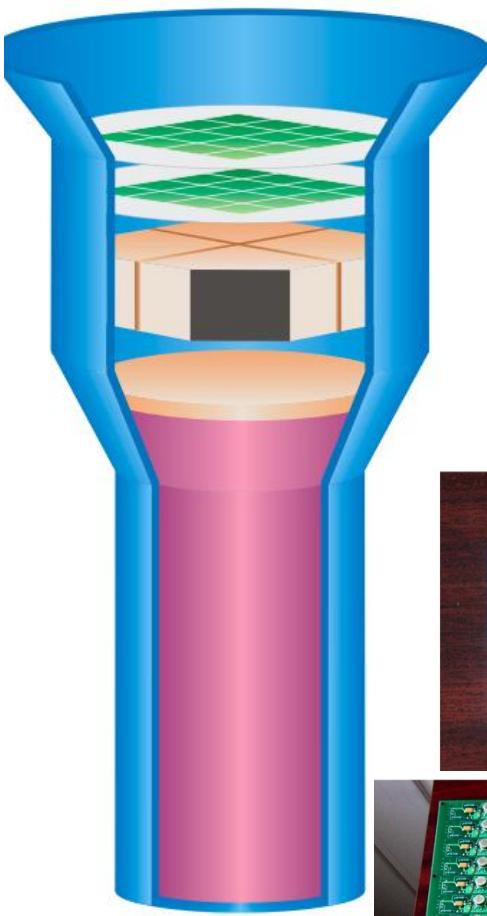
$f = 8 \dots 33 \text{ MHz}$

STEP-F instrument on the board of *CORONAS-Photon* satellite, 2009

1

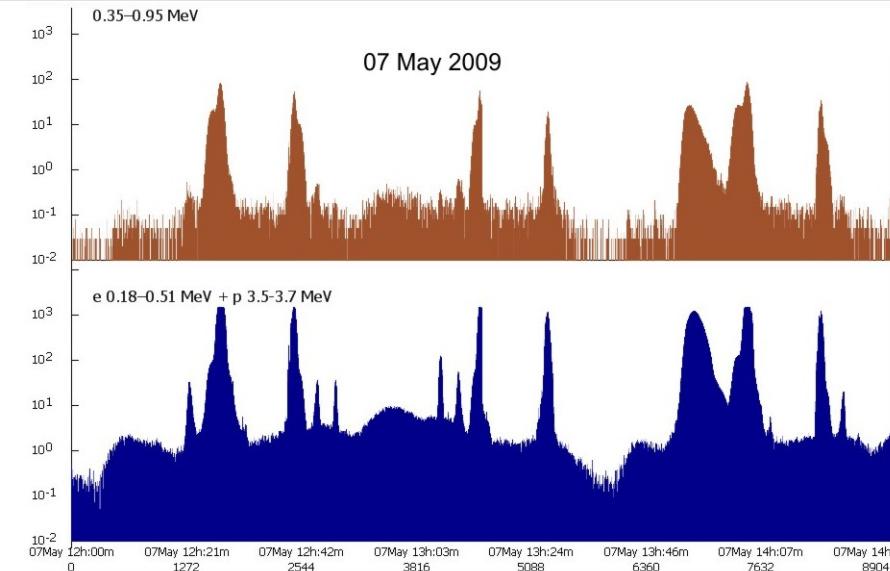
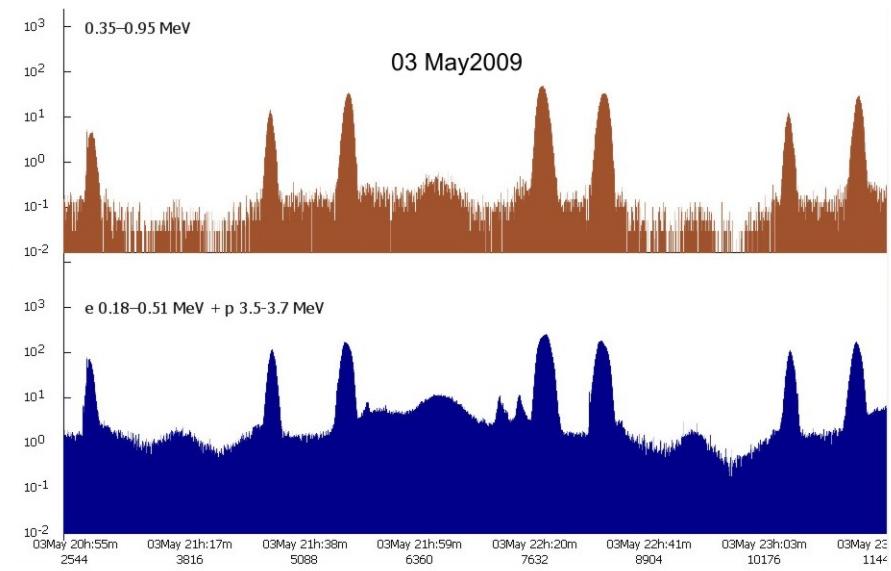
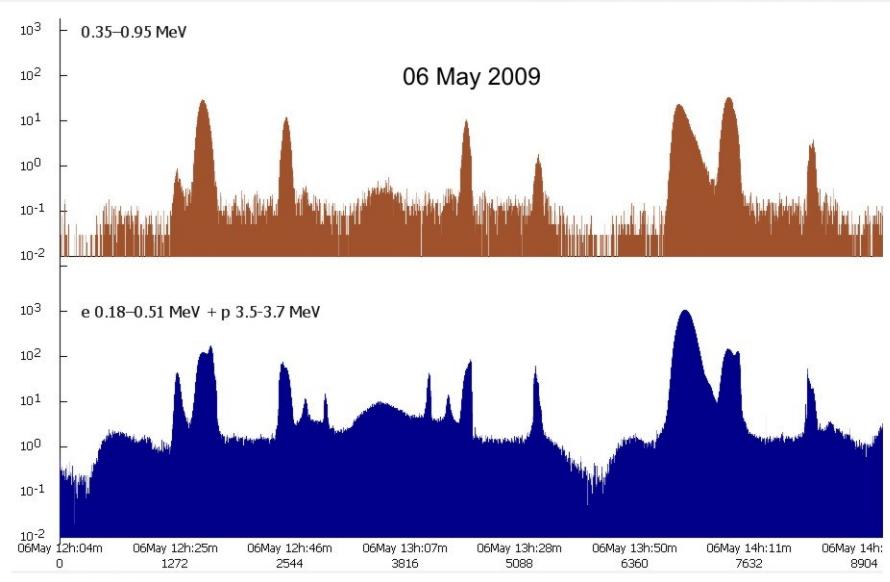
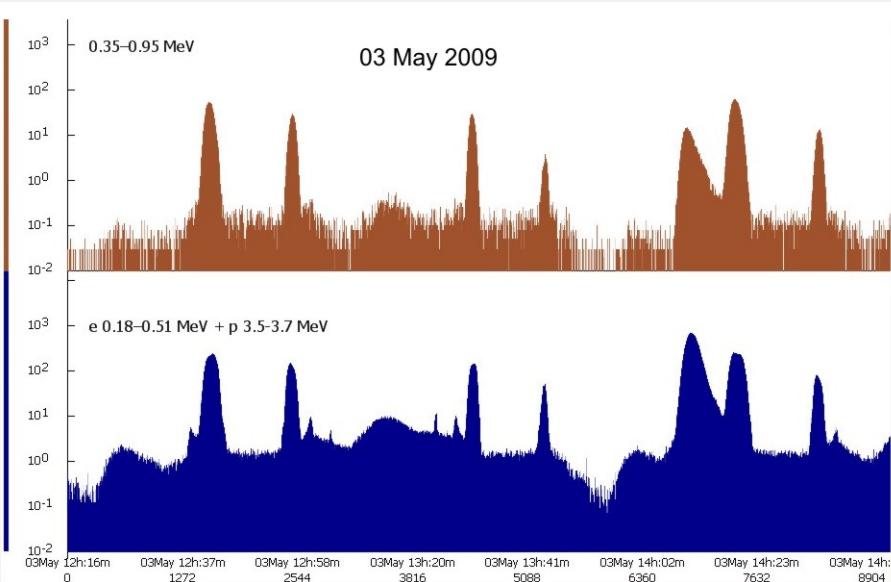


Satellite telescope of electrons and protons STEP-F

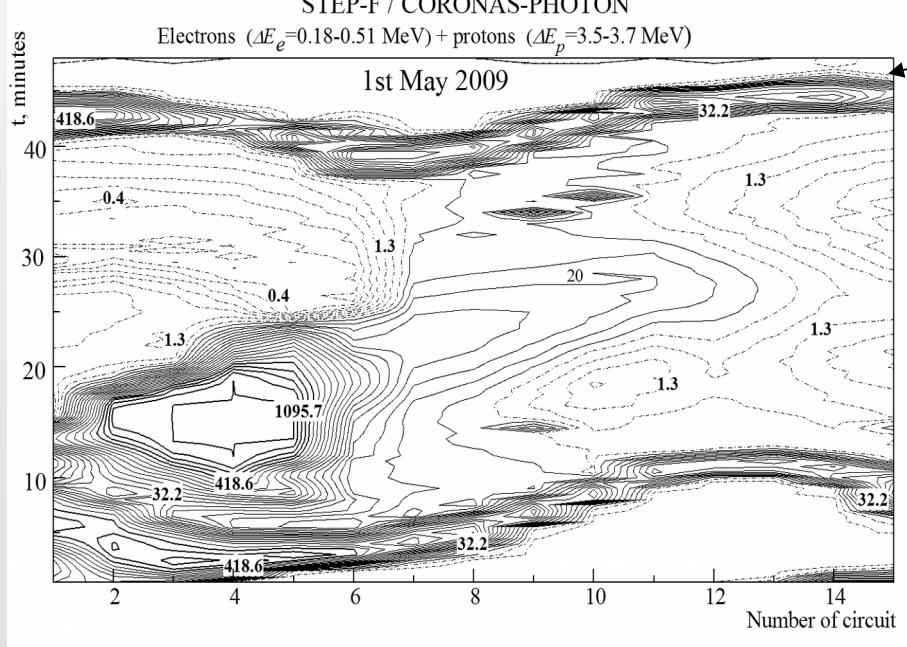


Measurements of electron fluxes of energies from $E_e = 0.18$ to ~ 1.2 MeV and protons of energies $E_p = 3.5$ to ~ 55 MeV

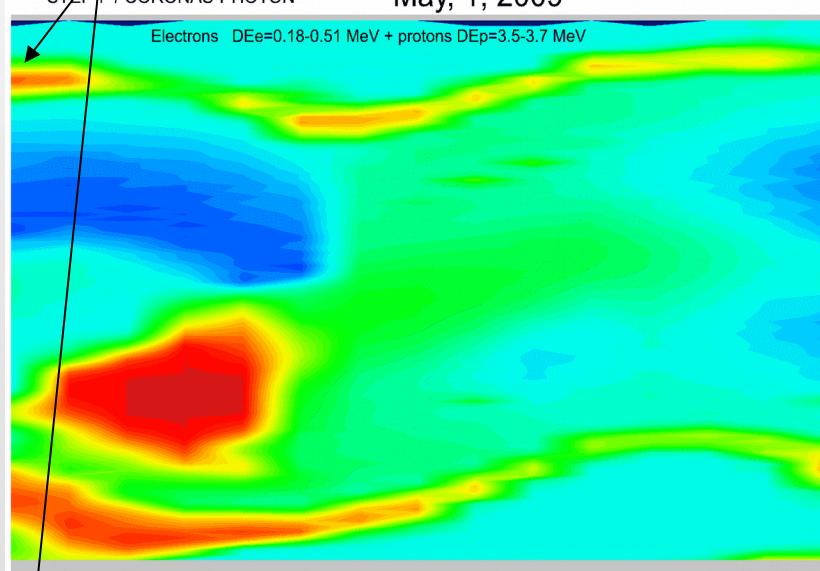
Typical initial temporal variations of electron fluxes of intermediate energies in May, 2009 from the data of the STEP-F/CORONAS-Photon device. Time resolution: 2 seconds



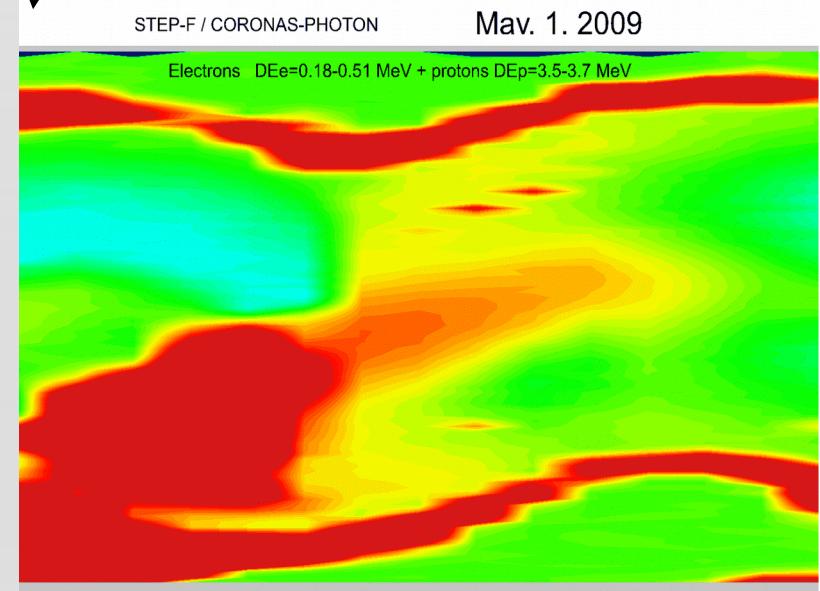
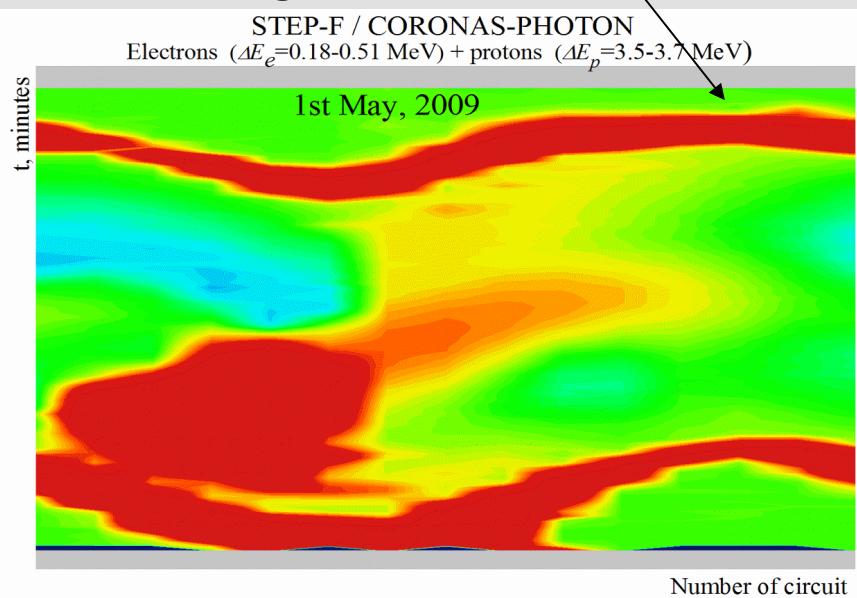
Mapping at the altitude of ~ 550 km in terms of electron fluxes, $E_e = 0.18\text{-}0.51 \text{ MeV}$; May 1, 2009



Ascending nodes of the orbit



Descending nodes of the orbit



Technical parameters of the STEP-F device

Mass:

STEP-FD — 15.4 kg;
STEP-FE — 2.7 kg;

Power consumption:

STEP-FD — 40 Watt;
STEP-FE — 8 Watt;

Dimensions:

- 1) STEP-FD:
337 mm x 395 mm x 293 mm;
- 2) STEP-FE:
95 mm x 287 mm x 160 mm;

Full angle of view:

— 108 x 108° for lowest energies;
— 98 x 98° for highest energies

Active areas:

semiconductor detectors — 20 cm²;
scintillation detectors — 36 и 49 cm²;

Geometry factors:

from 21.7 cm² · sr for low energies of particles,
to 12.4 cm² · sr for high energies of particles

Temporal resolution:

2 seconds — 12 values in each half of minutes;
30 s: 24 s collection + 6 s transmission information



Energy ranges: Electrons:

- 1) 0.35 – 0.95 MeV;
- 2) 1.2 – 2.3 MeV;
- 3) > 2.3 MeV



Protons:

- 1) 7.4 – 10.0 MeV;
- 2) 15.6 – 17.5 MeV;
- 3) 17.5 – 19.6 MeV;
- 4) 19.6 – 22.2 MeV;
- 5) 22.2 – 25.4 MeV;
- 6) 25.4 – 29.3 MeV;
- 7) 29.3 – 33.2 MeV;
- 8) 33.2 – 38.9 MeV;
- 9) 38.9 – 46.5 MeV;
- 10) 46.5 – 55.2 MeV;
- 11) > 55.2 MeV

Channels of mixed particle registration:

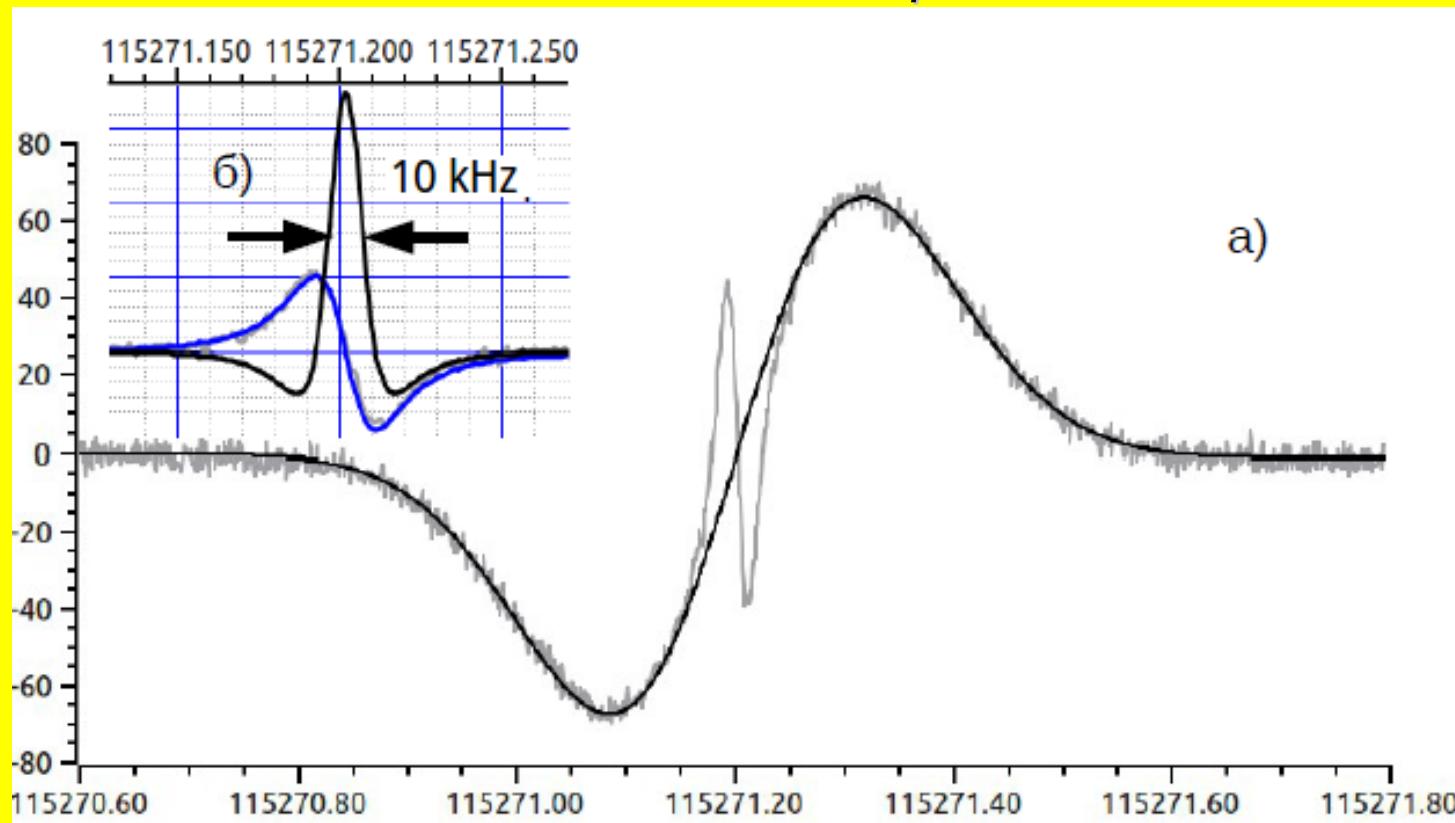
- 1) e (E_e=0.18-0.51 MeV) + p (E_p=3.5-3.7 MeV);
- 2) p (E_p=3.7-7.4 MeV) + e (E_e=0.55-0.95 MeV);
- 3) α (E_α=15.9-29.8 MeV) + p (E_p=7.4-10.0 MeV);

High-accuracy measurements: Lamb-dip record of CO transition

115271,2019±0,0005 MHz *Our measurements*

115271,2018±0,0005 MHz *Winnewisser et al*

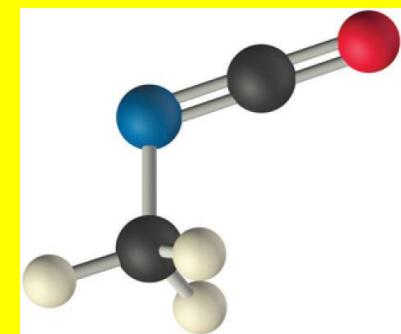
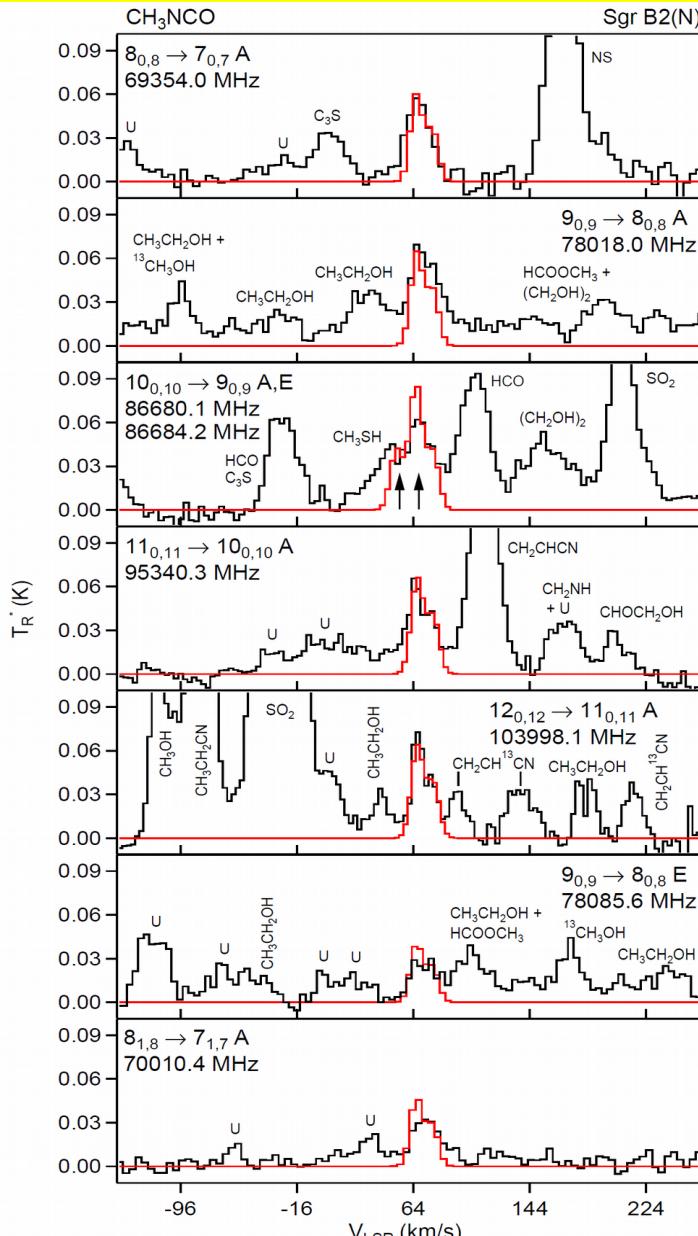
115271,2021±0,0004 MHz *Lapinov et al*



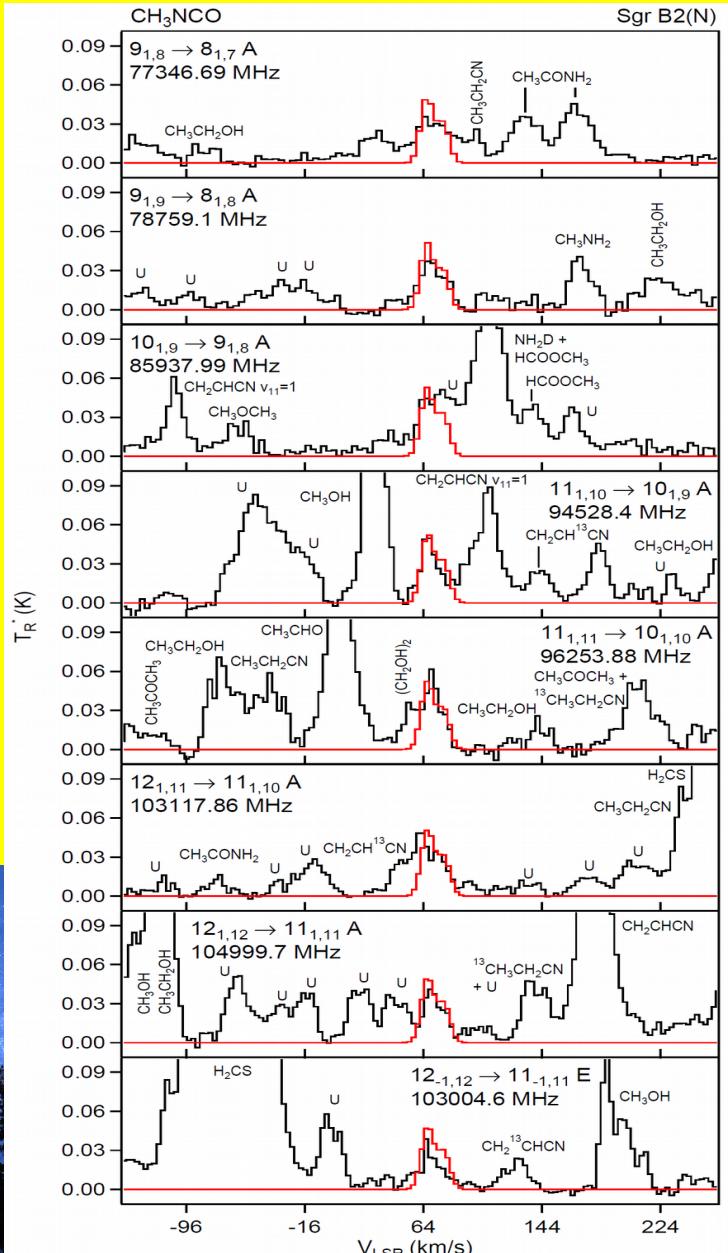


Detection of interstellar

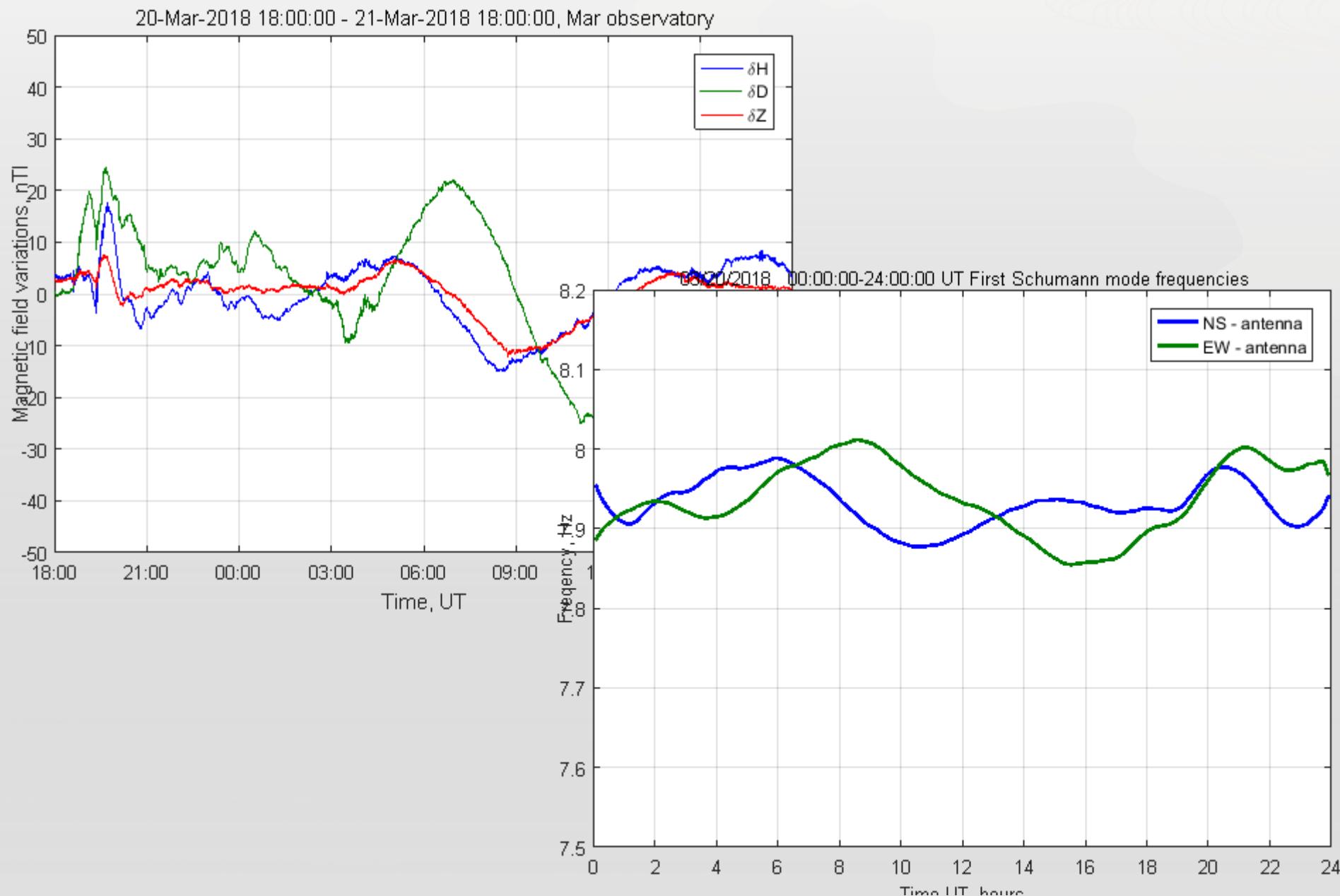
methyl isocyanate CH₃NCO in Sgr B2 (N)



3 of ~ 190 molecules in space were identified in IRA NASU



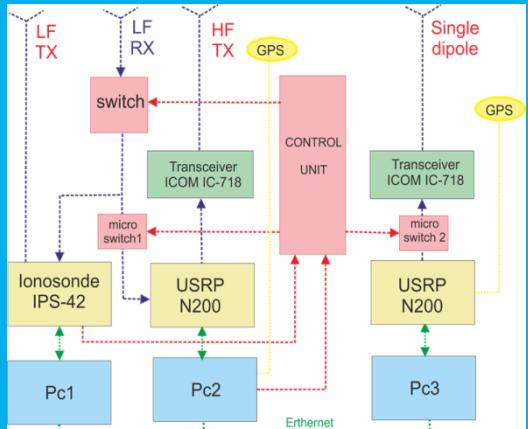
Department of Radiophysics of Geospace



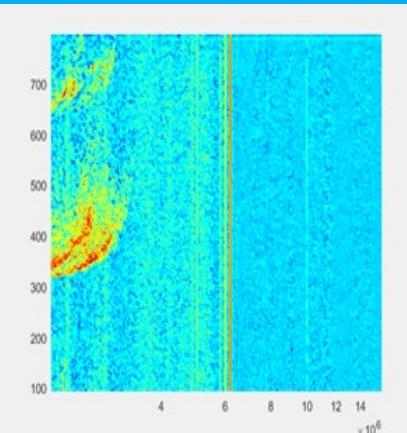
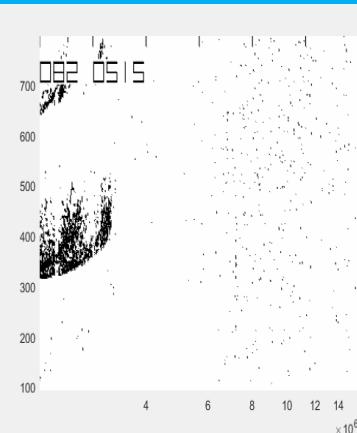


Facilities of the Antarctic station “Academician Vernadskii”

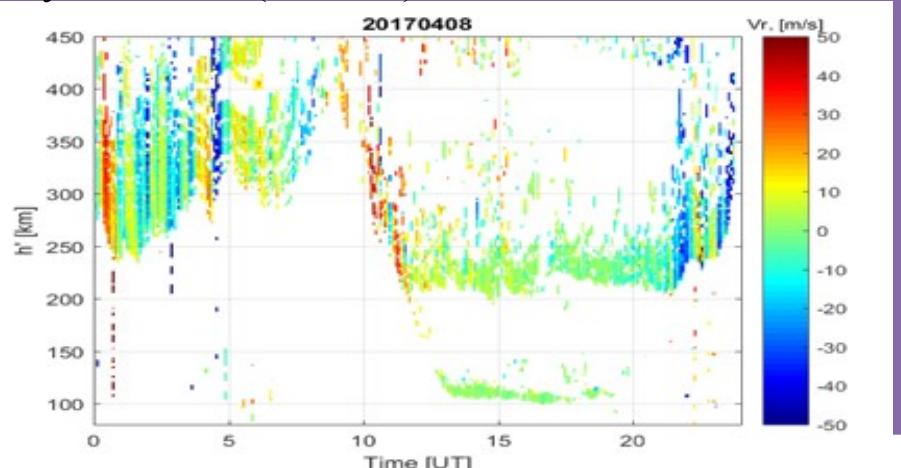
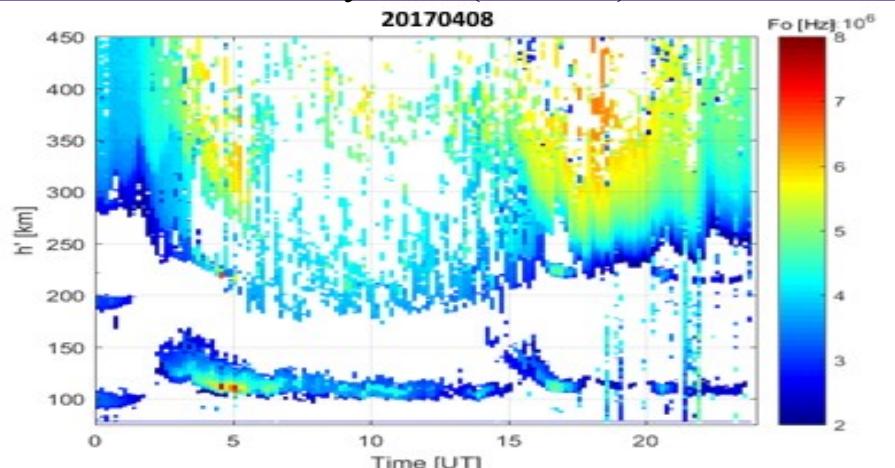
Transmitter/Receiver for ionosphere sounding



Data analyses (IPS-42 , left side) and new facilities of IRA NASU design (right side)



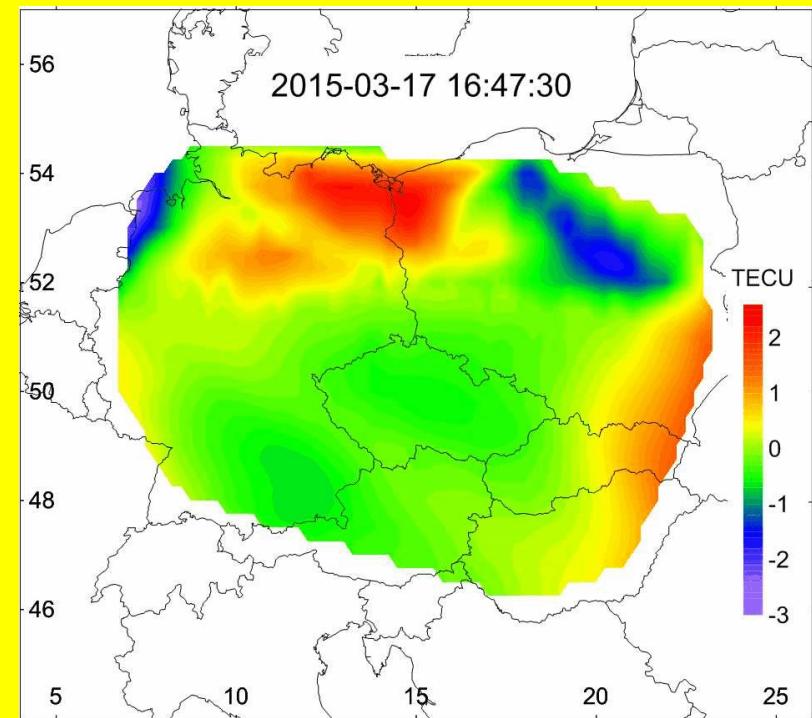
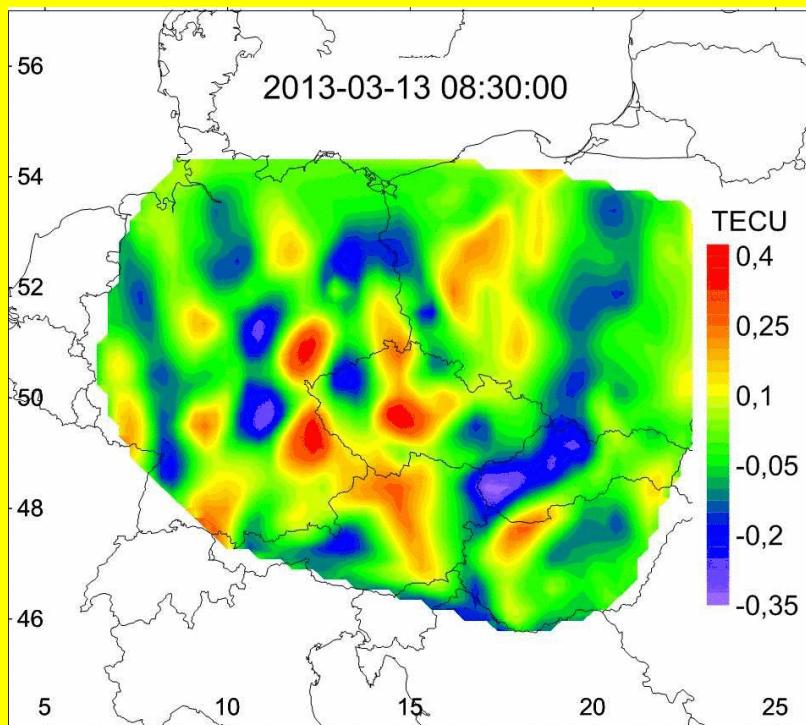
Daily TEC (left side) and Vertical velocity variations (left side)



Detection of plasma irregularities in near-Earth space

Two frequency sounding by single GPS satellite and with net of receivers:

Time delay between signals with different frequencies was measured
Quiet ionosphere Disturbed ionosphere



Department of Microwave Electronics IRA NASU



Meteorological Radar
8-mm range.

Products - Mozilla Firefox

Products > Meteorological Radars > Ka-band Doppler Polarimetric Scanning Radar (MRIAN-Ka-S-1.2)

Radar

Meteorological Radars

Antennas

Digital systems for data acquisition and processing

Transmitters

Digital receivers

Magnetrons

Power supplies

Generators

Radar positioners



International collaboration of IRA NASU



"ЗАТВЕРДЖУЮ"

Директор Радіоастрономічного інституту
Національної академії наук України



2016 р.

"ЗАТВЕРДЖУЮ"

Ректор Європейського національного
педагогічного університету ім.
Олександра Довженка



2016 р.

ДОГОВІР про наукове співробітництво

Договір укладено з метою розвитку і налагодження плідного співробітництва установами Національної академії наук України та установами Міністерства науки України, координації їх спільнотої діяльності та виконання у найбільш обсязі Закону України "Про вищу освіту" та Закону України "Про наукову та технічну діяльність".

Національна академія наук
України
РADIOАСТРОНОМІЧНИЙ
ІНСТИТУТ
ПІ НАНУ
вул. Мистецтв 4, м. Харків, 61002
тел/факс: +38(057)706-14-15
E-mail: rian@rian.kharkov.ua
Web: <http://ri.kharkov.ua>
Код СДРІОУ 02772020



National Academy of Sciences of
Ukraine
INSTITUTE OF RADIO
ASTRONOMY
IRA NASU
4, Mysletsiv St, Kharkiv, 61002, Ukraine
tel: +38(057)315 20 92
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E-mail: rian@rian.kharkov.ua

09.11. 2017 р. № 287-574/11

На
від
GURT mini array specification

Financial Information

The detailed cost break down is given in the Table 2

	#	€/each (incl. VAT*)	Total € (incl. VAT*)	Remarks
1. Antenna array				
Low band antennas	25	€ 210	€ 5250	
Coaxial cable	2500 m	€ 1	€ 2500	RG-58u or better
Equipment cabinet with construction materials	1	€2100	€2100	Incl. power and connec coax cable
2. Amplification system				
Antenna pre-amplifiers	60	€ 100	€ 6000	incl. 10
5-channels amplifiers	3	€ 750	€ 2250	incl. 1
Base amplifiers	3	€ 200	€ 600	incl. 1
Output amplifiers	3	€ 250	€ 750	incl. 1
Power suppliers	12	€ 240	€ 2880	incl. 2
3. Phasing & control				
Phase shifters	14	€ 1300	€ 18200	incl. 2
Beam former & control board	1	€ 1500	€ 1500	
4. Back-ends facility				
Digital receiver ADR incl. software	1	€ 24800	€ 24800	
Total			€ 66 830	

Table 2. Financial overview.

The total costs of GURT mini array is 66 830 € *

* Currently there is 0% VAT rate in Ukraine for the export of goods and services.

All local infrastructure related items and works as well as transport costs are excluded labour hours for the system integration and tests at the local site are excluded. The documentation will be delivered. All GURT system components are electronically tested before shipment. The quotation includes a one year guarantee for all electronical

Yours sincerely,

Director

Dr. Sc.



V. V. Zakharenko

Deputy director on science
Academician NASU

O. O. Konovalenko

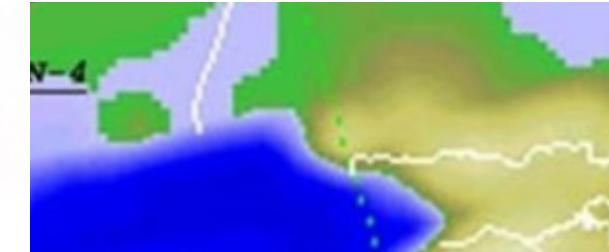


УГОДА ПРО СПІВРОБІТНИЦТВО МІЖ
ВЕНТСПІЛЬСЬКИМ УНІВЕРСИТЕТОМ
ПРИКЛАДНИХ НАУК (ЛАТВІЯ) ТА
Радіоастрономічним
Інститутом НАН України

Вентспільський Університет Прикладної
Наук – юридична особа, яка зареєстрована та діє
відповідно до Закону про Конституцію
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Професор д-р Карліс Крелінс (далі - Ректор), та

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