





Hunting for extreme blazars in the TeV band

Luca Foffano, Elisa Prandini, Simona Paiano, Cornelia Arcaro On behalf of the MAGIC collaboration

and Alberto Franceschini

AGN13, Milano October 9th, 2018

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Overview

- \geq Blazars and extreme blazars
- Extreme blazars hunting program in MAGIC \succ
- Looking for new TeV extreme blazars candidates \succ
- PGC 2402248: a new TeV extreme blazar detected by the MAGIC telescopes \succ





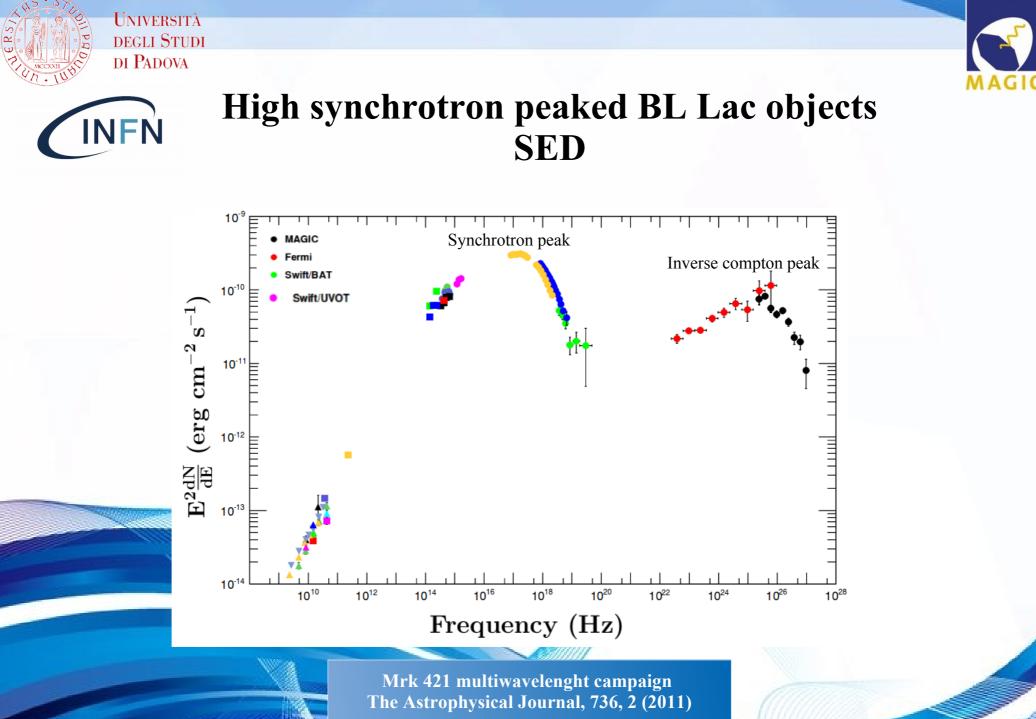




Blazars and extreme blazars

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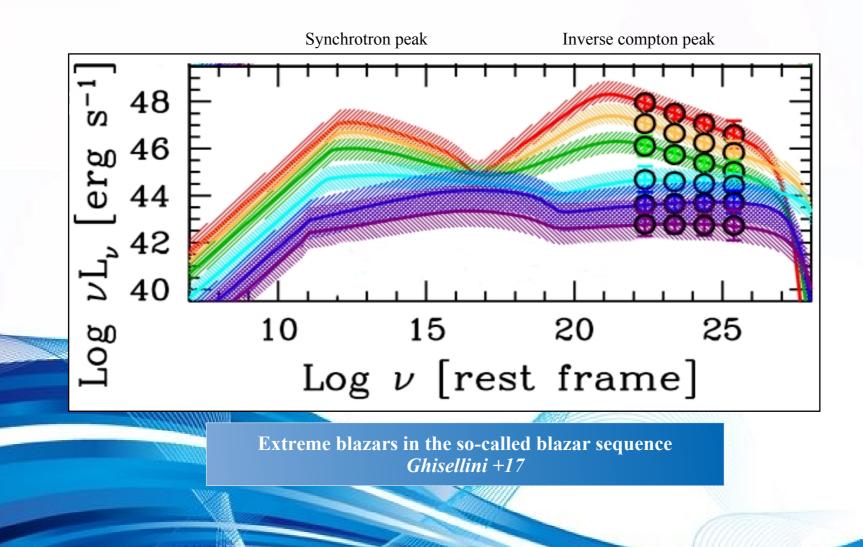
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Blazars and extreme blazars



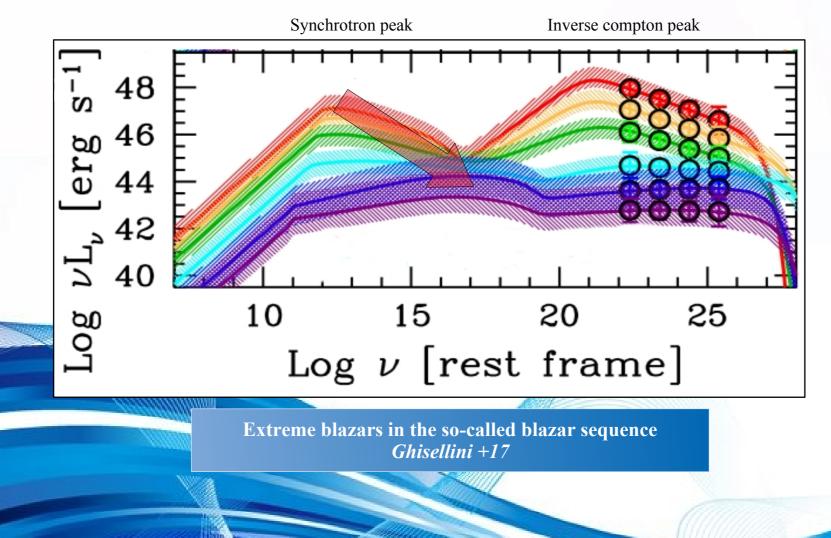
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Blazars and extreme blazars



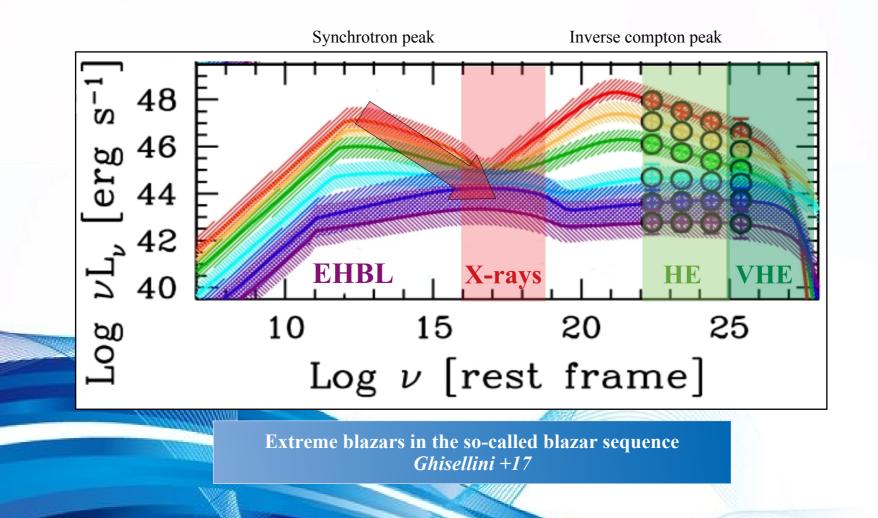
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Blazars and extreme blazars



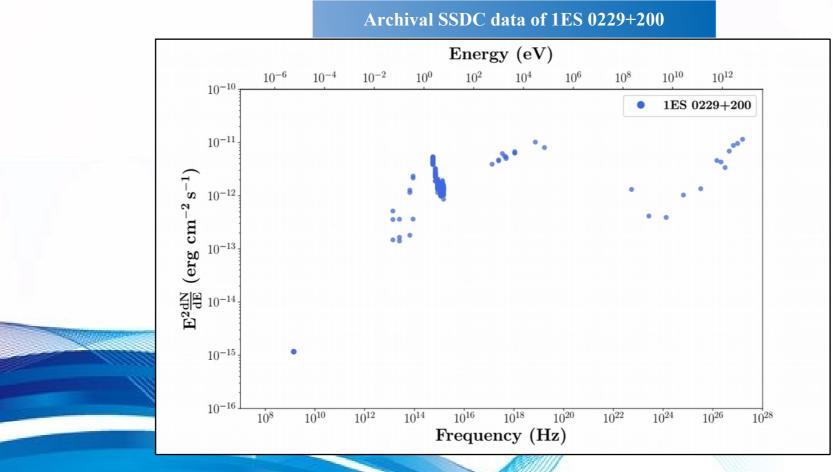
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Extreme blazars



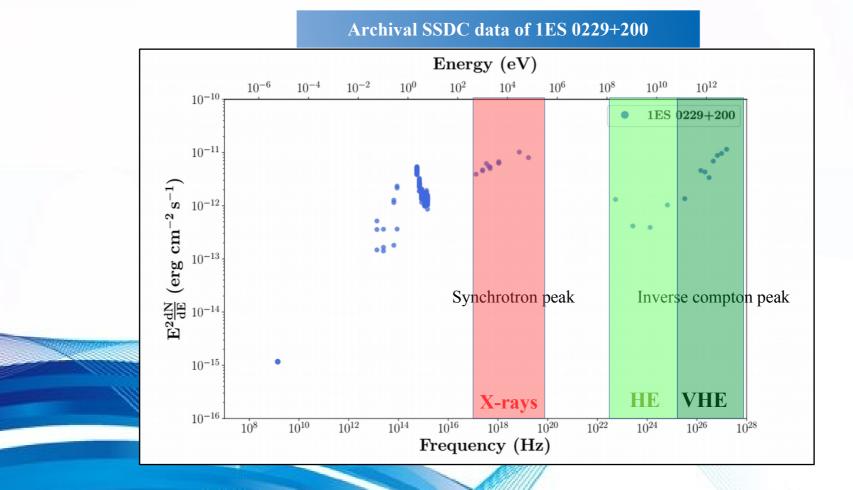
We are looking for the most extreme blazars in the VHE gamma-ray band

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Extreme blazars



We are looking for the most extreme blazars in the VHE gamma-ray band

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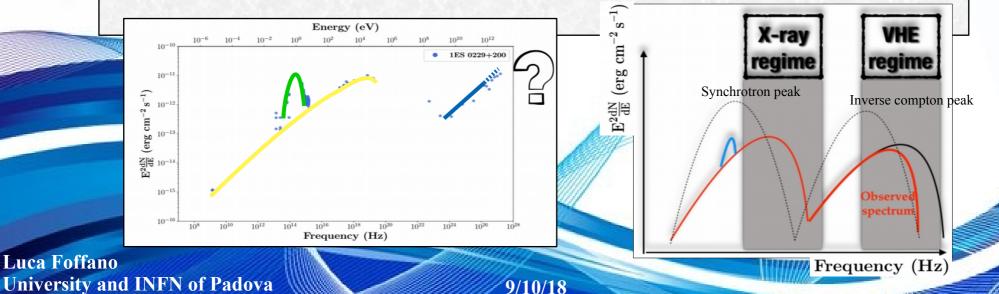
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Extreme blazars

- High synchrotron peak frequency $> 10^{17}$ Hz: hard X-rays
- High **"inverse compton" peak** frequency > 10²⁶ Hz: VHE
- Hard spectrum in X-rays and gamma-rays (important for cosmology wrt EBL, extragalactic studies IGMF, and neutrinos)
- Relatively low luminosity with respect to other blazars
- Potentially absorbed in the VHE due to interaction with optical-infrared diffuse background (EBL)
- **Galaxy** should be **detectable** in the optical range
- Not well detected in the HE gamma-ray band (Fermi-LAT)





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Extreme blazars

- Synchrotron Self-Compton (SSC) model (only leptonic) works but with extreme parameters
 - \rightarrow extremely low magnetization in the emission zone, very high doppler factor or minimum Lorentz factor (e.g. see Tavecchio+09 and Cerruti+15)
- Lepto-hadronic emission models are favoured, but with a lot of free parameters...
- Should photo-hadronic processes be preferred?
- HBL are good candidates for neutrino emission (e.g. see Padovani+16 and Resconi+17)
 → what about EHBLs? Could EHBLs be neutrino sources?
- Only few sources are known TeV EHBLs...

1ES 0229+200, 1ES 0347-121, RGB J0710+591, 1ES 1101-232

We need more sources!

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Extreme blazars with the MAGIC telescopes

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Extreme blazars with MAGIC

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Paper in preparation

	Source	Z	Log V _{sync,peak}
	TXS 0210+515	0.049	17.3
	BZB J0809+3455	0.083	16.6
	RBS 0723	0.198	17.8
	1ES 0927+500	0.187	17.5
	RBS 0921	0.236	17.9
	1ES 1426+428	0.129	18.1
	1ES 2037+521	0.053	n.a.
	RGB J2042+244	0.104	17.5
	RGB J2313+147	0.163	17.7
	1ES 0229+200	0.140	18.5
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- Nine new targets
- Extreme synchrotron peak (from 2WHSP Catalog)
- **Low redshift**: 0.049 < z < 0.236
- More than 200 hours of data
- All the targets are **TeV undetected** (except for 1ES 1426+428) (Aharonian, F. et al.) (HEGRA Collaboration 2003, A&A, 403, 523)
- 1ES 0229+200 considered here for comparison
- Now performing **SED modelling** of these sources

4 detections + 1 hint



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Extreme blazars with MAGIC

Paper in preparation

	Source	Z	Log V _{sync,peak}	$z^{s}_{a} = 160 \begin{bmatrix} TXS 0210+515 \\ 140 \end{bmatrix} TXS 0210+515 \\ Preliminary \\ N_{ex} = 91.8 \pm 16.8 \\ Significance (Li&Ma) = 5.95\sigma \end{bmatrix}$
	TXS 0210+515	0.049	17.3	
	BZB J0809+3455	0.083	16.6	
	RBS 0723	0.198	17.8	
	1ES 0927+500	0.187	17.5	
	RBS 0921	0.236	17.9	θ^{2} [deg ²]
	1ES 1426+428	0.129	18.1	≝ 350
	1ES 2037+521	0.053	n.a.	350 RBS 0723 Time = 45.31 h 300 Preliminary Non = 500; Non = 378.2 ± 8.7 Nex = 121.8 Significance (Li&Ma) = 5.40σ
	RGB J2042+244	0.104	17.5	
	RGB J2313+147	0.163	17.7	
	1ES 0229+200	0.140	18.5	
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Extreme blazars with MAGIC

Paper in preparation

	Source	Z	Log V _{sync,peak}	2 60 50 50 50 50 50 50 50 50 50 50 50 50 50 5
	TXS 0210+515	0.049	17.3	40
	BZB J0809+3455	0.083	16.6	
	RBS 0723	0.198	17.8	
	1ES 0927+500	0.187	17.5	0 0.1 0.2 0.3 0.4
	RBS 0921	0.236	17.9	θ ² [deg ²] ∞ 240
	1ES 1426+428	0.129	18.1	$\begin{array}{c} 5 \\ 220 \\ 200 \\ 2$
	1ES 2037+521	0.053	n.a.	180 Significance (Li&Ma) = 7.49σ
	RGB J2042+244	0.104	17.5	
	RGB J2313+147	0.163	17.7	
	1ES 0229+200	0.140	18.5	$\begin{array}{c} 20 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 0 $
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Extreme blazars with MAGIC

Paper in preparation

	Source	Z	Log V _{sync,peak}	$\sum_{n=1}^{100} 250 \begin{bmatrix} RGB J2042+244 \\ Preliminary \end{bmatrix} $ Time = 52.51 h Non = 404; Noff = 322.6 ± 10.4 Nex = 81.4 ± 22.6 Significance (Li&Ma) = 3.74\sigma
	TXS 0210+515	0.049	17.3	
	BZB J0809+3455	0.083	16.6	
	RBS 0723	0.198	17.8	
	1ES 0927+500	0.187	17.5	
	RBS 0921	0.236	17.9	$\theta^{0} 0 0.1 0.2 0.3 0.4 \theta^{2} [deg^{2}]$
	1ES 1426+428	0.129	18.1	= 117.46 h Non = 117.46 h Non = 1182; N _{eff} = 857.0 ± 16.8
	1ES 2037+521	0.053	n.a.	Z 700 IES 0229+200 Non = 1182; Noff = 857.0 ± 16.8 Preliminary Nex = 325.0 Significance (Li&Ma) = 8.96σ
	RGB J2042+244	0.104	17.5	
	RGB J2313+147	0.163	17.7	
	1ES 0229+200	0.140	18.5	200
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Extreme blazars with MAGIC

Paper in preparation

SED modelling

(following Asano+14)

SED modelling performed on the four new detections:

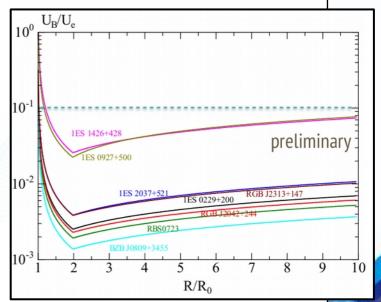
- Steady outflow
- Electron injection during the dynamical scale

Some considerations:

- SSC emission efficiency is low
- Relatively high electron energy density
- Lower magnetization compared to other usual blazars
- Synchrotron and IC peaks are generally not well constrained
- Precise TeV and X-ray data are crucial

Main physical processes considered:

- Electron injection
- Synchrotron emission and cooling
- Inverse Compton emission and cooling
- Adiabatic cooling
- Photon escape
- No electron escape!



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Looking for new TeV extreme blazars

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Looking for new TeV EHBL candidates



We used the Swift-BAT 105-months hard X-ray catalogue

- Selecting blazars with good detection in the hard X-ray band
- Checking the synchrotron peak location above 10¹⁷ Hz (taking care of misclassifications and errors in peak estimations in the catalogues)
- Performing a new *Fermi*-LAT analysis over 10 years of data to verify the detection in the HE gamma-ray band
- Extrapolating the spectrum up to the **TeV gamma-ray** band to check detectability by Cherenkov telescopes
- Studying the final sample of 34 candidates, and looking for relations between already TeV detected and TeV undetected sources

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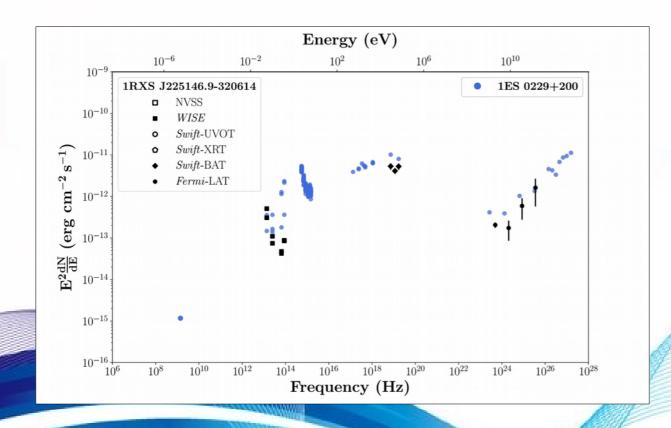


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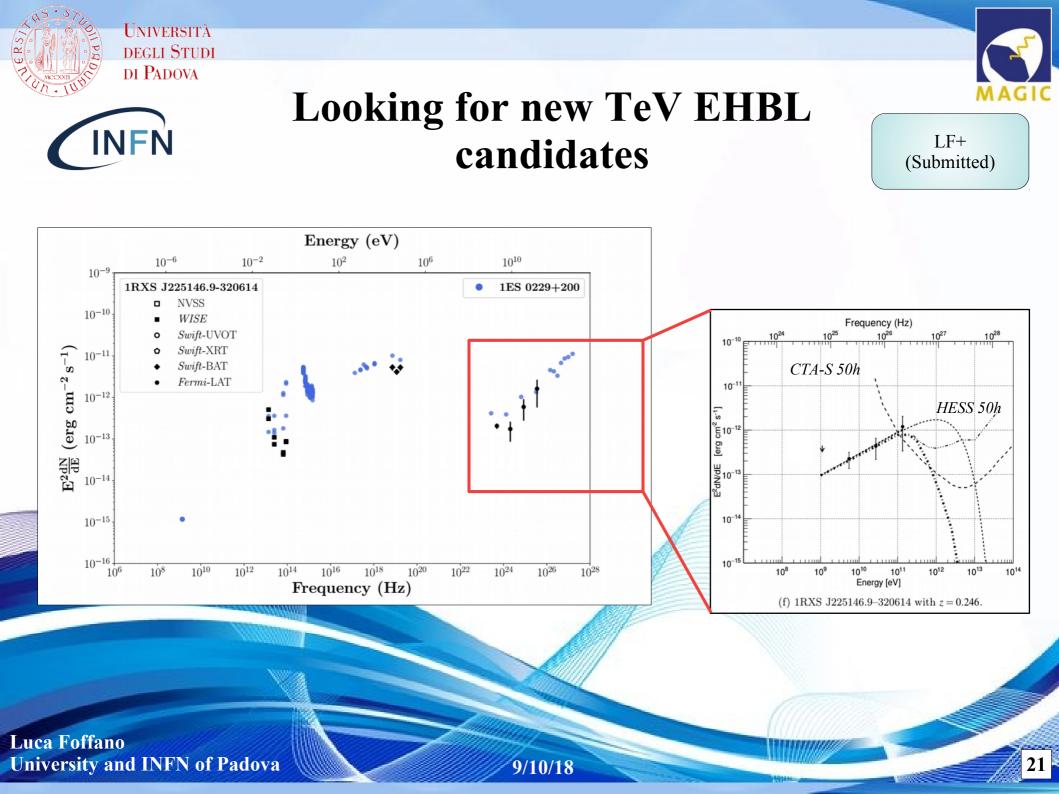
DI PADOVA

Looking for new TeV EHBL candidates

LF+ (Submitted)



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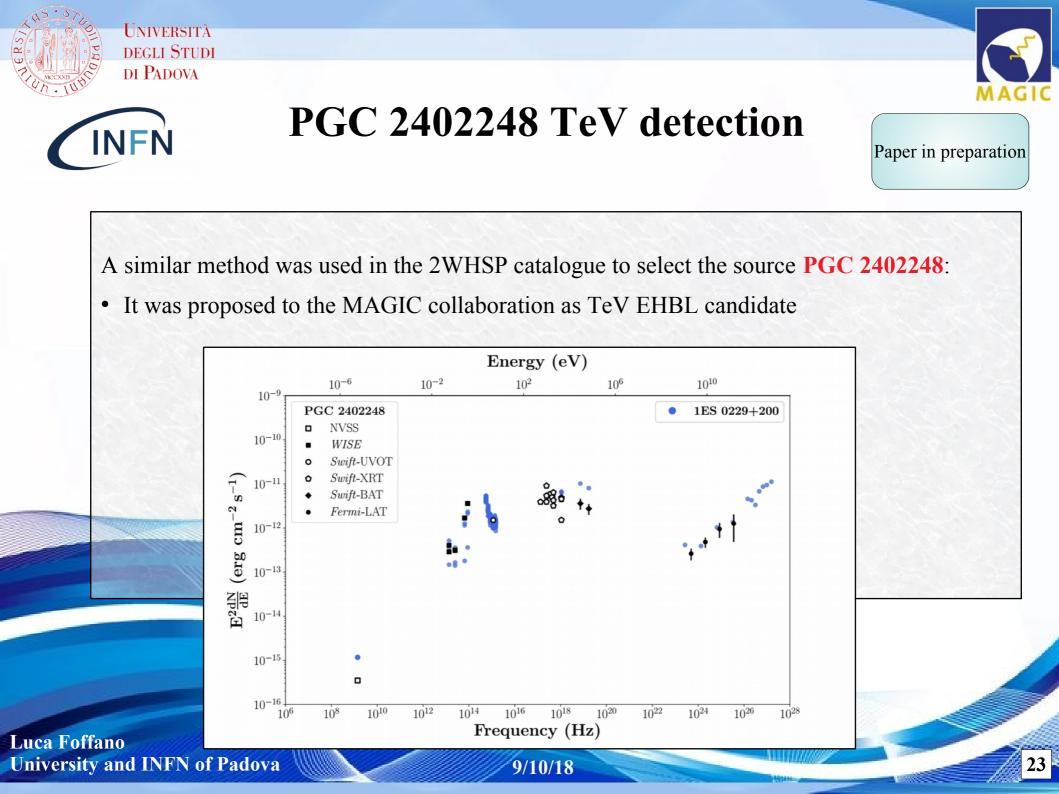






PGC 2402248 TeV detection with the MAGIC telescopes

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PGC 2402248 TeV detection

Paper in preparation

A similar method was used in the 2WHSP catalogue to select the source PGC 2402248:

- It was proposed to the MAGIC collaboration as TeV EHBL candidate
- It was accepted as first source to be observed
- It was **successfully** detected by MAGIC after 23 h

First detection of very-high-energy gamma-ray emission from the extreme blazar PGC 2402248 with the MAGIC telescopes

ATel #11548; Razmik Mirzoyan (Max-Planck-Institute for Physics, Munich), on behalf of the MAGIC collaboration on 19 Apr 2018; 15:30 UT Credential Certification: Razmik Mirzoyan (Razmik.Mirzoyan@mpp.mpg.de)

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Conclusions

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Take home message

- Extreme blazars could be a new category of blazars with extreme spectral parameters
- We need more TeV gamma-ray detected EHBLs
- The MAGIC collaboration detected **5 new TeV EHBLs** + **1 hint**
- We are extracting **new TeV EHBL candidates** studying their proprierties in the **hard X-ray** and **HE gamma-ray** bands

 \rightarrow This method was successful in detecting PGC 2402248 in the VHE gamma-ray band by the MAGIC telescopes

• This work will continue with other catalogues

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tre/Mel9

22-25 January 2019 Padova

Topics:

- The blazar family
- Mechanisms of particle acceleration and radiation in jets
- Extreme blazars as neutrino factories •
- Extreme blazars and the connection with the highest energy cosmic rays.
- Cosmology and fundamental physics
- Future observatories



Scientific Organising Committee

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email extreme19@dfa.unipd.it

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Abstract welcome by October 31st







