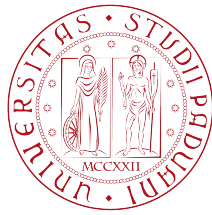




**INAF**  
ISTITUTO NAZIONALE  
DI ASTROFISICA



**AQUEYE+IQUEYE**

# Data Reduction

## Aqueye+/Iqueye/IFI+/EFI

**Michele Fiori**

***INAF – Astronomical Observatory of Padova***

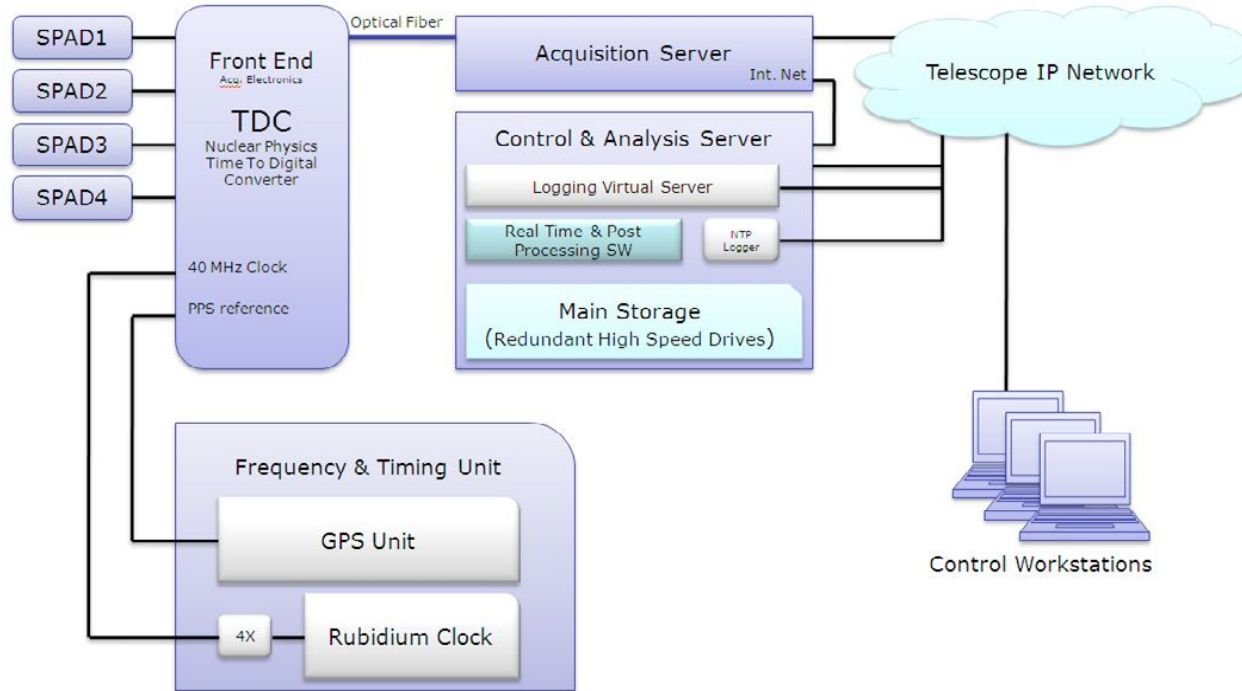
L. Zampieri, G. Naletto, A. Spolon, S. Conforti, C. Barbieri,  
U. Munari, L. Lessio, T. Forte, A. Frigo, M. Mosele,  
L. Trasverso, P. Ochner  
and the **Aqueye+Iqueye Team**

**Workshop on Fast Optical Bursts - Asiago - 04-07/05/2026**



# Iqueye data acquisition system (IDAS)

<https://www.aanda.org/articles/aa/pdf/2009/46/aa12862-09.pdf>



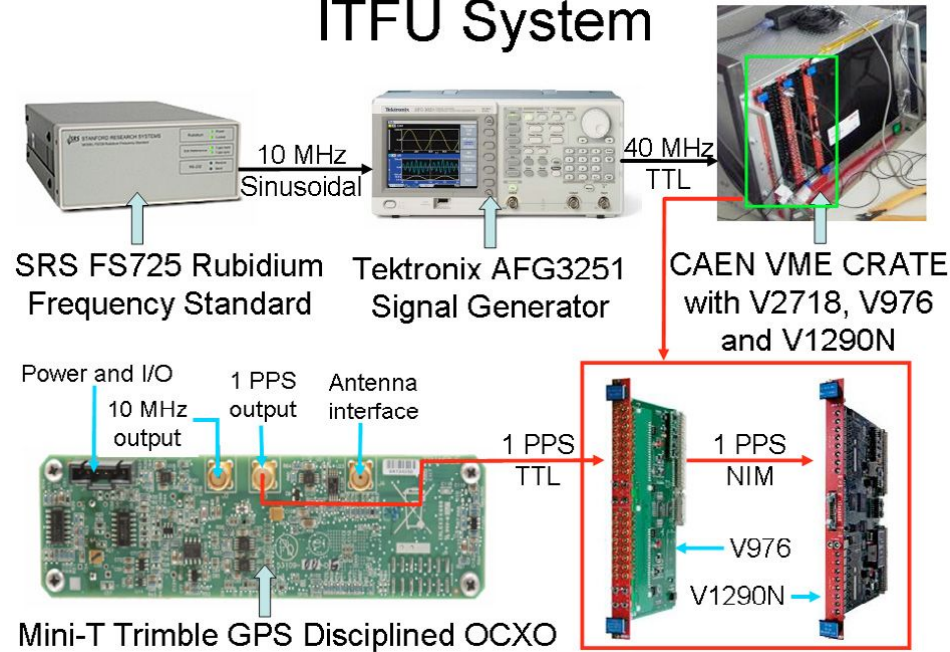
IDAS collect pulses produced from each SPAD when a photon is detected, assign a very accurate absolute time tag to each event, and store the time tags in the external memory.

# Iqueye time and frequency unit (ITFU)

<https://www.aanda.org/articles/aa/pdf/2009/46/aa12862-09.pdf>



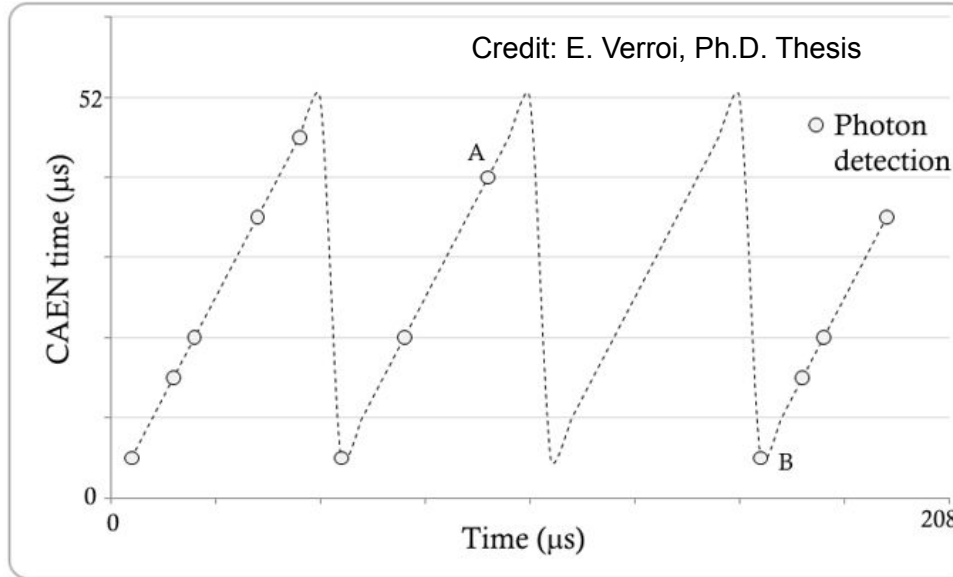
## ITFU System



ITFU needed to correct the long-term drift of the reference frequency provided by the rubidium oscillator by means of a post-processing algorithm that uses the PPS signal provided by a GPS receiver. The GPS also provides the synchronization to UTC.

# Rollover

<https://www.aanda.org/articles/aa/pdf/2009/46/aa12862-09.pdf>



Data acquisition electronics is fed with an external signal to avoid to lose any rollover

The times of arrival are stored in a buffer using 21 bits for the time tag. This means that after  $2^{21} \cdot 25 \text{ ps} \approx 52 \text{ } \mu\text{s}$  the counter is reset and the counting re-starts from zero, so the time tags coming from the TDC are actually number digitalized at 25 ps spanned in the range  $[0 \div 52 \text{ ms}]$ . That implies the non-uniqueness of each tag. We called **rollover** each reset of **these 52  $\mu\text{s}$  period**.

# Absolute and relative timing performances

<https://www.aanda.org/articles/aa/pdf/2009/46/aa12862-09.pdf>



## Absolute time accuracy

$$\sqrt{25^2 + 35^2 + 50^2 + 50^2} \approx 100 \text{ ps}$$

Resolution TDC + Detector jitter + Stochastic residual noise rubidium clock + Residual jitter in electronic chain

## Time accuracy wrt UTC

$$\sqrt{100^2 + (25000/\sqrt{3600})^2} \approx \sqrt{100^2 + 400^2} \lesssim 500 \text{ ps}$$

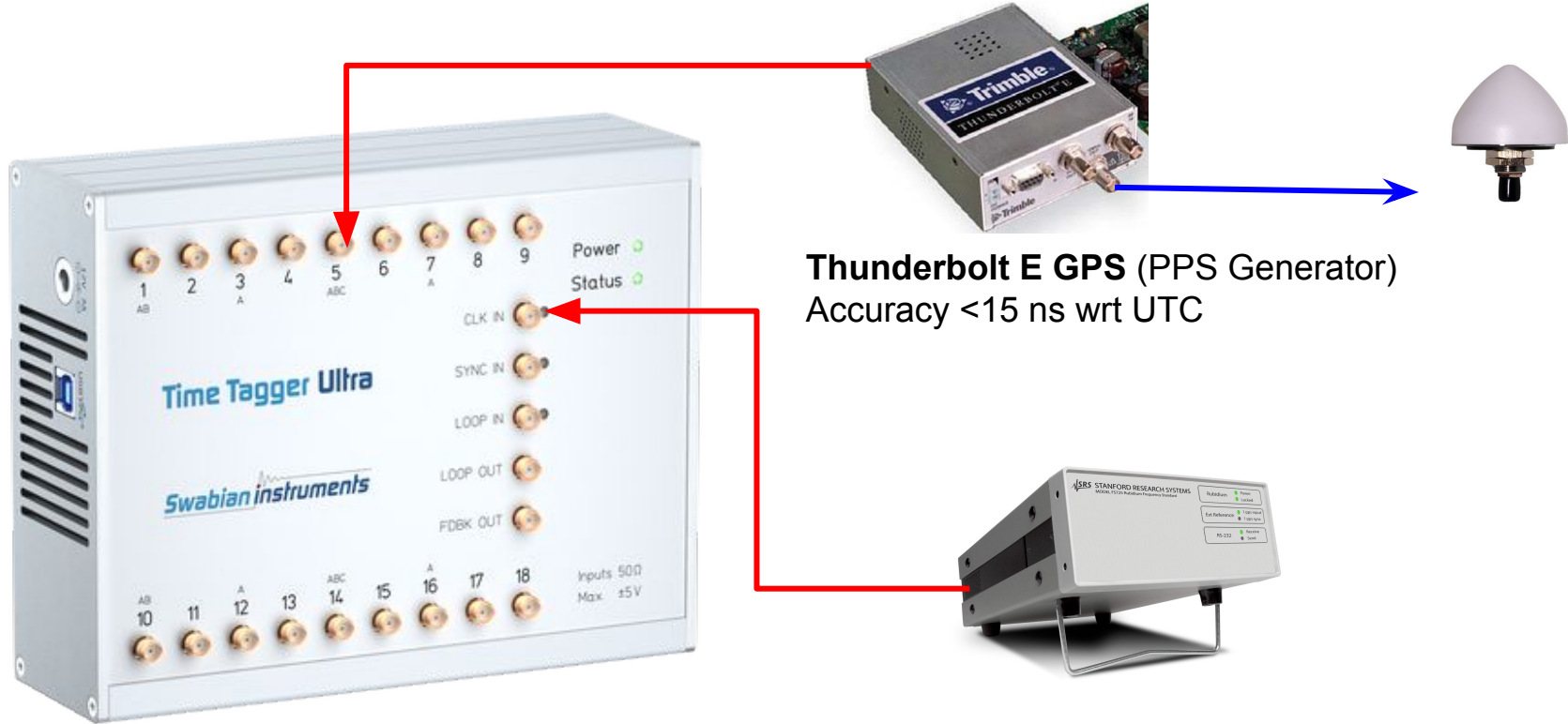
PPS rms uncertainty approximately  $\pm 25$  ns  $\rightarrow$  Interpolation reduces the error associated with the starting time in a way proportional to the number of acquired PPS  $\rightarrow$  After one hour the uncertainty become  $\sim 400$  ps.

# Example (from ASTRI-SI<sup>3</sup> BEE laboratory test)

Accuracy achievable using a TDC + GPS receiver + Rubidium clock



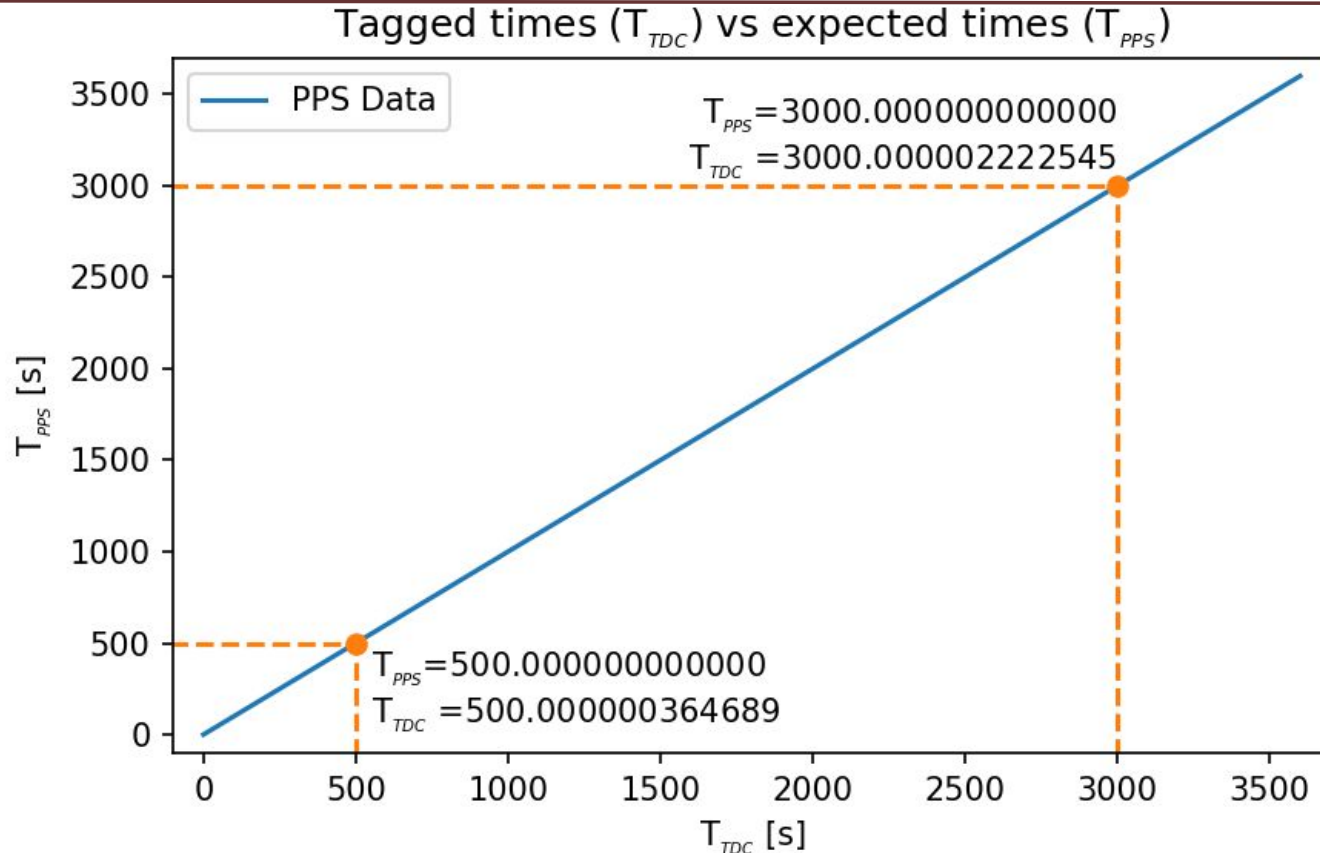
## Setup





## Example (from ASTRI-SI3 laboratory test)

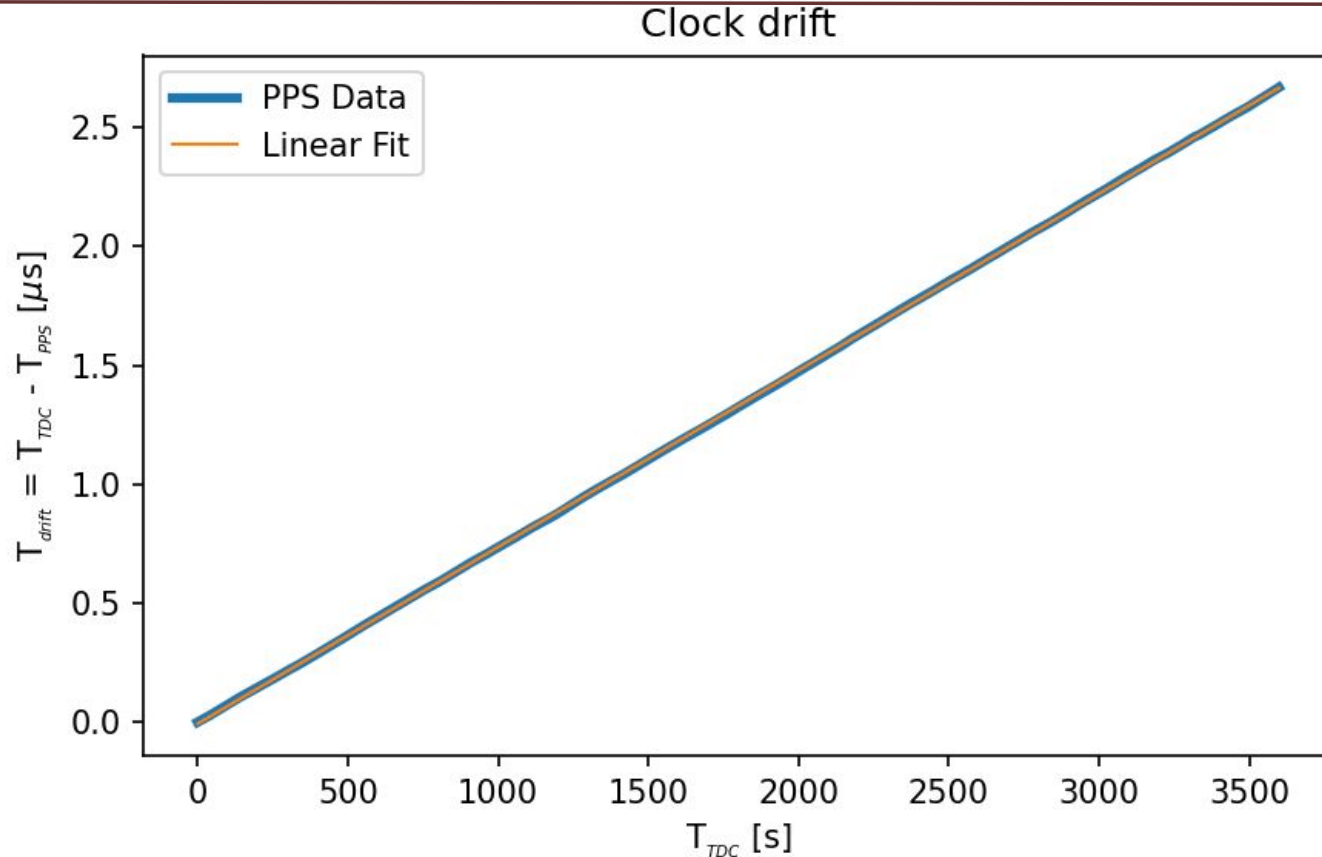
Comparison between the times in output from the Time Tagger and the “expected” times of the PPS.





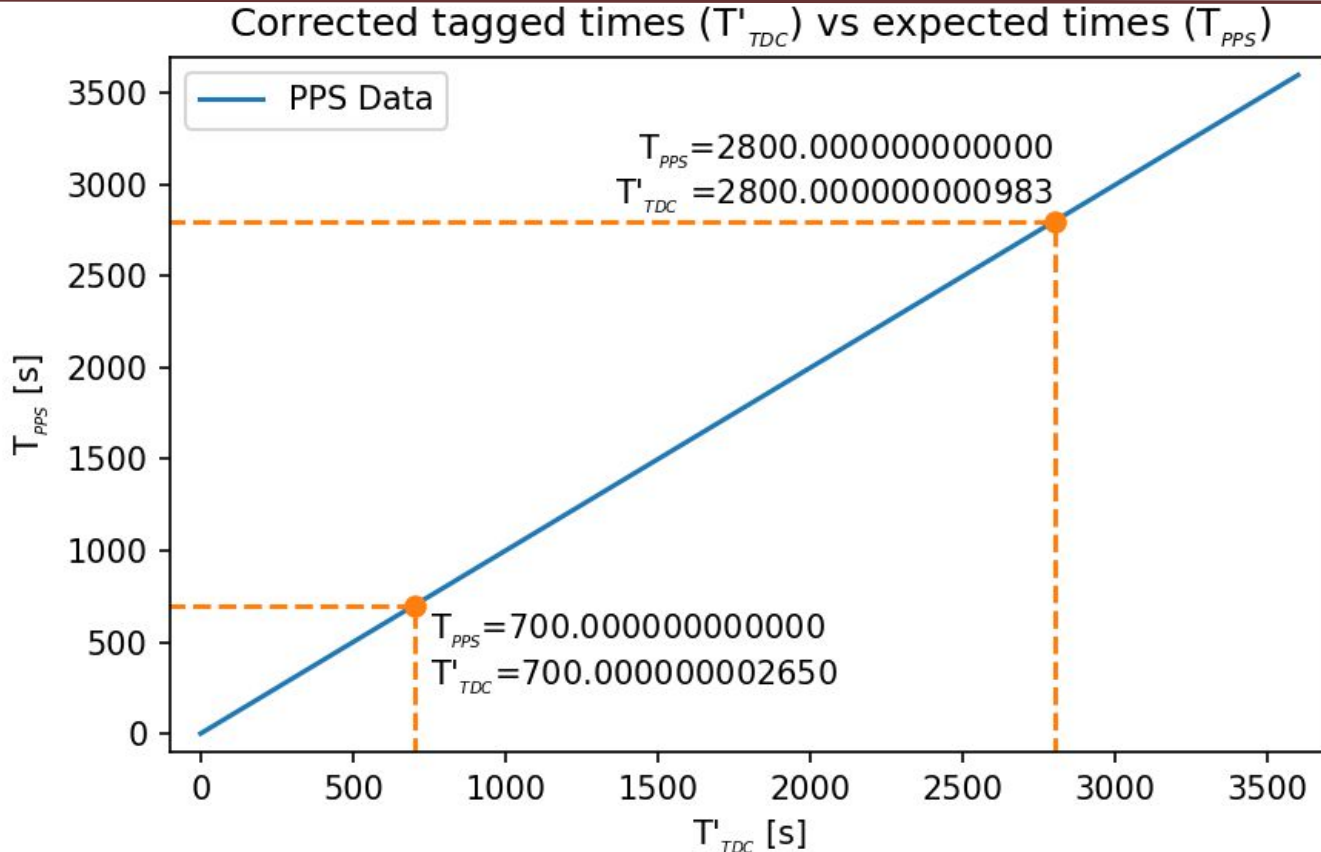
## Example (from ASTRI-SI3 laboratory test)

We fit  $T_{\text{drift}}$  (the difference between TTDC and TPPS)  $\rightarrow T_{\text{drift,fit}} = f(T_{\text{TDC}}) = a + b \cdot T_{\text{TDC}}$



# Example (from ASTRI-SI3 laboratory test)

Removing  $T_{\text{drift,fit}}$  to  $T_{\text{TDC}}$  we correct for the linear drift of the clock.

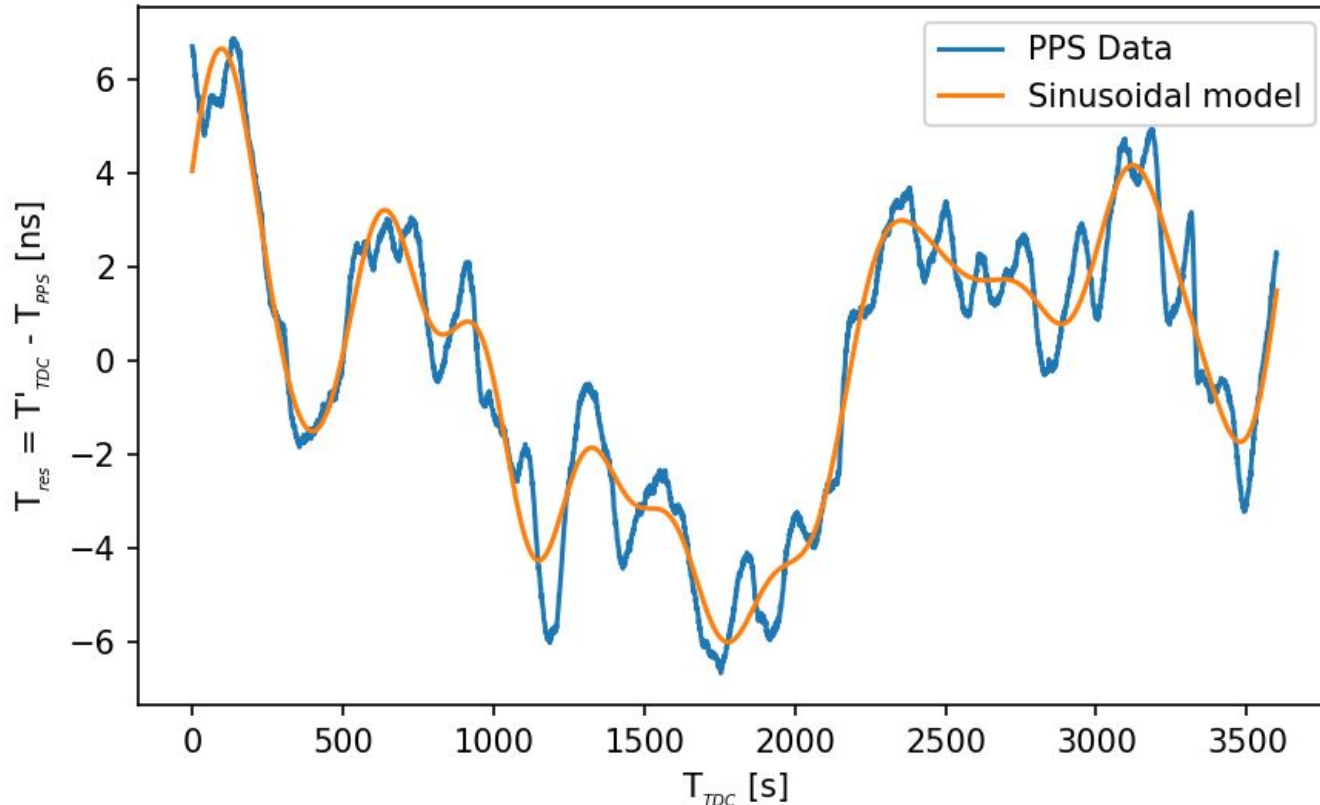




## Example (from ASTRI-SI3 laboratory test)

PPS signal residuals ( $T'_{TDC} - T_{PPS}$ ) after removing the drift of the clock using a linear fit  $\rightarrow T_{res,fit} = g(T'_{TDC})$

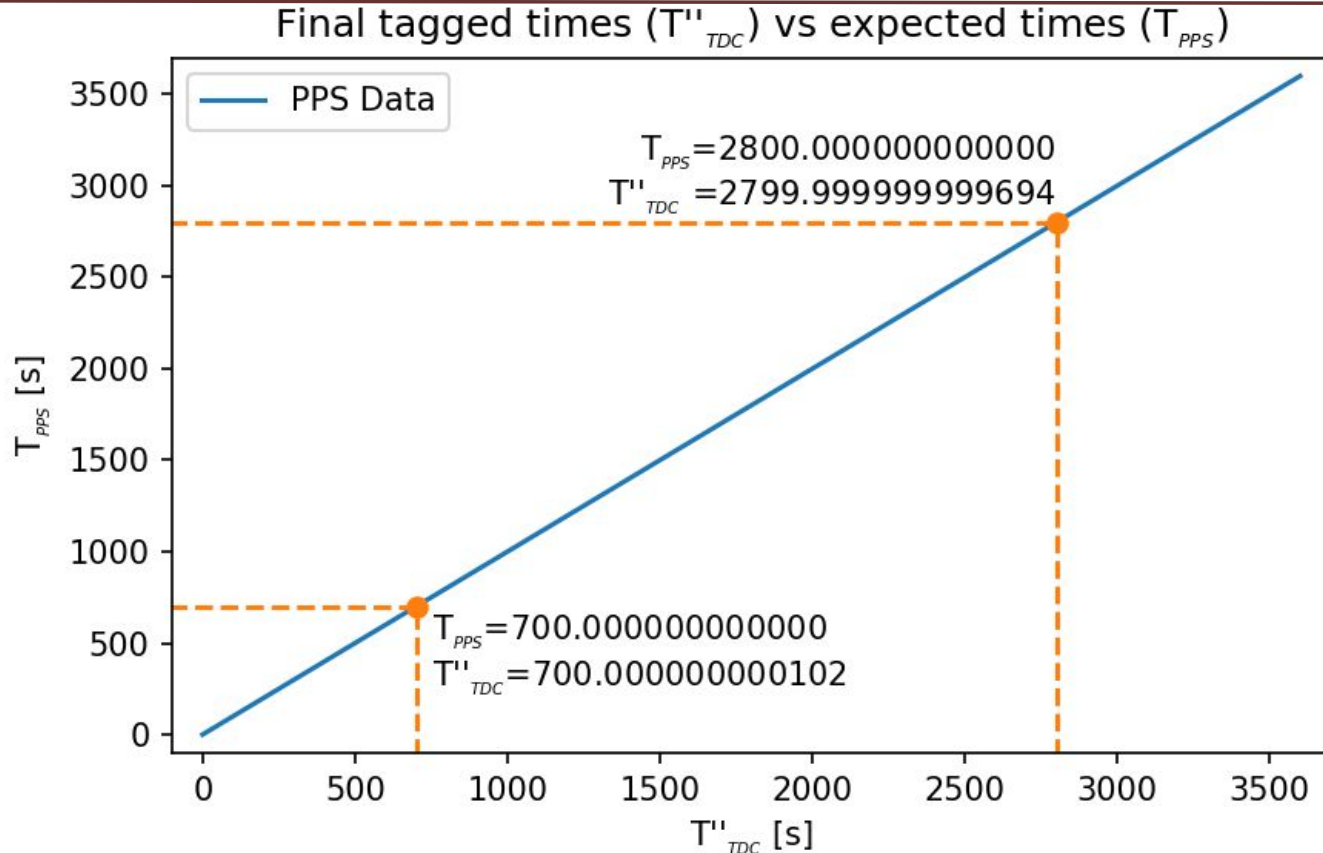
Residuals after removing clock drift (linear corr.)





## Example (from ASTRI-SI3 laboratory test)

Subtracting  $T_{res,fit}$  to  $T'_{TDC}$  removes much of the sinusoidal pattern.

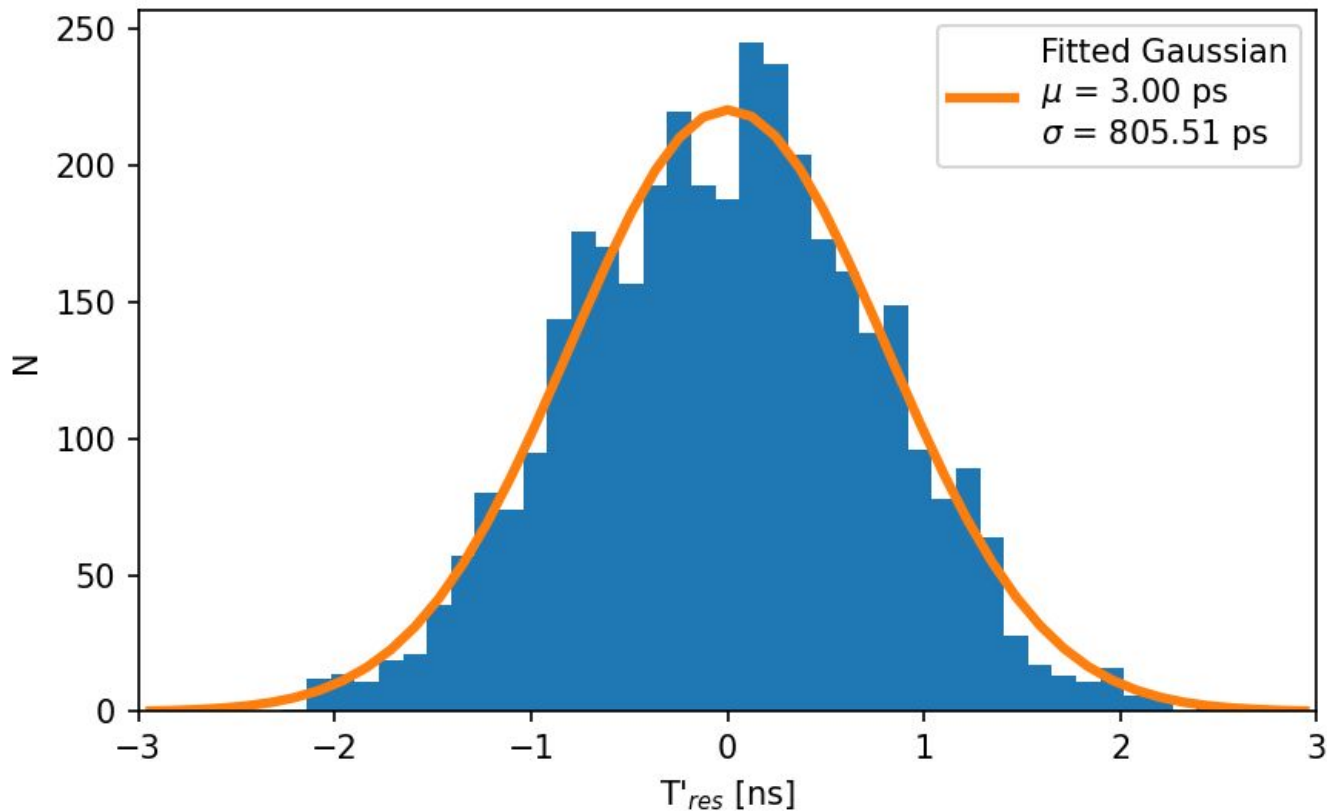


## Example (from ASTRI-SI3 laboratory test)

Distribution of the PPS signal residuals after all the corrections.



Distribution of residuals after corrections

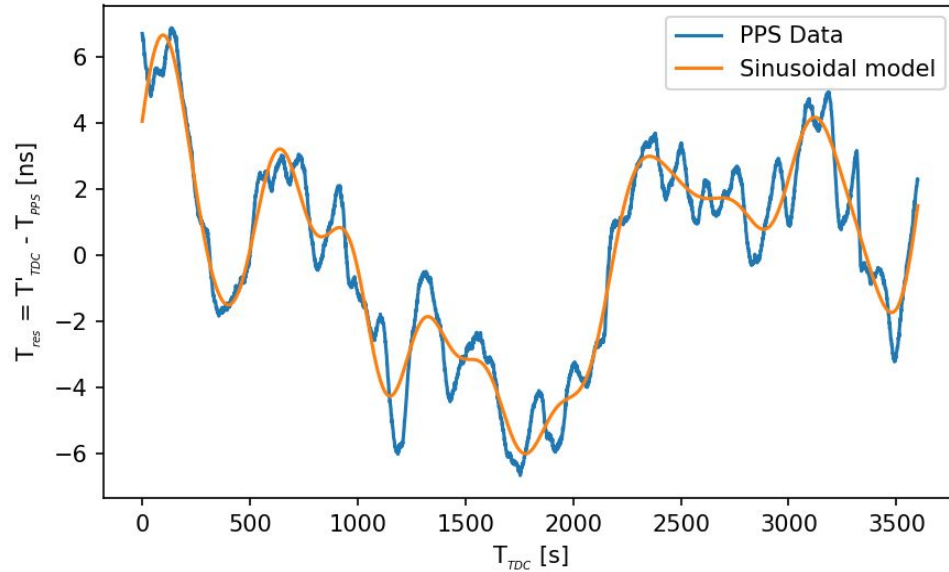


# Example (from ASTRI-SI3 laboratory test)

Comparison using the internal clock of the TDC

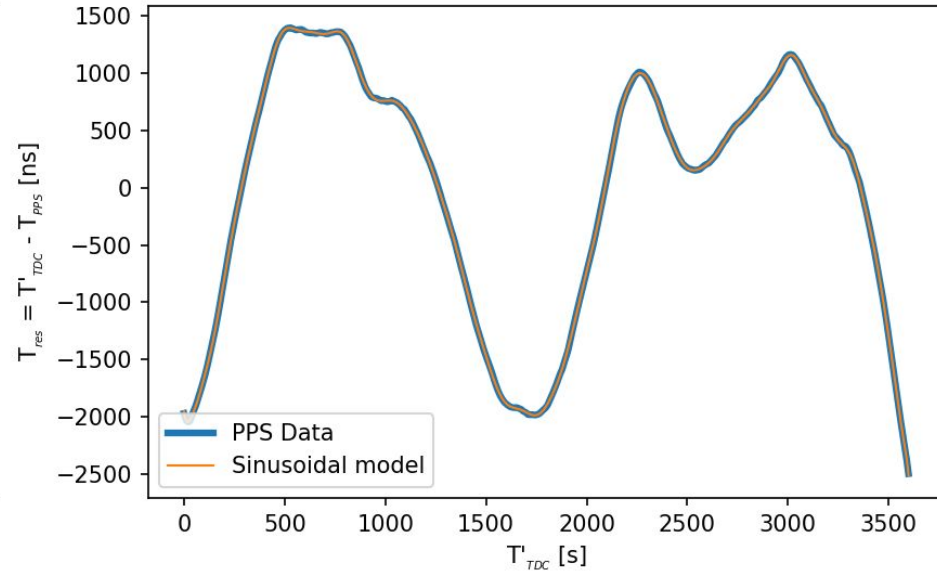


Residuals after removing clock drift (linear corr.)



With Rubidium clock

Residuals after removing clock drift (cubic corr.)



With internal TDC clock

# How to accurately measure the real timing performances?

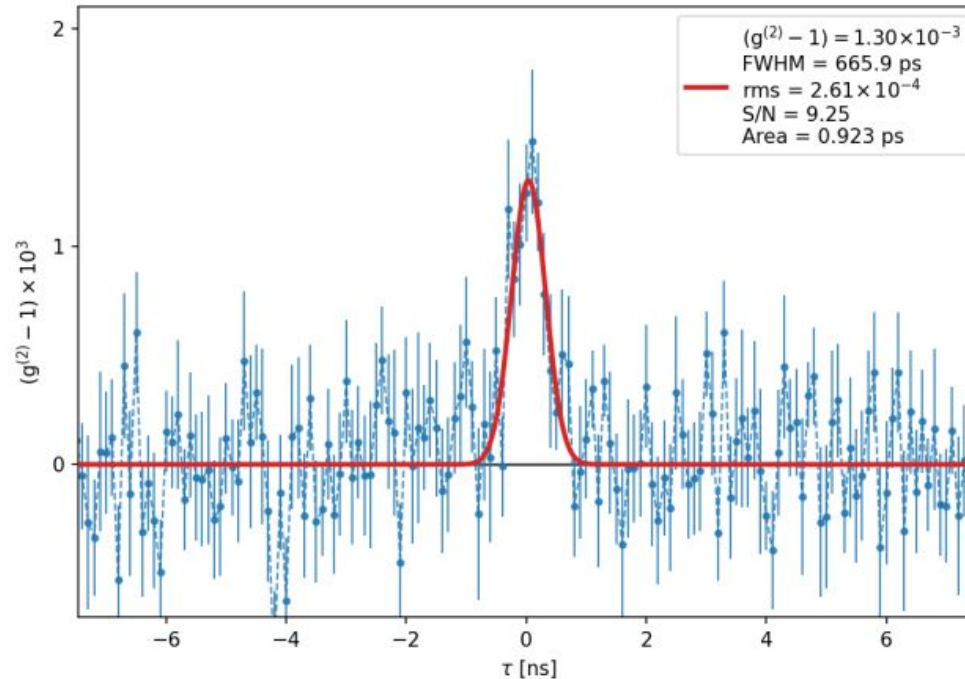


- Picosecond pulsed laser? → Very expensive!!
- Timing of (ms) pulsars? → Not enough to reach the ps level...

# How to accurately measure the real timing performances?



- Picosecond pulsed laser? → Very expensive!!
- Timing of (ms) pulsars? → Not enough to reach the ps level...
- **Intensity Interferometry measurements!!!**





**AQUEYE+IQUEYE**

Based on observations collected at the Copernicus telescope (Asiago, Italy) of the INAF-Osservatorio Astronomico di Padova and at the Galileo telescope (Asiago, Italy) of the University of Padova

Presently funded by INAF (Research Grants “Uncovering the optical beat of the fastest magnetised neutron stars (FANS)”, “Coordinated multiwavelength exploration of Fast Radio Bursts (COMEFAR)”) and the University of Padova (DOR grant)