

TORRCH

*Field-level reconstruction of
reionization morphology from galaxy tracers*

Soumak Maitra

Postdoctoral Fellow

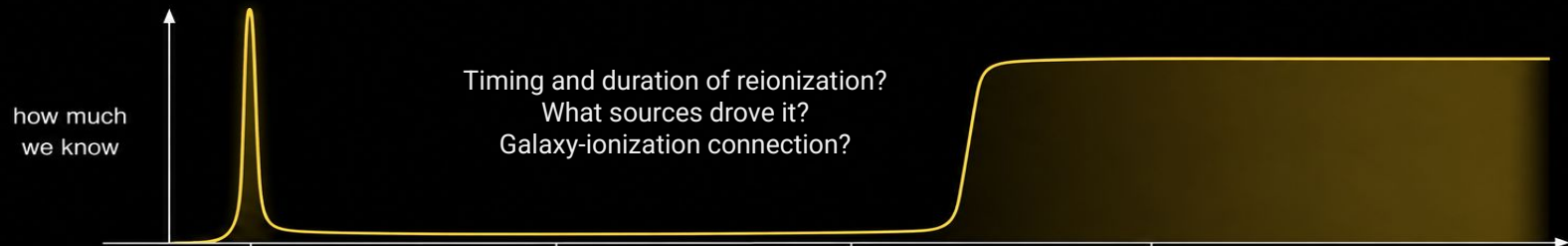
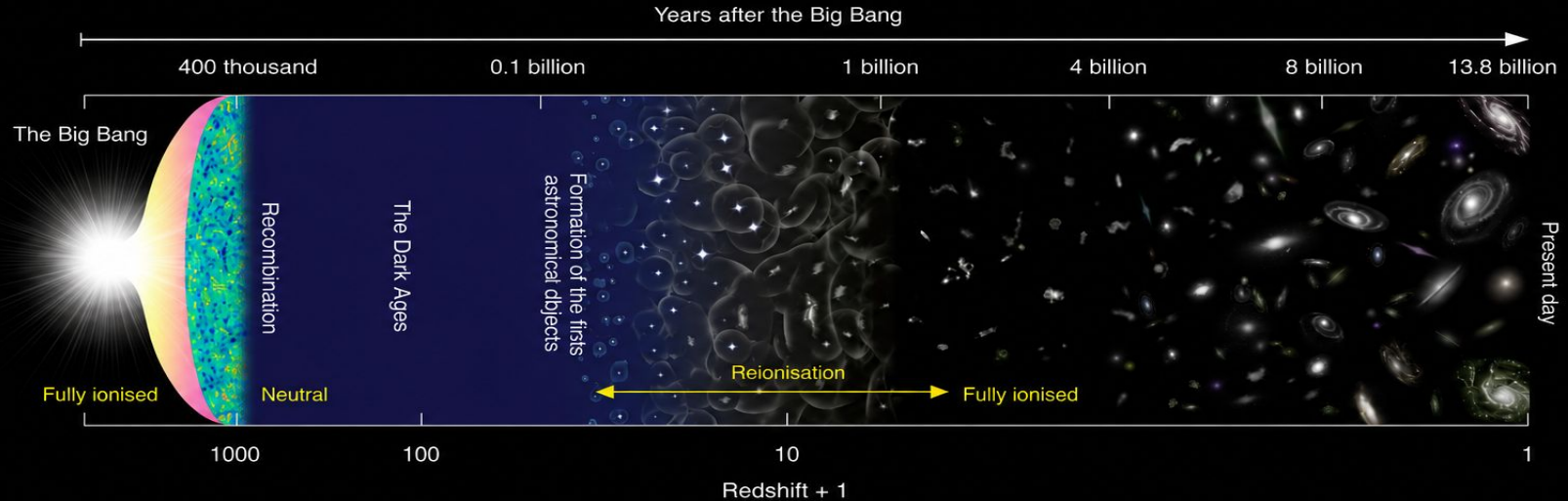
Tata Institute of Fundamental Research, Mumbai, India



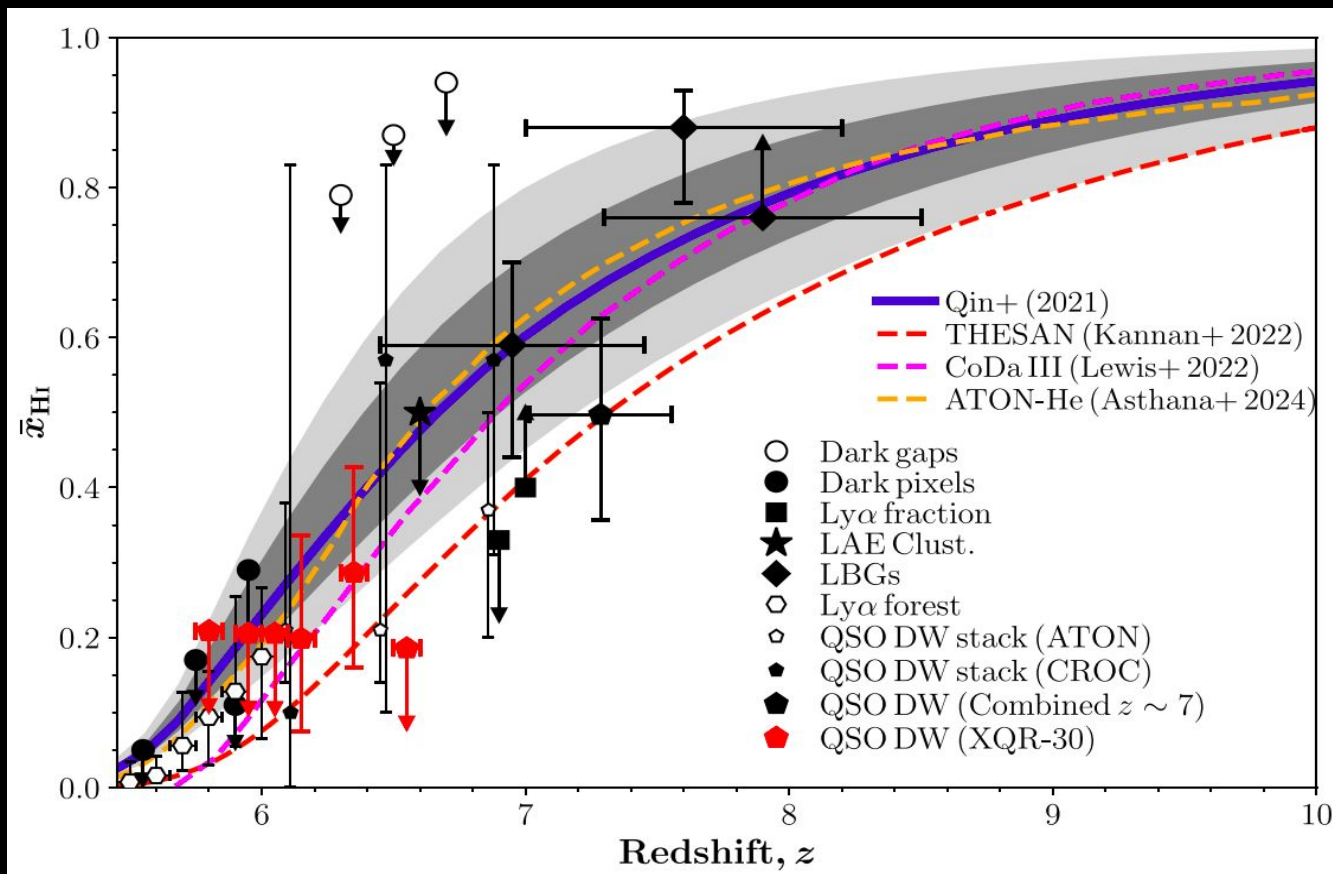
with Girish **Kulkarni** (TIFR), Vipul **Arora** (KU Leuven),
Matteo **Viel** (SISSA), Shikhar **Asthana** (KICC), James S. **Bolton**
(U.Nottingham), Martin G. **Haehnelt** (KICC), Laura **Keating** (U.Edinburgh)

ML4ASTRO3 , 31 Aug - 4 Sep 2026

Reionization: A pivotal but poorly understood phase transition

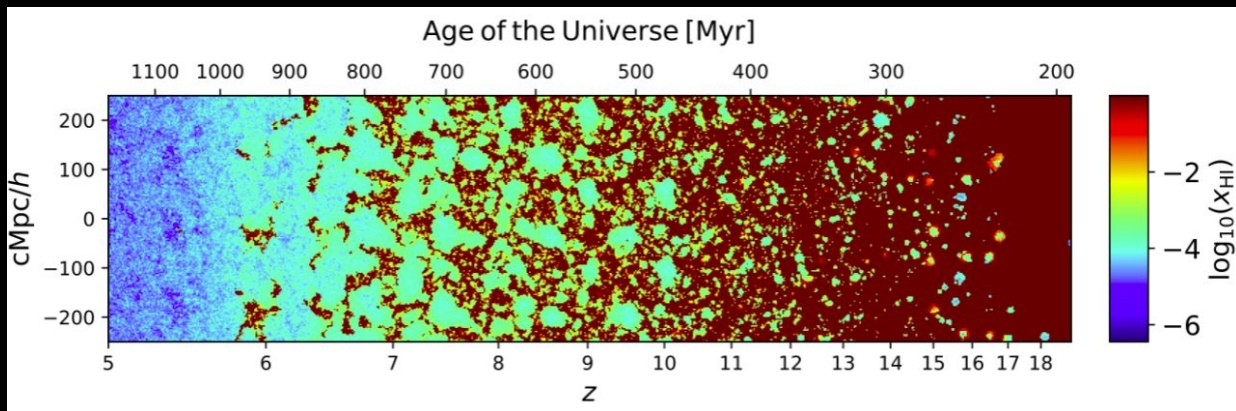


Our view of Reionization is still mostly global



Greig+2024

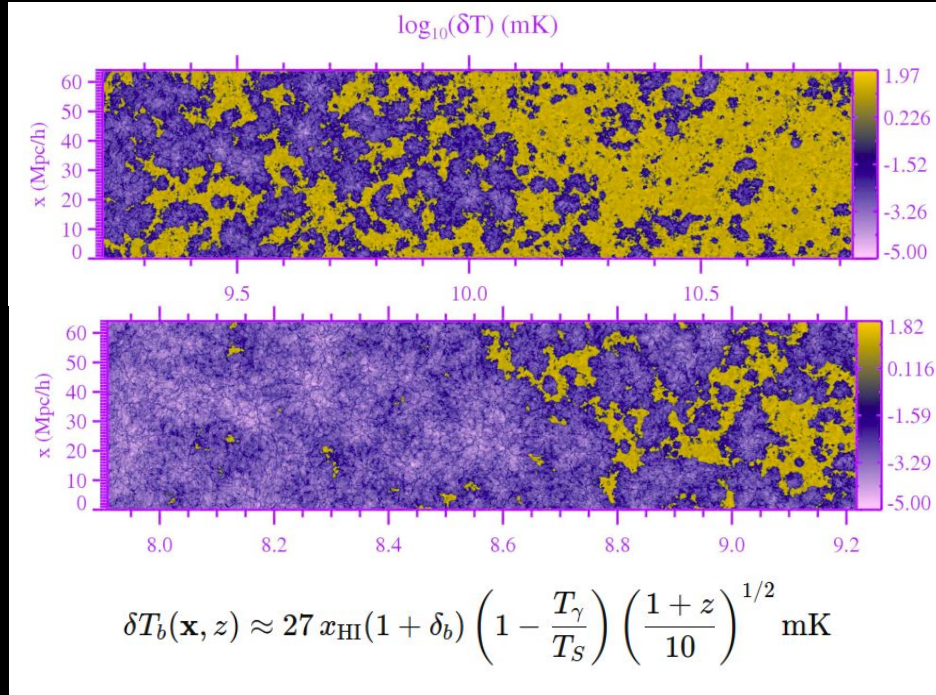
The Global Average Hides a Patchy Reionization Field



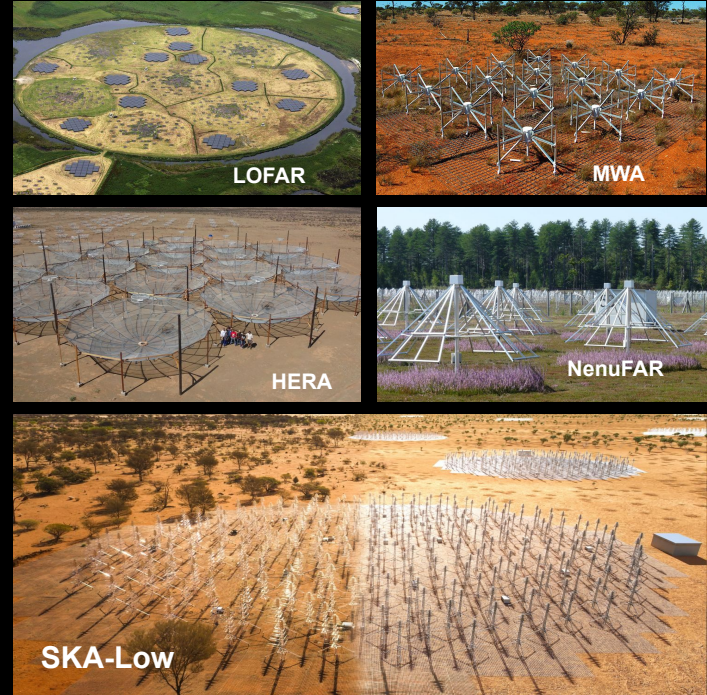
Asthana et al. 2024

Morphology of the patchy reionization process is crucial in understanding what drove Reionization

The Next Frontier is Reionization Tomography

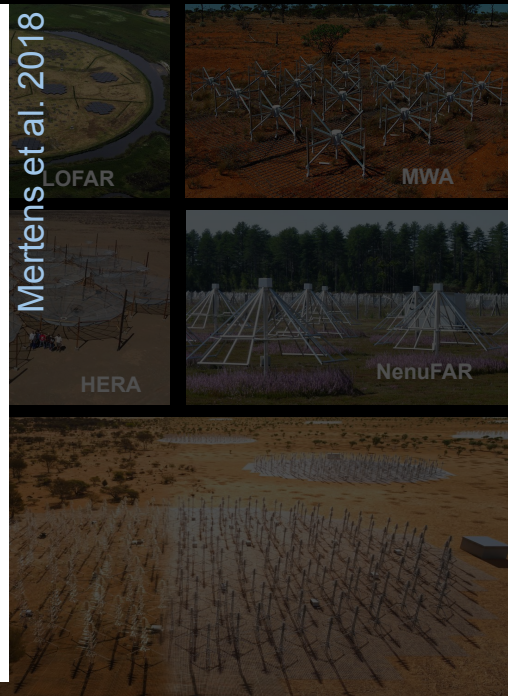
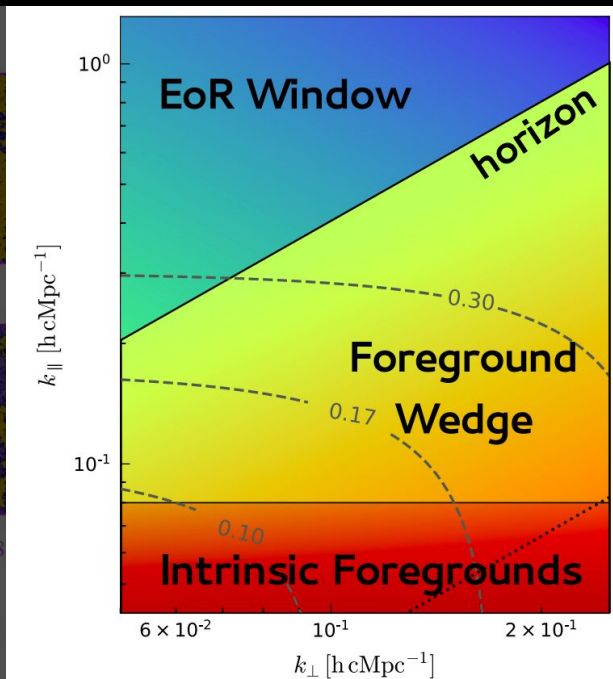
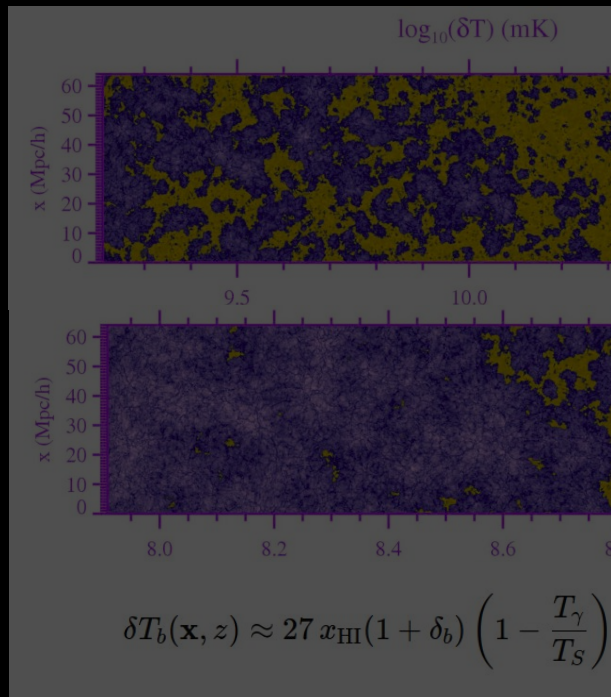


Furlanetto et al. 2009



Major 21-cm experiments are pushing reionization studies beyond global averages and toward morphological studies and spatial mapping.

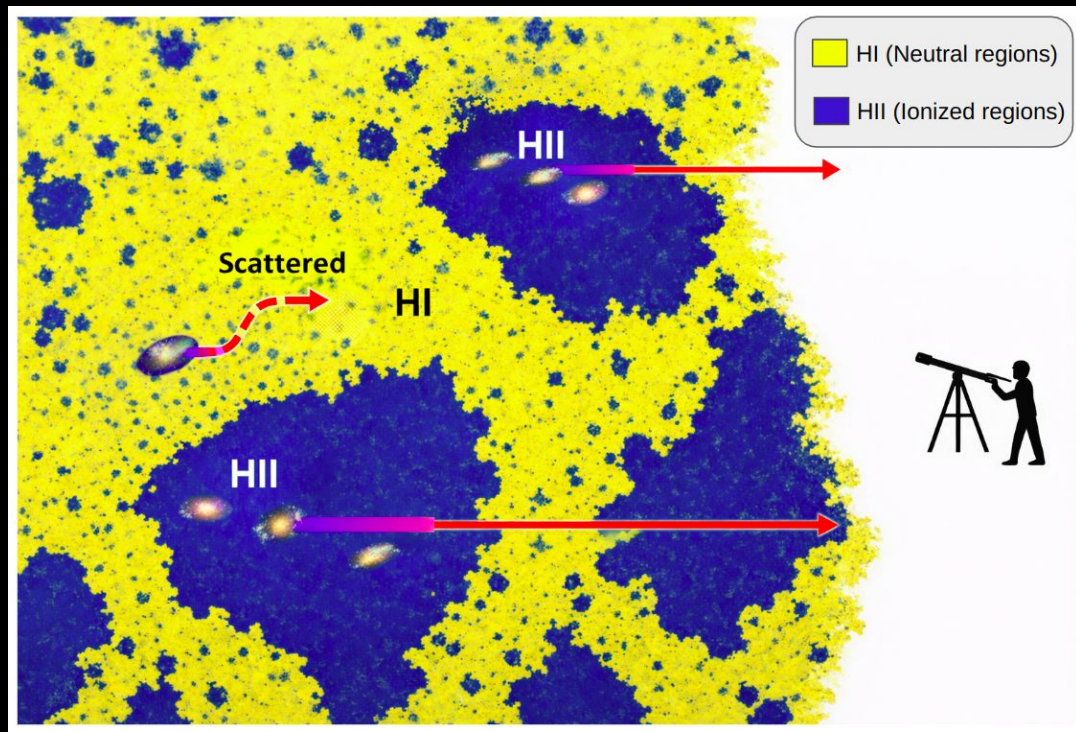
The Next Frontier is Reionization Tomography



Foreground Challenge

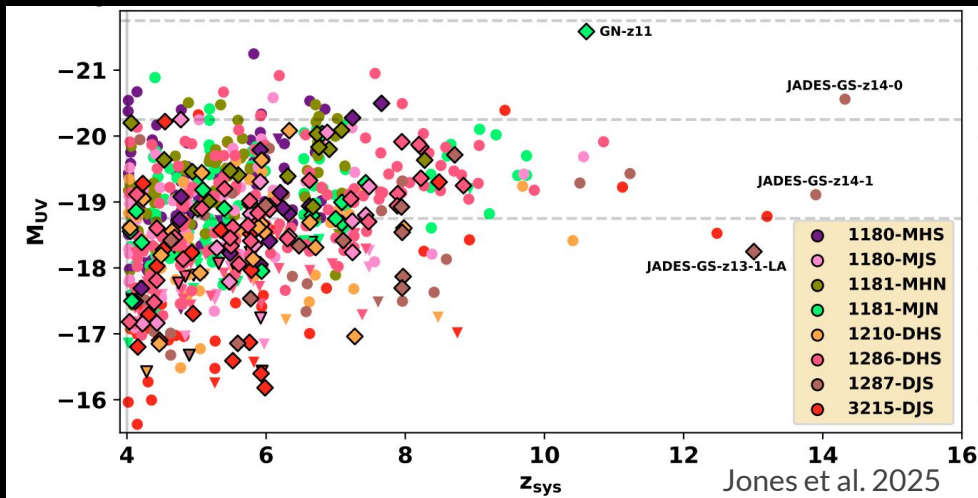
Major 21-cm experiments are pushing reionization studies beyond global averages and toward morphological studies and spatial mapping.

Galactic tracers of the Ionization Field



Lyman- α Emitters (LAEs)

Large sample of LAEs now available for Reionization studies!

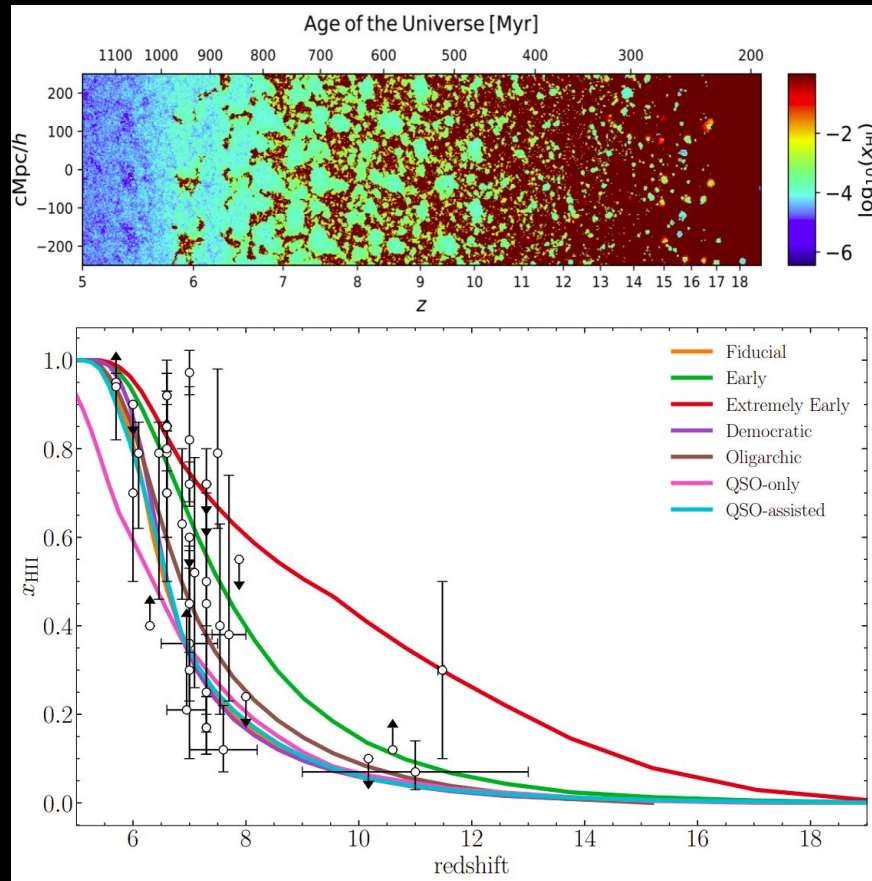


Survey	Instrument	Redshift Range	Area (deg ²)
SILVERRUSH	Subaru/HSC (NB)	2.2–7.3	~25
PFS-SSP	Subaru/PFS	5.7–6.6	~15
DAWN	NEWFIRM (NB)	7.7	~1
MUSE-Wide	VLT/MUSE (IFU)	2.9–6.7	~0.03
MUSE-HUDF	VLT/MUSE (IFU)	2.9–6.7	~0.003
LAGER	DECam (NB964)	6.9	~3
CIDER	NB964/NB1008	6–7.5	-
WERLS	Keck/MOSFIRE+LRIS	6–8	~0.7
MOONS	VLT/MOONS (Multi-object)	6–8.5	~1–2
CEERS	JWST/NIRSpec	6–9	~0.03
JADES	JWST/NIRSpec+NIRCam	5.8–8.0	~0.05
COSMOS-Web	JWST/NIRCam	6–11	0.54
Roman Grism	Roman Space Telescope	6–10+	~16

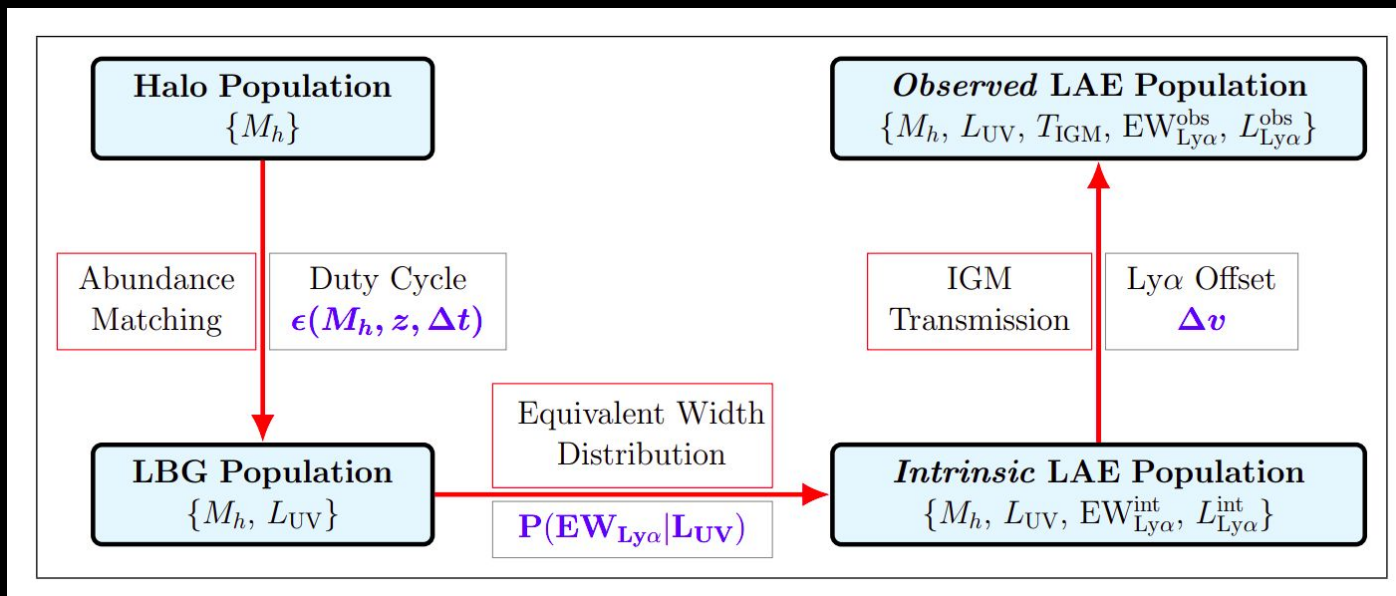
Field-level reconstruction of reionization from LAEs?

Step 1: Reionization Simulations

- P-GADGET3 for the hydrodynamic simulation (SHERWOOD Relic): $(160h^{-1}\text{cMpc})^3$, 2×2048^3 particles
- ATON (Aubert & Teyssier 2008) and ATON-HE (Asthana et al. 2024b): Radiative transfer code for the thermal and ionization fields.

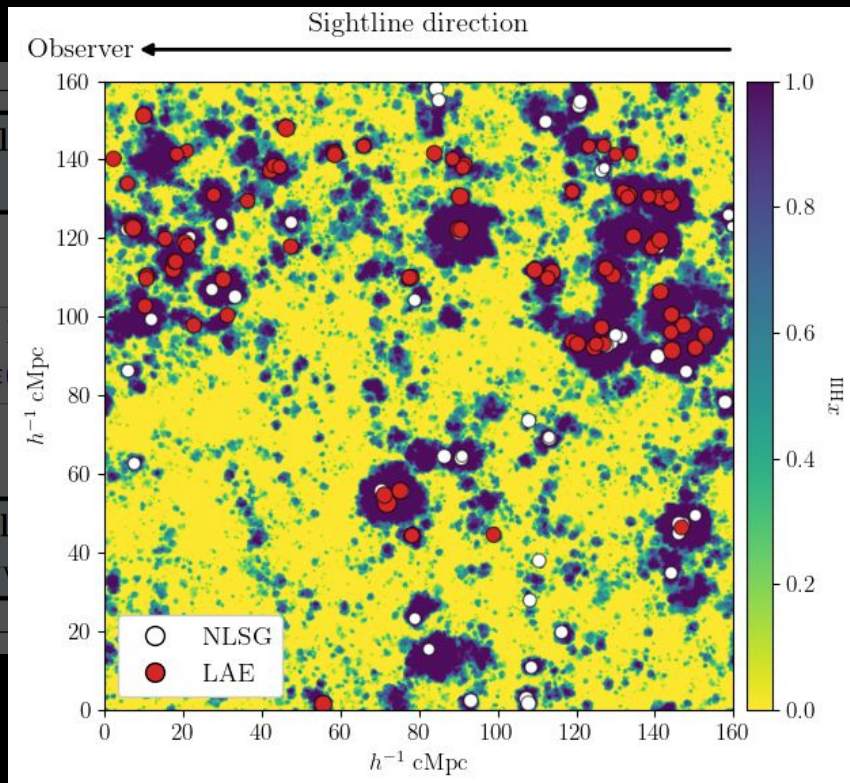


Step 2: Forward model for the LAEs



Weinberger+2019, Maitra+2025b

Step 2: Forward model for the LAEs



Halo Popul

$\{M_h\}$

Abundance
Matching

LBG Popul

$\{M_h, L_{\text{UV}}\}$

Population

$\{L_{\text{Ly}\alpha}^{\text{obs}}, L_{\text{Ly}\alpha}^{\text{obs}}\}$

α Offset

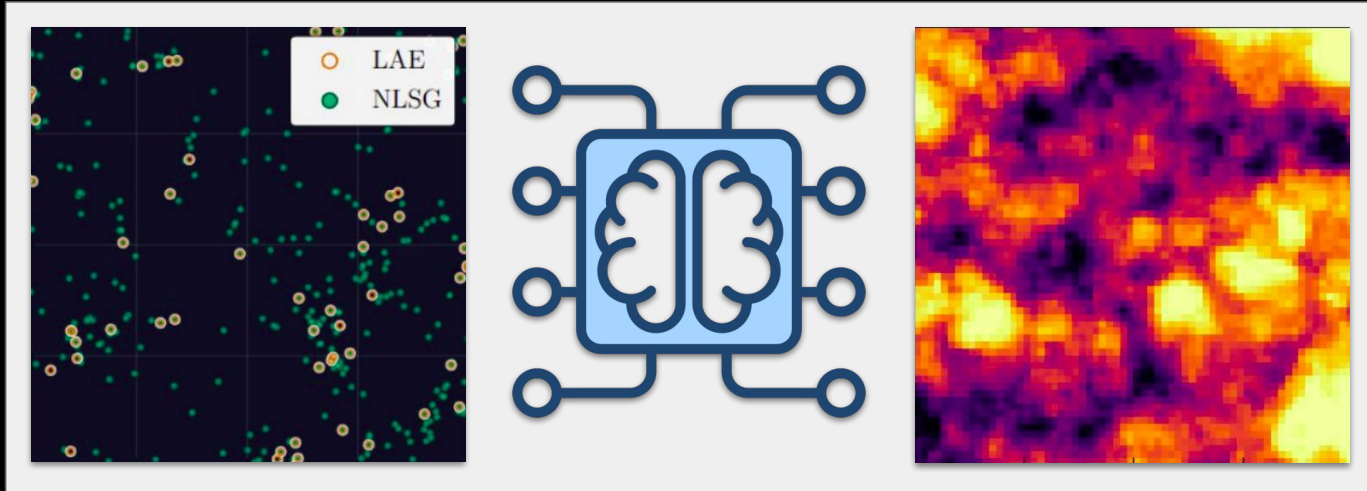
Δv

Population

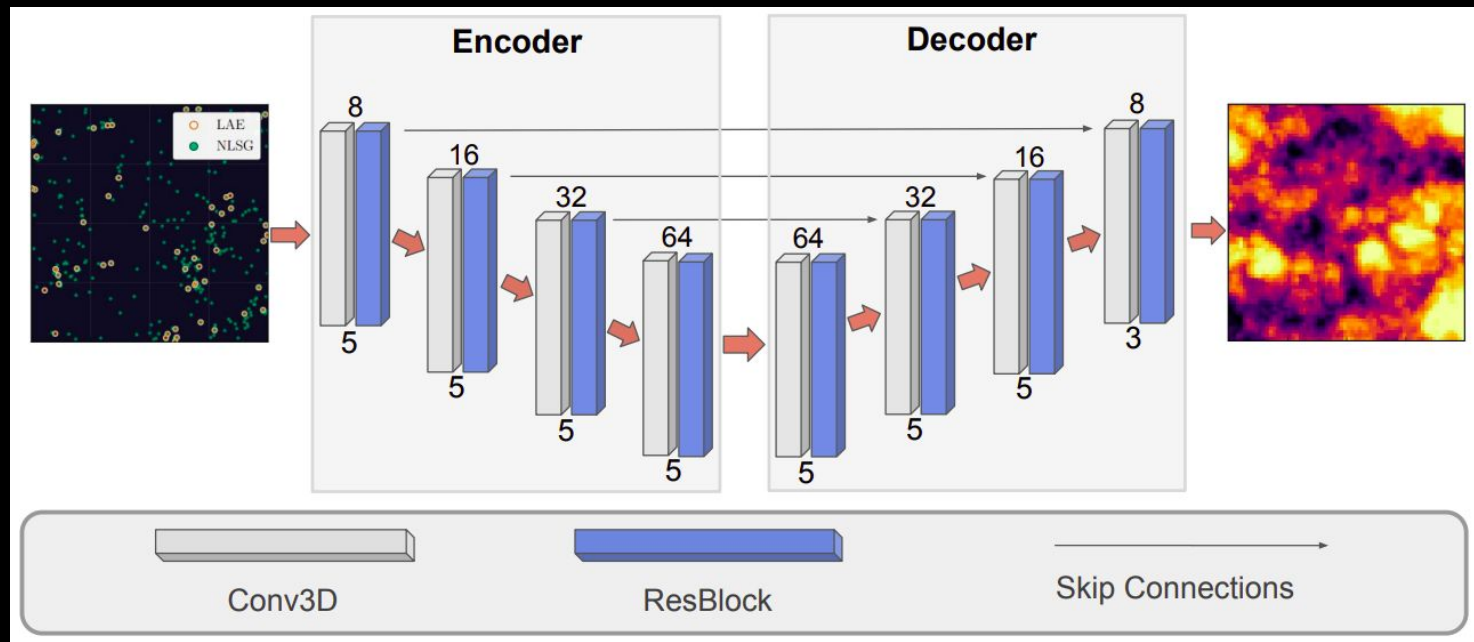
$\{L_{\text{Ly}\alpha}^{\text{int}}, L_{\text{Ly}\alpha}^{\text{int}}\}$

ger et al. 2019

Step 3: An intelligent tool for the tomographic inversion

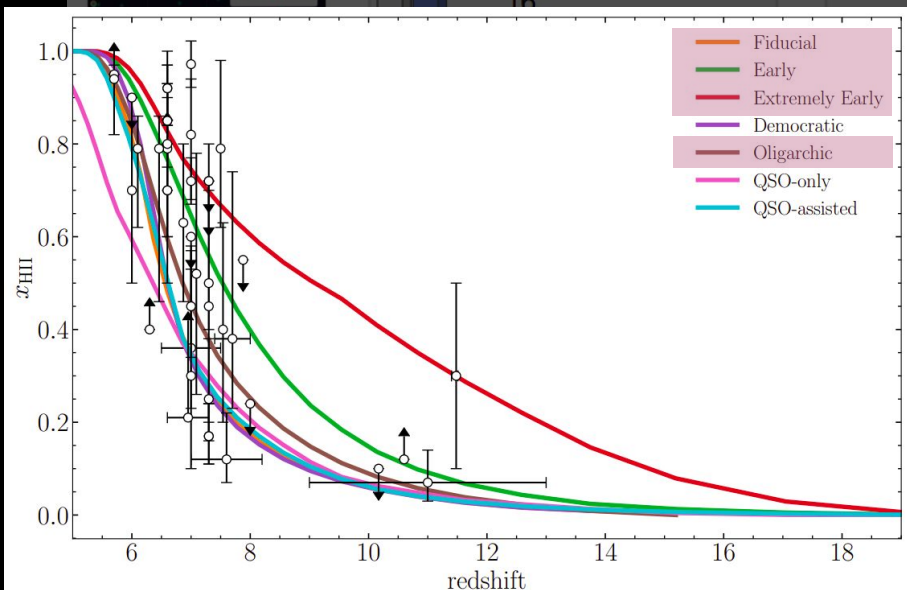


TORRCH: TOmographic Reconstruction of the Reionization of Cosmic Hydrogen with Ly α emitters and non-Ly α -selected galaxies (z=7.14) [Maitra et al. 2026]



TORRCH: TOmographic Reconstruction of the Reionization of Cosmic Hydrogen with Ly α emitters and non-Ly α -selected galaxies (z=7.14) [Maitra et al. 2026]

Training sample:



• $\hat{\mu}(\mathbf{x}) \equiv \hat{x}_{\text{HI}}(\mathbf{x})$ (ionization field) • $\log \hat{\sigma}^2(\mathbf{x})$ (learned variance)

Table 1. Configurations of the three galaxy surveys considered in this work.

Survey type	$\log(L_{\text{Ly}\alpha}/\text{erg s}^{-1})$ cut	M_{UV} cut	Tracers included
Deep	> 41.5	< -18	LAEs + NLSGs
Shallow	> 42.5	< -19	LAEs + NLSGs
LAE-only	> 42.5	–	LAEs only

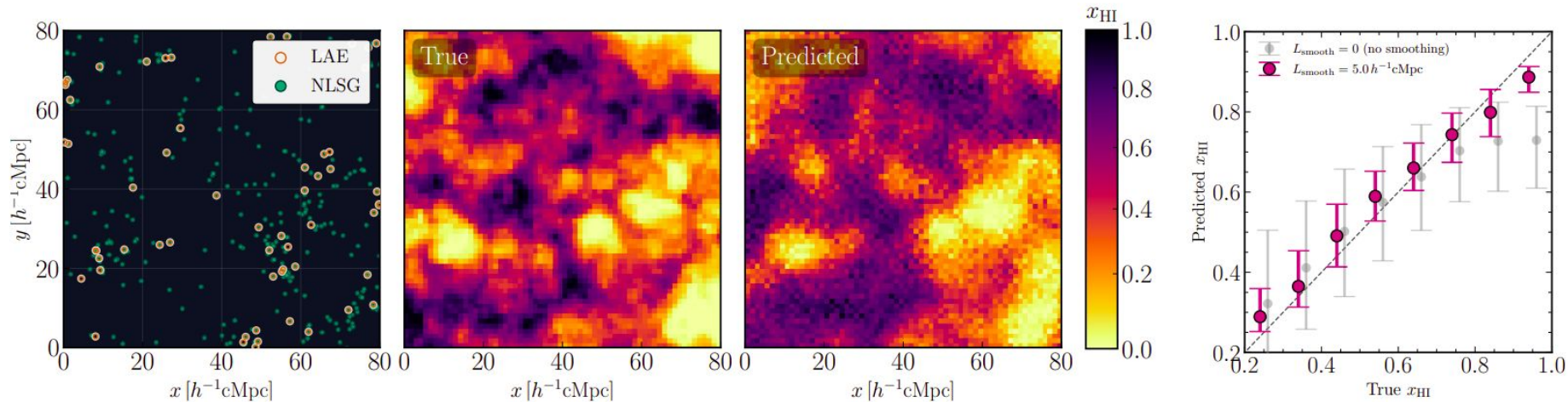
Loss | Heteroscedastic β -NLL

$$\mathcal{L}_{\beta\text{-NLL}} = \text{sg}[\sigma^{2\beta}] \left[\frac{(x_{\text{HI}} - \mu)^2}{2\sigma^2} + \frac{1}{2} \log \sigma^2 \right]$$

TORRCH: TOmographic Reconstruction of the Reionization of Cosmic Hydrogen

with Ly α emitters and non-Ly α -selected galaxies (z=7.14) [Maitra et al. 2026]

Early Model: Shallow Survey

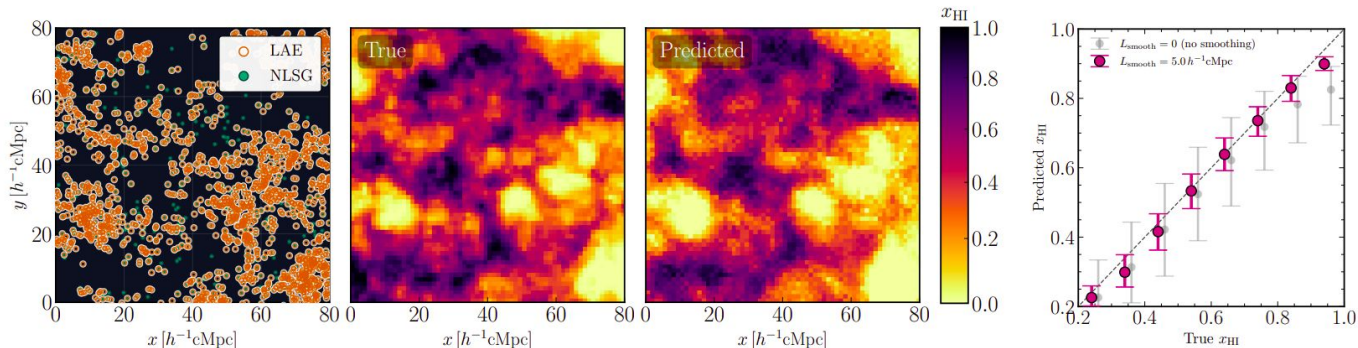


2D Projection with $30h^{-1}\text{cMpc}$ thickness.

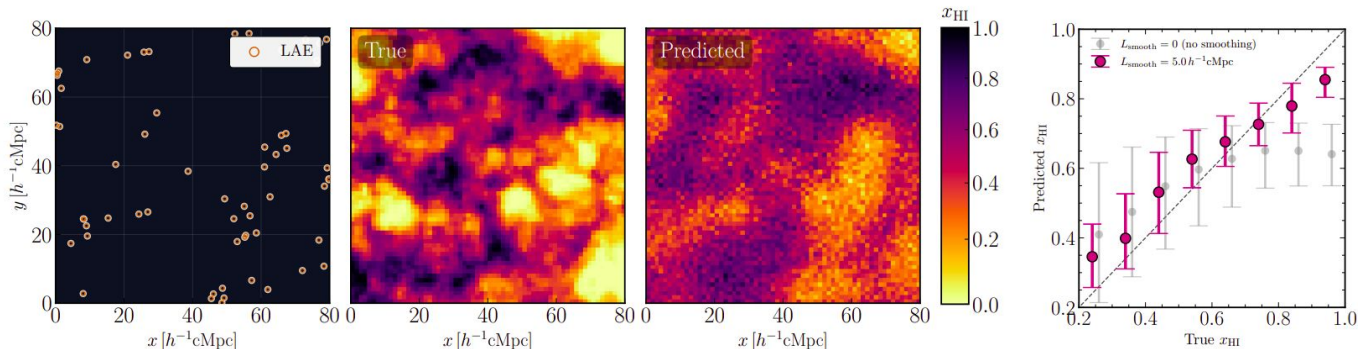
TORRCH: TOmographic Reconstruction of the Reionization of Cosmic Hydrogen

with Ly α emitters and non-Ly α -selected galaxies (z=7.14) [Maitra et al. 2026]

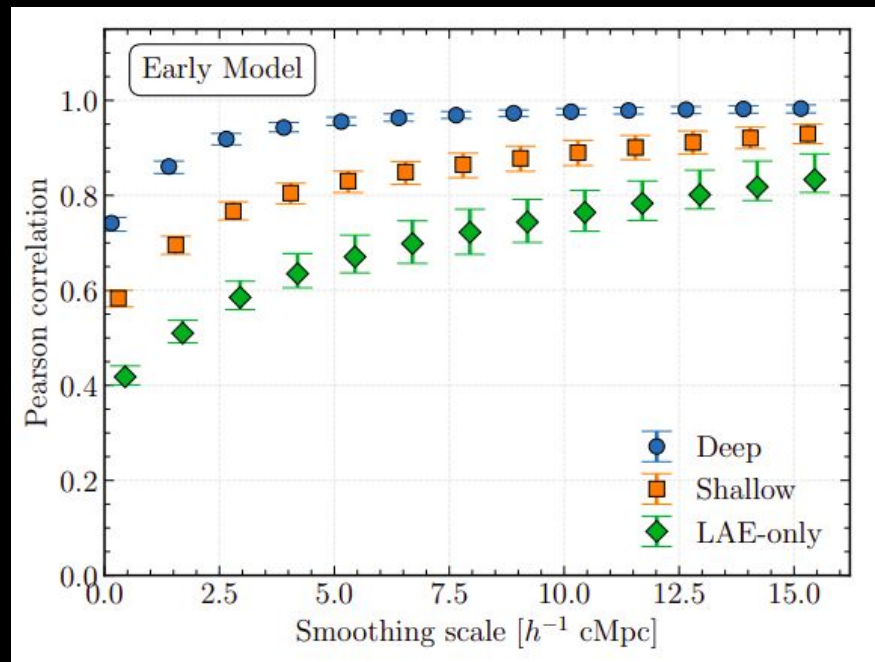
Early Model: Deep Survey



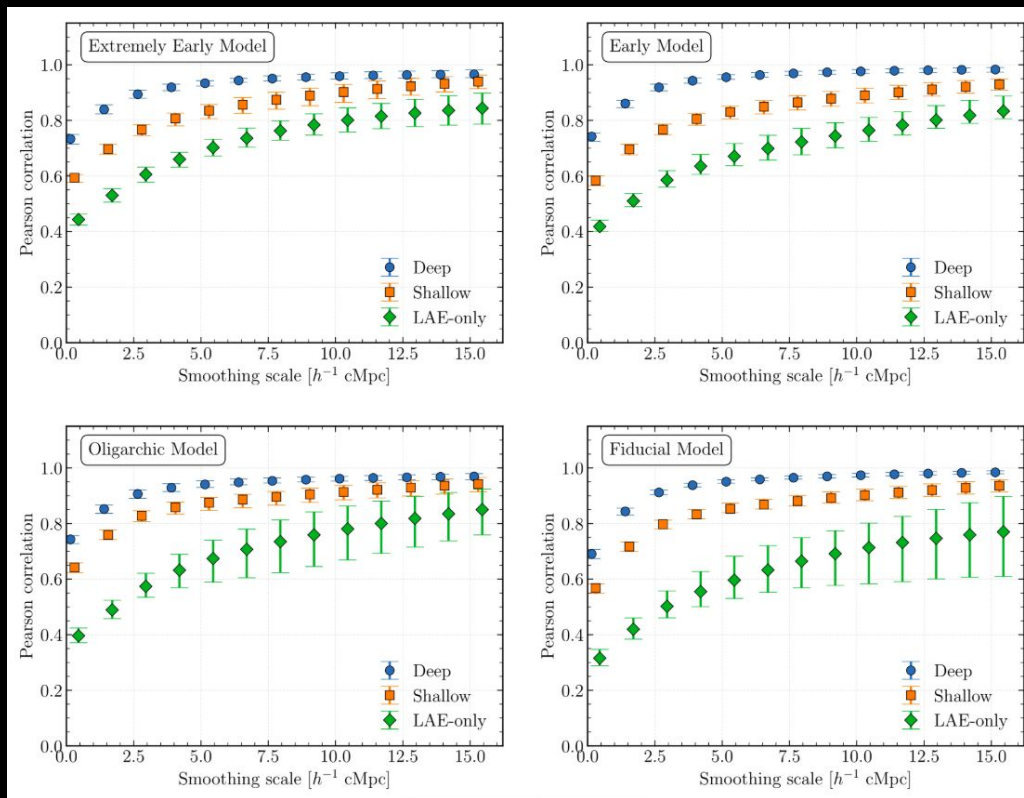
Early Model: LAE-only Survey



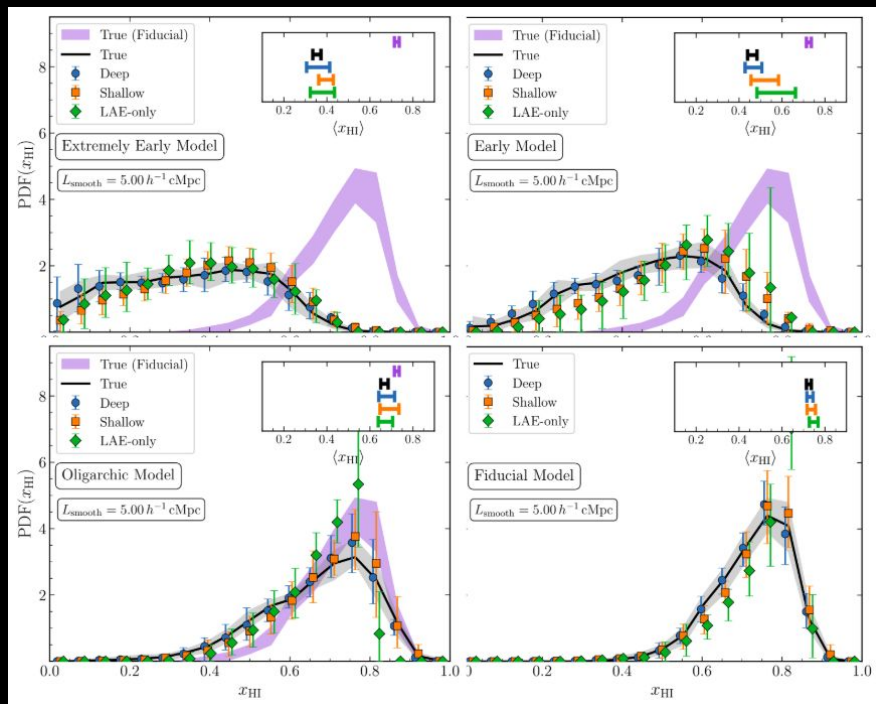
Reconstruction fidelity improves with scale



Reconstruction fidelity improves with scale

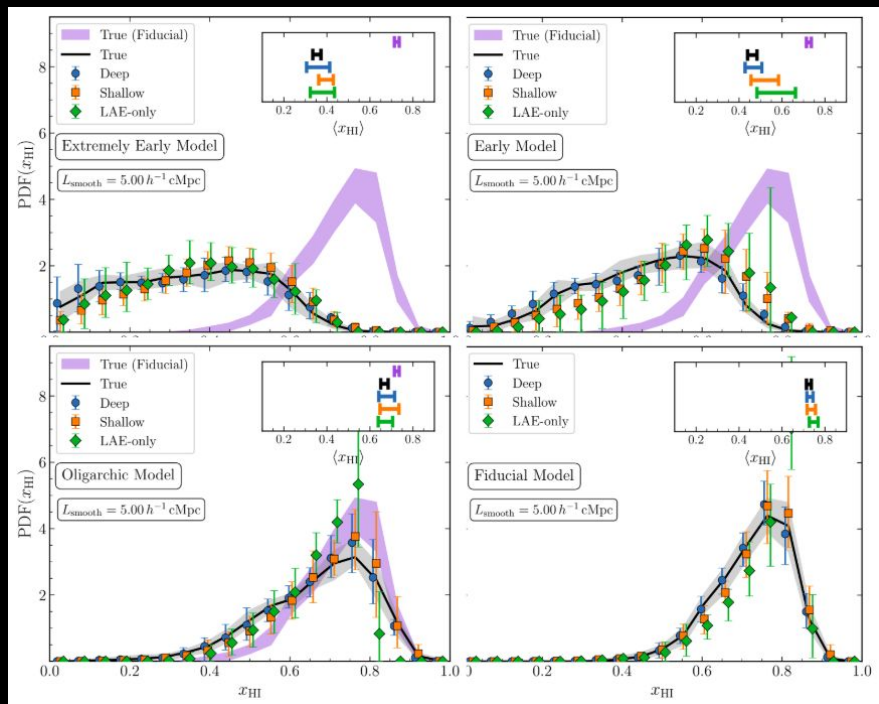


The summary statistics are well reproduced

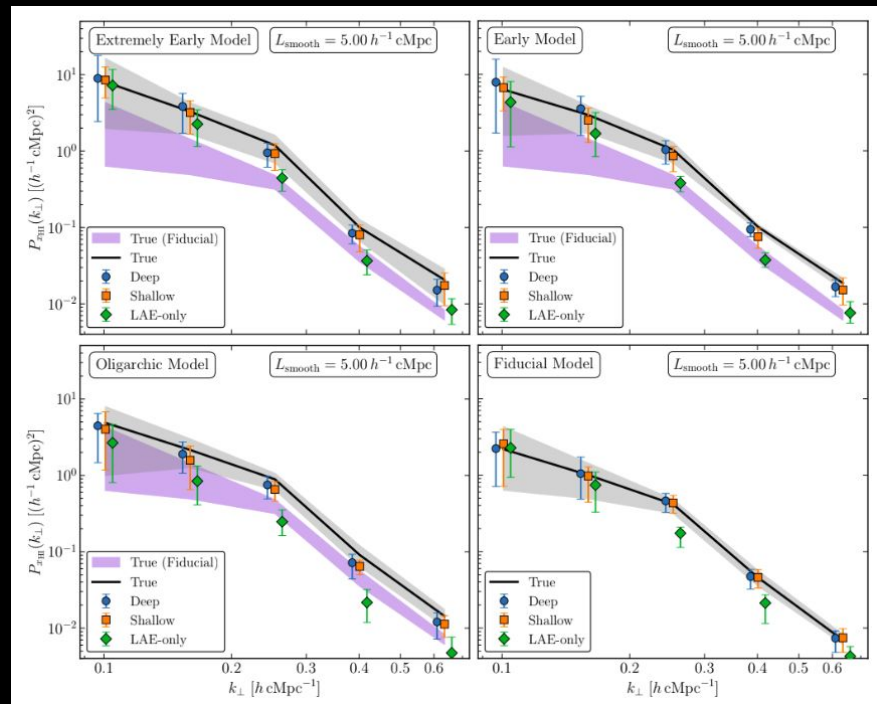


Probability distribution function

The summary statistics are well reproduced

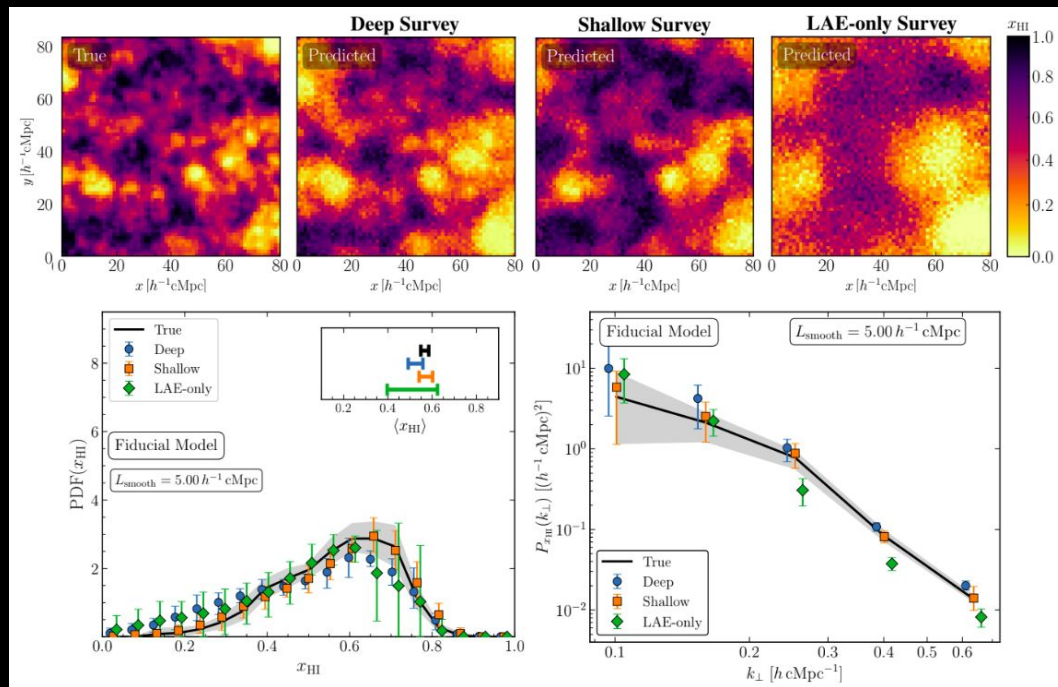


Probability distribution function



Power spectrum

Robustness check



Network trained at $z=7.14$ tested on $z=6.6$.

What we are currently working on

Modeling



Martin Haehnelt
KICC



James Bolton
U. Nottingham



Girish Kulkarni
TIFR



Laura Keating
U. Edinburgh



Matteo Viel
SISSA



Raghunathan
Srianand, IUCAA



Supriyo Saha
TIFR



Jahaan Thakkar
TIFR

- Intrinsic Ly α Luminosity/profile modeling.
- Larger simulation boxes.

Observation



Romain Meyer
Geneva Observatory



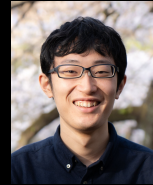
Koki Kakiichi
U. Copenhagen



Masami Ouchi
U.Tokyo



Hiroya Umeda
U.Tokyo



Yuta Kageura
U.Tokyo

- Data acquisition: JWST, Subaru
- Modeling the data systematics.

Machine Learning



Jatin Batra
TIFR



Dheeraj Nagaraj
Google DeepMind



Vipul Arora
KU Leuven



Parthe Pandit
IIT Bombay

- Reliable uncertainty estimation for predictions.
- Posterior with Generative algorithms.