

DISCLAIMER



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This talk contains jokes !

DISCLAIMER

gettyimages

Credit: bgblue

30 years of high-mass star formation
by Cesa & mosca
(mm interferometers & maser VLBI)



Consortium for Very Long Baseline
Interferometry in Europe



closest collaborators:

Maite Beltrán

Alberto Sanna

Ciriaco Goddi

Fabrizio Massi

HM SF Open Problems and Talk Outline

Low-mass ($\sim 1 M_{\odot}$) SF: disk-jets well characterized by observations.

In high-mass ($> 8 M_{\odot}$), prior of ALMA/JVLA observations:

A few (claims of) disks towards OB-type YSOs

A few thermal jets towards high-mass YSOs (VLA, rms ~ 0.3 mJy)

IRAS 20126+4104: an (almost) unique disk-jet system in HMYSO.

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IRAS 20126+4104: an (almost) unique disk-jet system in HMYSO.

Peculiar for high-mass SF:

Impact of radiation pressure and photoionization (thermal pressure from ionized gas) on the accretion/ejection.

The onset of the ionization: hyper-compact (HC) HII region.

The HC HII region inside core A1 in the SFR G24.78+0.08.

Cesa in Firenze

In the shadows of the towers.....

mosca in Siena



Firenze and Siena: bitter enemies in the Middle Age



Montaperti 1260

*Lo strazio e 'l grande scempio
che fece l'Arbia colorata in rosso,
tal orazion fa far nel nostro tempio».* *Inf.* X, 85-88

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San Romano 1432

Trittico di Paolo Uccello:
“Disarcionamento di Bernardino della Carda”
(Uffizi, Firenze)

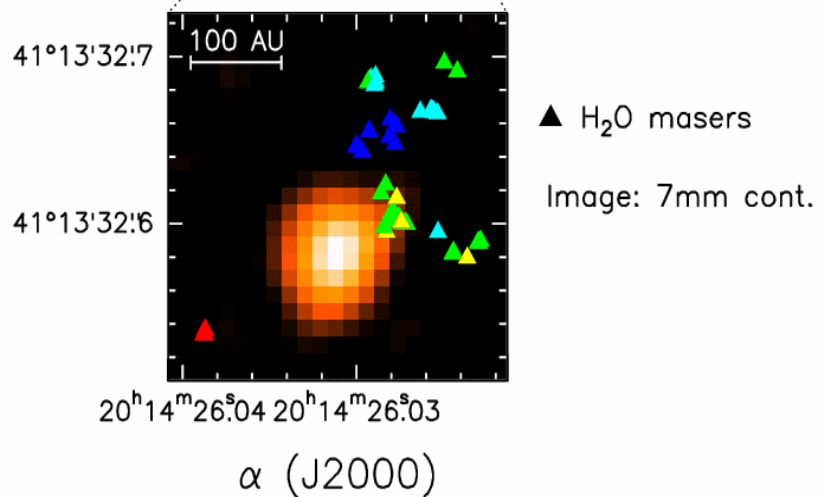
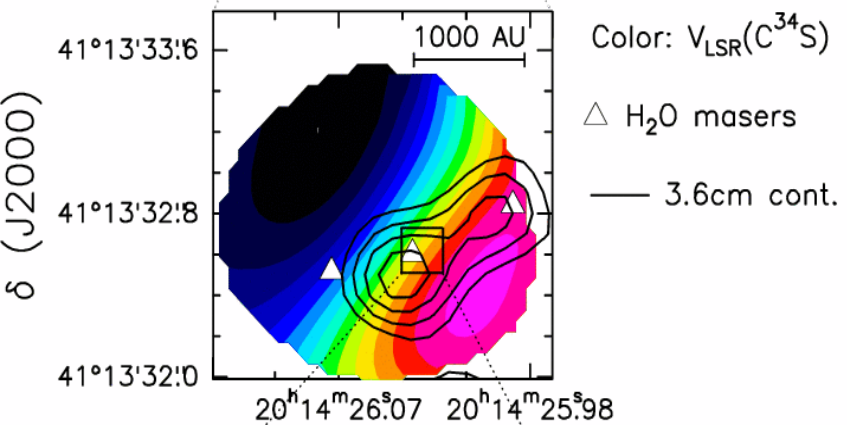
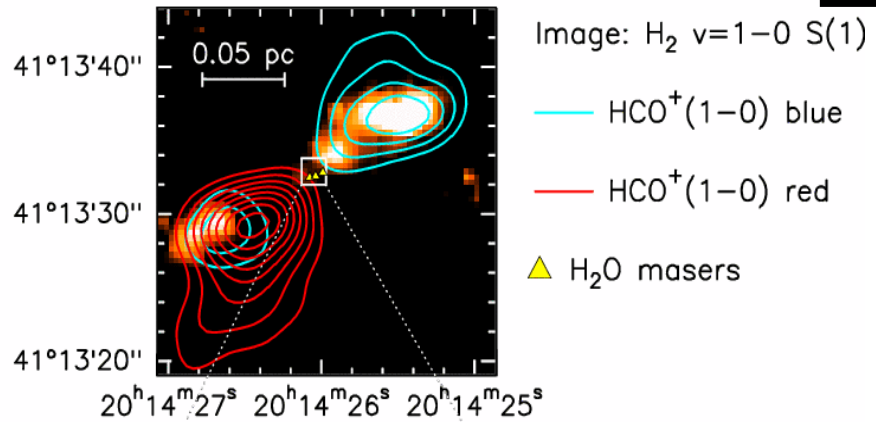


IRAS 20126+4104

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$M_* \sim 12 M_\odot$ $d = 1.64 \pm 0.05$ kpc

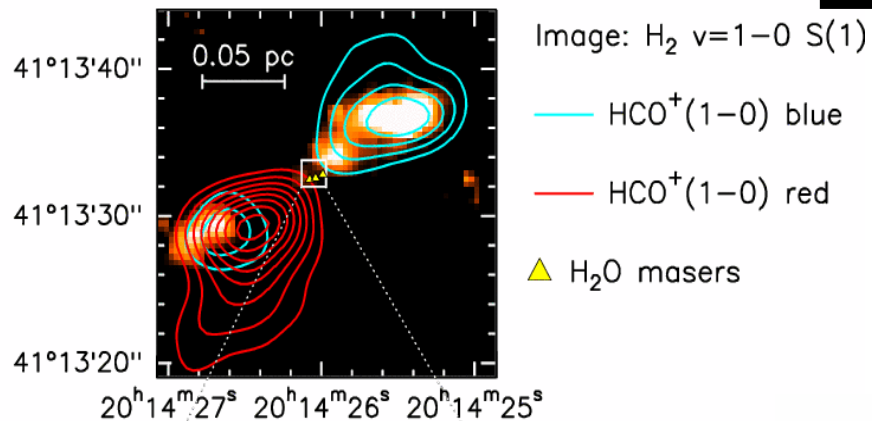
Cesaroni et al.
Hofner et al.
Moscadelli et al.



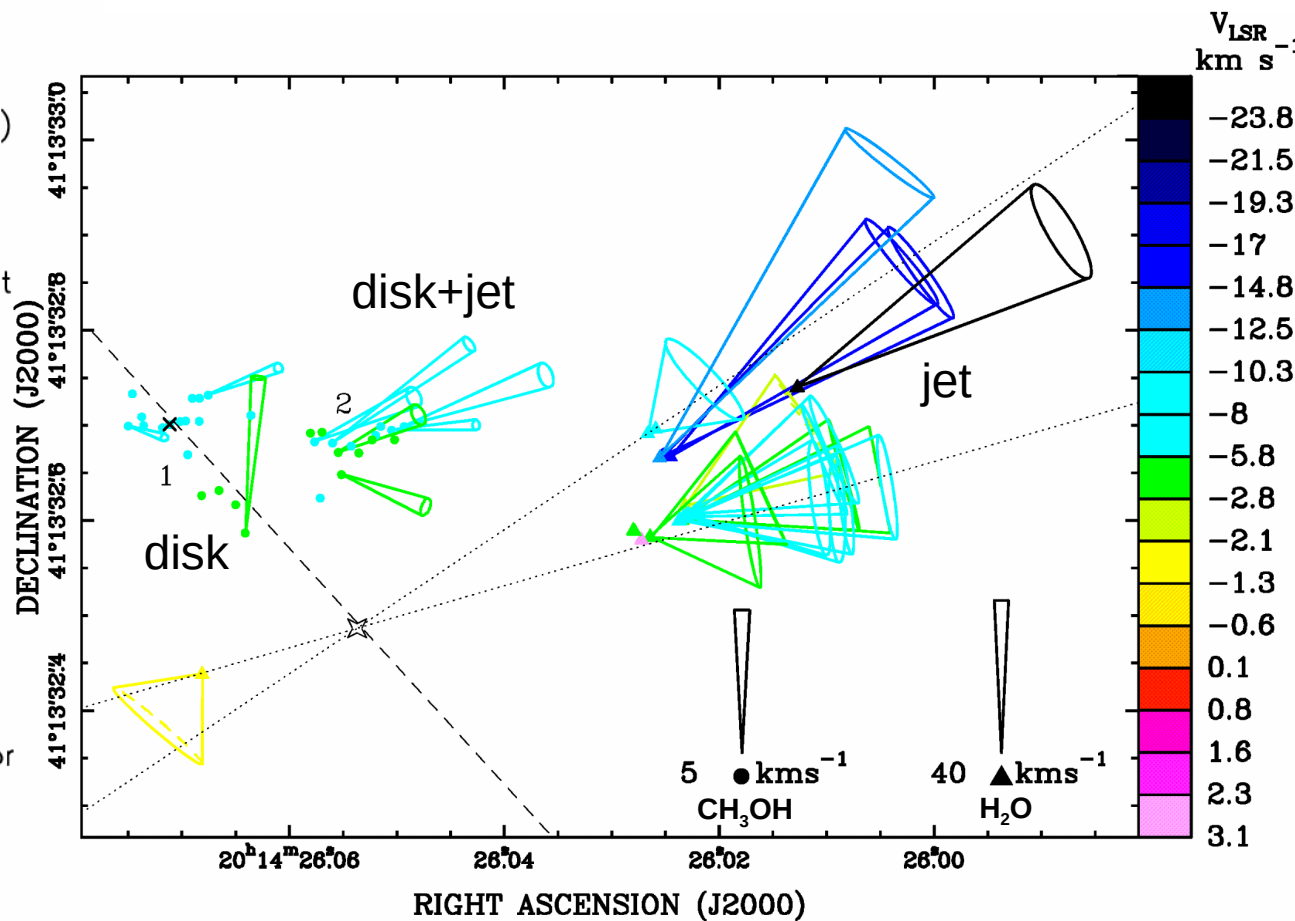
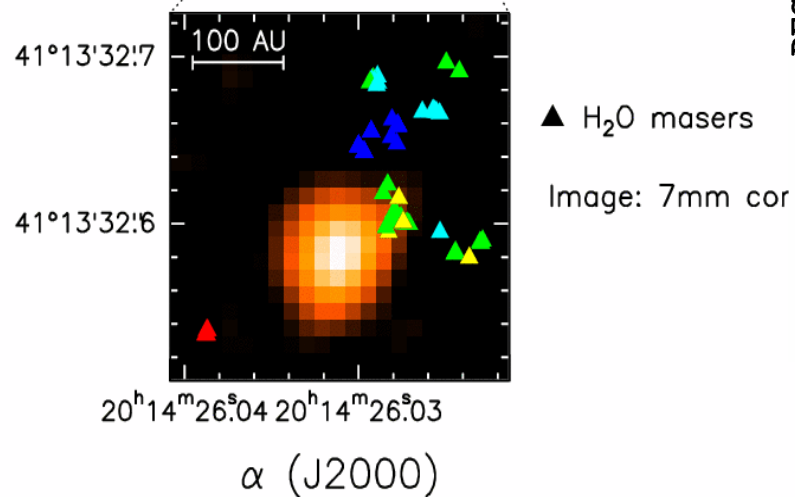
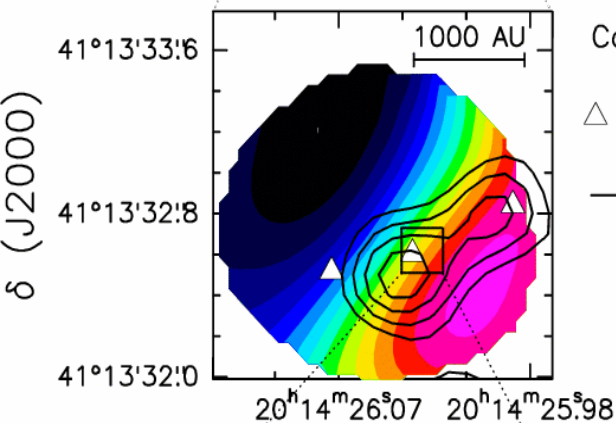
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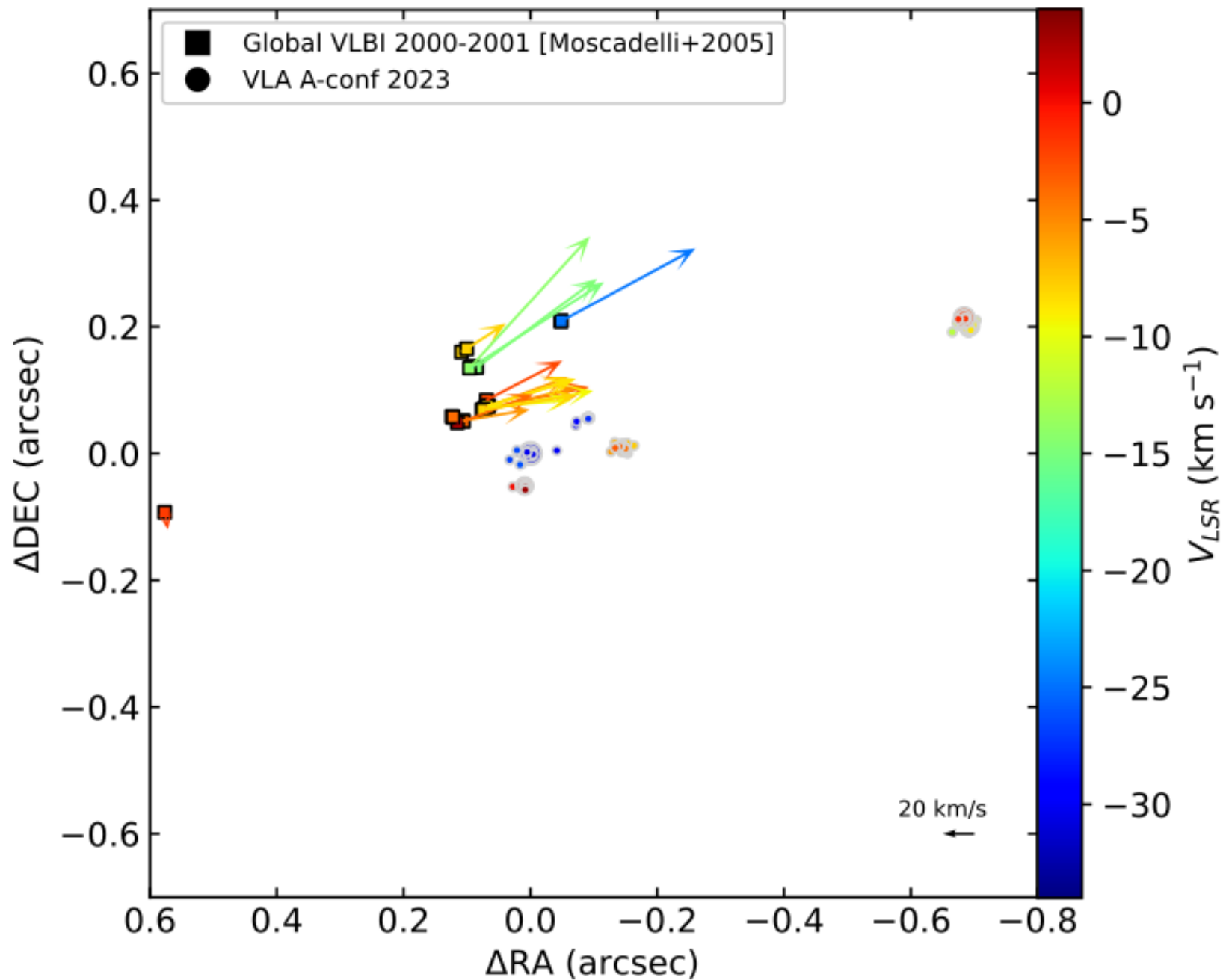
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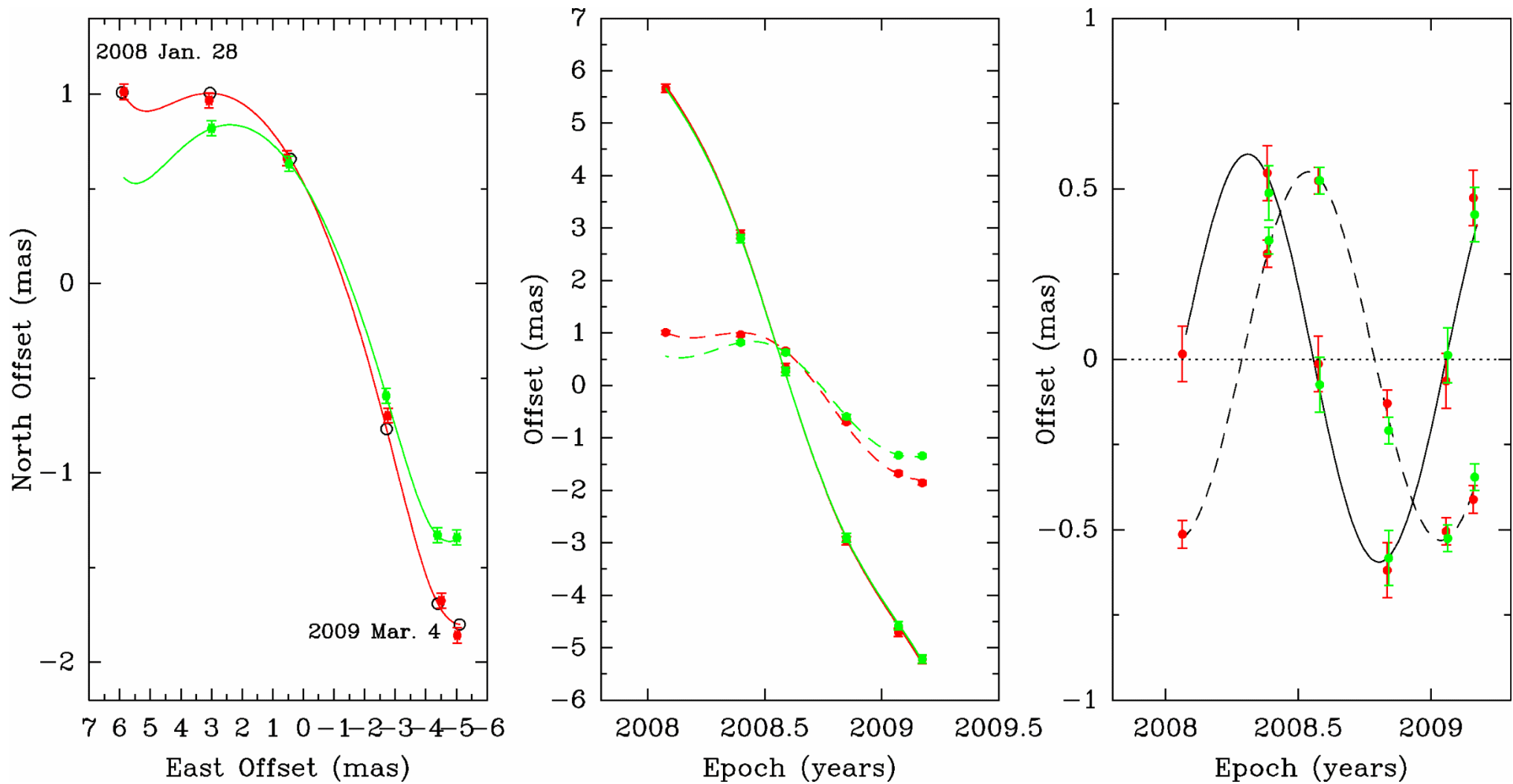
Moscadelli et al. (2011)

Hints at jet rotation

IRAS20126+4104: 22 GHz water maser

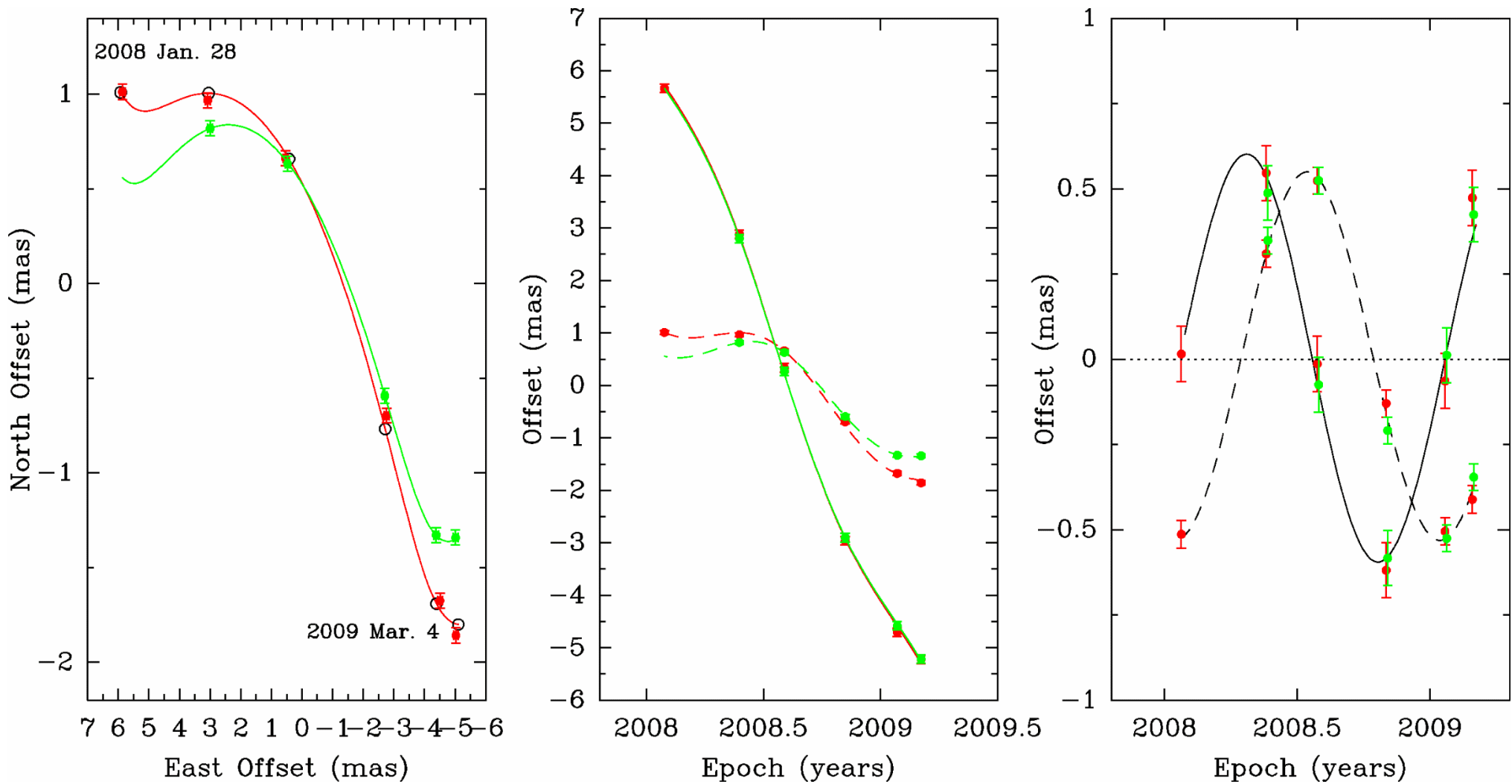


22 GHz maser parallax in IRAS 20126+4104



22 GHz maser parallax in IRAS 20126+4104

$\pi = 0.61 \pm 0.02$ mas \rightarrow $d = 1.64 \pm 0.05$ kpc



April 17, 1555:

The (glorious) Republic of Siena and the French forces of Biagio di Montluc, after a long siege – from January 1554,

surrendered to the army of Carlo V (Spain) & Cosimo I de' Medici (Firenze).

Because of the heroic defense, Carlo V accorded the honours of war and the possibility to leave the town to those we wished to do that.

Hundreds of families together with Biagio di Montluc and the remnant of the army left Siena and went to Montalcino. The “Republic of Siena in Montalcino” kept on fighting till 1559, when the treaty of Cateau-Cambrésis stated the end of the war between Spain and French in Italy with the victory of Spain.

“(Lo Stato di Siena) è mio et a me s'appartiene in tutto”

(Cosimo I de' Medici after the signature of the treaty of Cateau-Cambrésis)

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Summer 2016:

mosca, after a long siege, surrendered to Maite Beltrán and **Cesa**

and accepted to take part to the data reduction and analysis of

mm interferometric data,

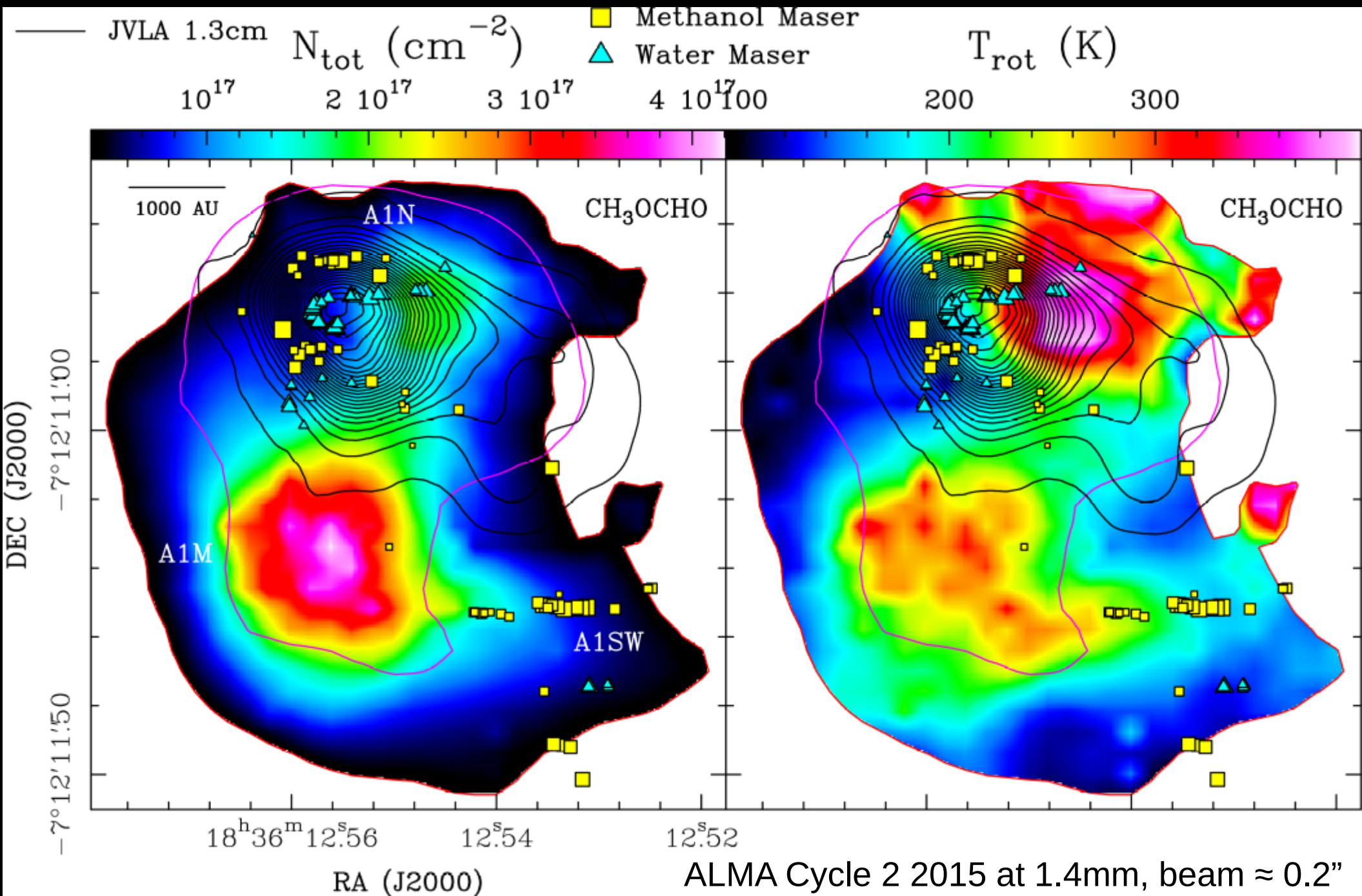
(specifically, the ALMA Cycle 2 observations of the source G24.78+0.08).

Because of the strenuous resistance, the honours of war were accorded,

and **mosca** was allowed to reduce and analyze the corresponding

maser VLBI data, too.

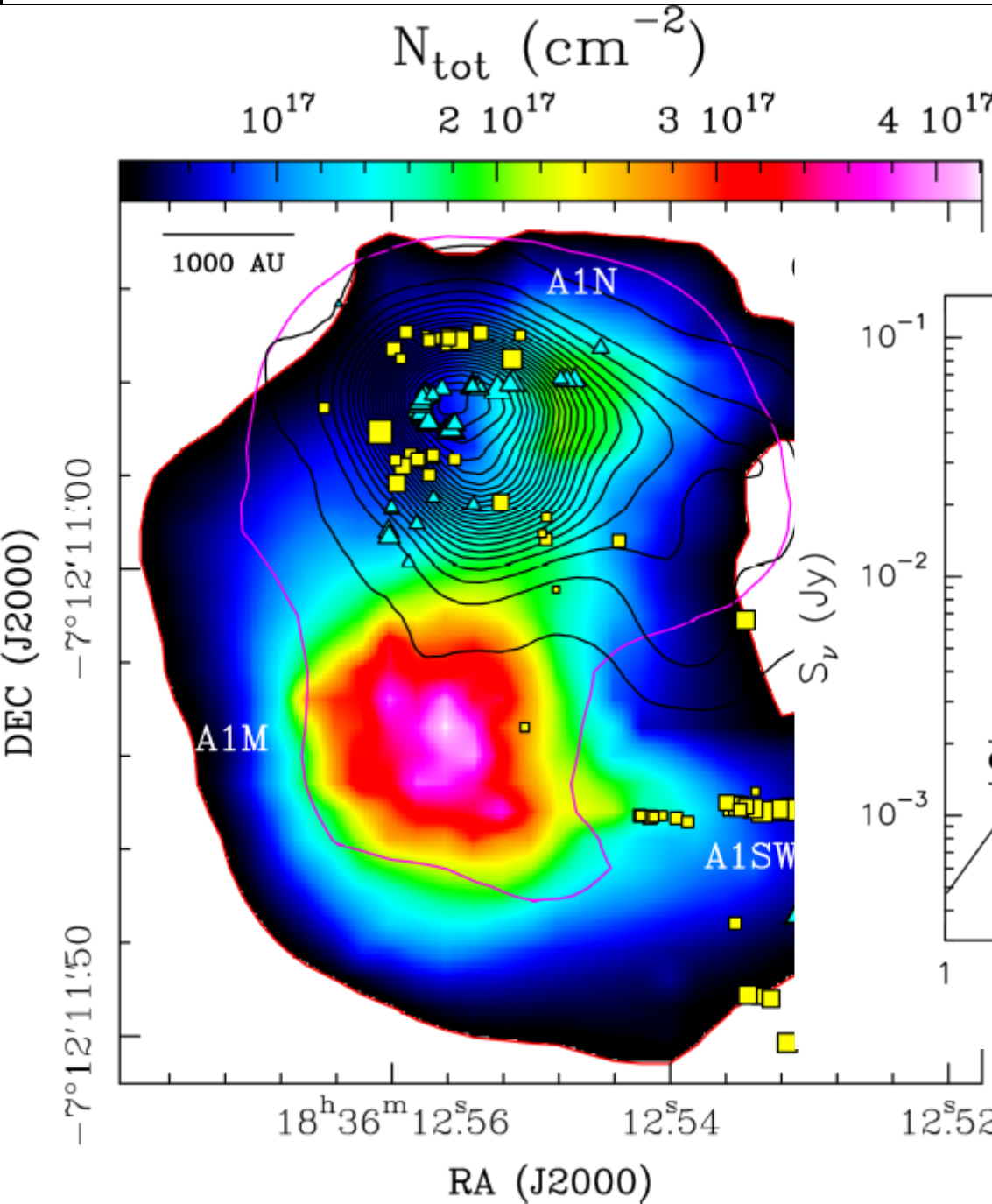
Core A1 in the high-mass SFR G24.78+0.08



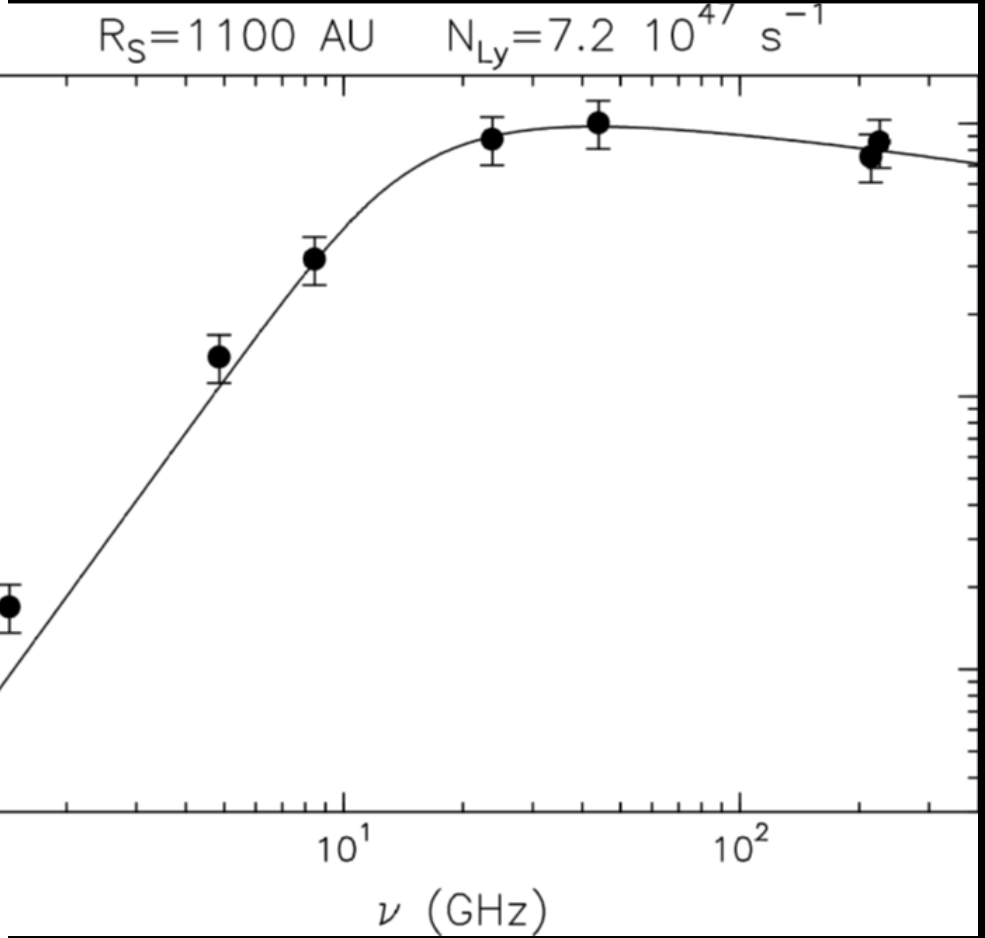
(Moscadelli et al. 2018)

Core A1 in the high-mass SFR G24.78+0.08

— : JVLVA 1.3 cm ; ■ : meth. maser ; ▲ : water maser



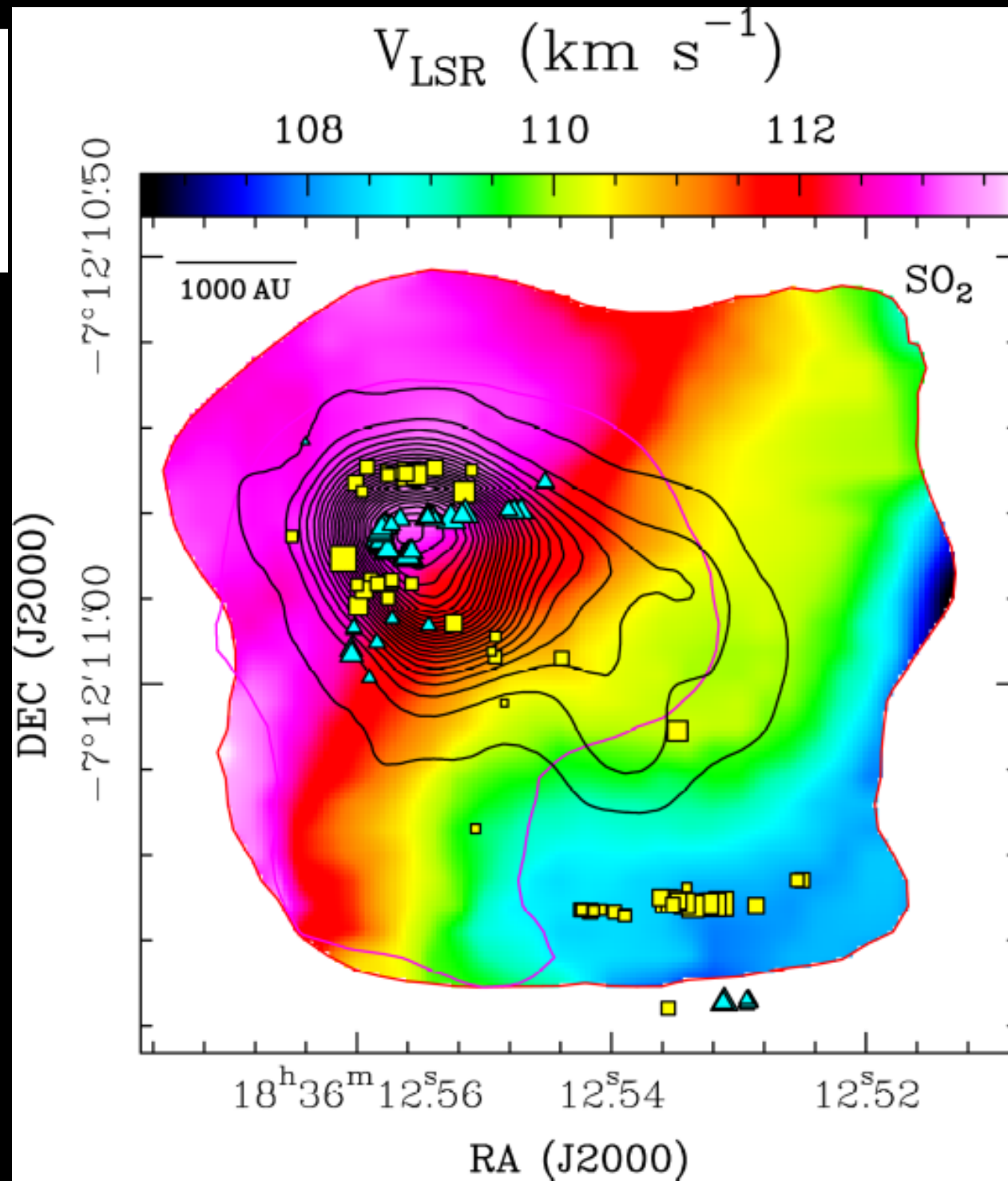
FREE-FREE EMISSION:
ZAMS TYPE O9.5, $M \approx 20M_\odot$



ALMA Cycle 2 2015 at 1.4mm, beam $\approx 0.2''$
(Moscadelli et al. 2018)

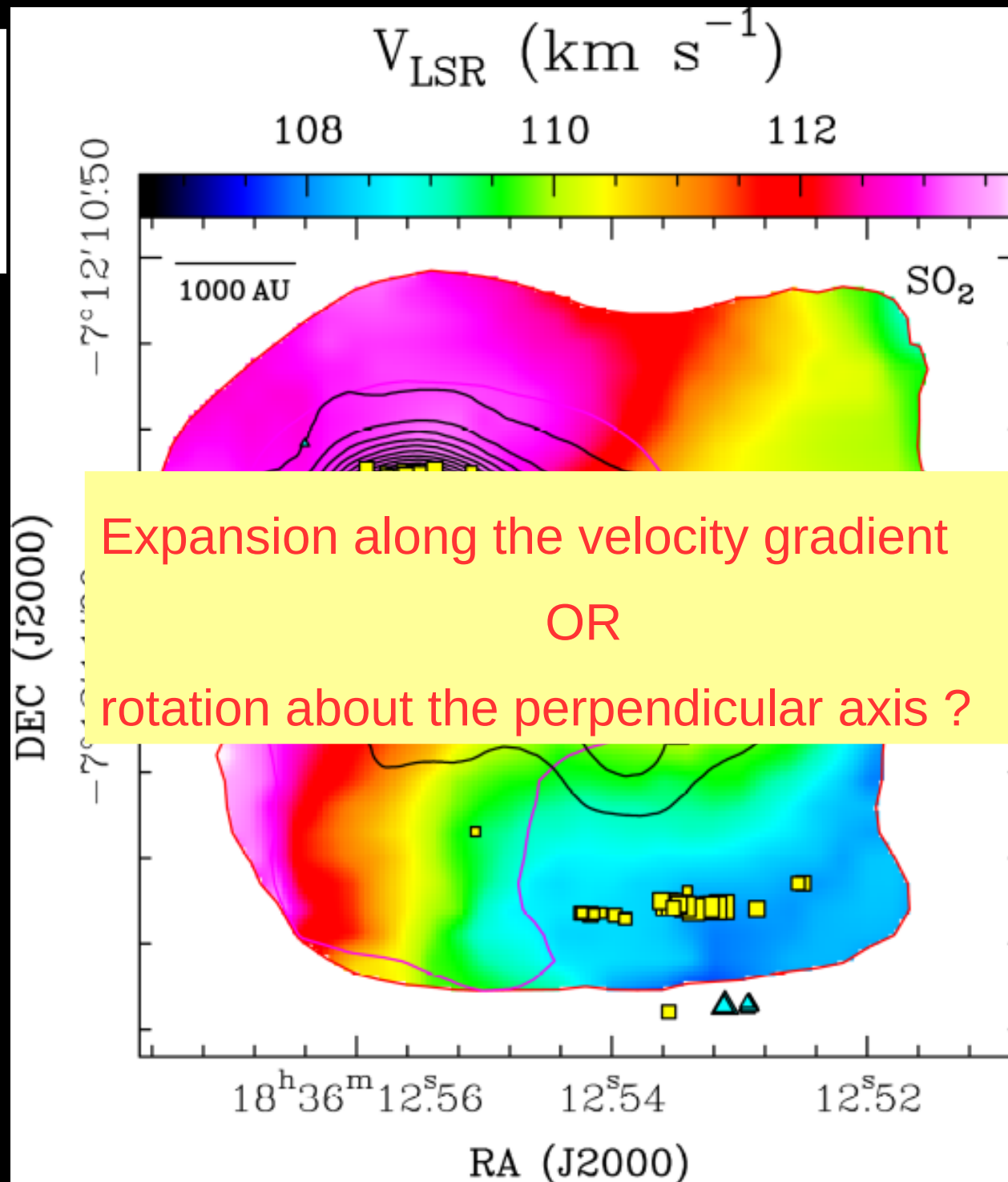
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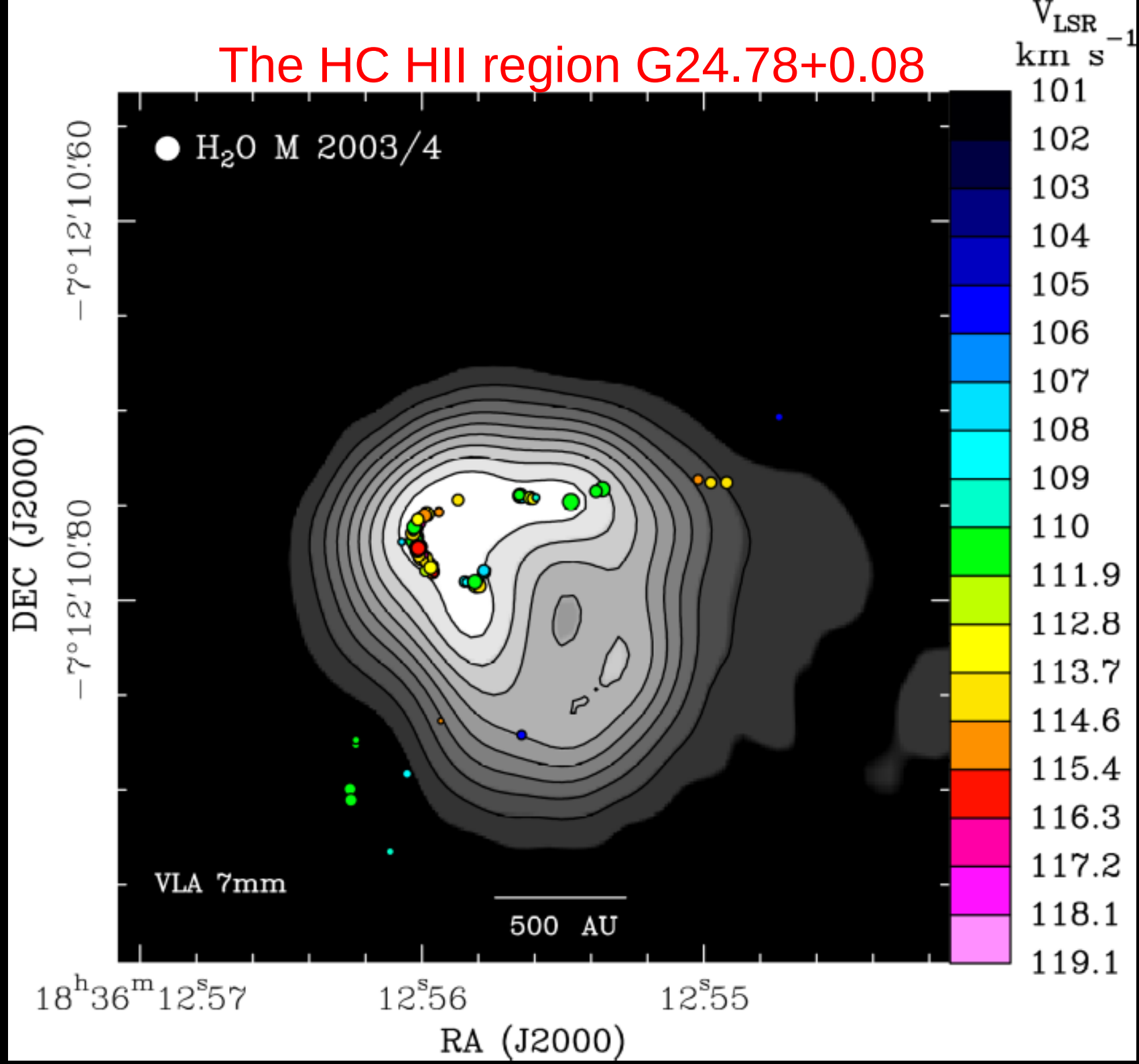


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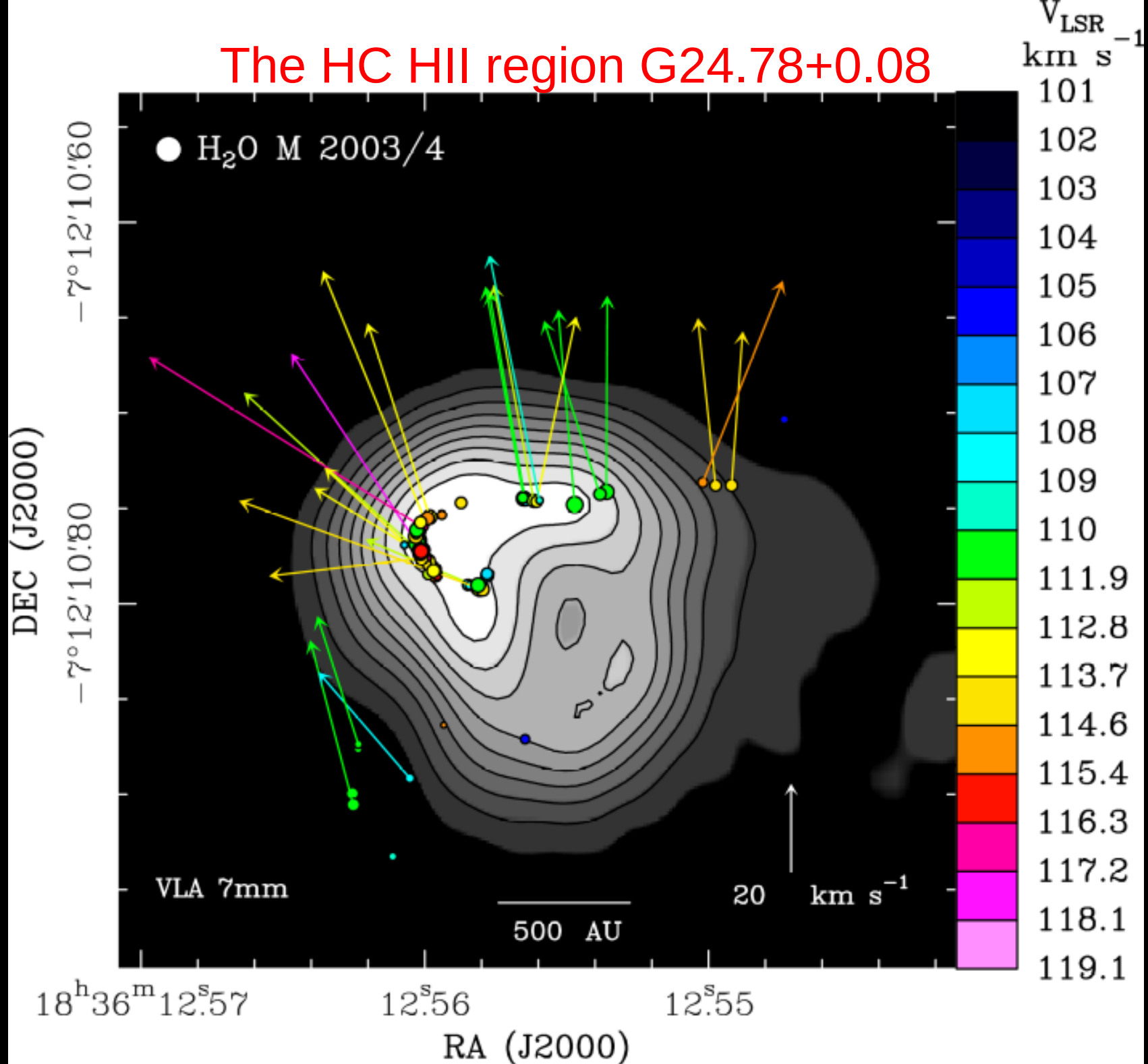
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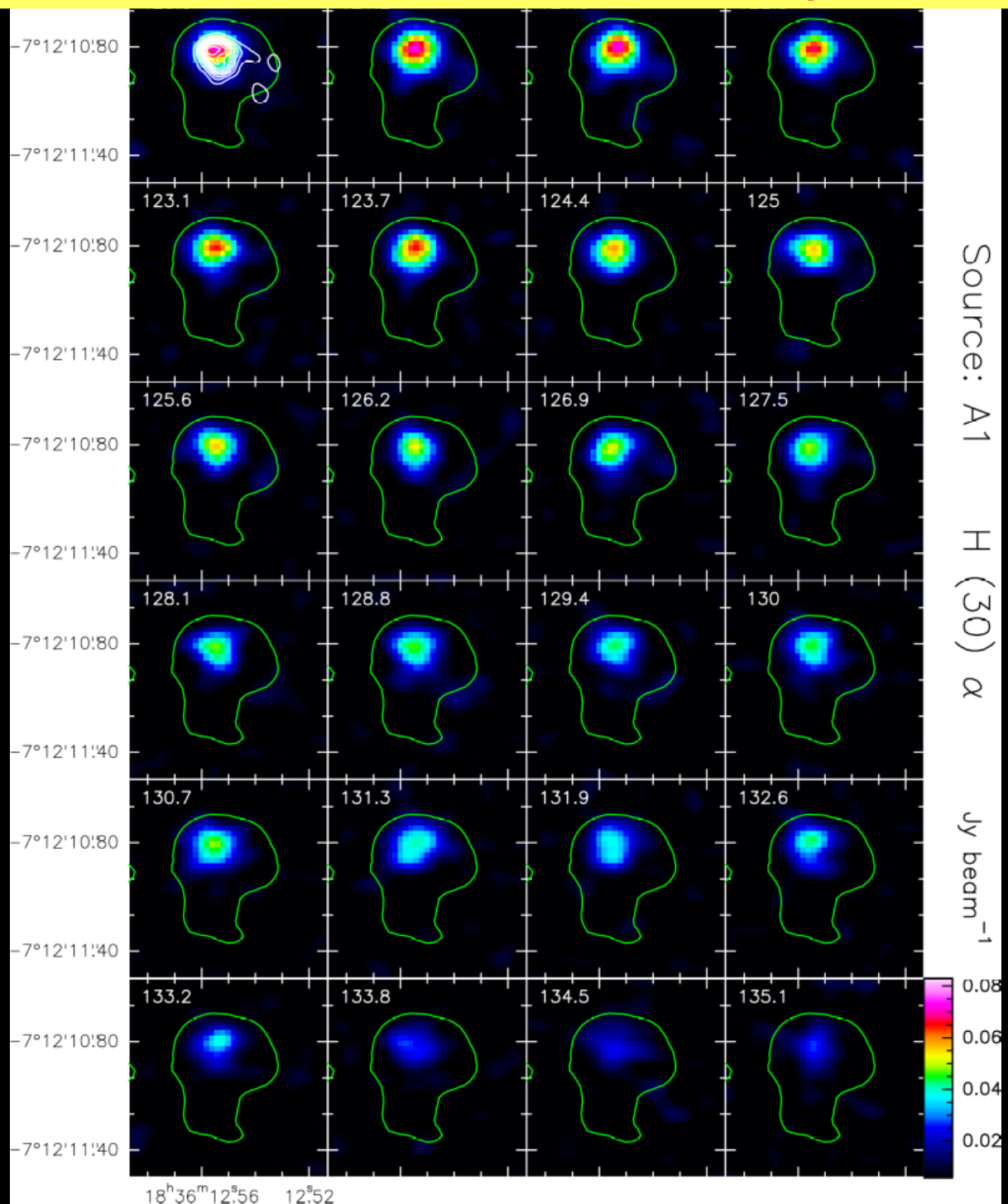
The HC HII region G24.78+0.08



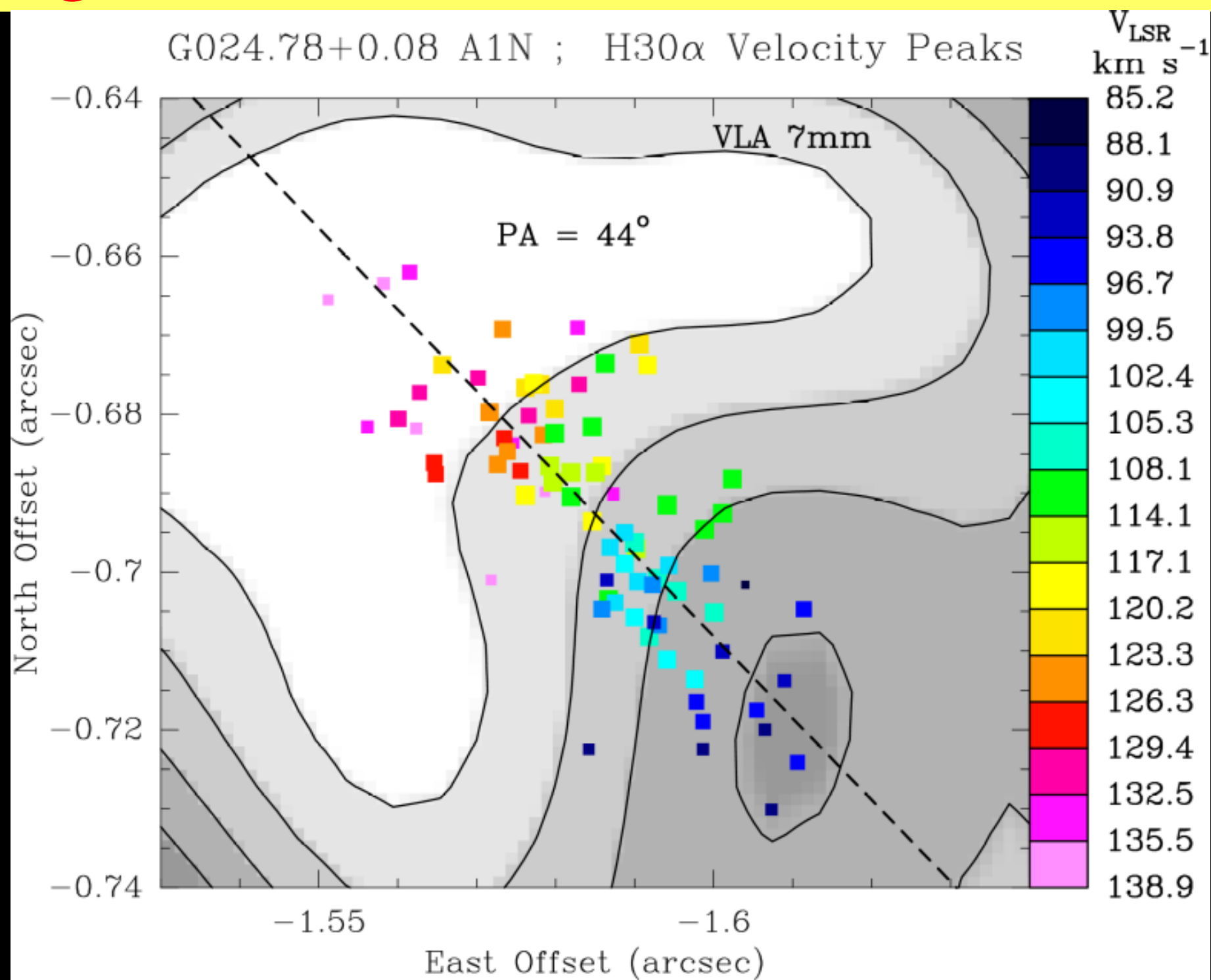
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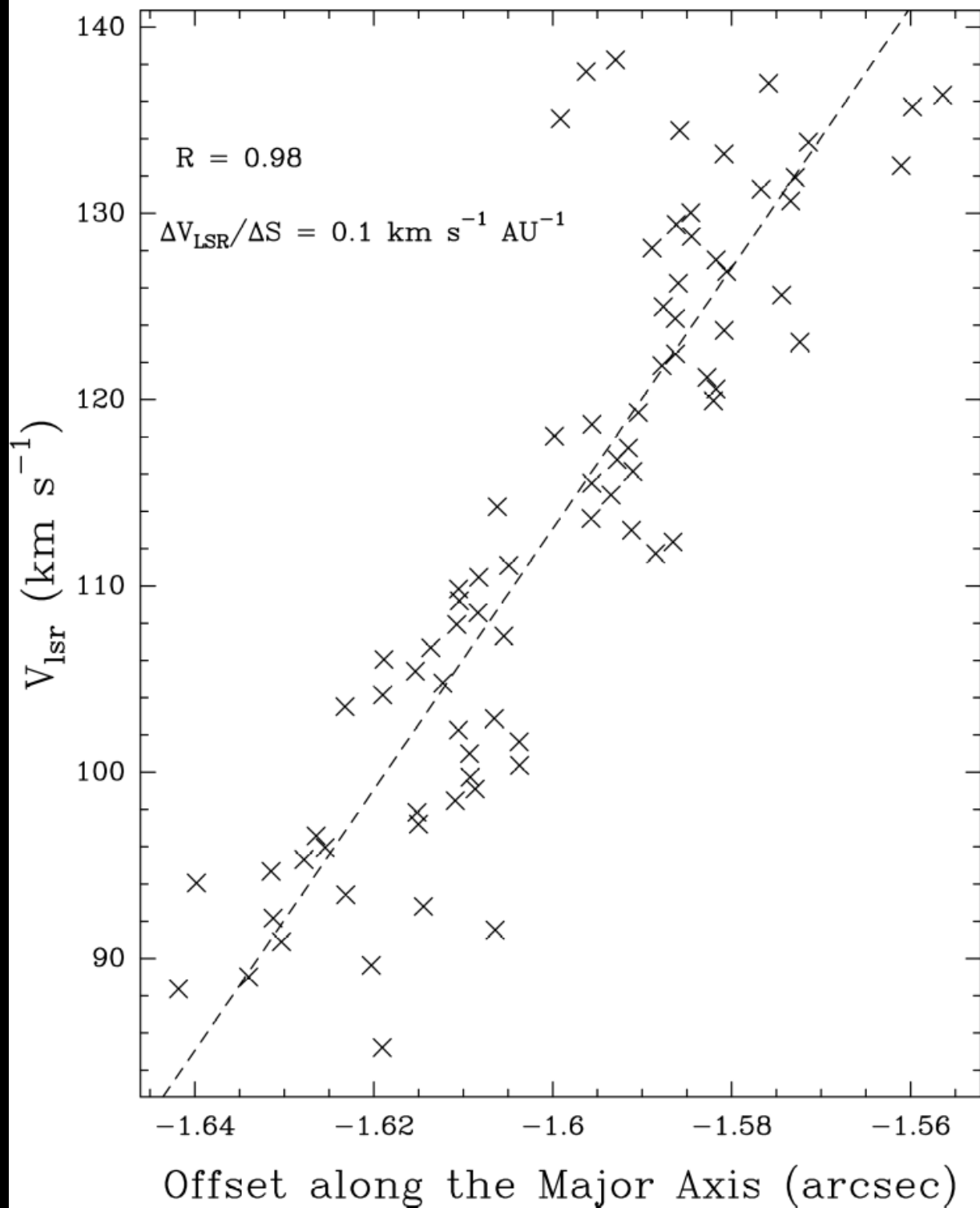


ALMA @ 1.4mm: H30 α channel maps towards core A1

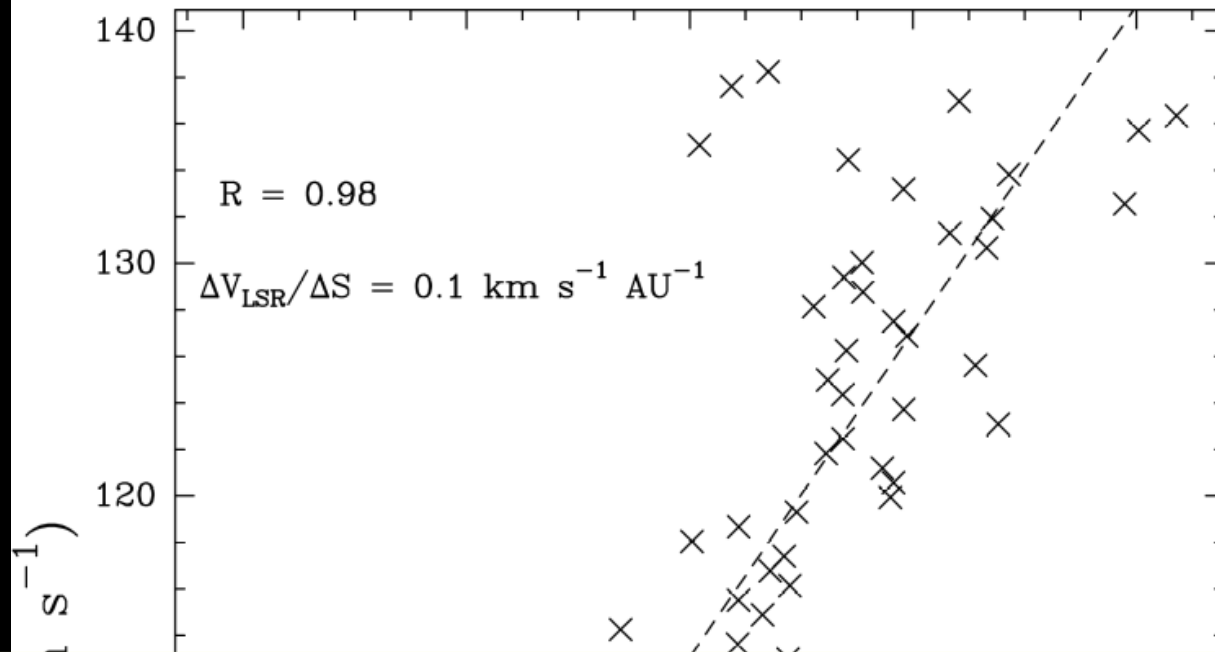


ALMA @ 1.4mm: H30 α Vel. Grad. towards A1N HCHII

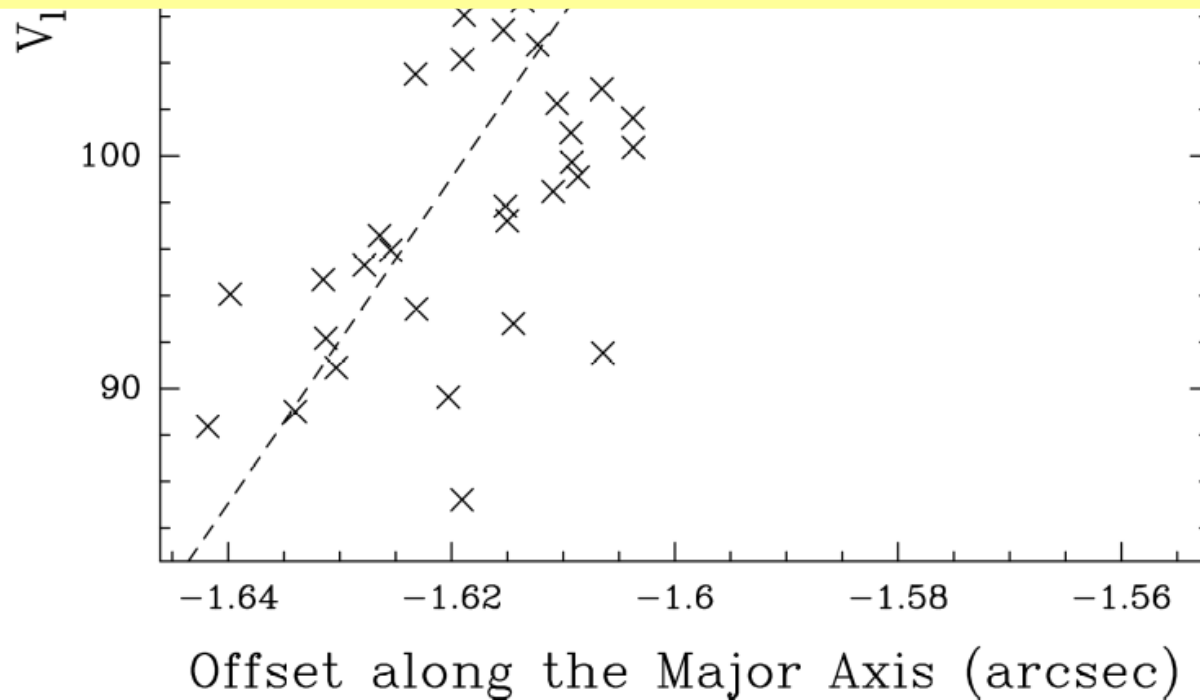


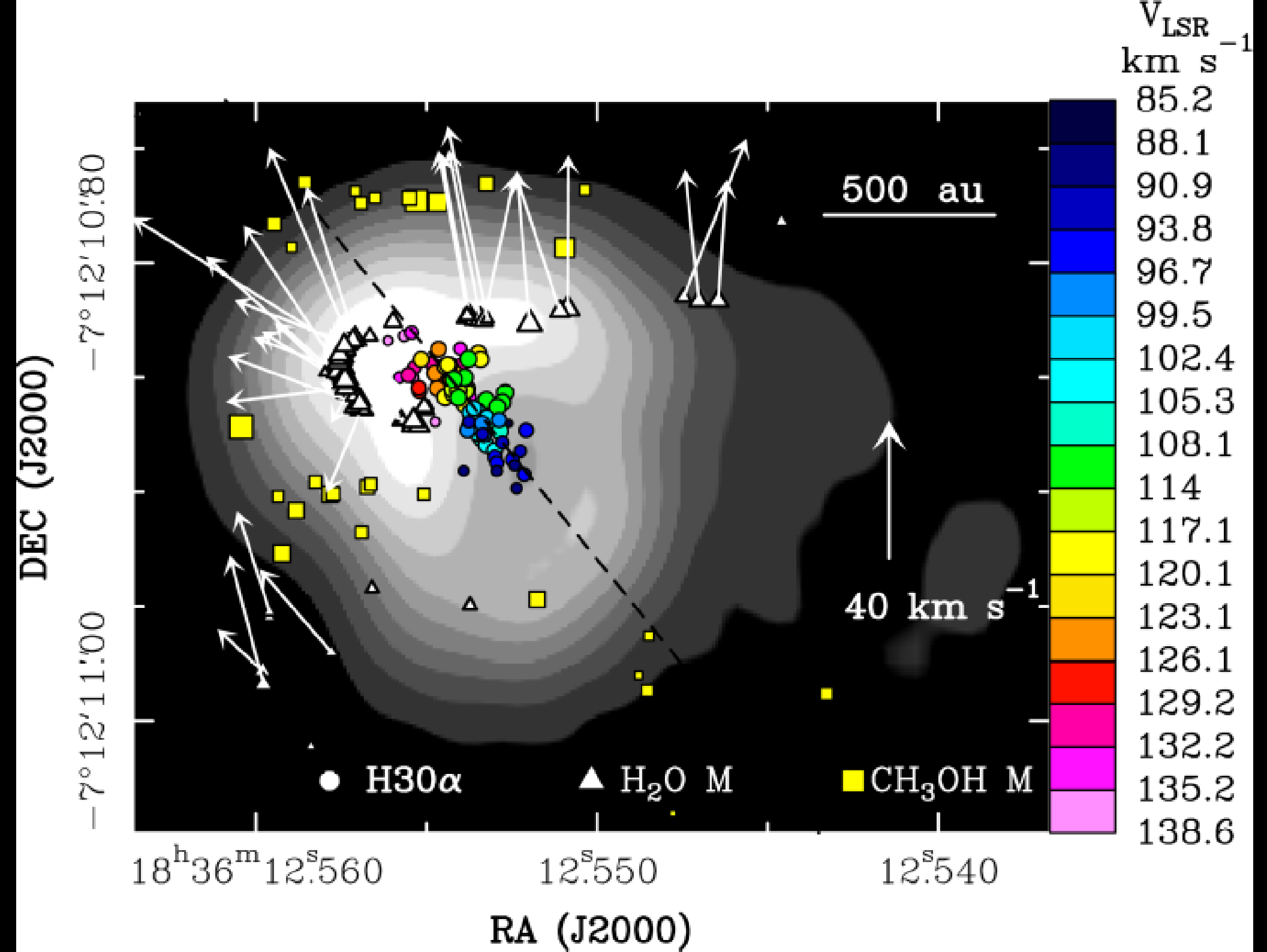


G24.78 A1N; H30 α V_{LSR} grad. @ PA = 44 $^\circ$

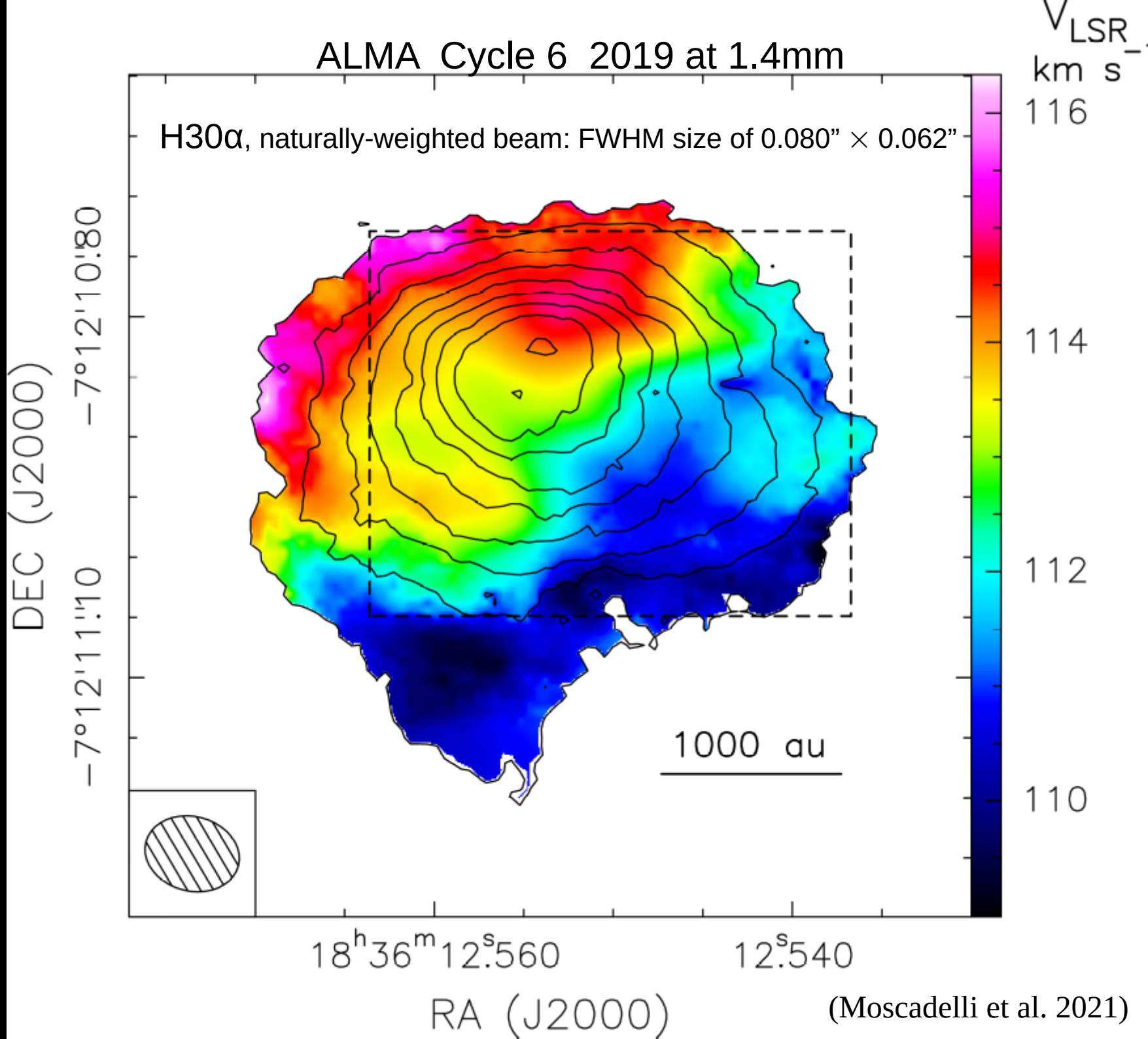


$$\Delta V_{\text{LSR}} / \Delta S \sim 50 \text{ km s}^{-1} / 500 \text{ AU} \Rightarrow M_{\text{dyn}} > 180 M_{\odot}$$



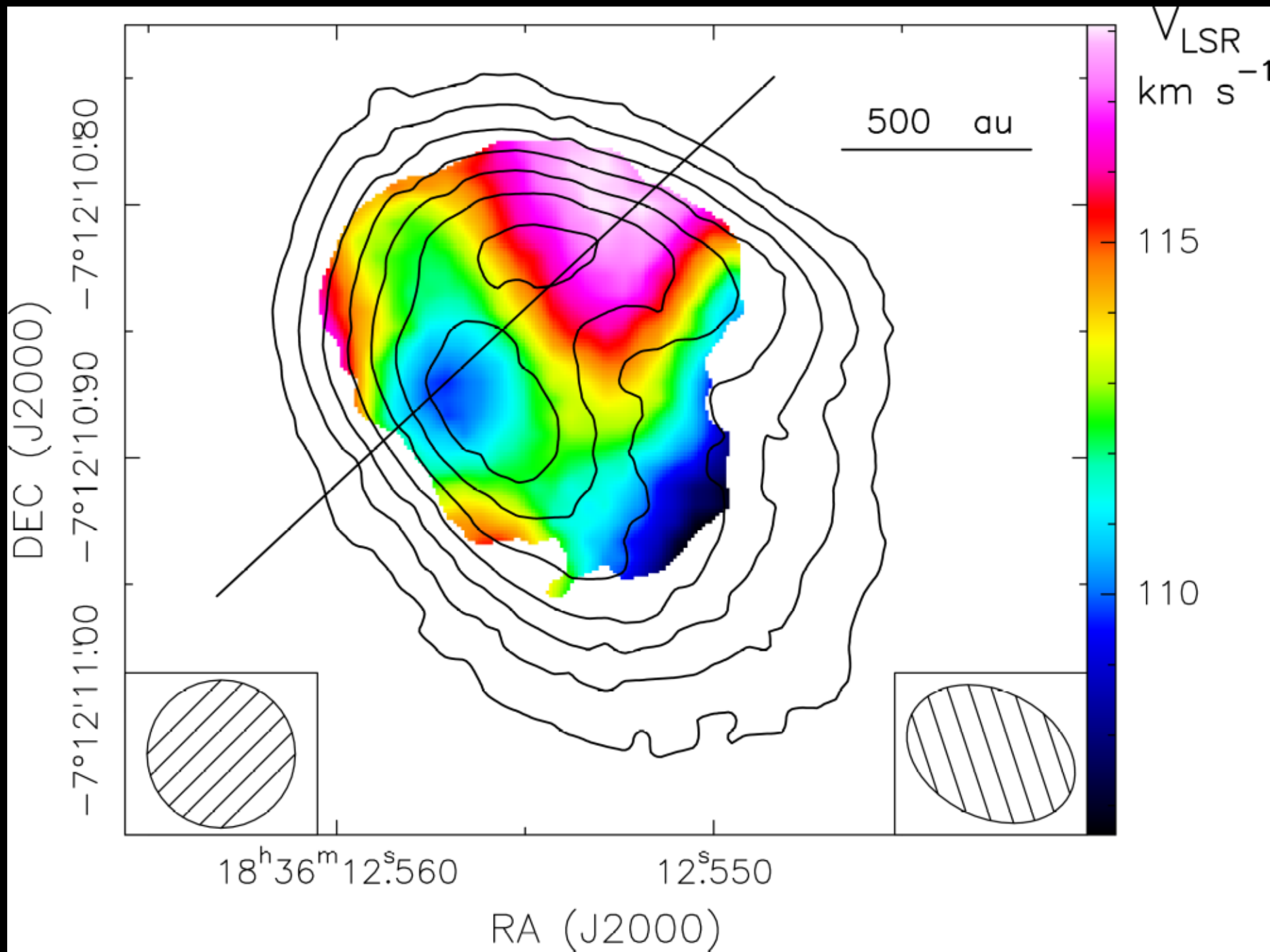


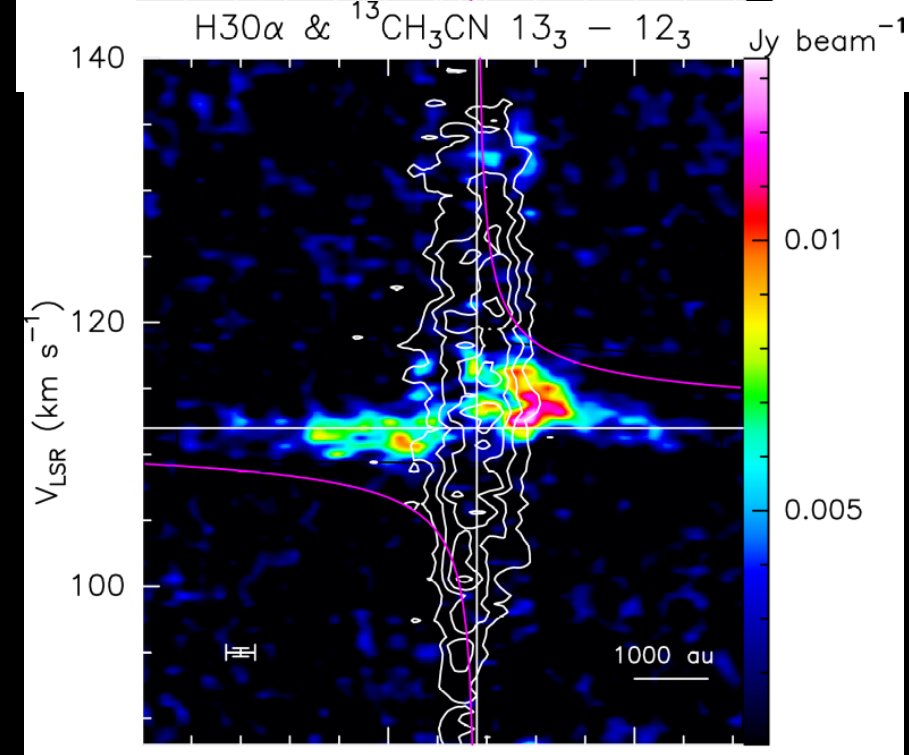
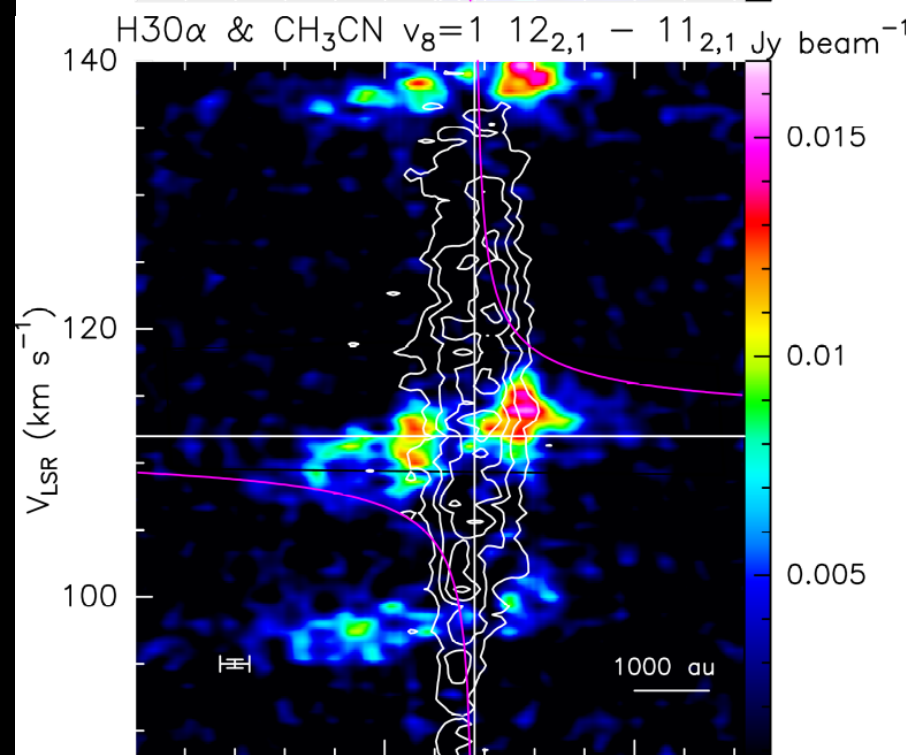
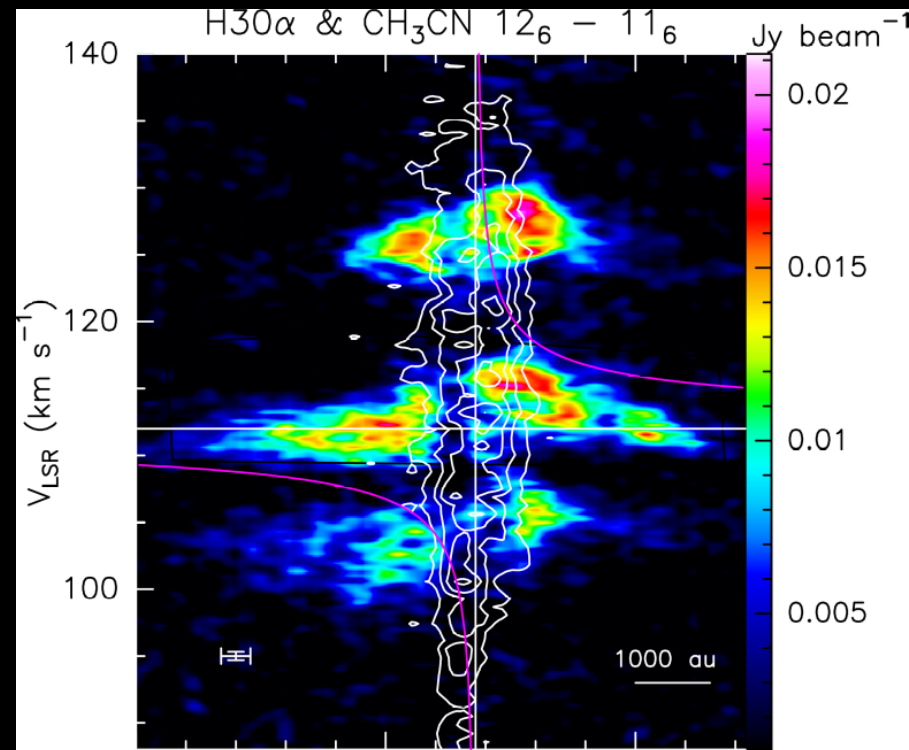
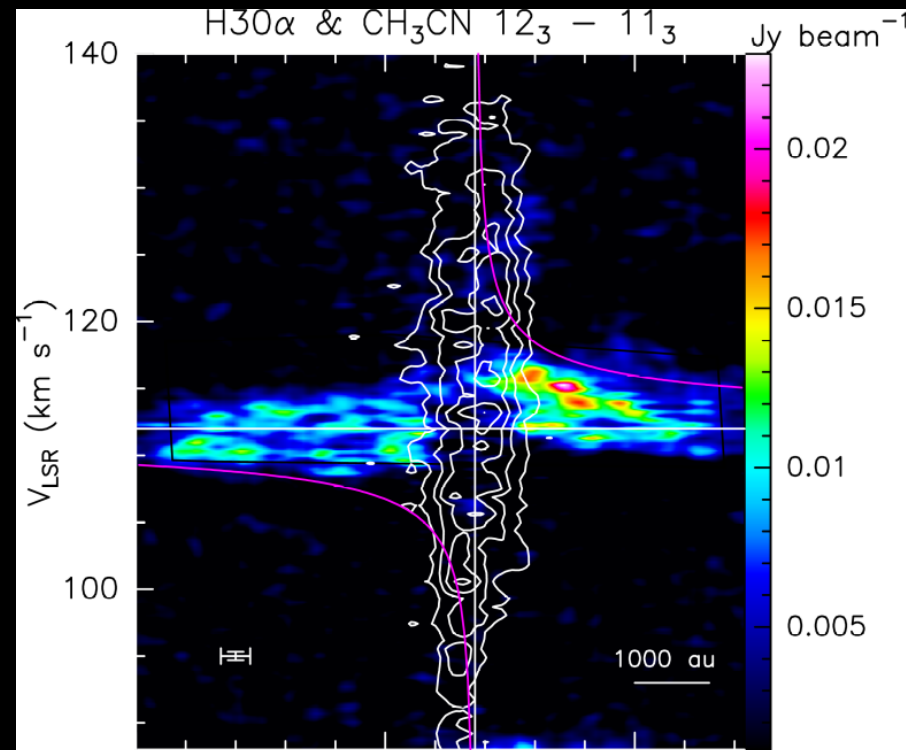
ALMA Cycle 6 2019 at 1.4mm



Color map: H30 α , Briggs robust=0.5 beam: FWHM size of 0.058"

Black contours: 1.4 mm ALMA continuum





0 -0.5 -1
Offset along axis @ PA = 133° (arcsecond)

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Offset along axis @ PA = 133° (arcsecond)

CONCLUSIONS

- 1) In G24.78+0.08, 22 GHz water maser 3D motion helps distinguish between molecular core (A1) expansion and rotation. A strong case for a massive ionized star surrounded by an outer molecular and inner ionized disk, which is likely still actively accreting from its parental core.

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- 2) In IRAS 20126+4104, 6.7 GHz methanol and 22 GHz water masers trace the disk-jet system at scales of 10-100 au. In particular, the water masers trace a fast collimated jet, with hints at rotation. Accurate distance from water maser trigonometric parallax.

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- 3) A very detailed view of high-mass (embedded) YSOs can be obtained combining JVLA + ALMA (continuum / line, 0.05"-10") thermal data with multi-epoch VLBI (milli-arcsec) maser observations.

Palio in Siena: suddenly the Middle Age “resurrects”

Nicchio



Contrade;
the quarters
of Siena



Palio in Siena: suddenly the Middle Age “resurrects”

Nicchio



Contrade;
the quarters
of Siena



Cesa: the new jockey ("fantino") of Nicchio



Overwhelmed by emotion, I warmly thank Riccardo for his friendship, generous contribution to our research and sense of humor that made everyday work more joyful.

