

# Commissioning SKA-Low

First results and future plans  
towards Science Verification

Giulia Macario (on behalf of the SKA-Low Science Commissioning team)

Fifth National Workshop on the SKA Project, Bologna, Italy  
24th November 2025

We recognise and acknowledge the Traditional Owners of the lands on which our facilities are located, and pay our respects to their Elders past and present.



# SKAO Science Commissioning in the big picture

*Internal* activity that focuses on tests of science capability, using astronomical observations



*Includes Short Functional Tests (SFTs)*

*Includes Test Cases, Verification Events, ...*

*Structured around a Commissioning Plan for each Array Assembly*

*Ensuring the telescope meets the needs of the science & operational users*



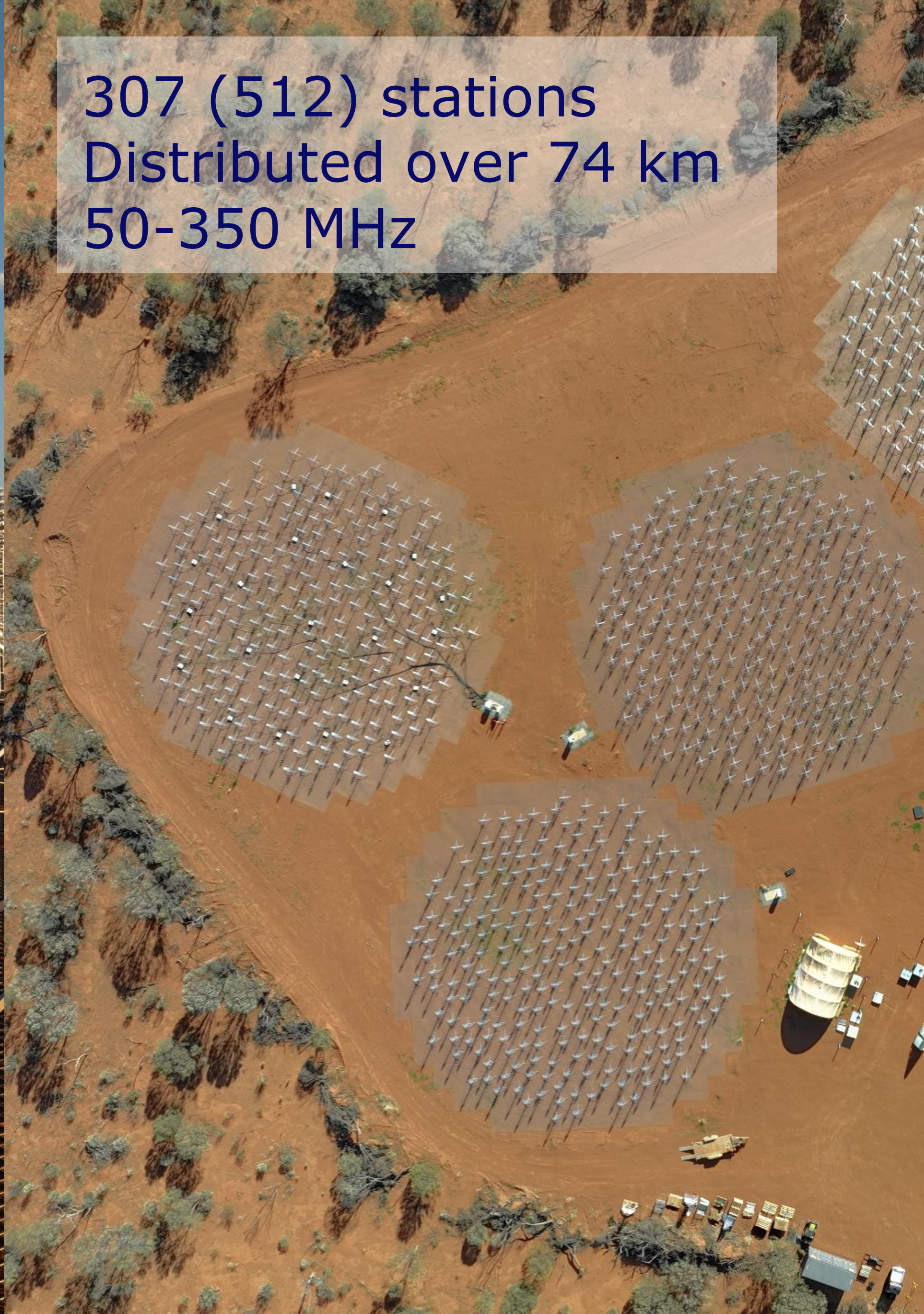








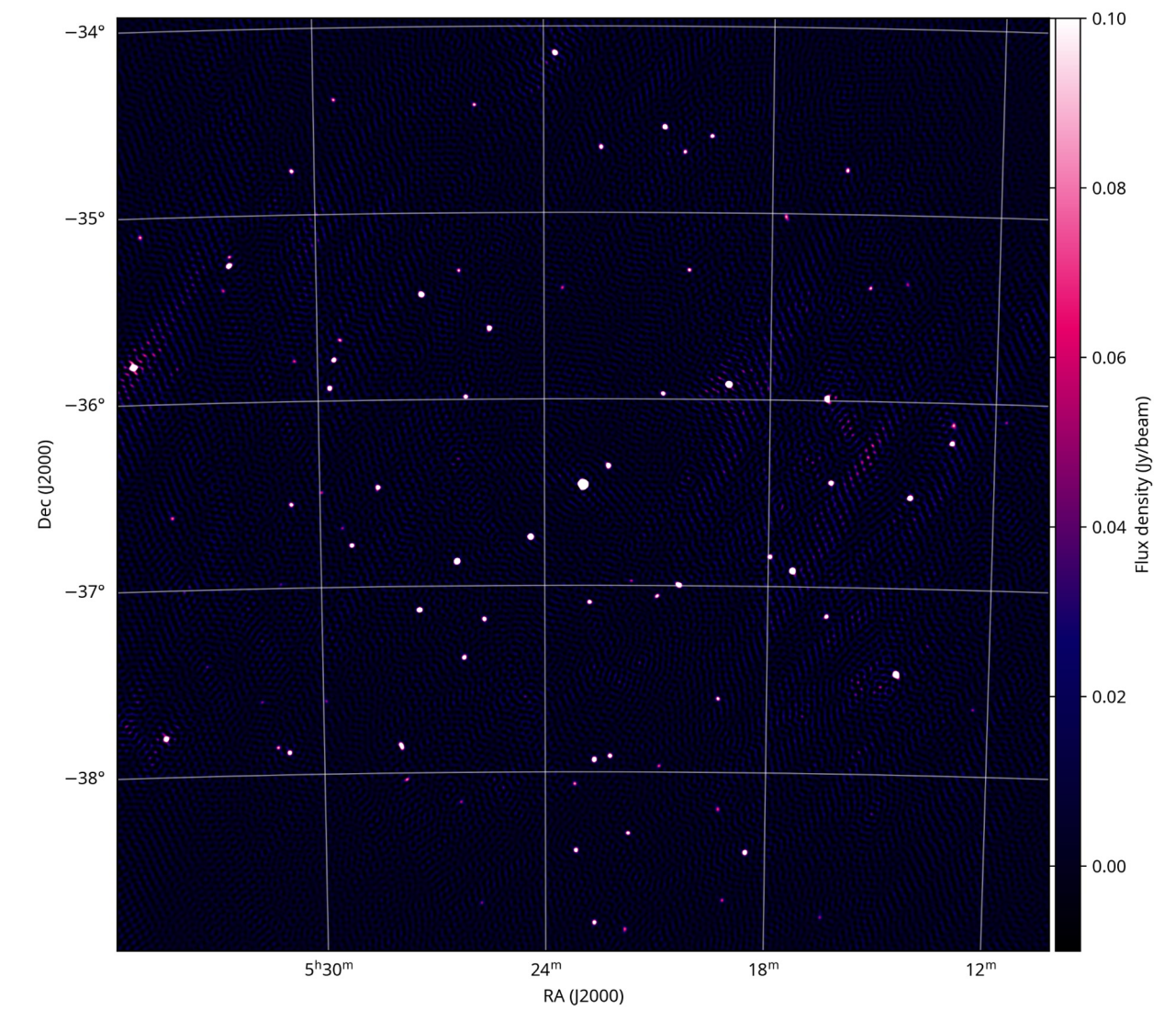
307 (512) stations  
Distributed over 74 km  
50-350 MHz



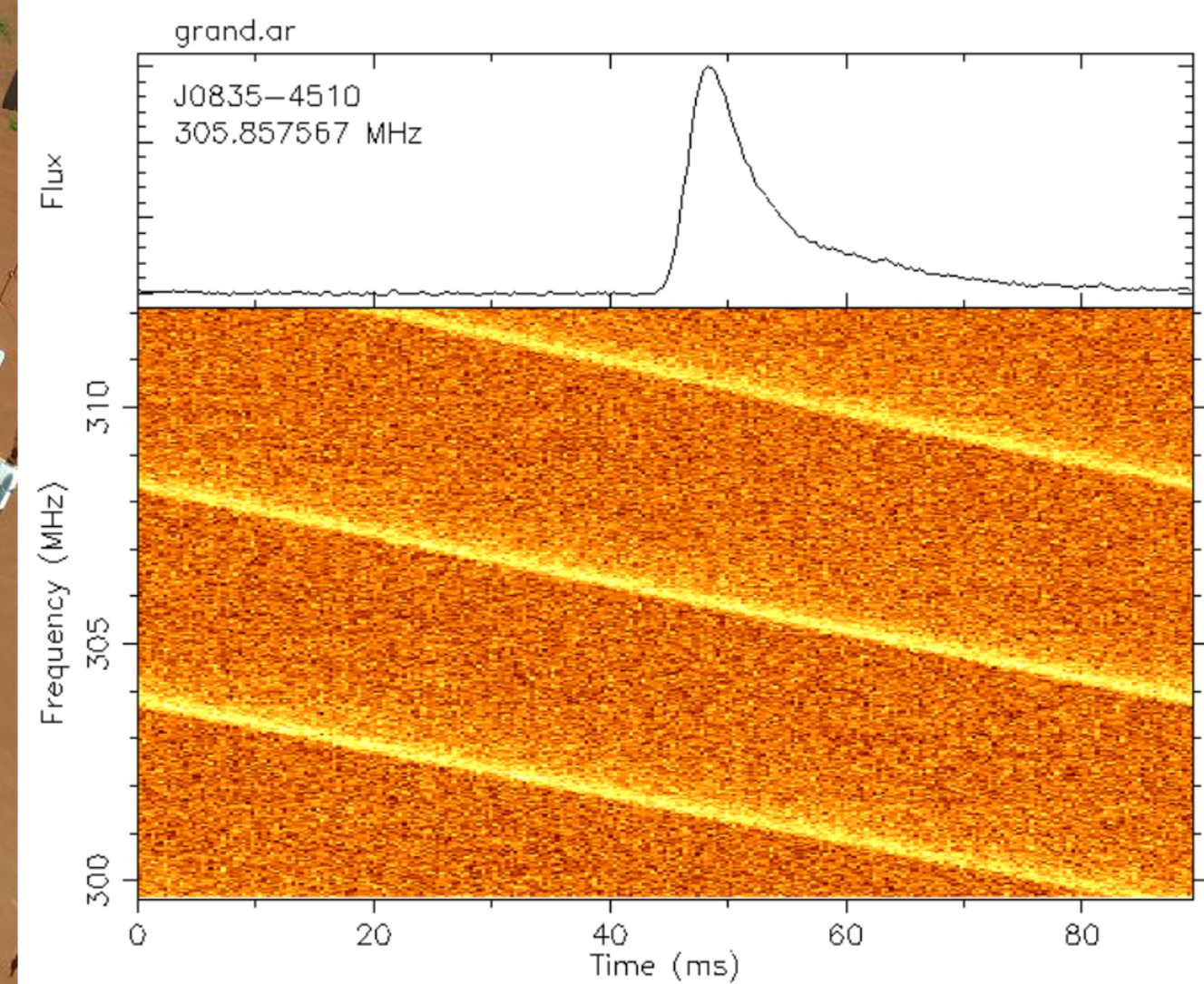




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**Images**  
**Tied-array beams**

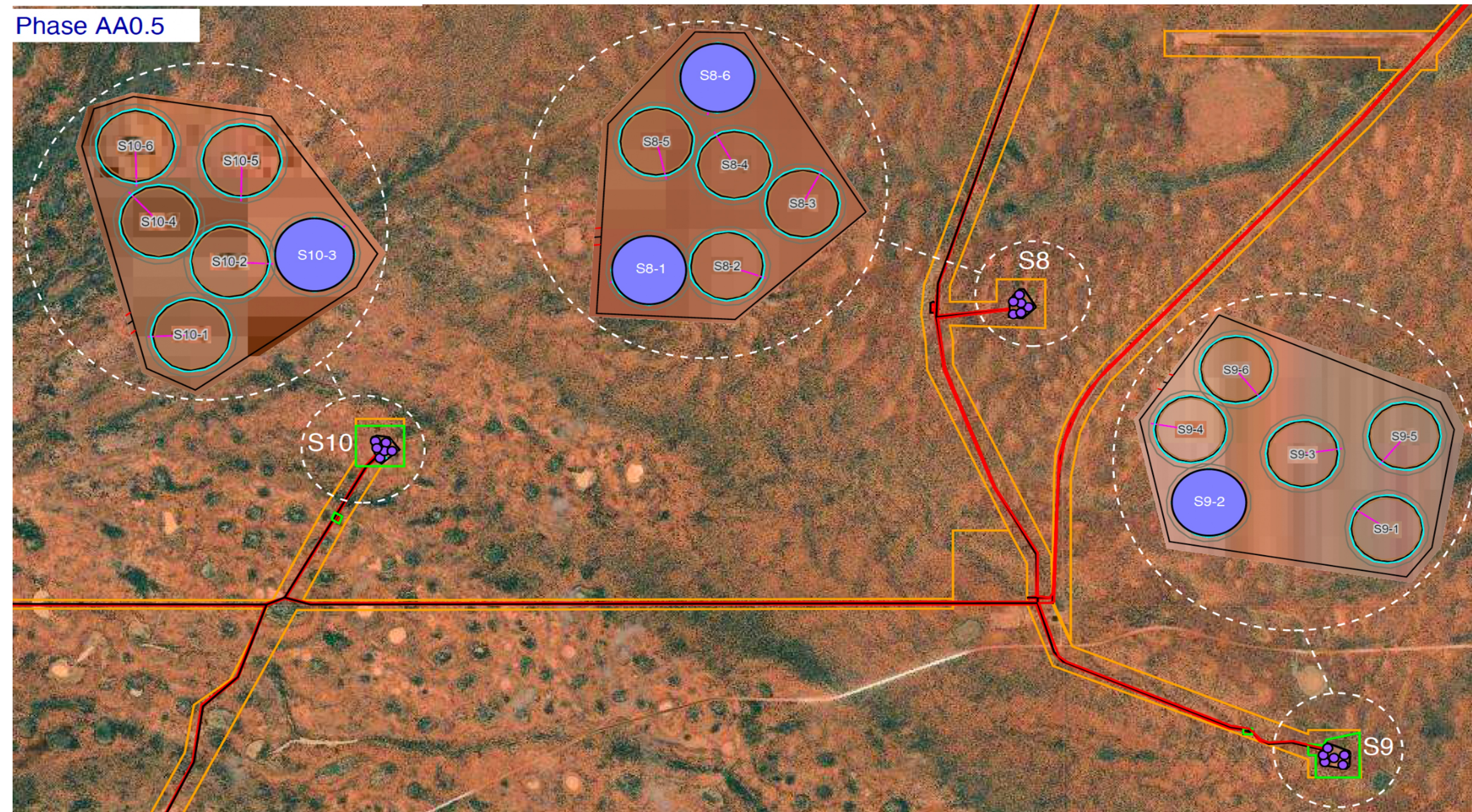




# SKA-Low Array Assembly 0.5 - layout

AA0.5 Complete!

- Four stations:  
S8-1, S8-6, S9-2 and  
S10-3
- Shortest baseline  
~120m, longest  
~5.5km
- Stations are rigidly  
rotated relative to  
each other
- Next milestone:  
AA1 (16 stations)



AA0.5: Four operational stations spread across the Southern spiral arm  
AA1: 16 operational stations in the same three clusters





# SKA-Low Science Commissioning main focus areas

**AA0.5** (40 tests, 95% completed); **AA1** (47 tests, already started)

**Single station**  
(system performance and  
pulsars detections)

**Array calibration and  
imaging performance**  
(continuum, polarization,  
spectral line)

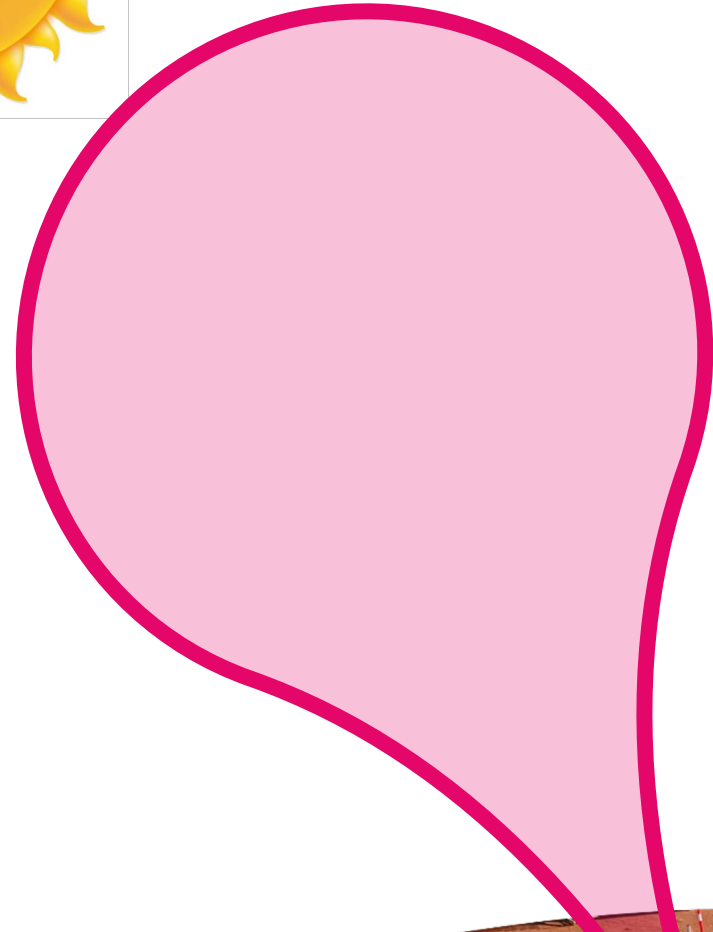
**Array beamforming**  
(tied array beams)

Development of **SKA-Low  
Calibrator Database**  
(and cal. survey plans)





# Single station: behaviour



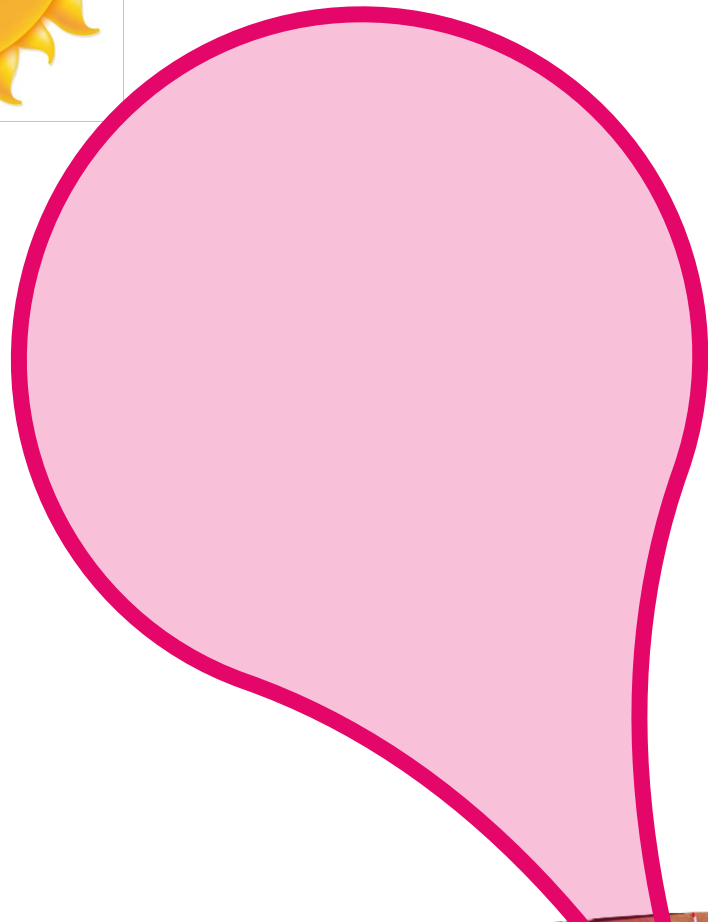
## Commissioning outcomes so far:

- Good calibration solutions for each antenna within a station
- Stable performance with time and varying environmental conditions
- Per-antenna delays for pointing and tracking are fit for purpose
- Beam shape and sensitivity are approximately as expected



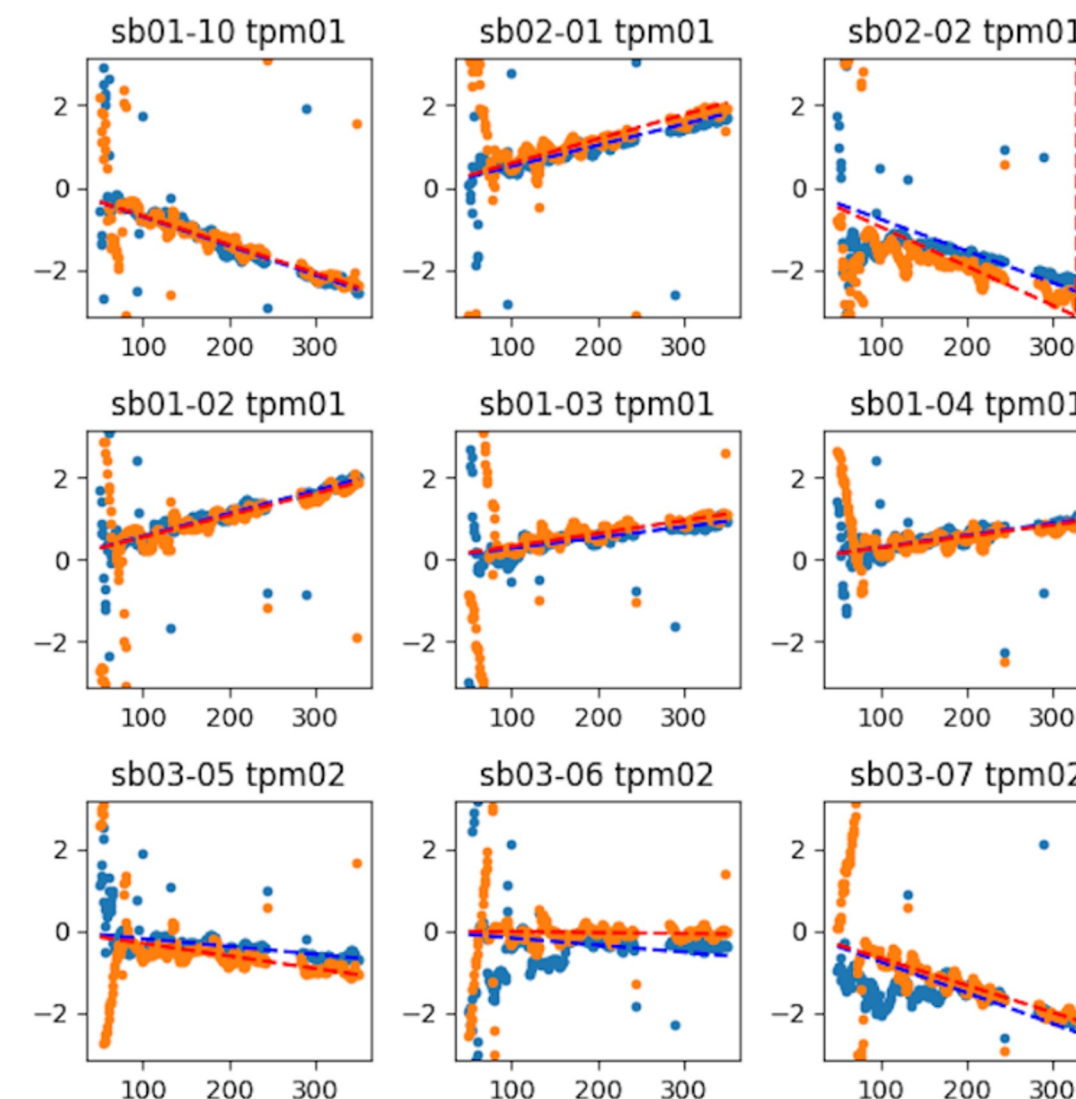


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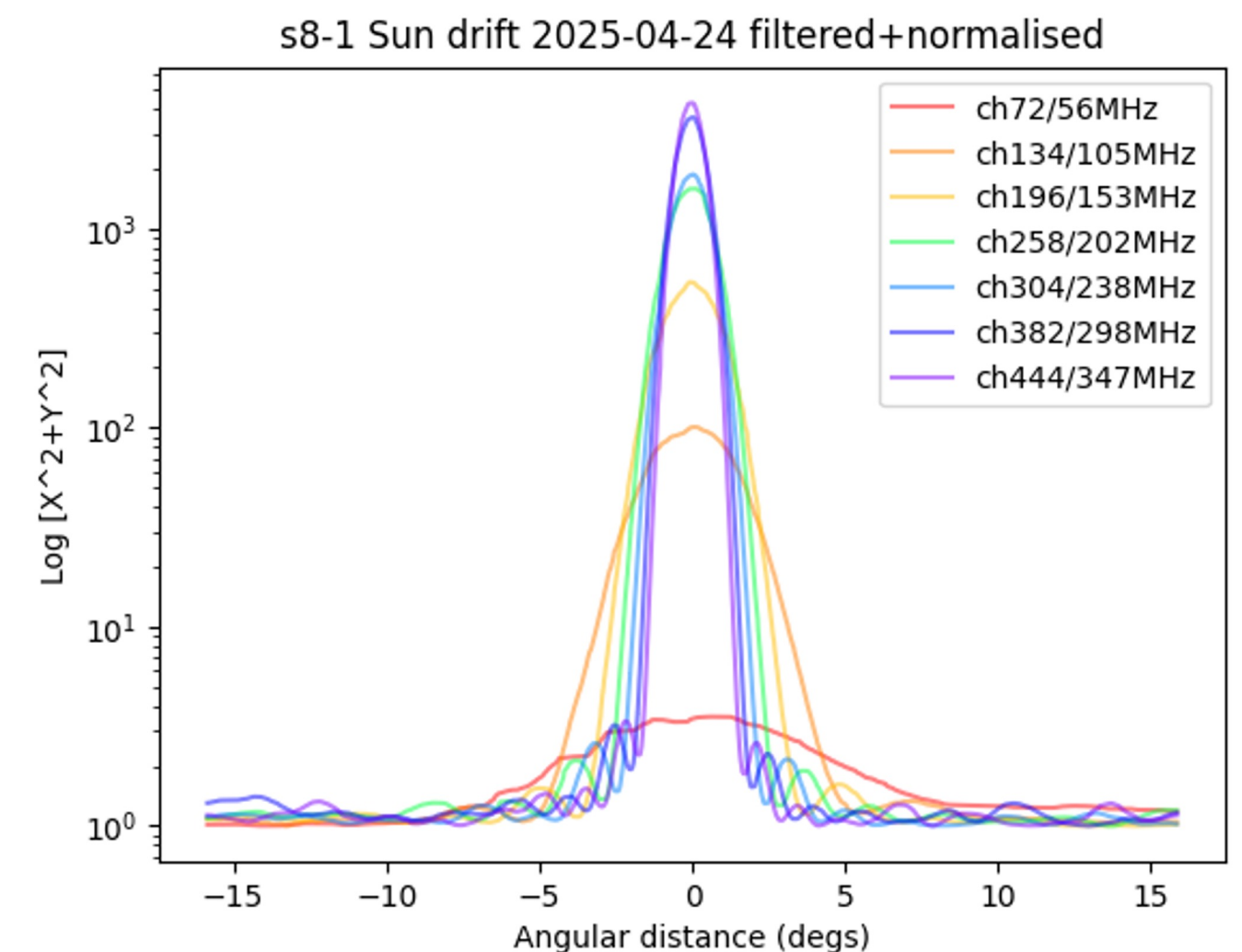
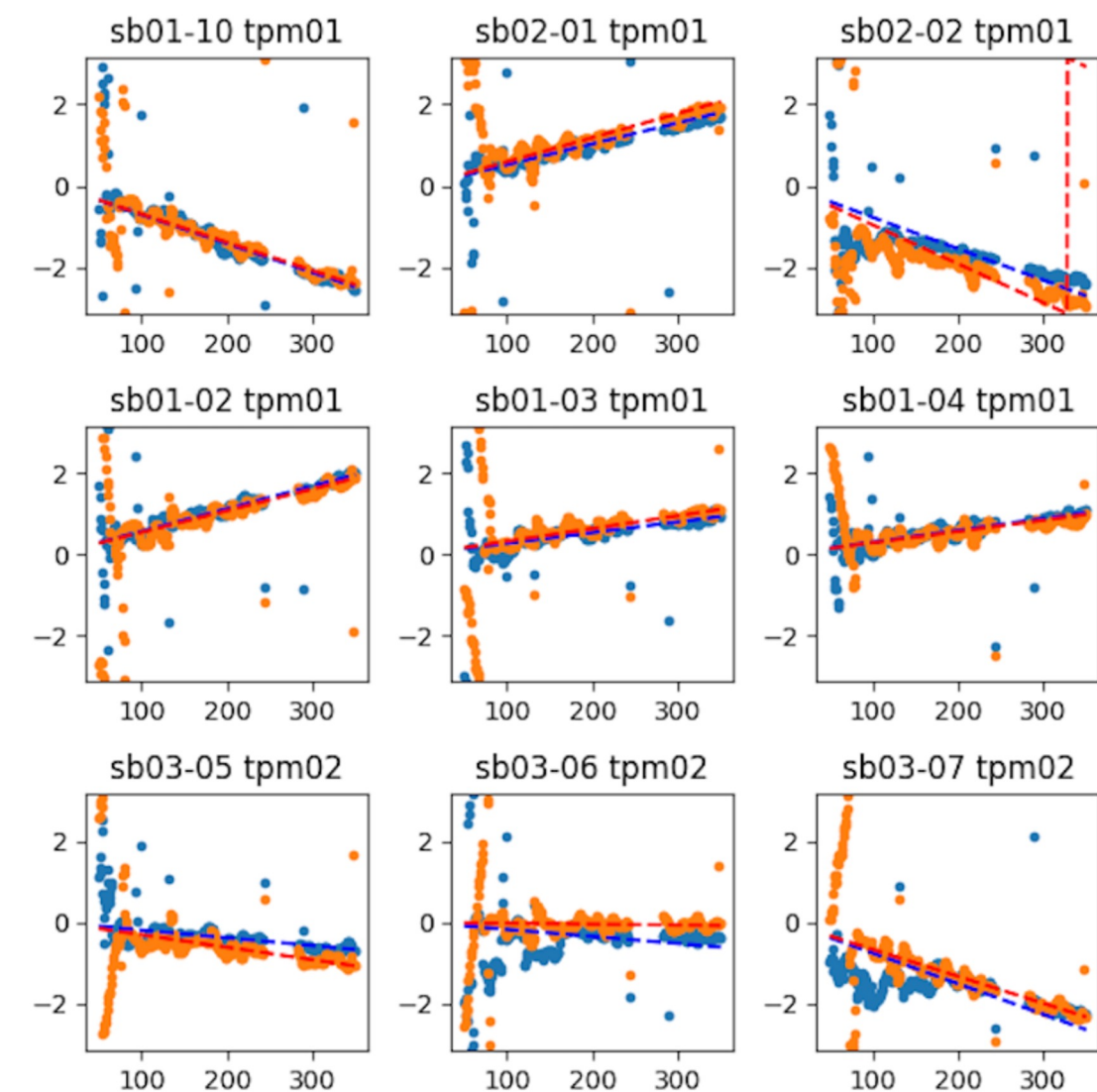


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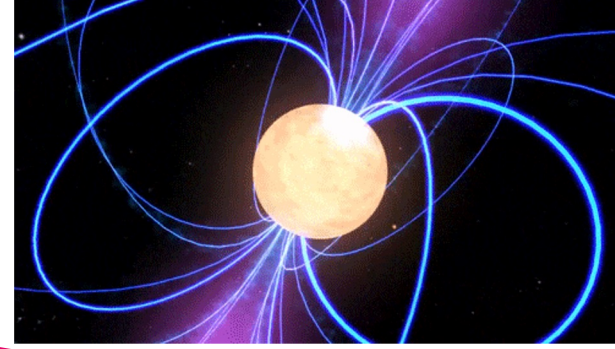
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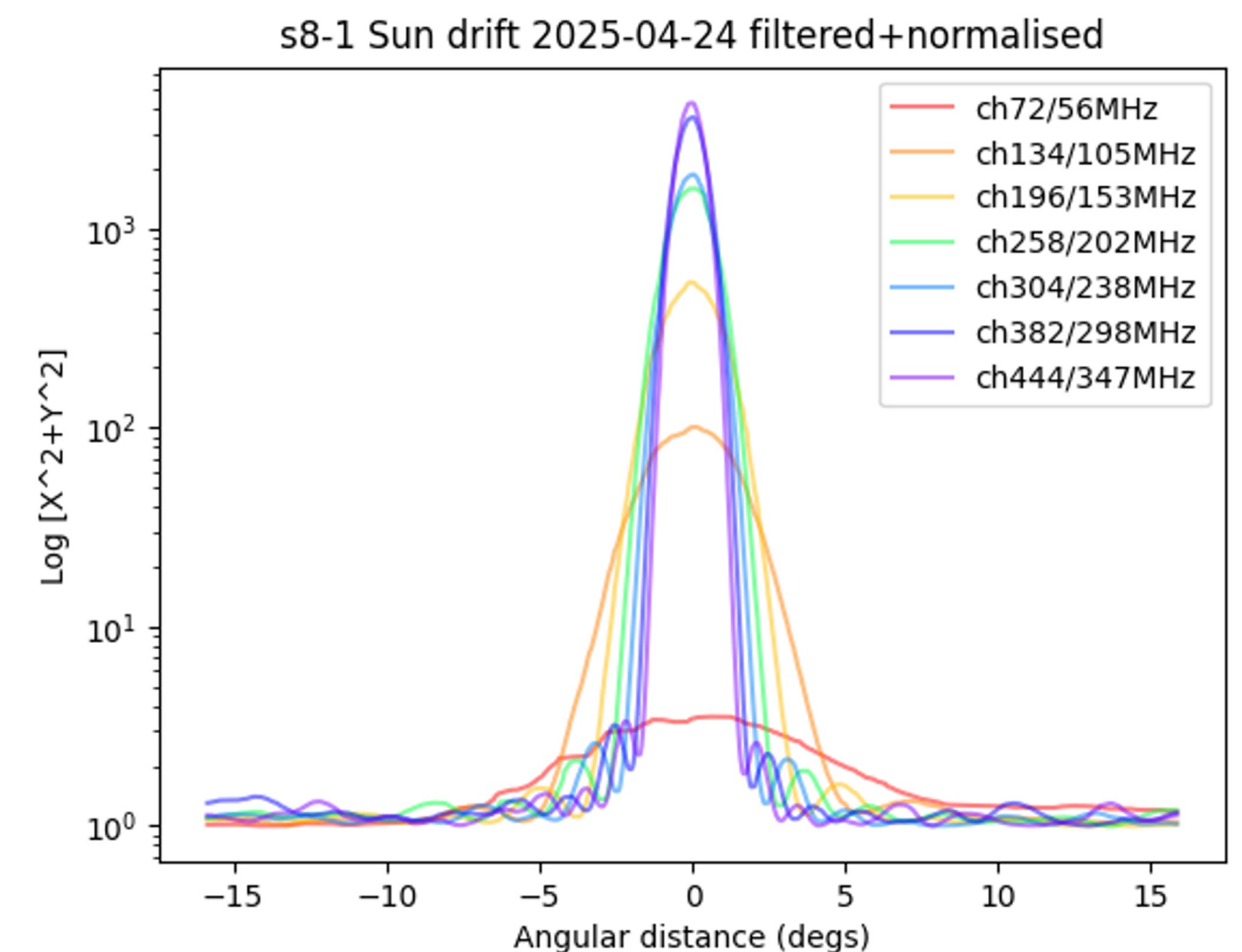
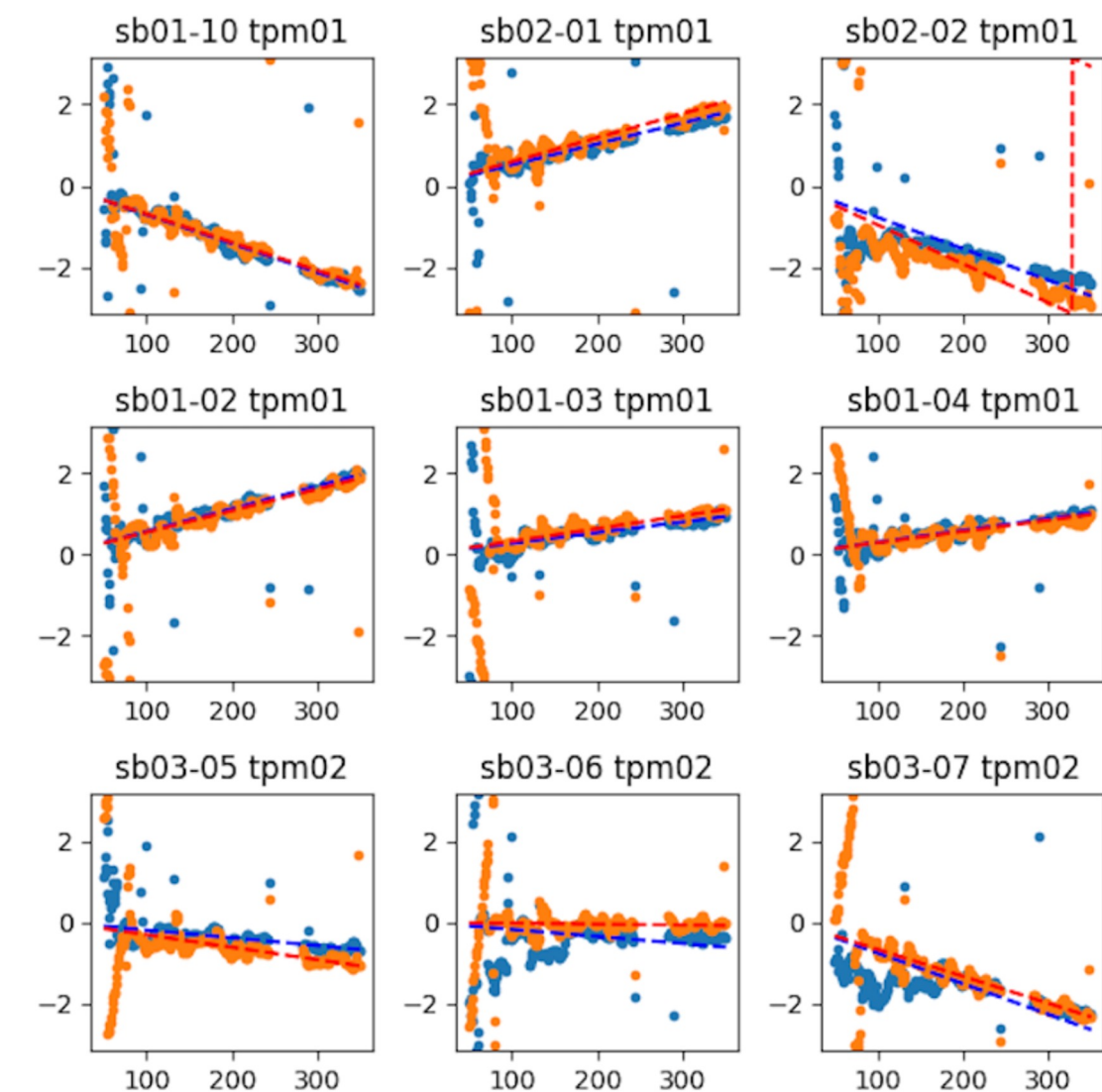


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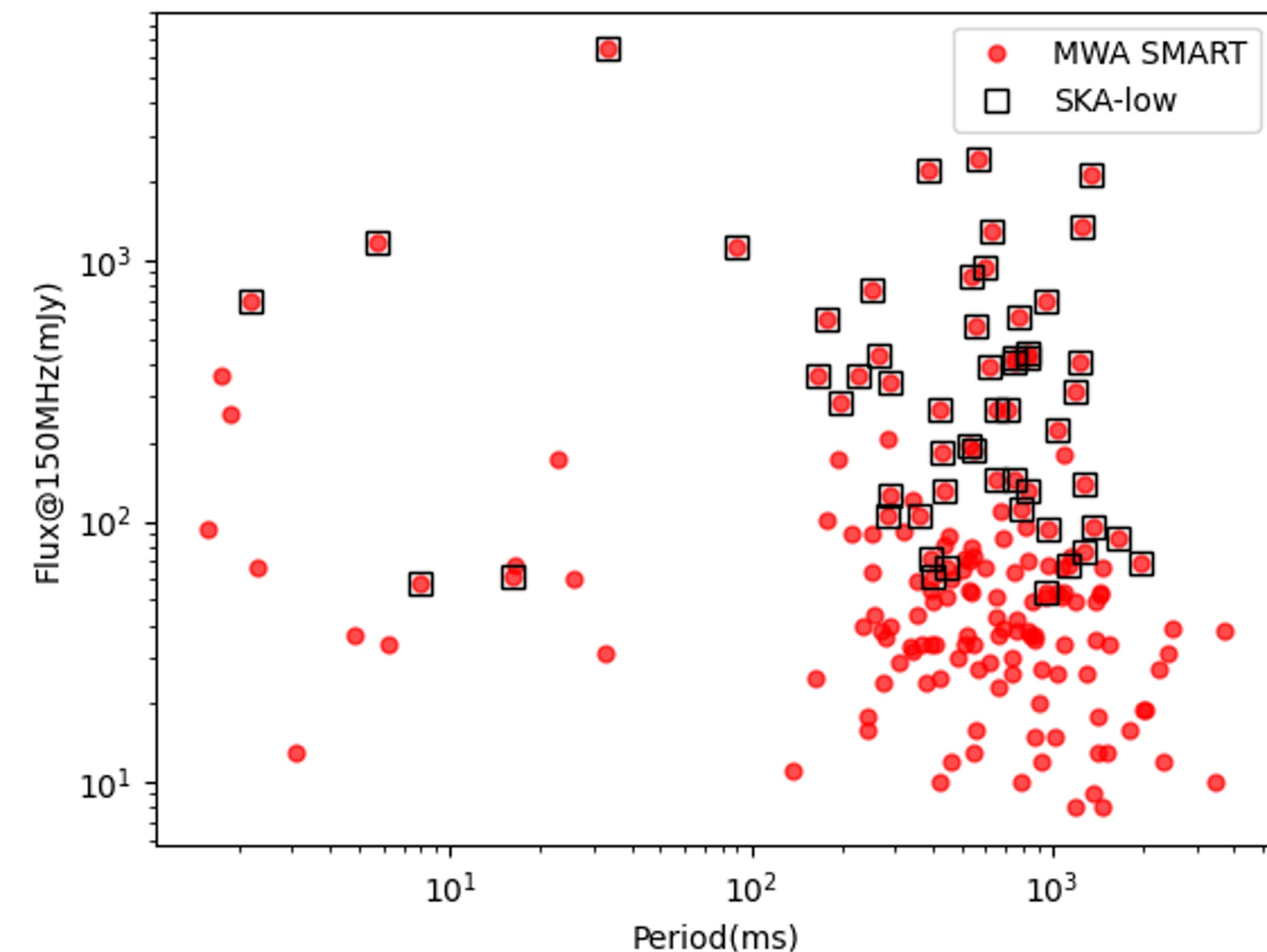
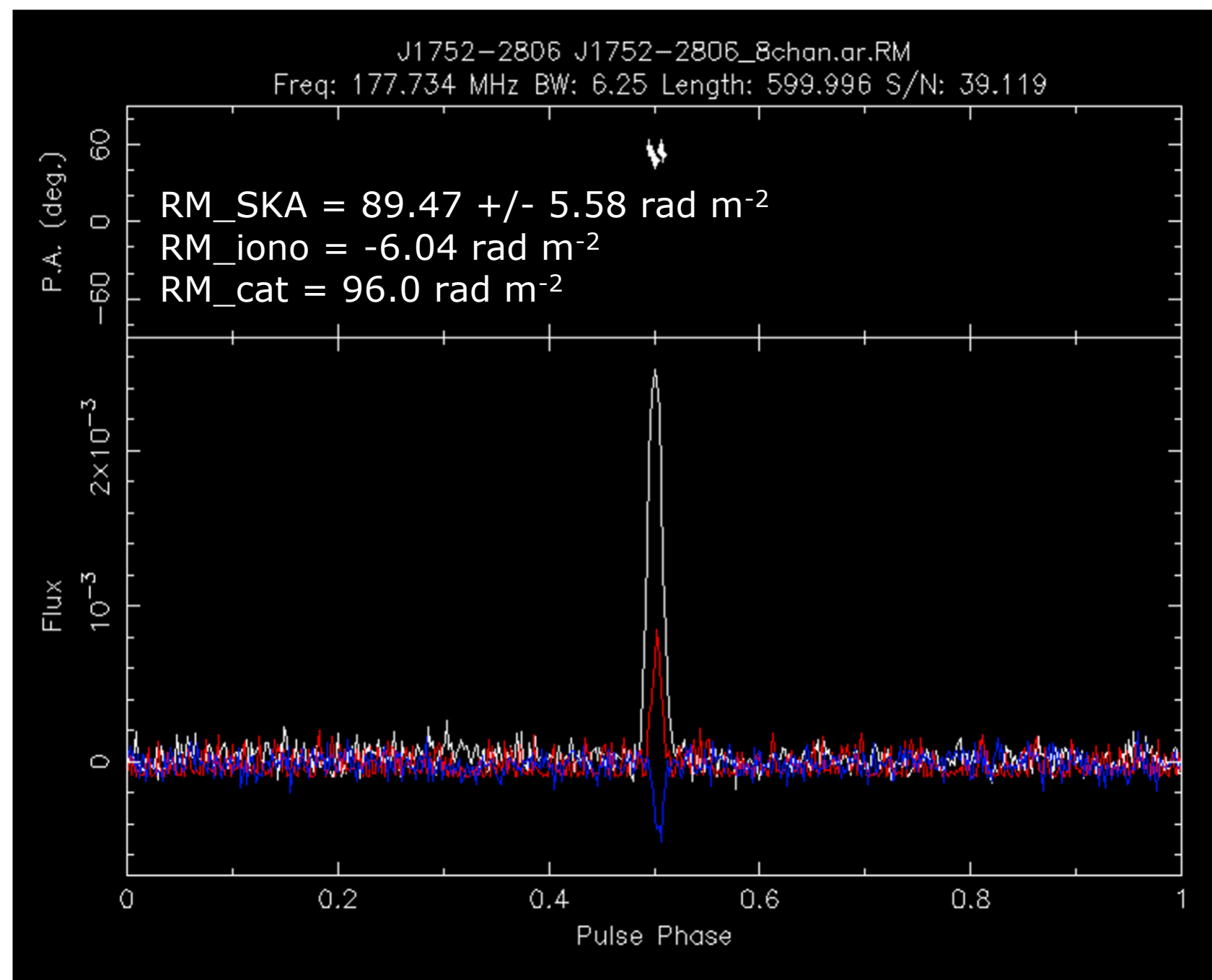
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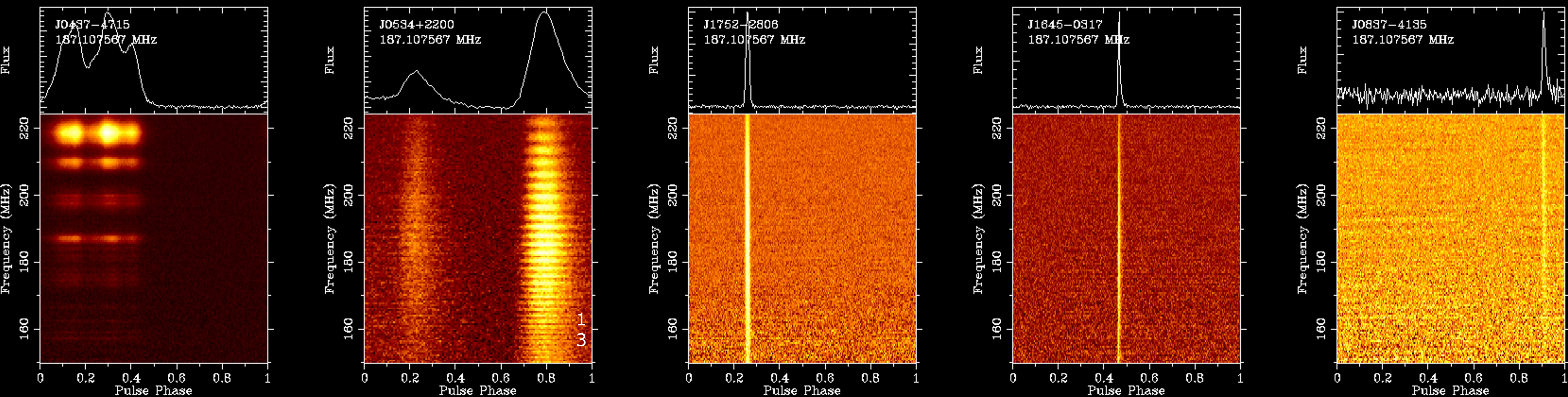


# Single station: pulsar detections

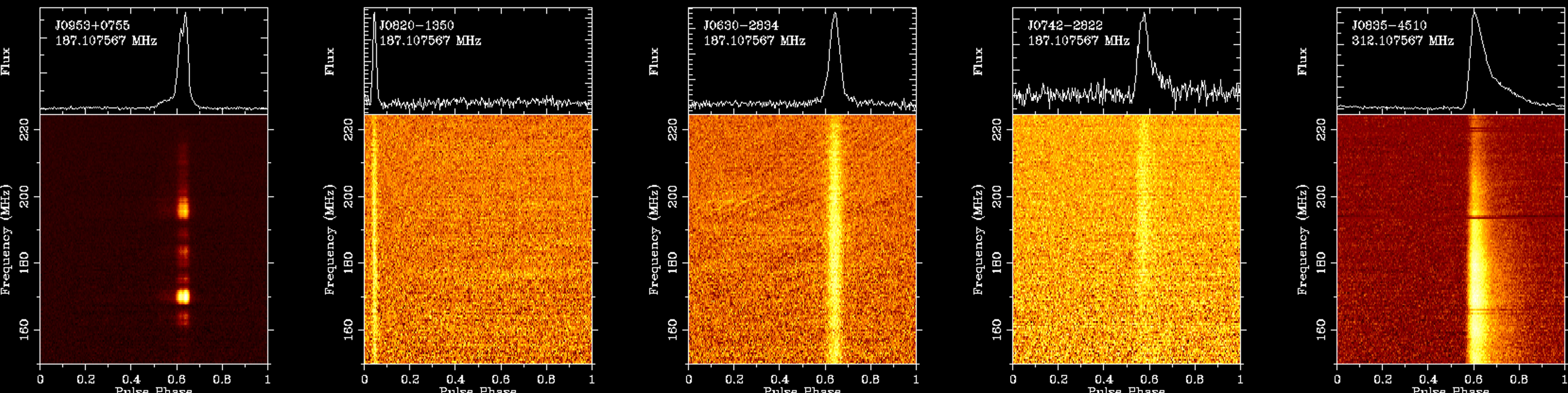
- Detected **56** known pulsars using *single* stations raw voltages (6.25 MHz bandwidth) and Pulsar Timing Subsystem (up to 75 MHz)
- Sneak peek at full polarization properties, with promising results
- Initial testing of tied-array performance is very encouraging - realtime calibration loop in development now, so full coherence is coming soon







**Testing tied array observation (2 stations), 75 MHz: excellent data quality !!**





# Basic array performance

## Commissioning outcomes so far:

- Interferometric fringes detected
- Per-station delays determined to be stable and match fibre lengths
- Closure phase verified (indicating that visibilities can be calibrated)
- Station sensitivities measured
- First images produced
- Station bandpasses determined

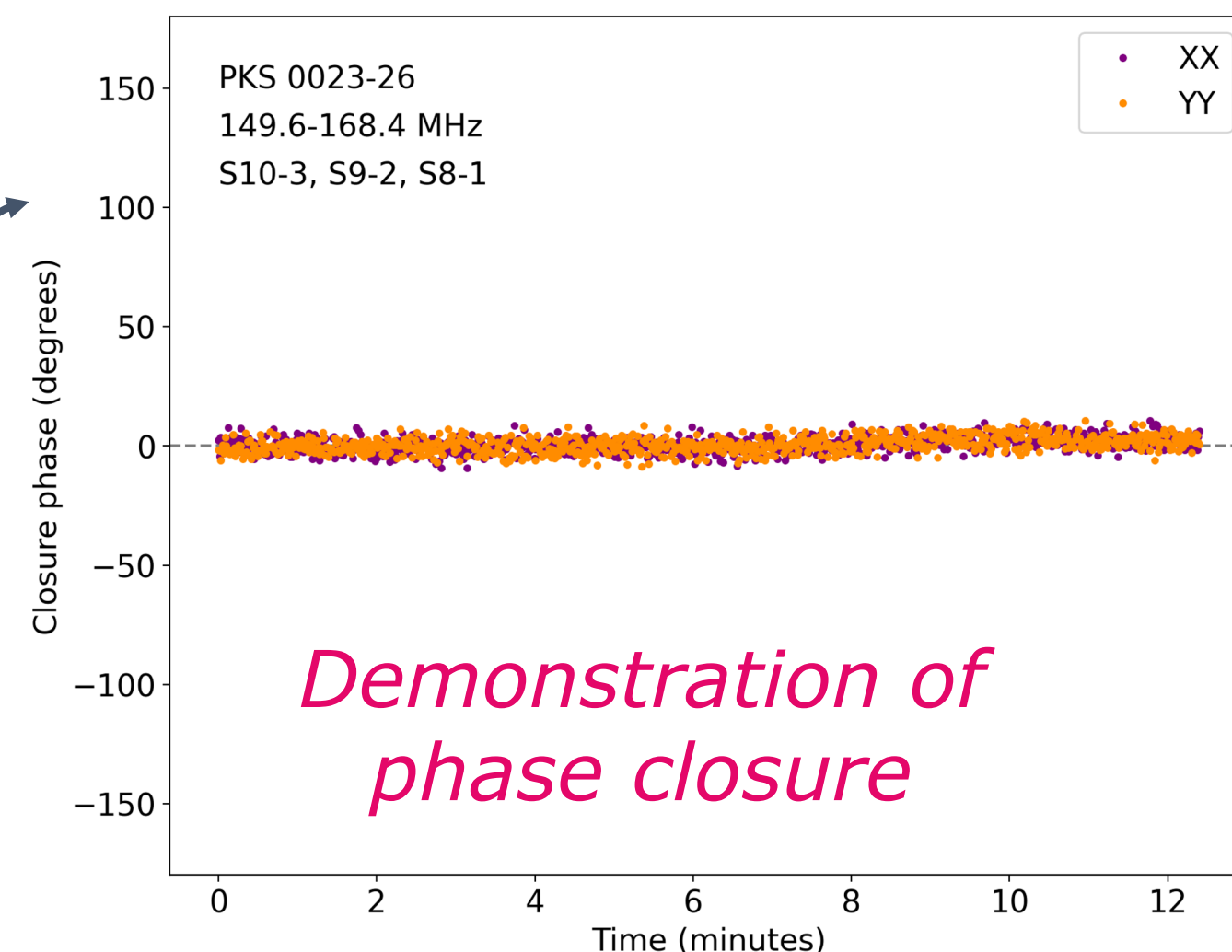




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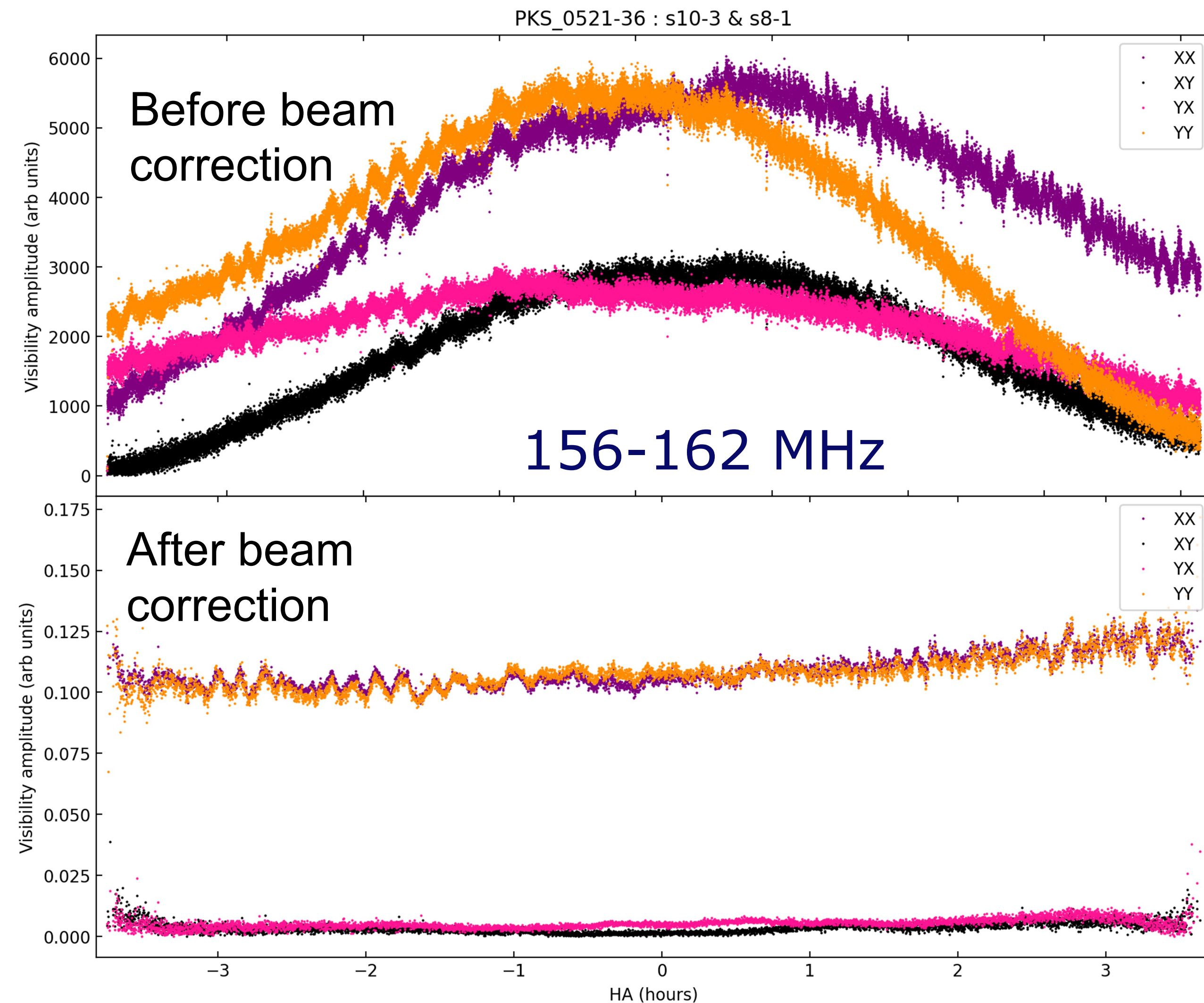
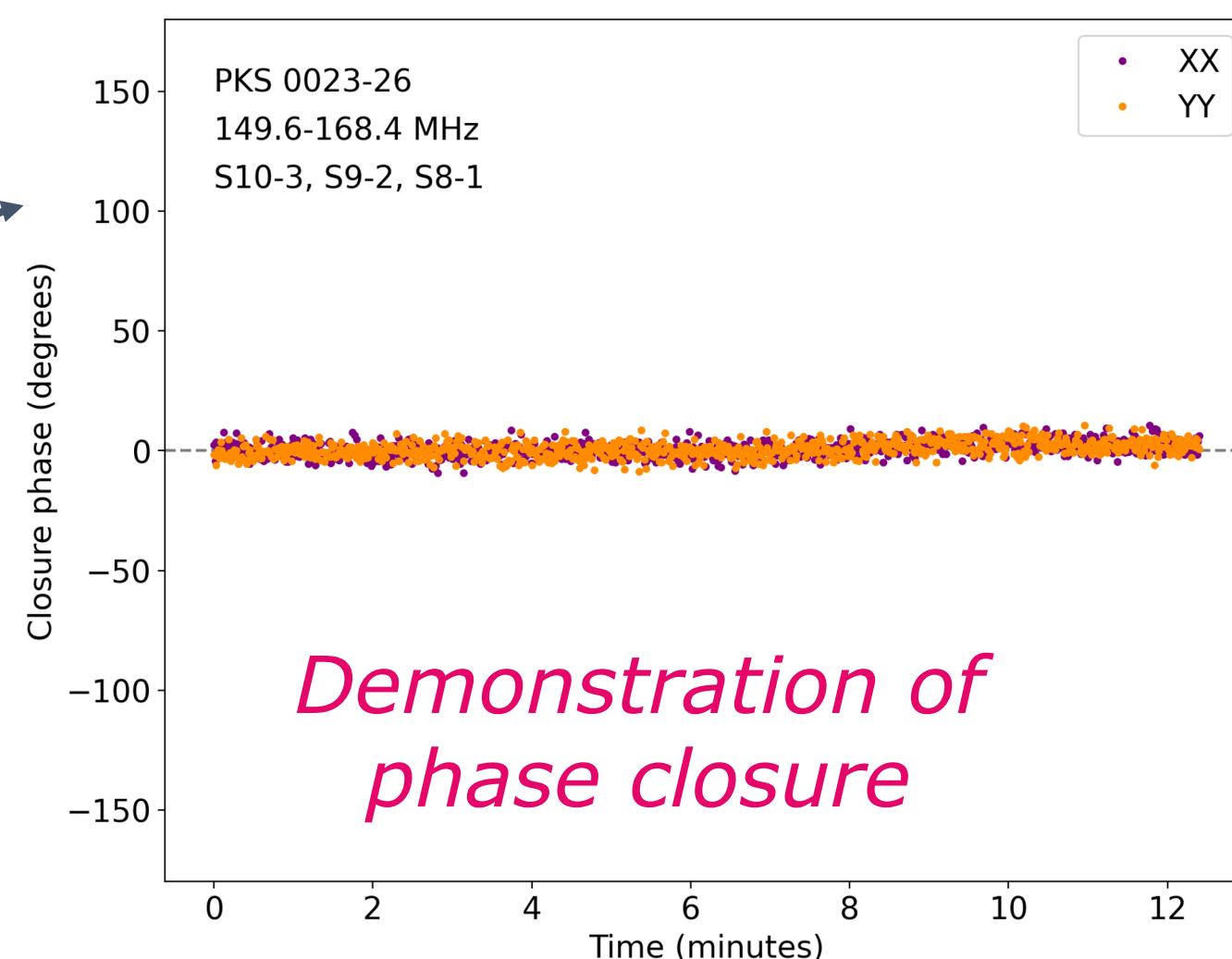




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*Unflagged(!), delay calibrated & averaged visibilities*

*Beam correction compensates for antenna response and relative station rotations*





# First SKA-Low AA0.5 image!

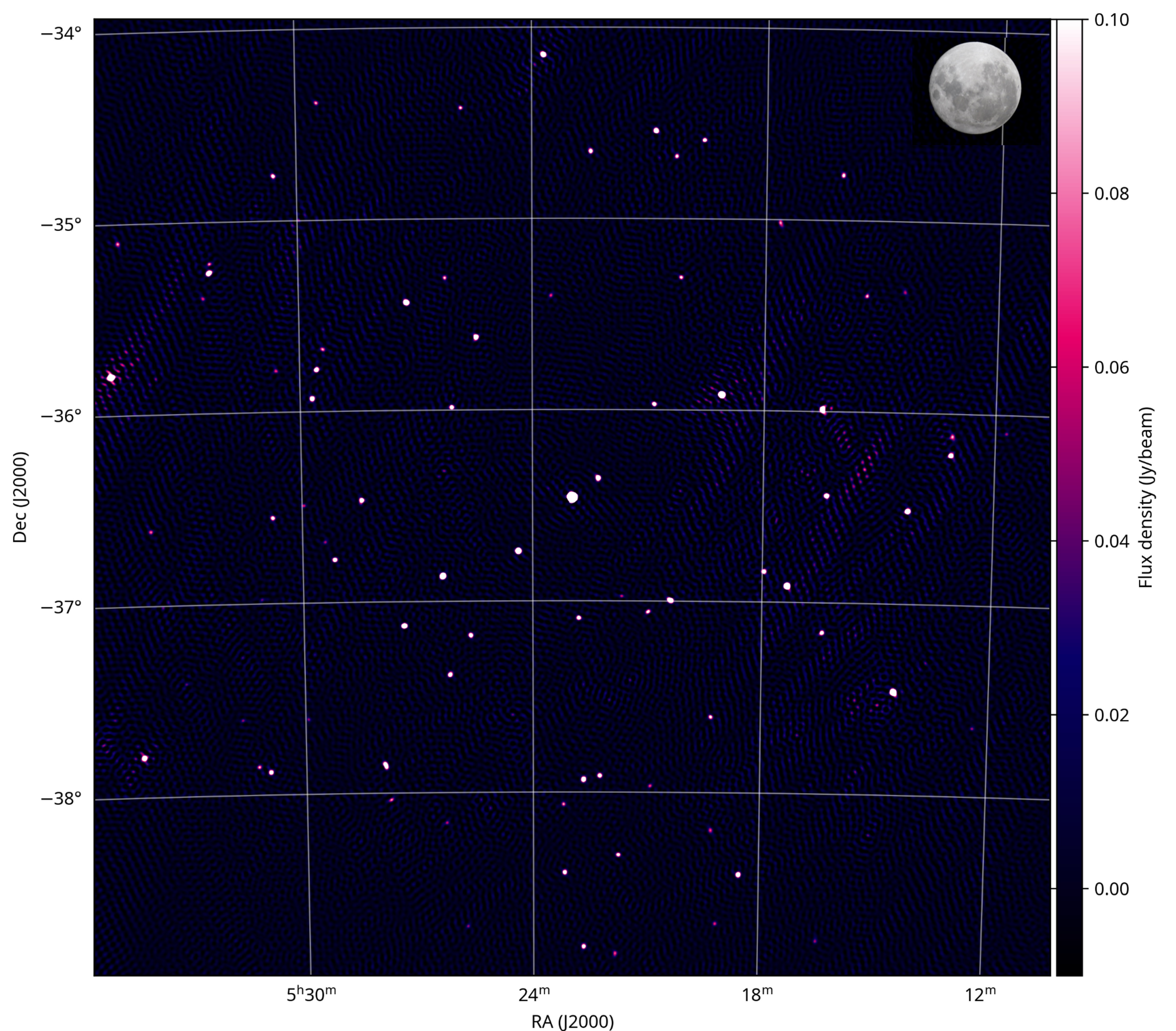
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- 150-175 MHz BW
- $\sim 7$ h effective int.





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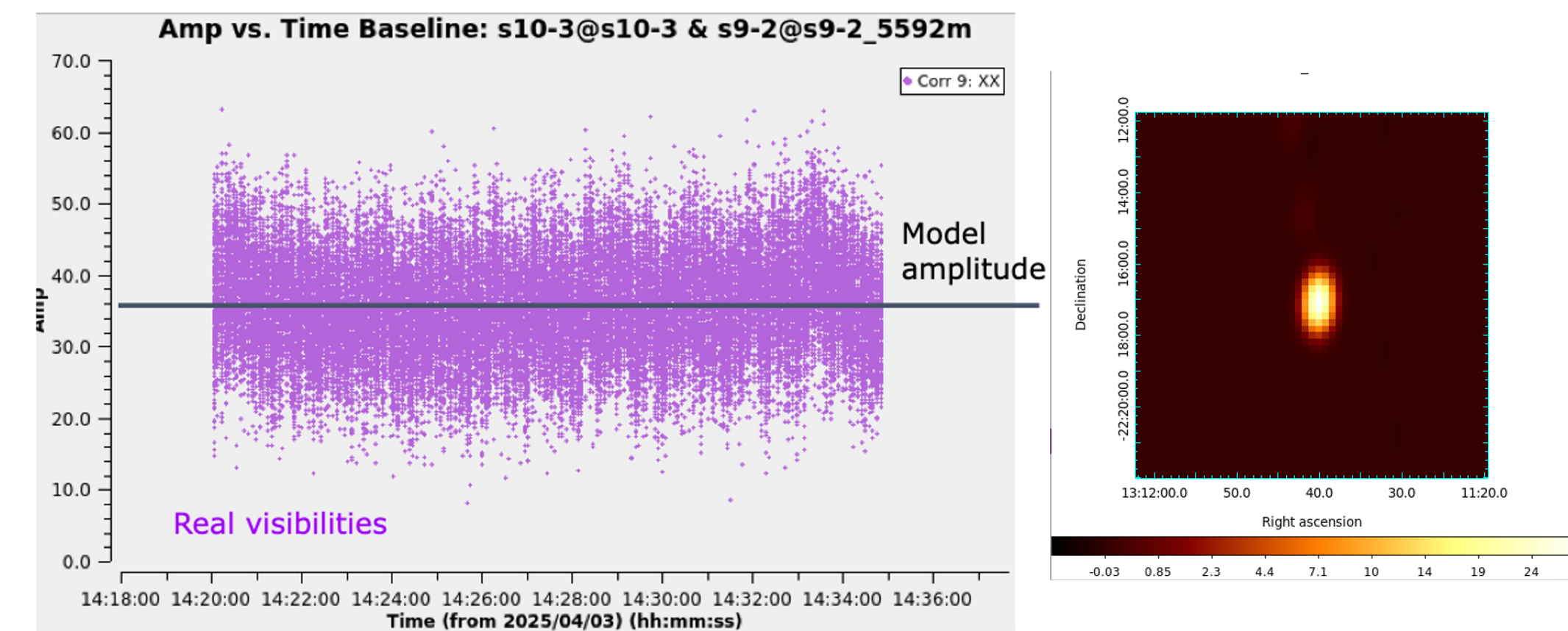
- Field surrounding PKS 0521-36 ( $\sim 56$  Jy at 150 MHz)
- 150-175 MHz BW
- $\sim 7$ h effective int.
- Image noise:  $\sim 5$ -6 mJy/beam at  $70''$  final resolution
- 85 sources detected
- Very good match to GLEAM-X catalog (with mostly unguided self-cal!)





# SKA-Low station SEFD

SEFD estimates from rms noise of point source calibrated visibilities

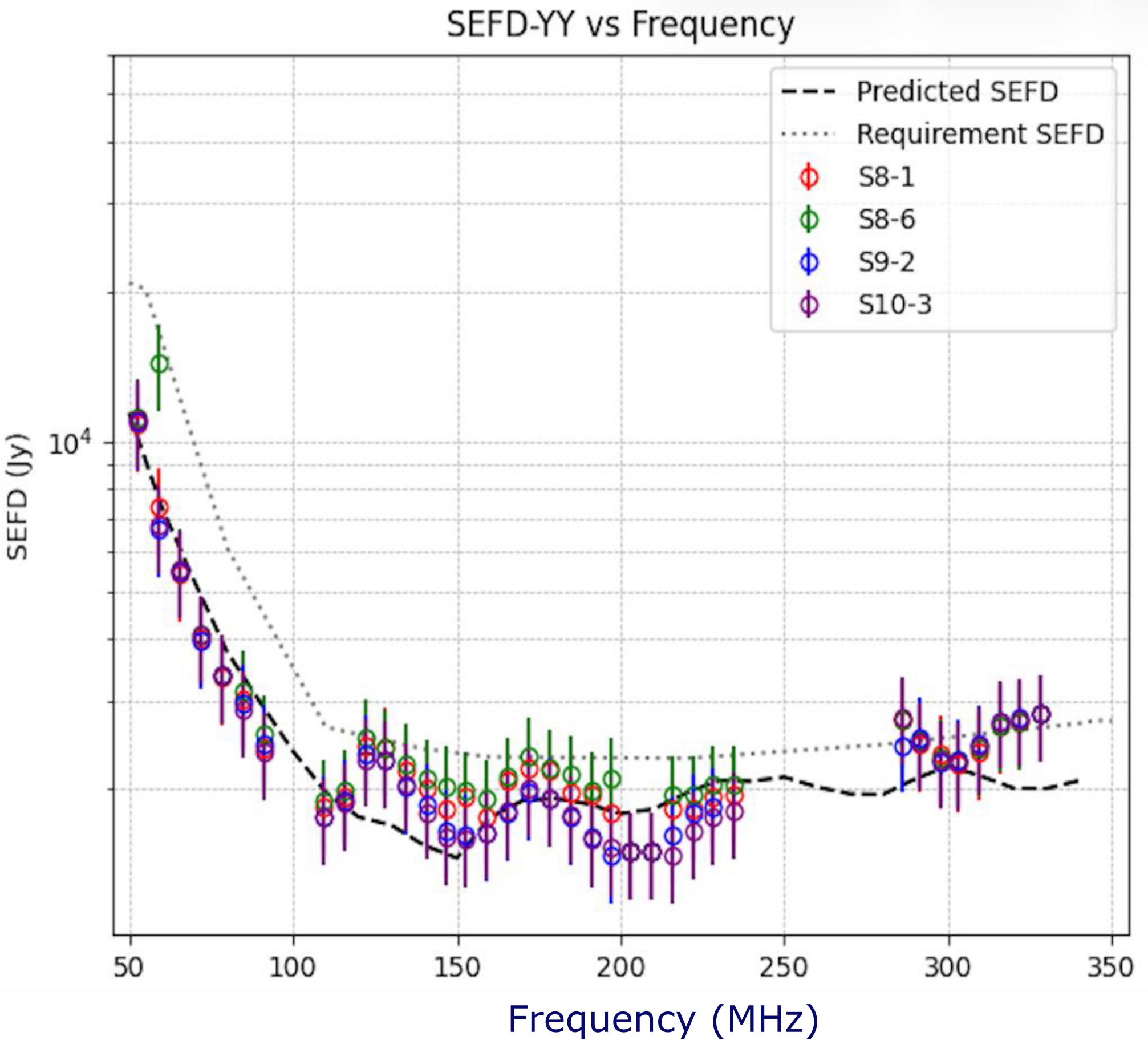
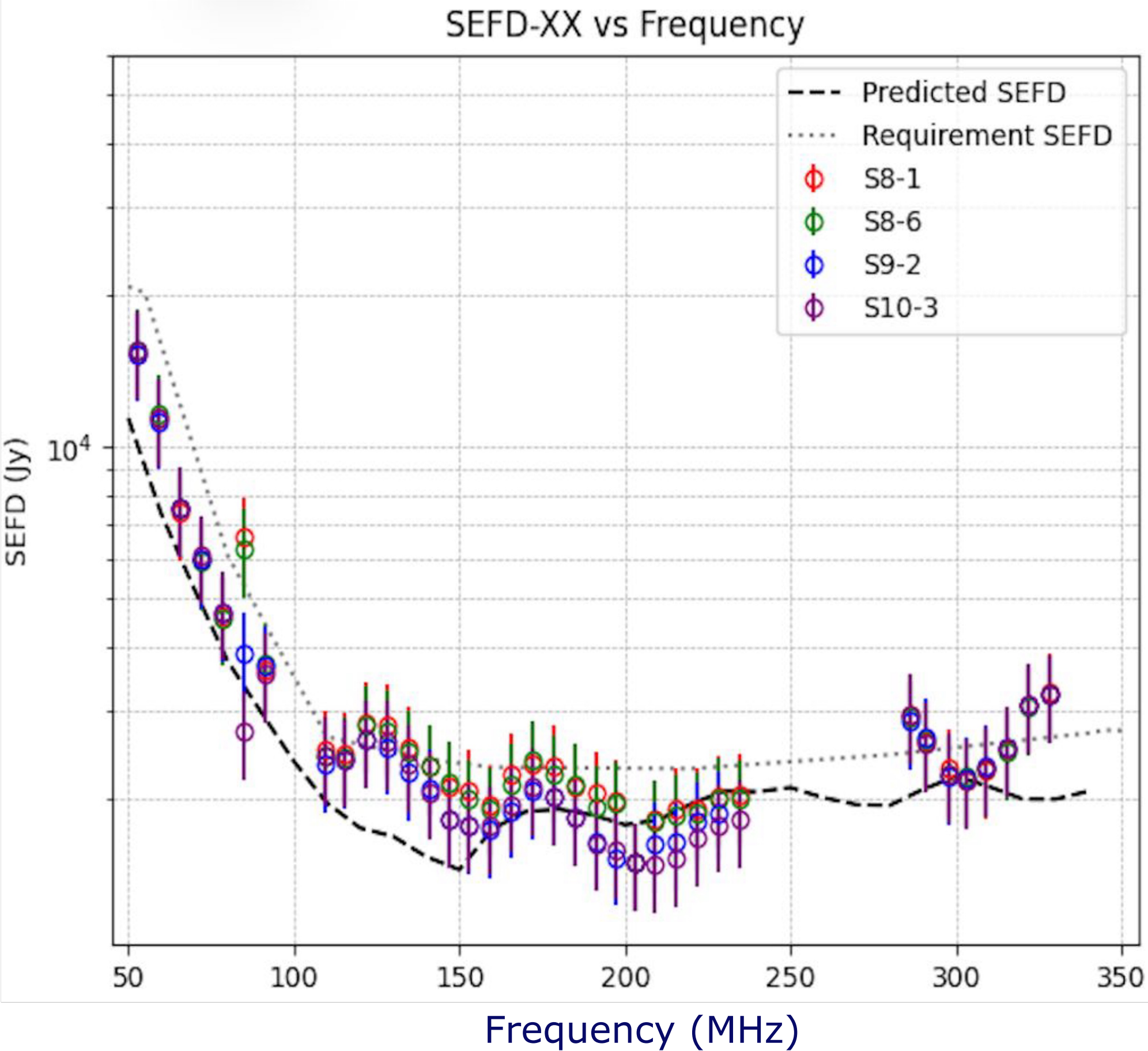
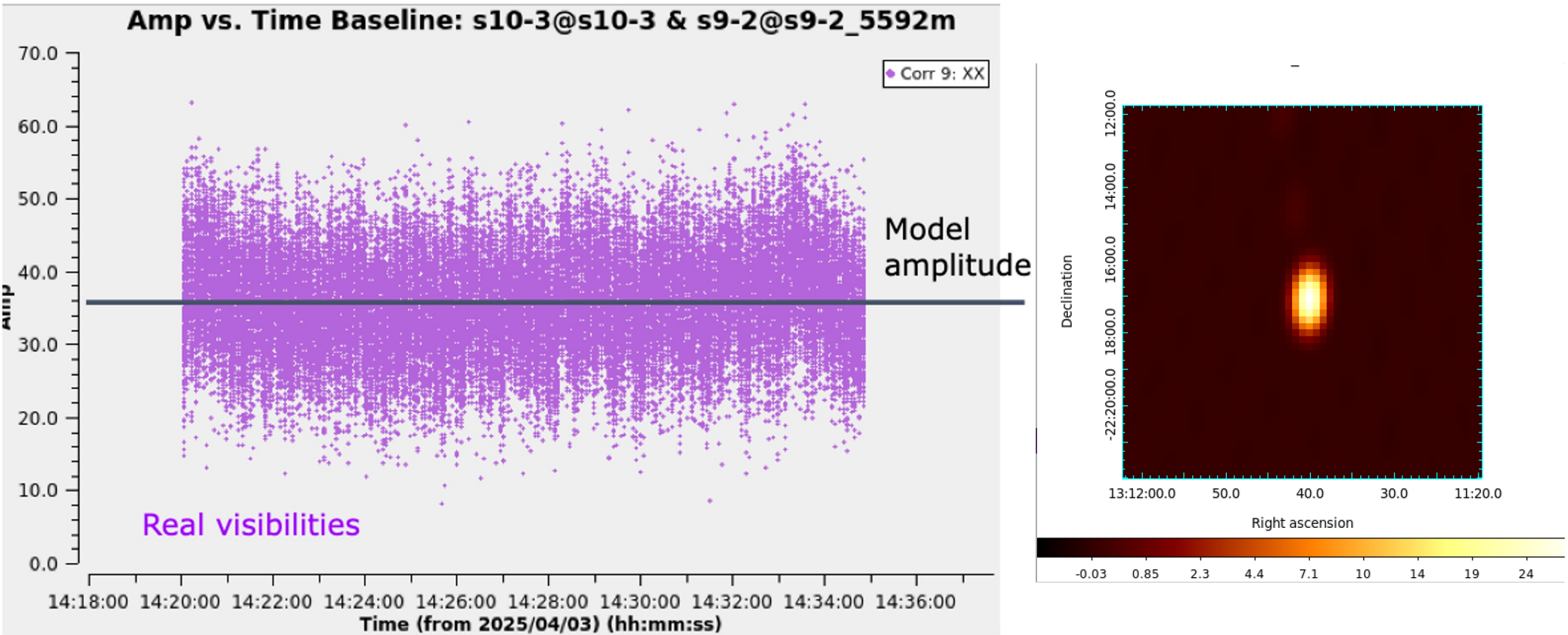




# SKA-Low station SEFD

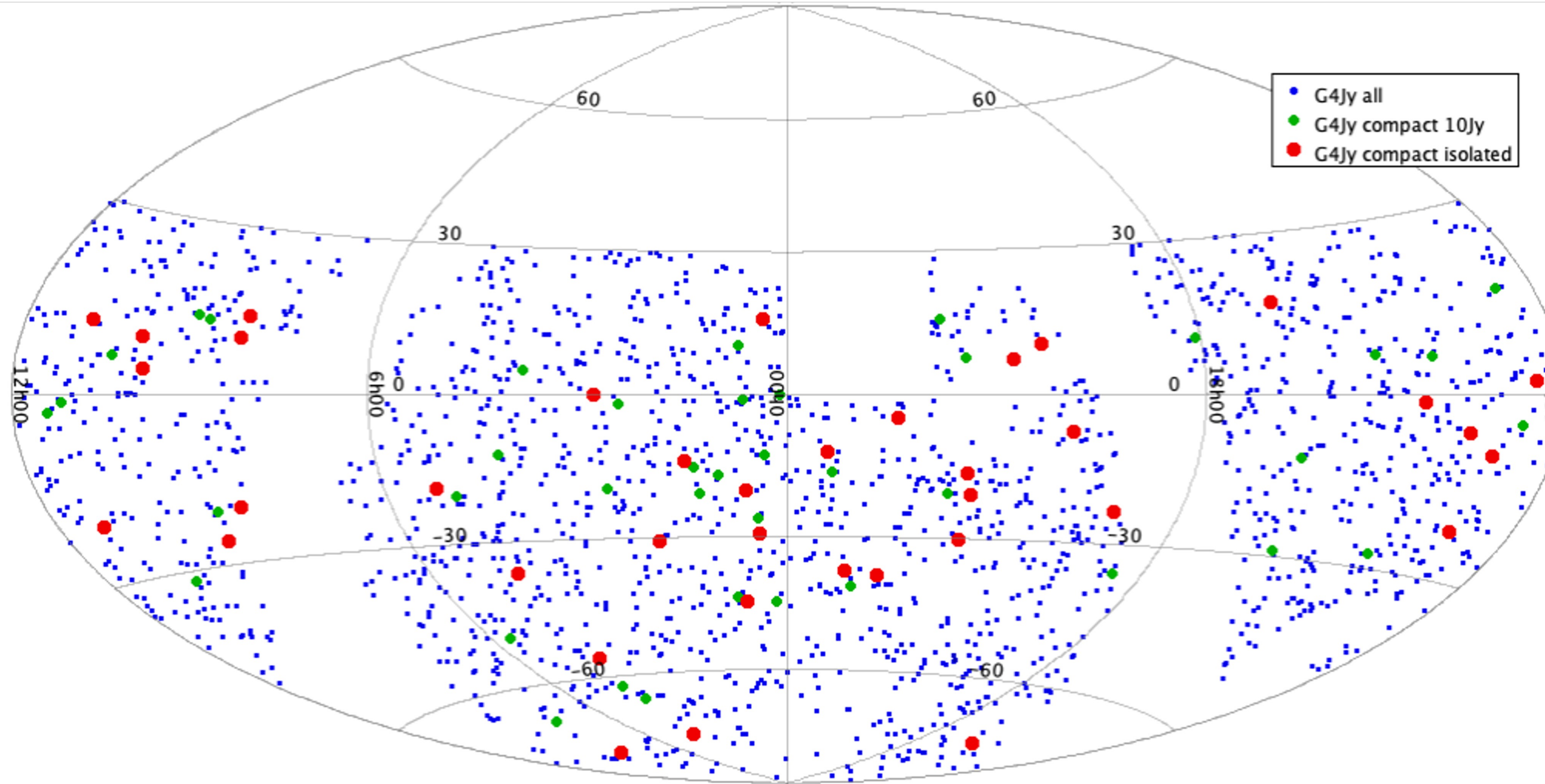
SEFD estimates from rms noise of point source calibrated visibilities

**Lower SEFD = better sensitivity!**





# SKA-Low Calibrator Database (ongoing effort)



**38** compact isolated sources identified from existing catalogs (GLEAM, TGSS...)

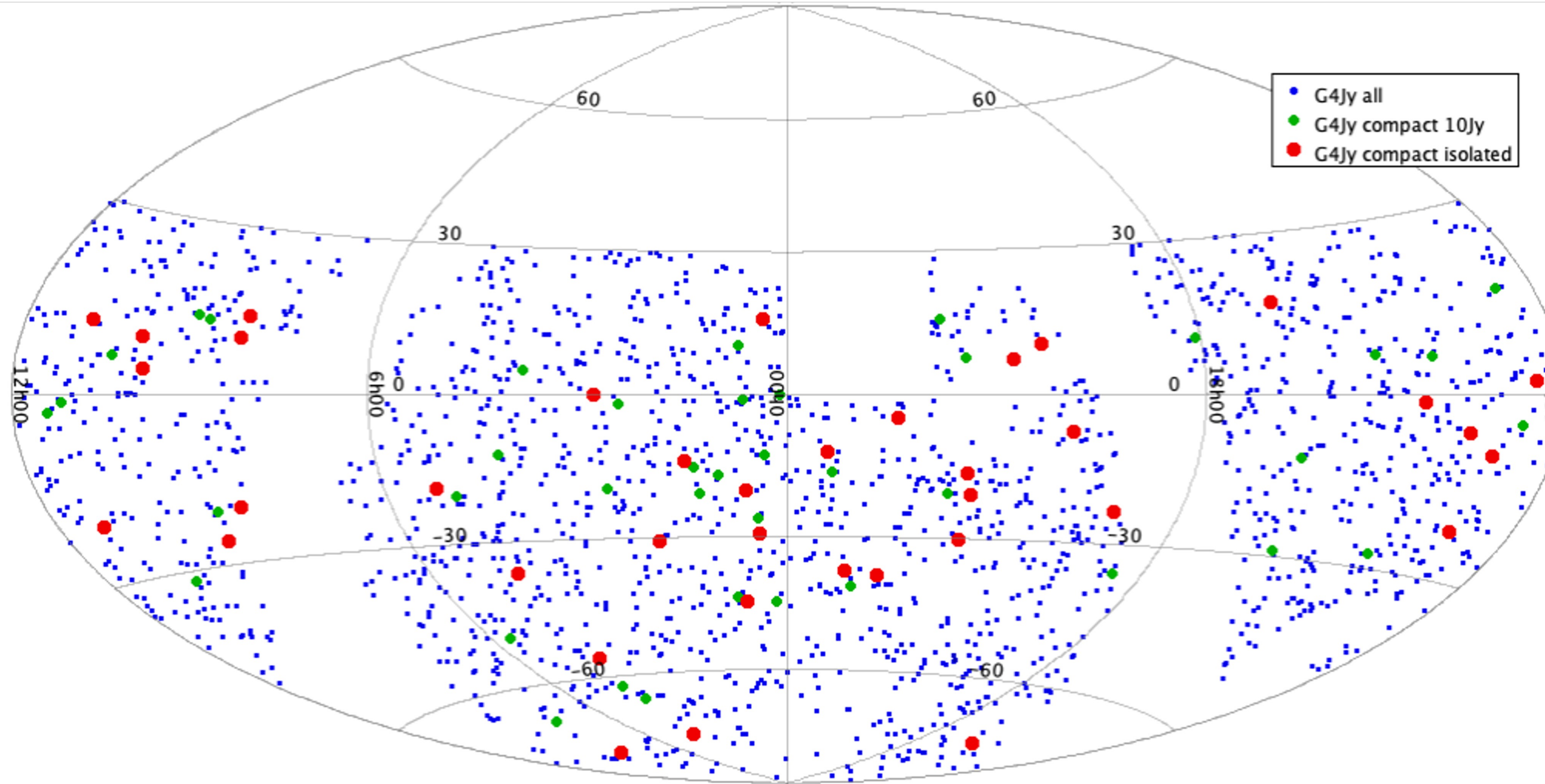
## Current status

- 23/38 observed across full 50-350 MHz
- Q&A plots of raw visibilities to support Operators
- Bandpass calibration tests for *all* sources (150-200 MHz only)
- Full bandwidth calibration tests
- Sky model tests with single point source, also with field sources
- Comparing legacy software (Miriad/CASA) and AO Calibrate
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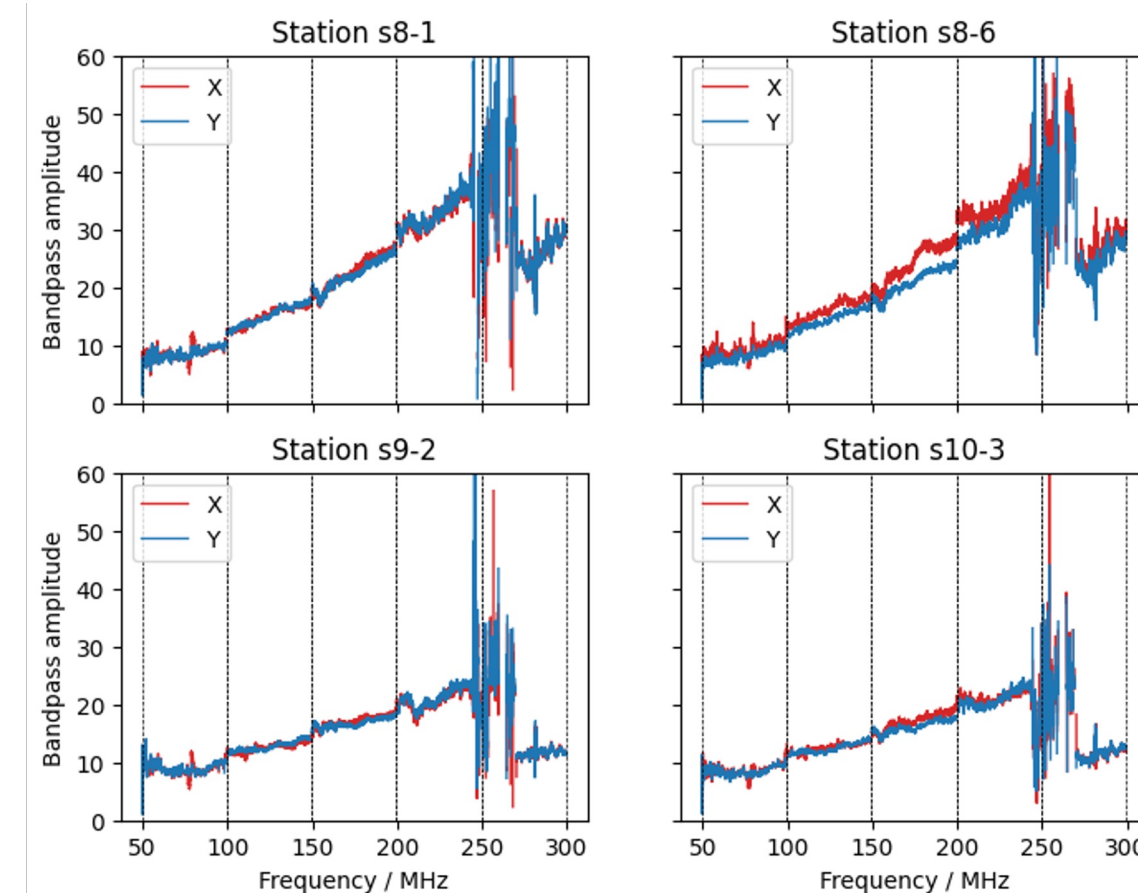
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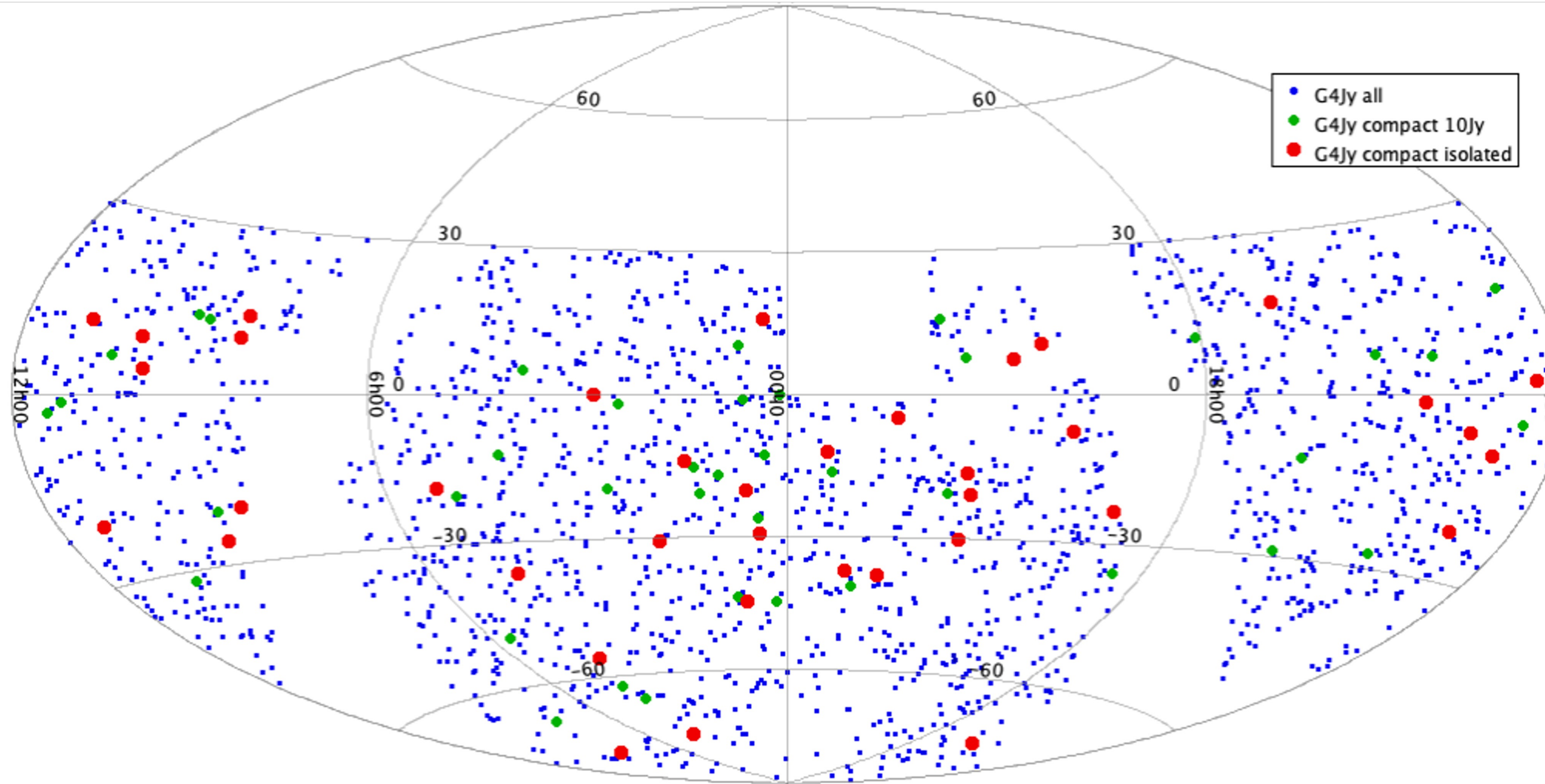
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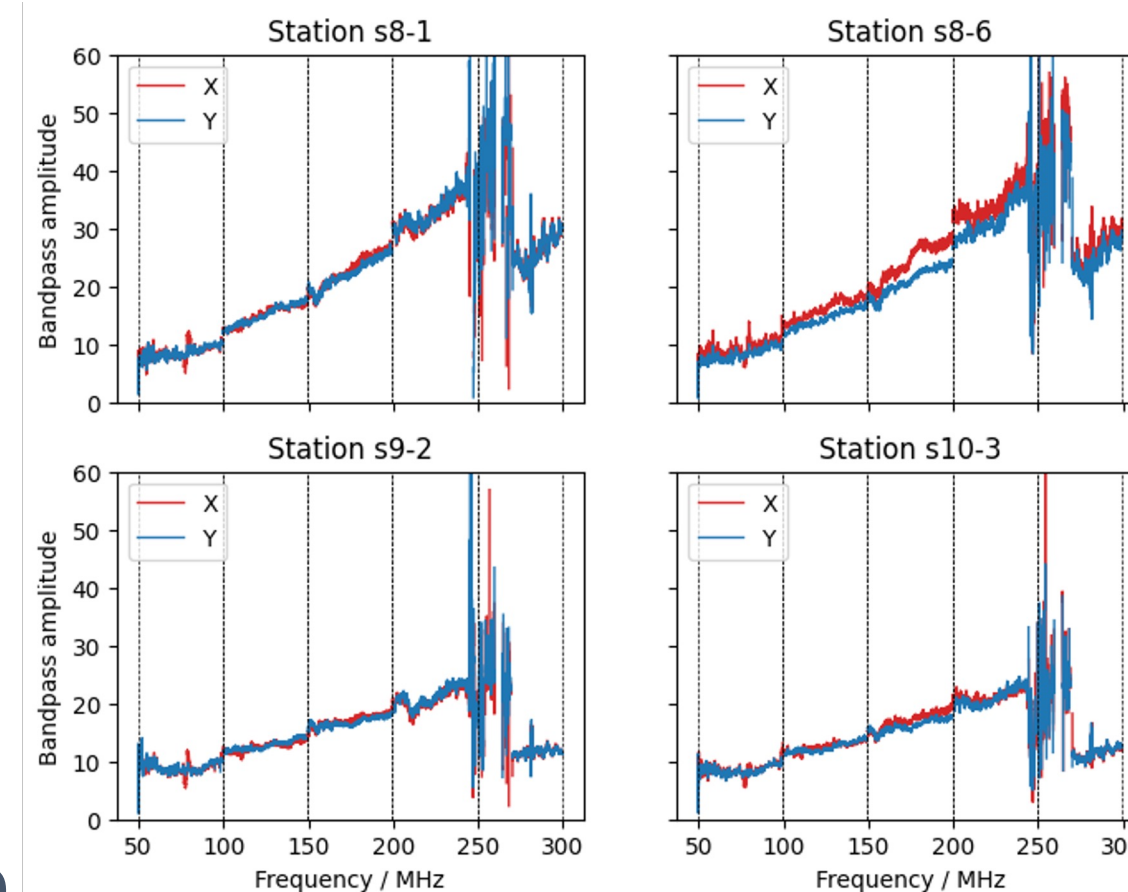
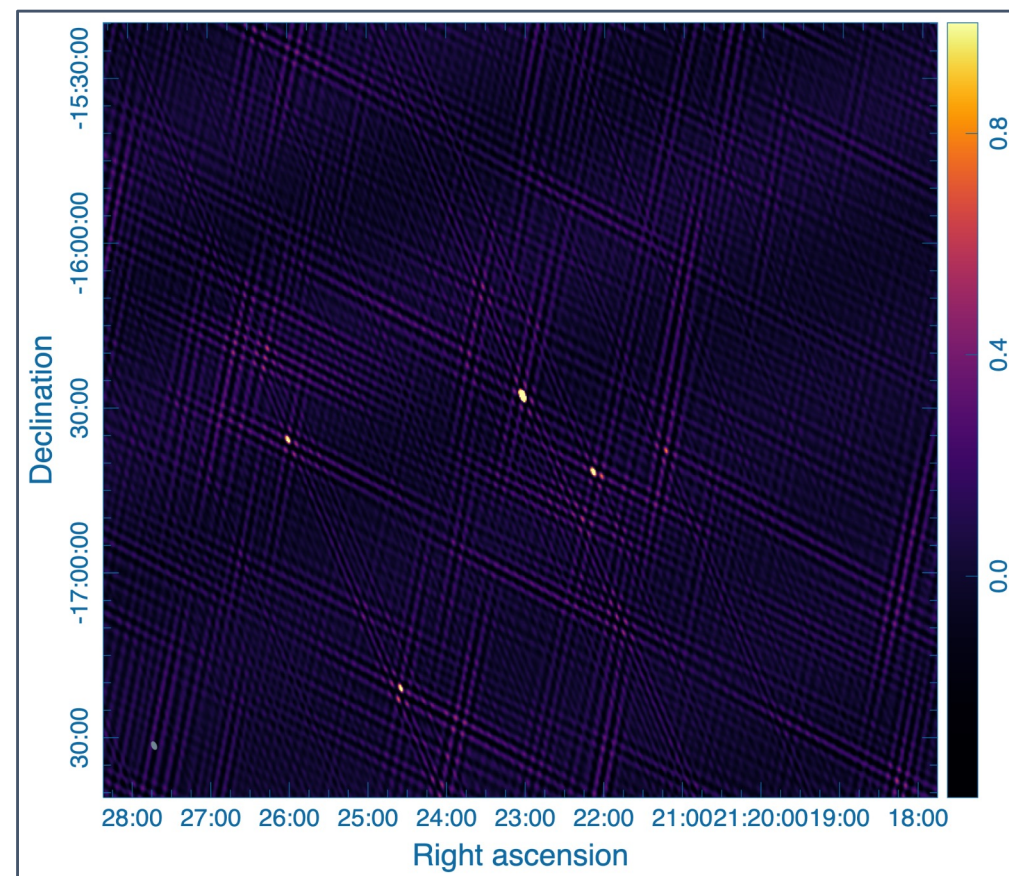
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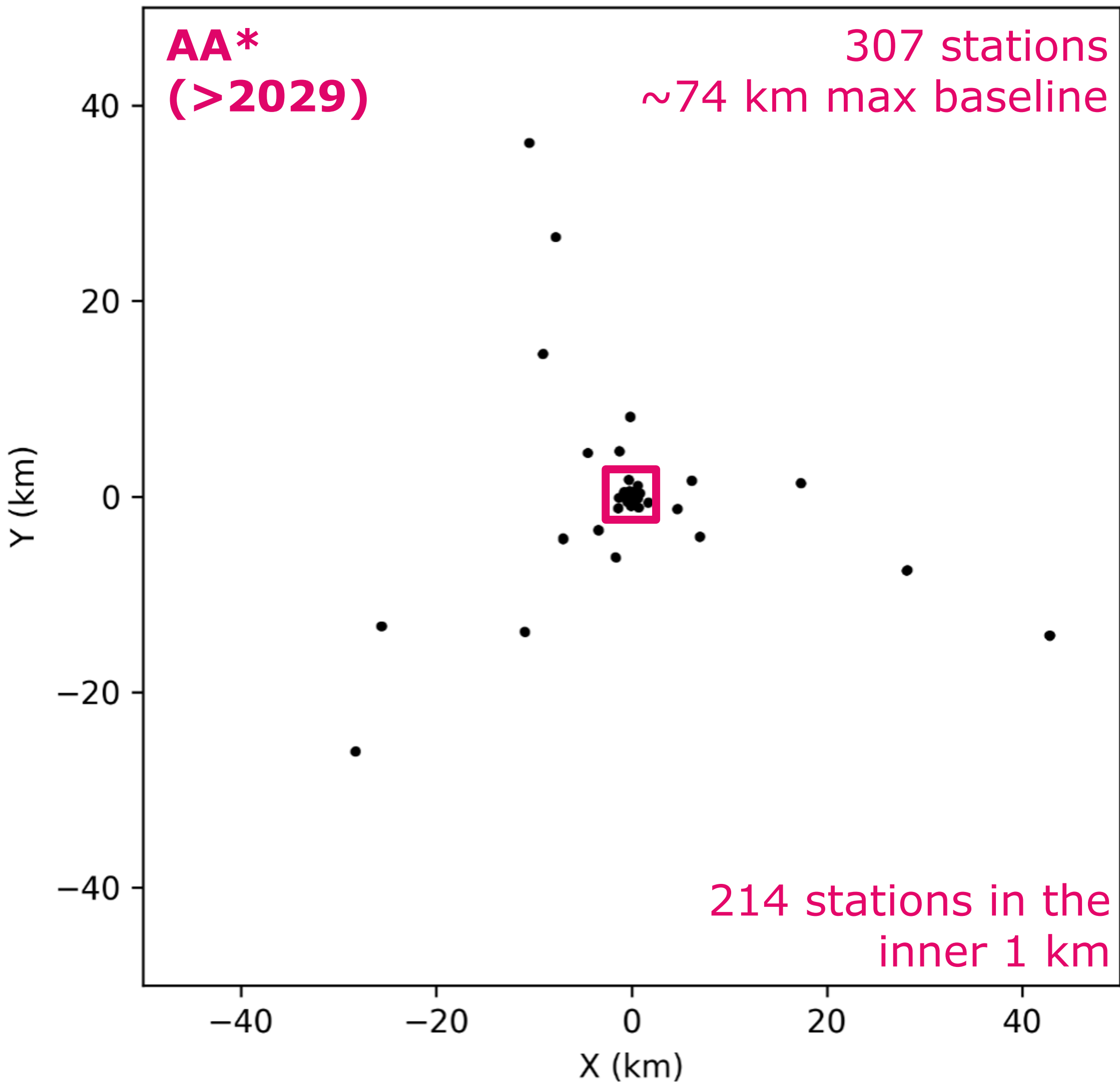
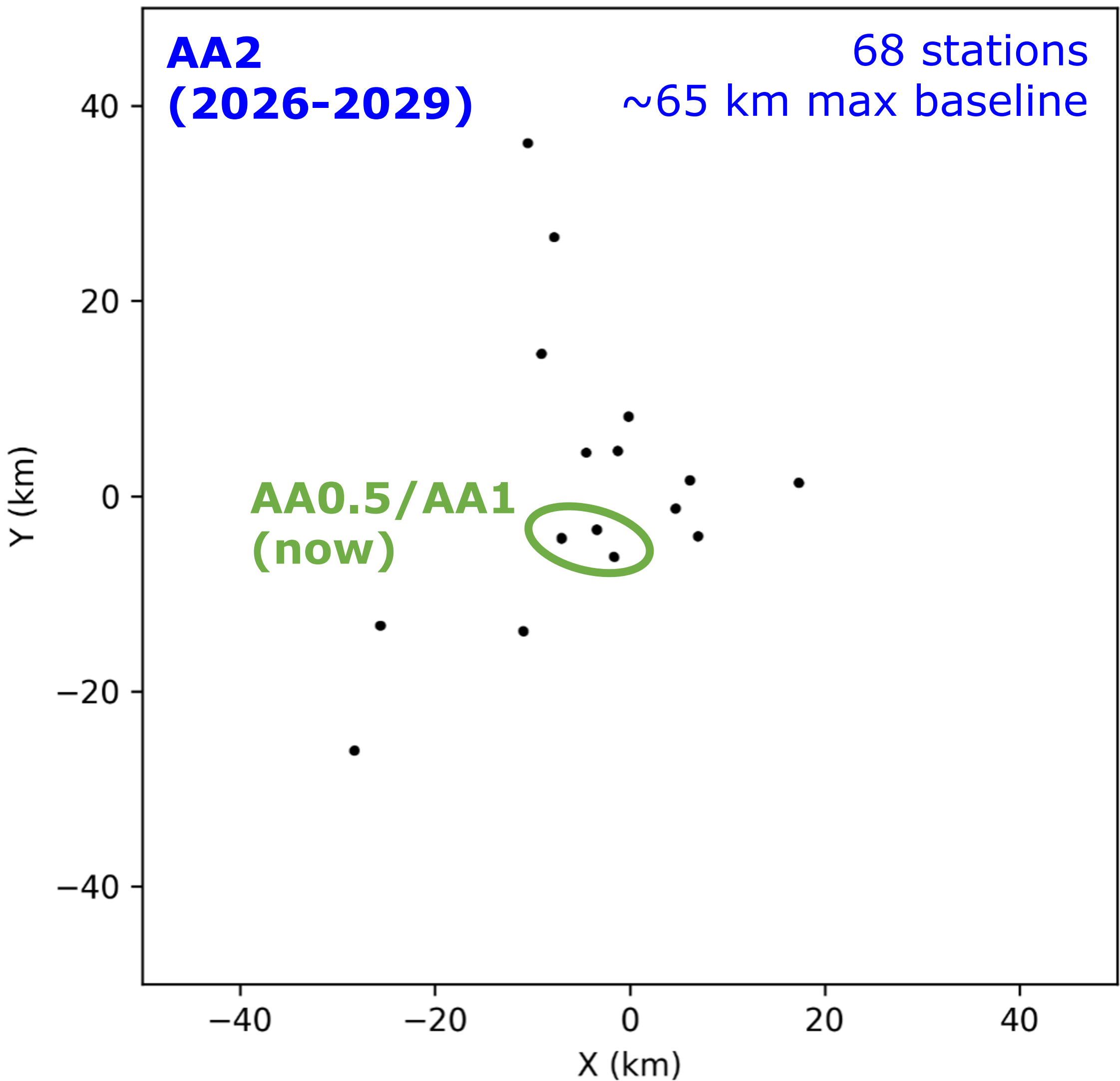


*Stage 1 of SKA-Low **Global Sky Model (GSM)** to kick off the **Science Data Processor** pipelines!*



# SKA-Low Array Assemblies

SKAO tools for  
your science!





# SKA-Low Calibrator Surveys

Progression of observational campaigns to establish calibrator database for basic array calibration, enabling Science Verification

These surveys are not designed for science, but resulting Observatory Level Data Products expected to be made public in Science Verification





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**AA0.5 / AA1**  
4 / 16 stations  
Within 3 clusters  
 $B_{\max} = 5.5$  km

- Compact point sources only
- Establish SEDs and field models
- Monitoring for variability

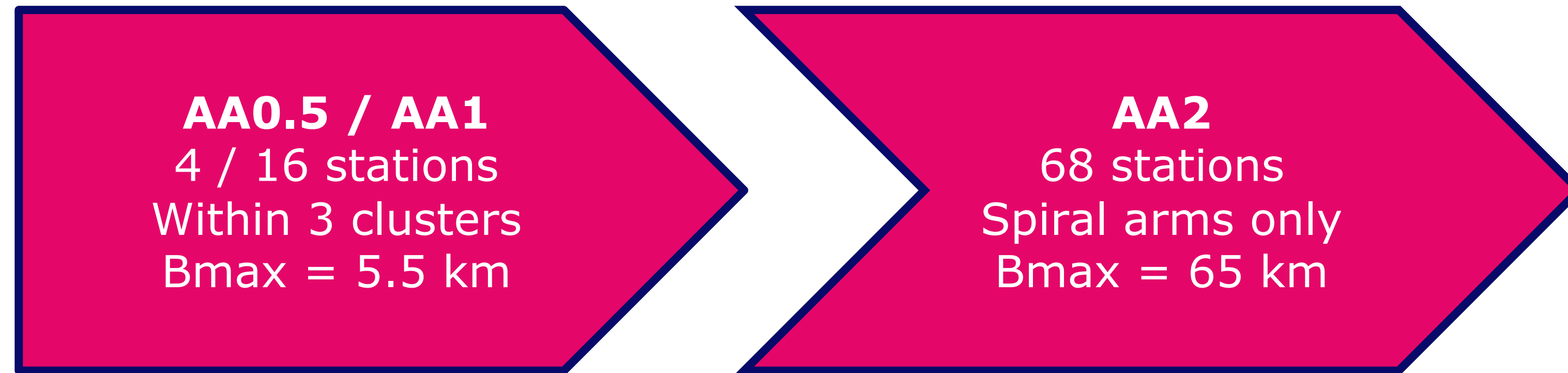




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- Compact point sources only
- Establish SEDs and field models
- Monitoring for variability

- Full sky coverage, but not expecting GSM entries in MW plane
- Focus: morphology, astrometry, flux scale
- Indicative depth: 1 mJy/beam (TBC)

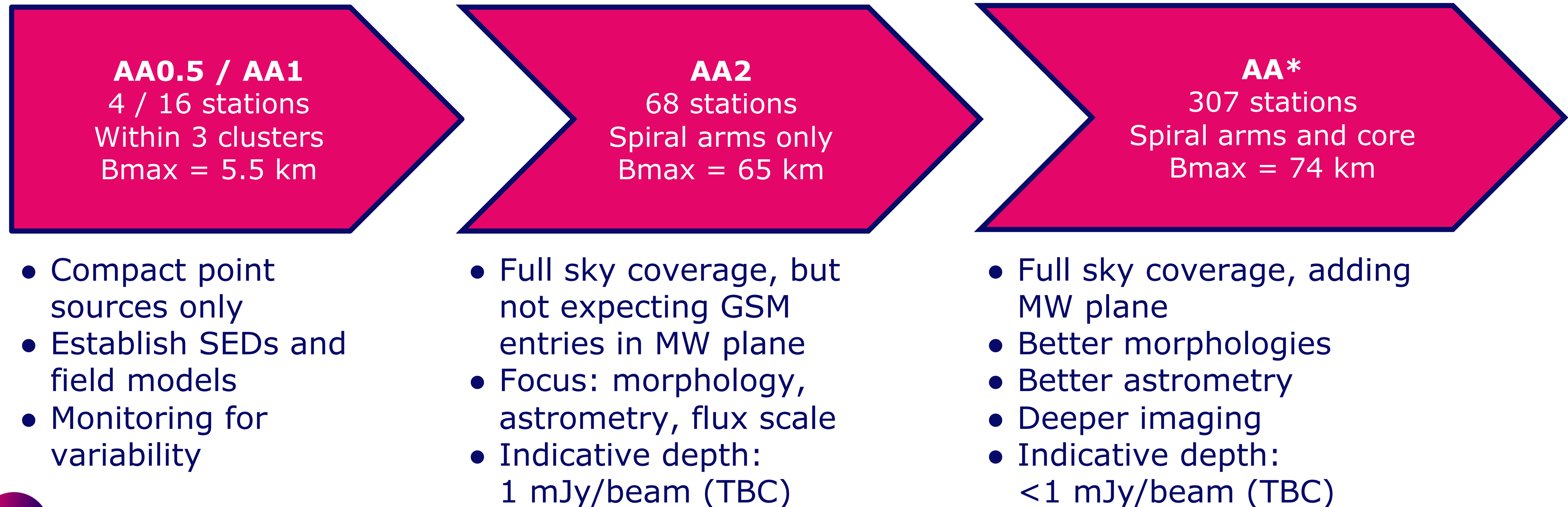




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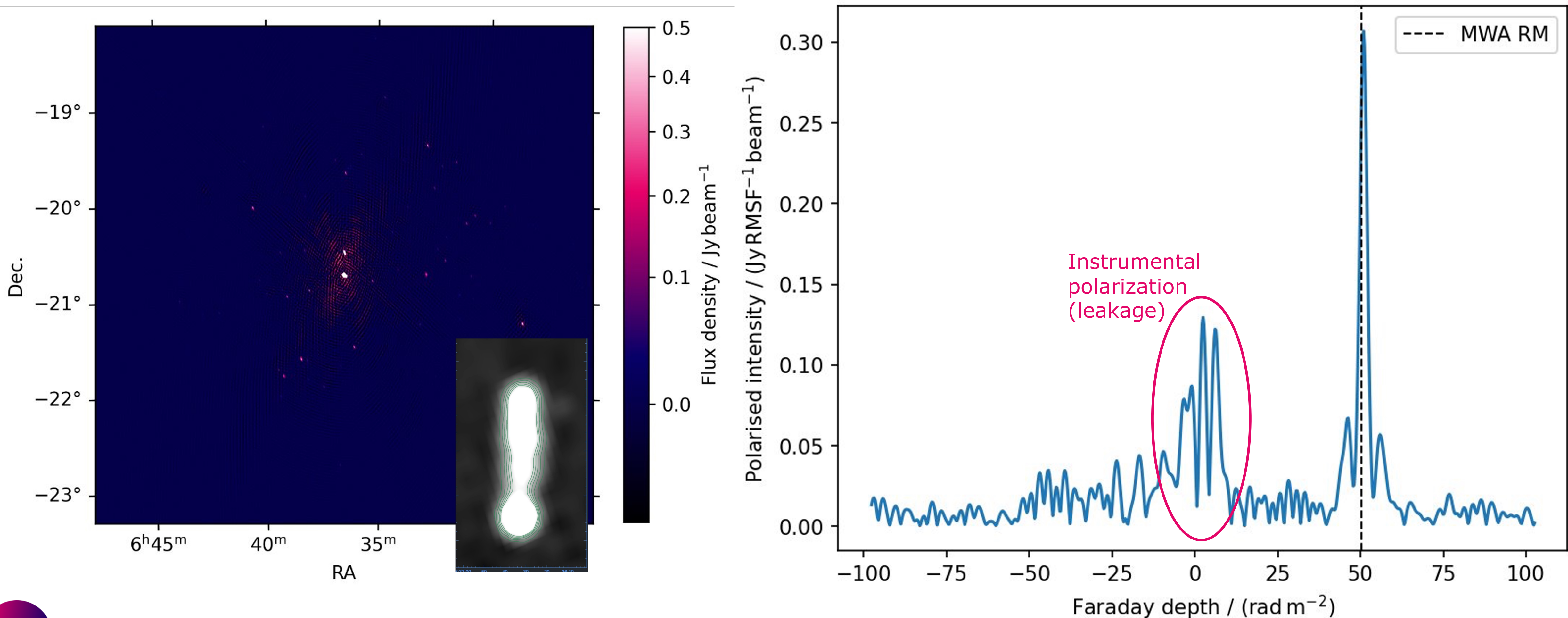
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# Polarization

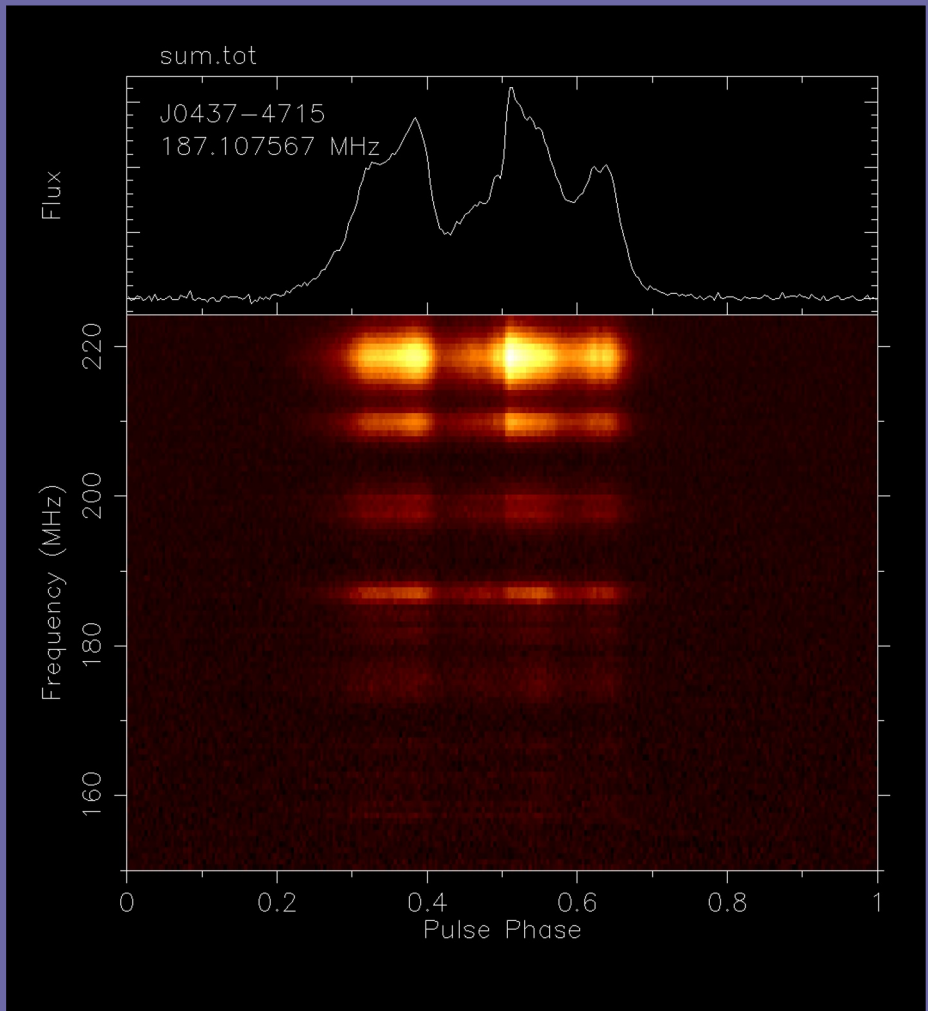
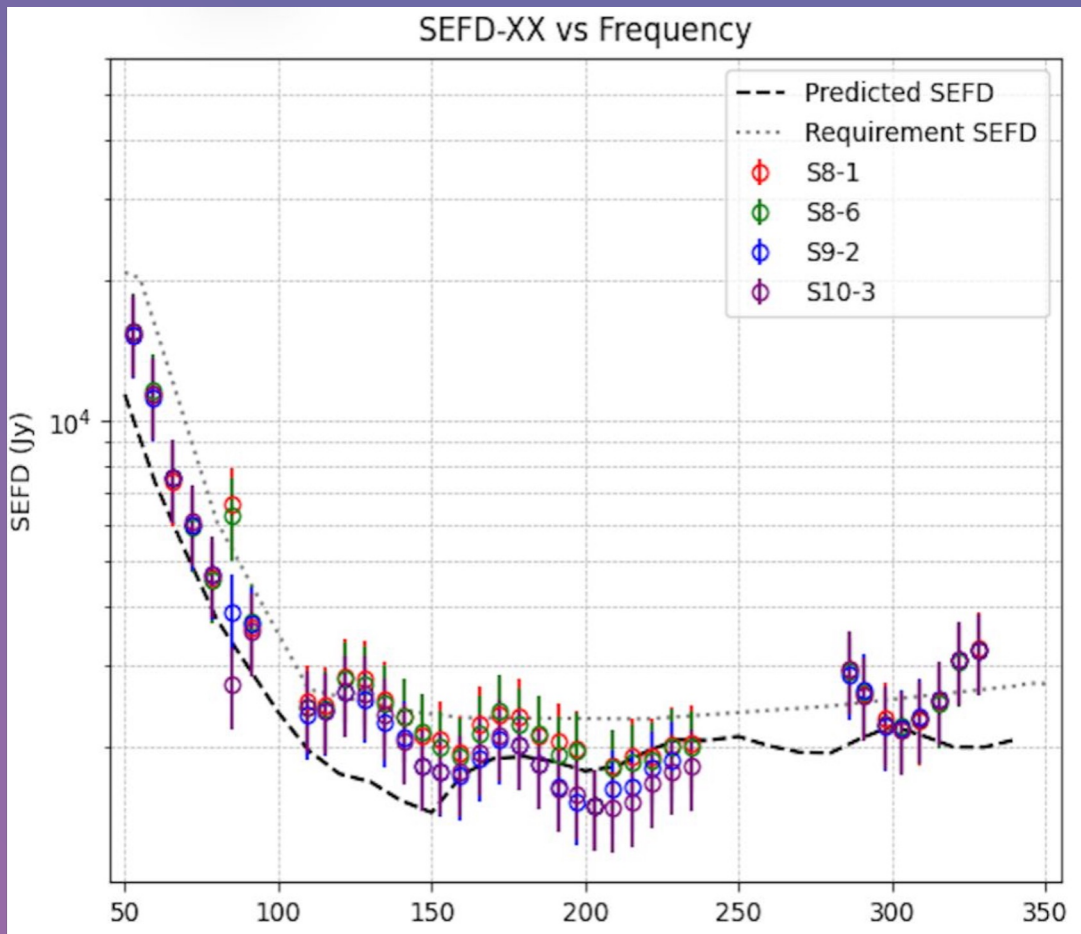
Initial polarization tests from interferometric imaging as well, with RM also aligned with MWA results





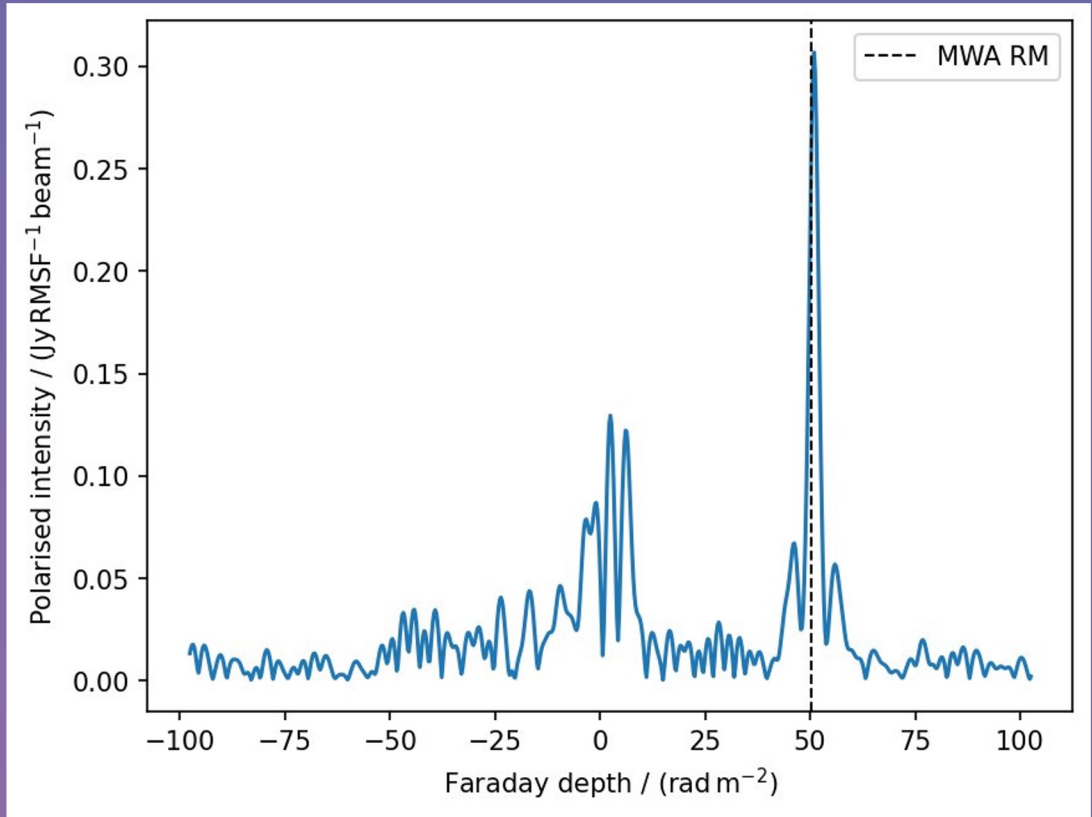
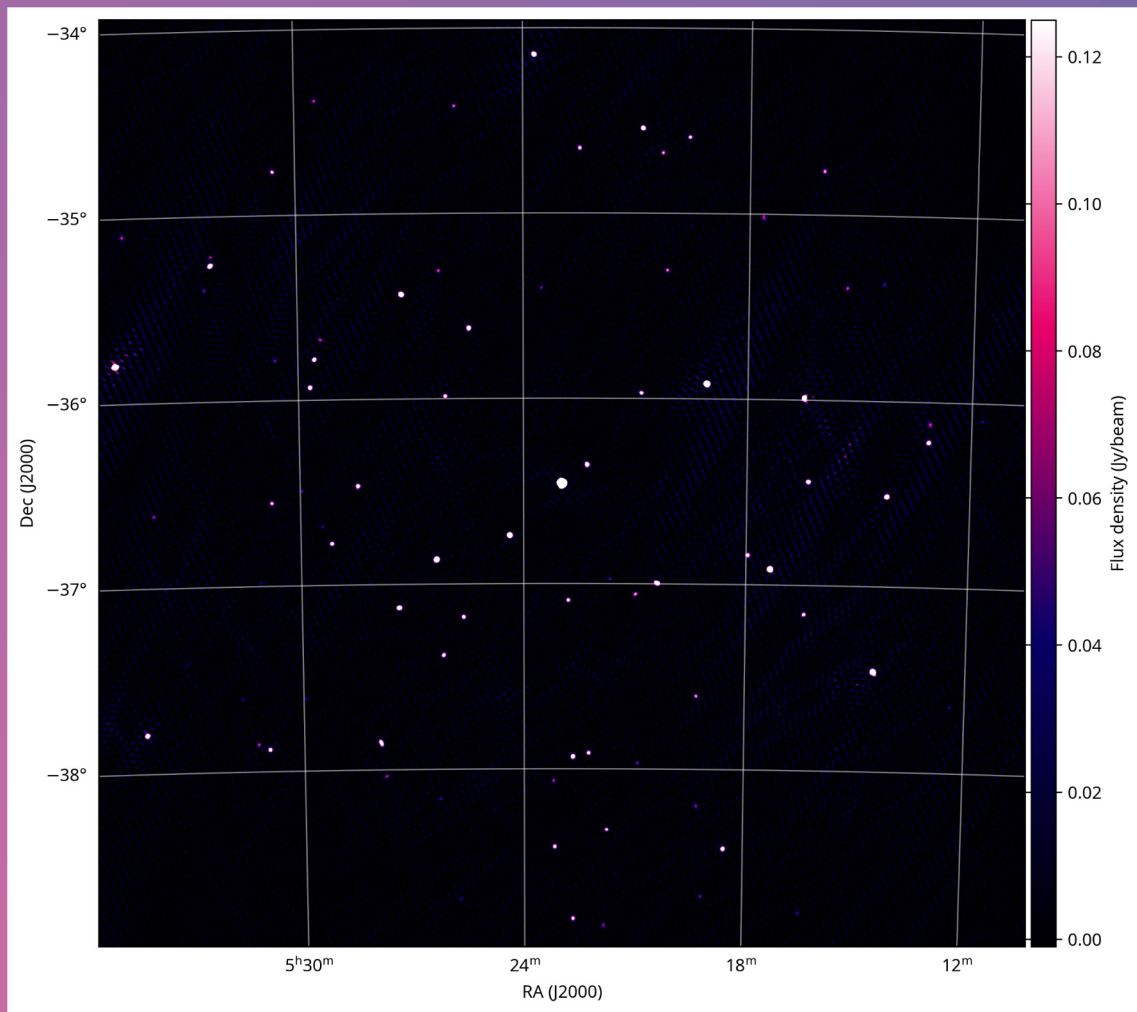
# Recap: Science Commissioning outcomes

Good single station calibration, pointing, tracking, sensitivity



Initial tied-array beam results in pulsar timing (PST) mode, with full coherence coming soon

Good array calibration demonstrated, with first images being produced



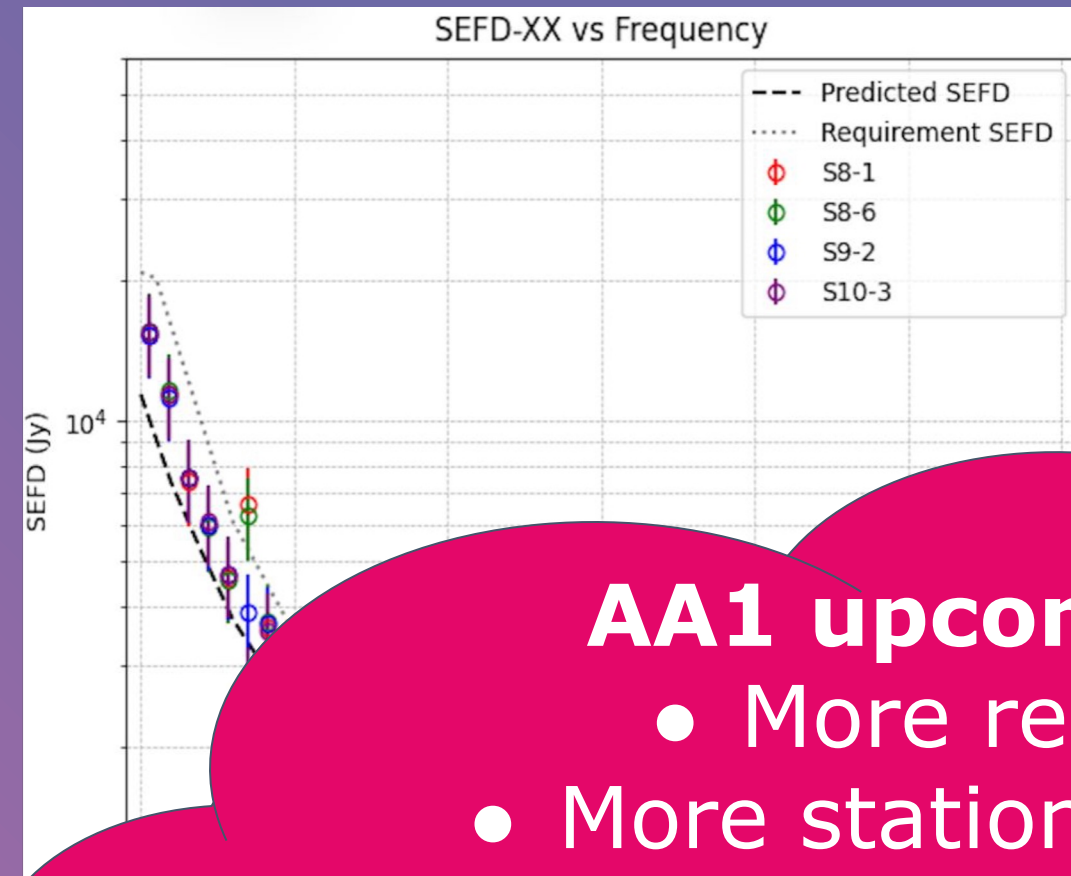
Good results for polarimetry, both from pulsar and imaging observations





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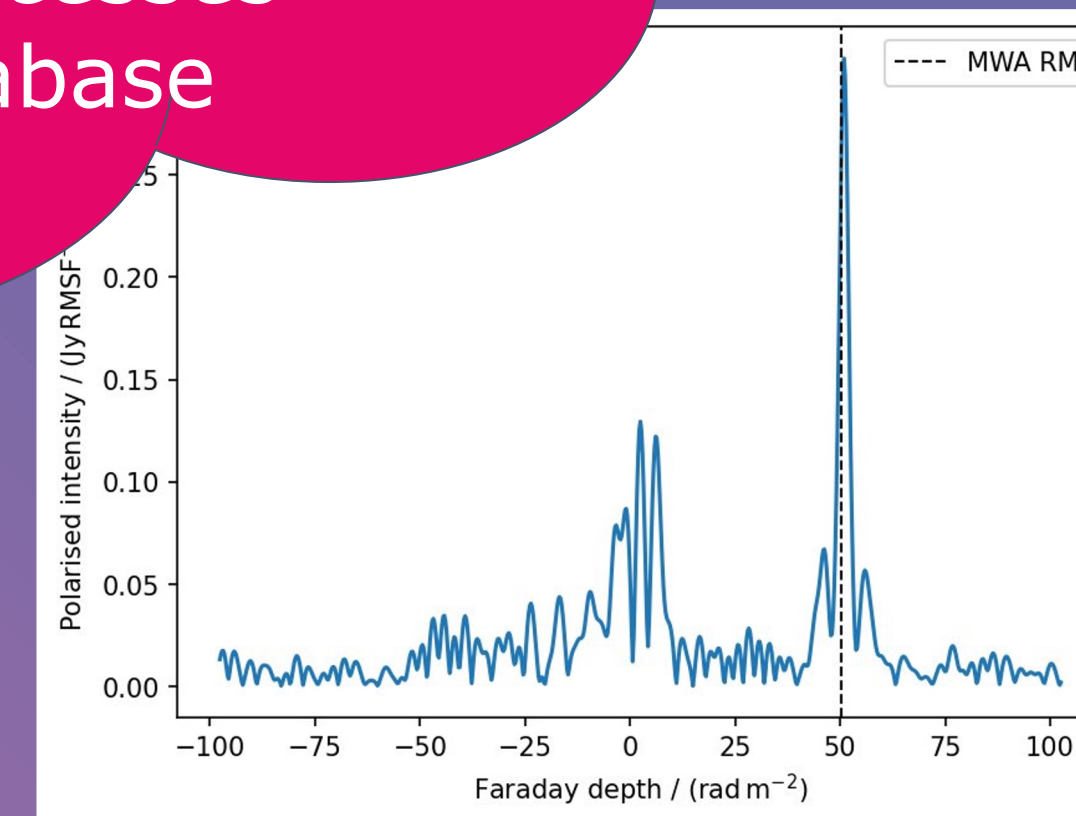
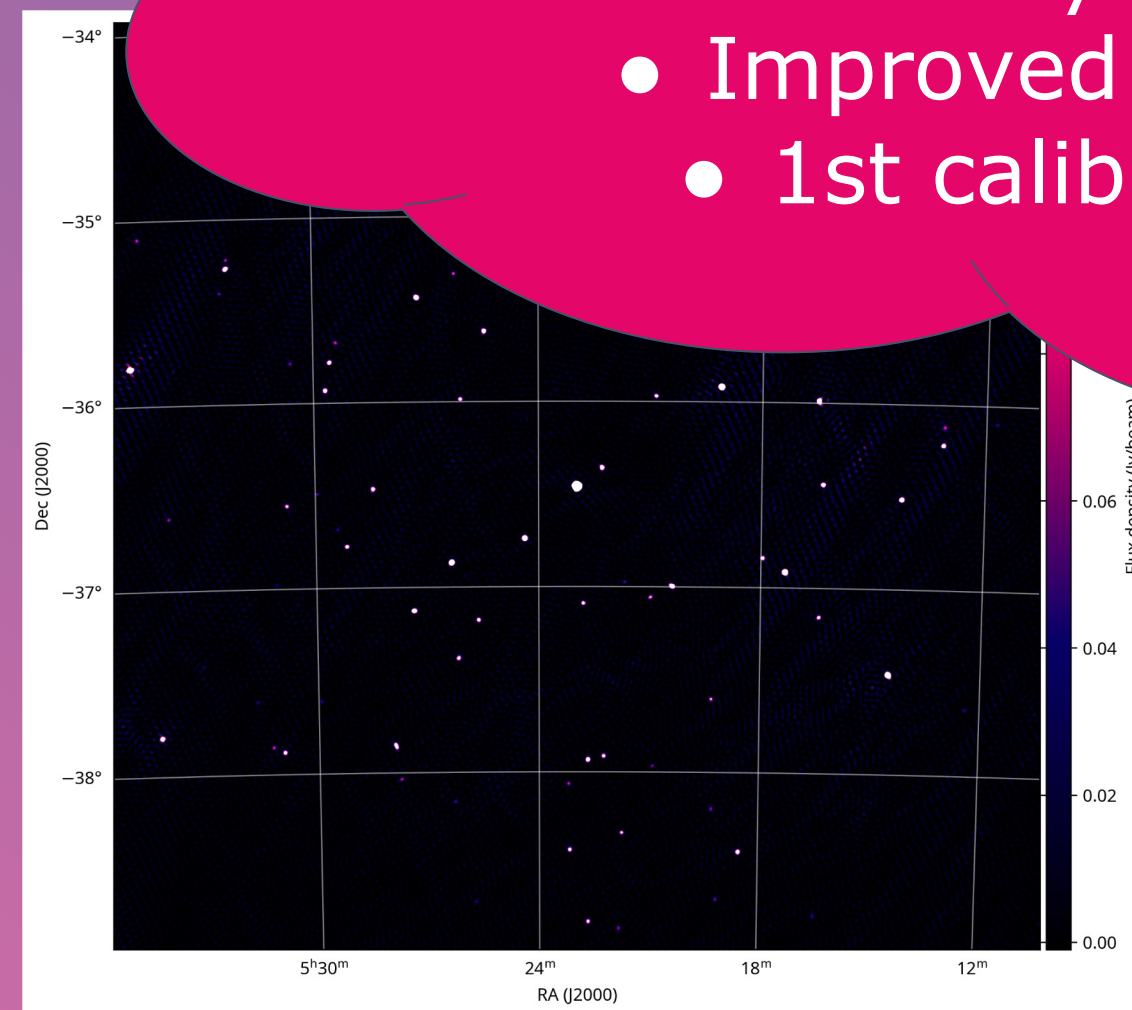
## AA1 upcoming shortly!

- More reliable uptime
- More stations → much better calibration & sensitivity
- Better understanding of systematics
- Improved team processes
- 1st calibrator database



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Good results for polarimetry, both from pulsar and imaging observations





**Feedback:**

[sciops@skao.int](mailto:sciops@skao.int)

**Science User Q&A :**



*We recognise and acknowledge the Indigenous peoples and cultures that have traditionally lived on the lands on which our facilities are located.*



**The Low Science Commissioning Team (left to right):**

Sam Mcsweeney, Giulia Macario, George Heald (team leader), Alec Thomson,  
Shivani Bhandari, Randall Wayth and Xiang Zhang

[www.skao.int](http://www.skao.int)



*We recognise and acknowledge the Traditional Owners of the lands on which our facilities are located, and pay our respects to their Elders past and present.*

*Australia's Indigenous people are the first scientists and have long standing knowledge of the Universe that we continue to build on today.*

*We acknowledge the Wajarri Yamaji as the Traditional Owners and native title holders of Inyarrimanha Ilgari Bundara, the CSIRO Murchison Radio-astronomy Observatory, where we are building the SKA-Low telescope in Australia.*

*We acknowledge the Whadjuk Noongar as the traditional owners of the land where our Science Operations Centre is situated in Perth, and the Southern Yamatji as the traditional owners of the land where our Engineering Operations Centre is situated in Geraldton.*

*I also pay my respects to all First Nations people in attendance.*



A collaborative painting from Aboriginal Yamaji artists from WA for the SKAO *Shared Sky* exhibition. Credit: Yamaji Arts Centre.

