

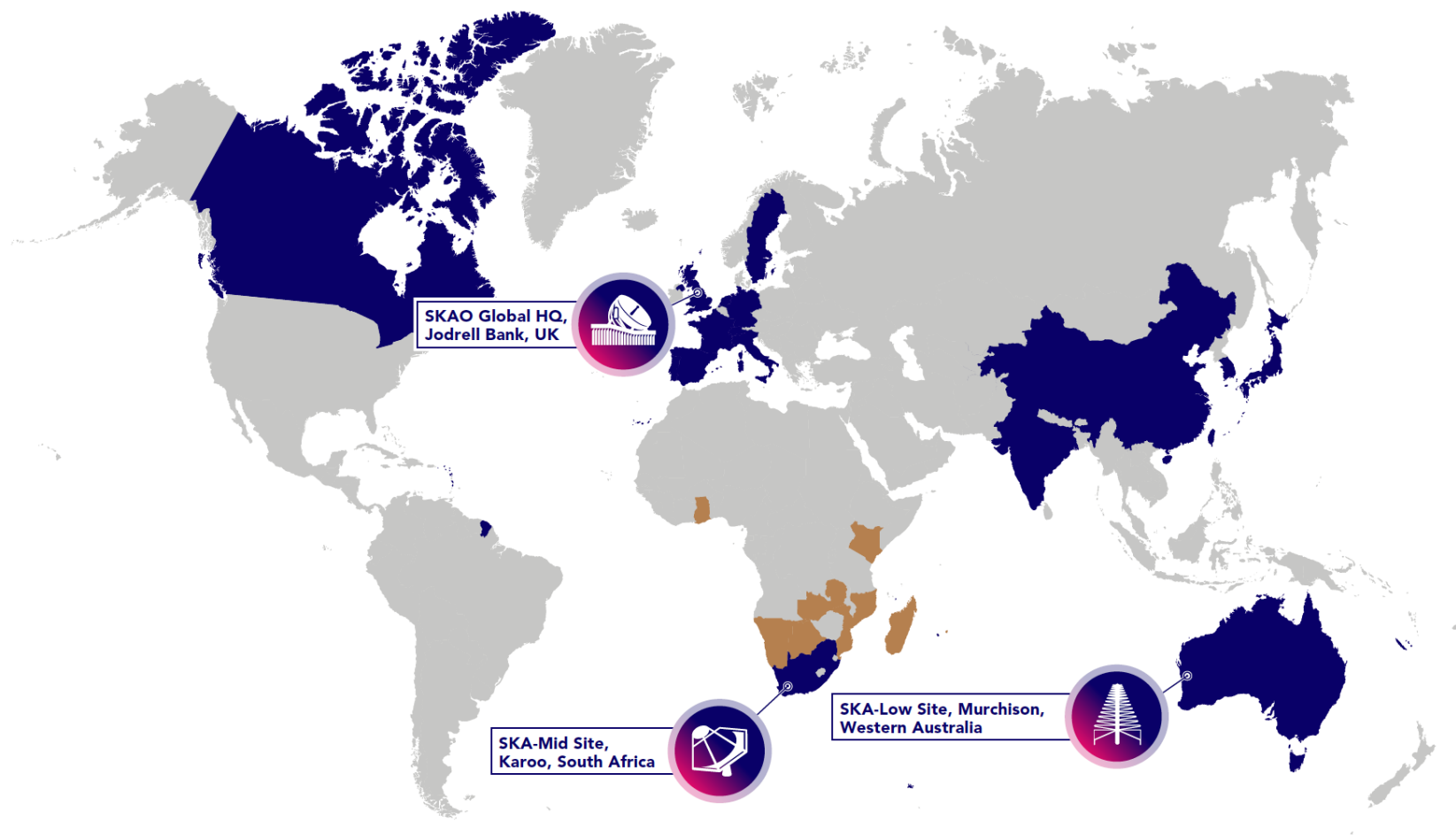


# **The SKA Observatory: scientific perspectives and timeline**

Anna Bonaldi– Head of Scientific Services

The Fifth National workshop on the SKA project  
Monday 24 November 2025





SKAO Partnership - includes SKAO Member States\* and SKAO Observers (as of June 2022)



African Partner Countries



# Italy as a valued Member of SKAO

- . Italian scientists participated in early definition of the SKA in 1990s
- . Italian engineers participated in all SKA design activities in 2000s and 2010s
- . Italy participated in all governance activities from 1999 onwards
- . Italy bid to host the SKAO HQ in 2015
- . Italy convened and chaired the SKAO Convention negotiations; Convention signed at Italian Ministry of Foreign Affairs in March 2019.
- . Italy contributing €120M to SKAO construction and first 10-years of operations.
- . Italian industry and INAF contributions
  - . Feed indexers for SKA-Mid dishes; antennas for SKA-Low; SKA-Low Signal Processing System; local monitor and control software for SKA-Mid dishes; subsystems for SKAO Time scale (critical for an interferometer); software for pulsar search and timing; support from Italian National Metrology institute
- . Italian scientists prominent in all SKAO Science Working Groups; major SKA Science meeting held in Catania in 2014



## Key Science Projects (KSPs)

- Truly transformational & exciting science
- Significant resources and observing time (>1 observing cycle)
- Leadership team restricted to people from member countries
- Co-Is from non-member countries are possible

Reminder: SWGs are NOT KSPs

## Principal Investigator (PI) Projects

- Observing projects that require less time and resources than a KSP
- Typically allocated over a single cycle

## Director-General's Discretionary Time (DDT)

- Timely projects that could not have been submitted in a standard cycle

>500 hr

50-70%

<500 hr

30-50%

## Access to SKA

**Observing time** will be allocated on the basis of **scientific merit** and **technical feasibility** while maintaining access in proportion to **member shares**

Building to a **small amount of Open Time**

### + Proposal attributes

🕒 Long Term Projects

🎯 Target of Opportunity

👥 Joint SKA projects

🔗 Coordinated Projects

+ **Projects based on archival data**  
(that will still require compute resources)



# Updated SKAO Science Book

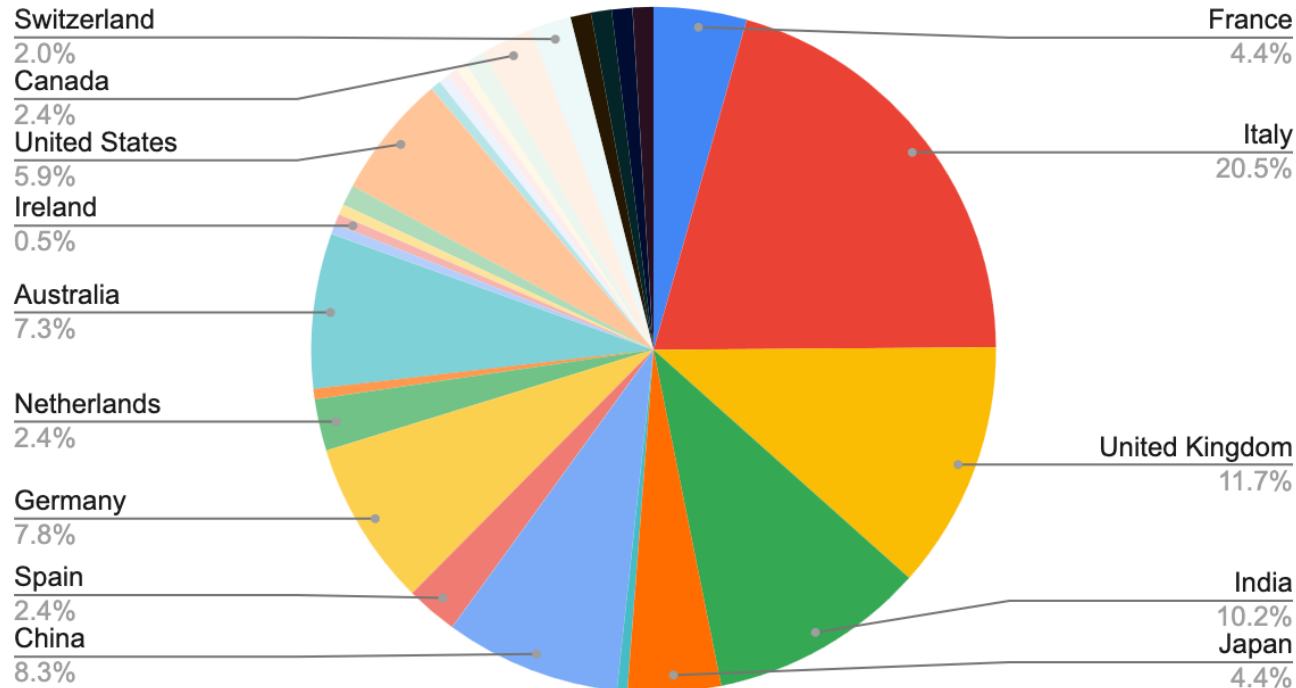
*"Advancing Astrophysics II"*

- 10 years from the publication of *Advancing astrophysics with the SKA*
- Up-to-date coverage of the science enabled by the SKA *Baseline Design* (AA4)
  - To what extent can the science be achieved with AA\*?
  - What enhancements beyond AA4 can further support it?
- Community-driven
  - Science cases from members of our scientific community
  - Coordination and consolidation of cases from the Science Working Group chairs
  - Good representation of all our science areas
- Up-to-date
  - Changes in the scientific landscape
  - Mature set of telescope specifications and expected performance
- Support strategic direction of the Observatory, by providing detailed science requirements for
  - The path from AA\* to AA4
  - The Observatory Development Plan

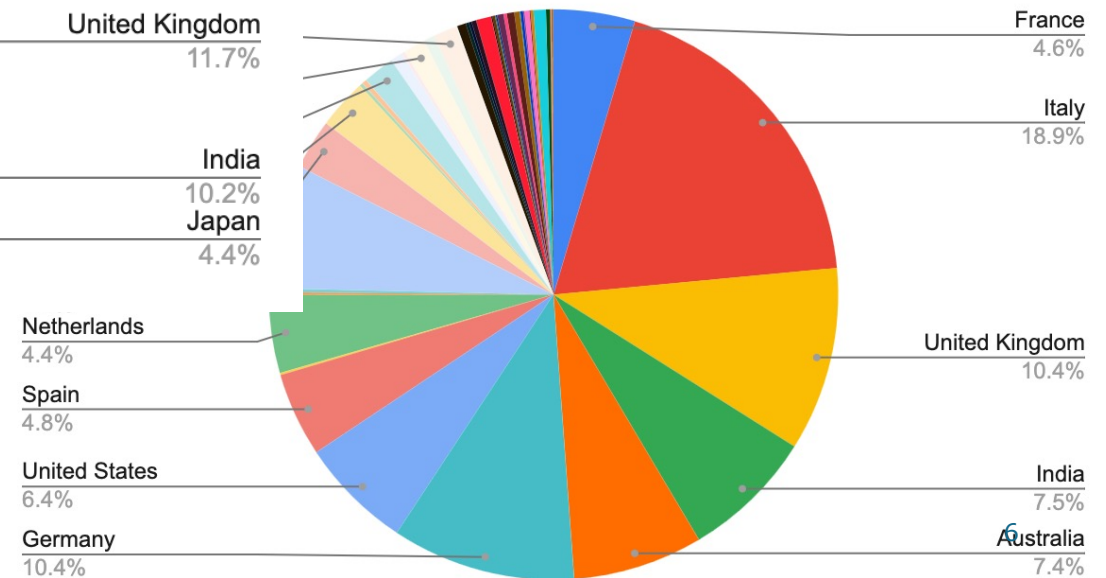


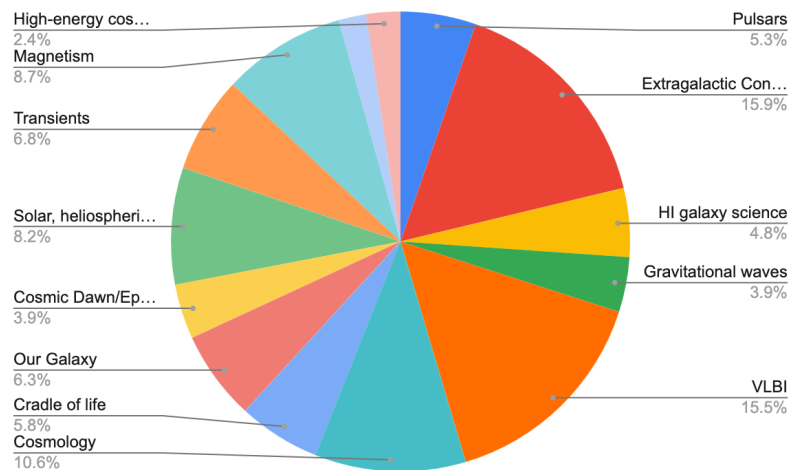
# Italian participation to AASKAII book

## First Authors



## Co-authors



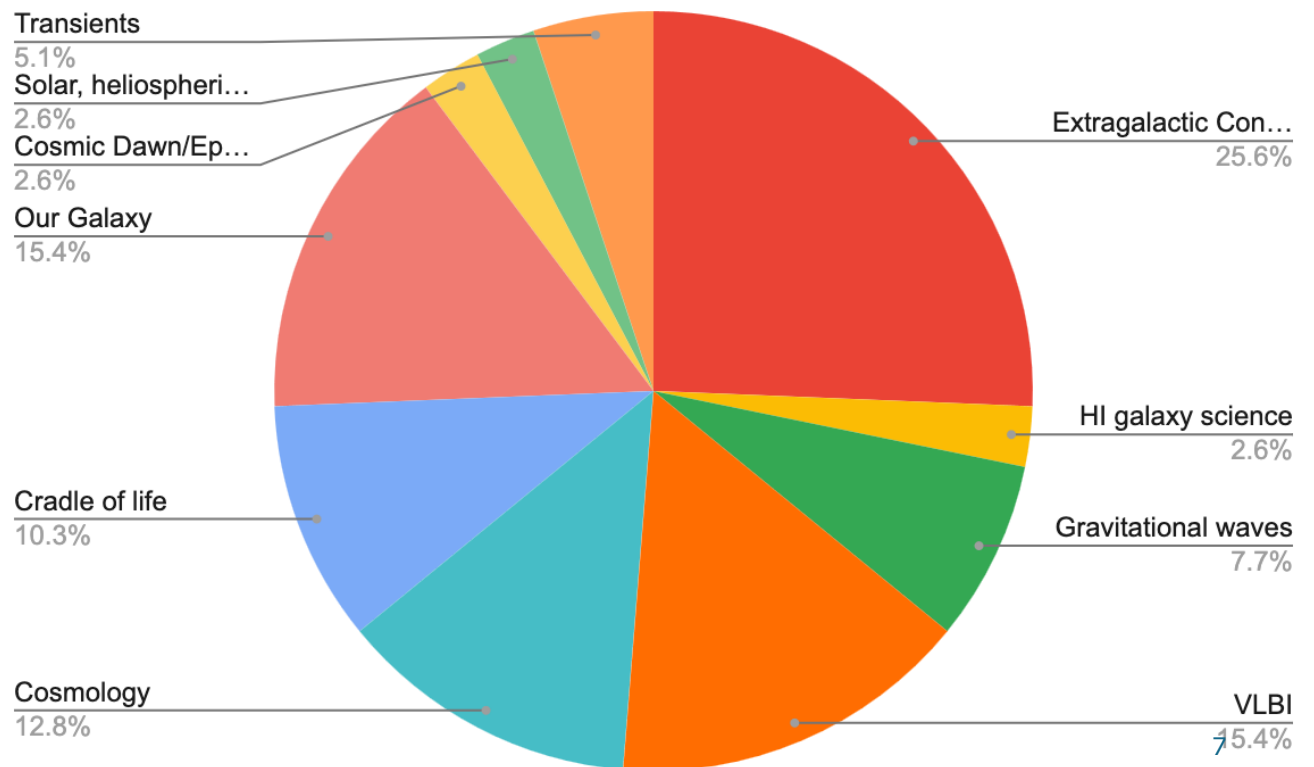


*Submitted papers by SWG category - all*

Enhanced participation to:  
 Extragalactic Continuum  
 Our Galaxy  
 Cradle of Life  
 Gravitational waves  
 Cosmology



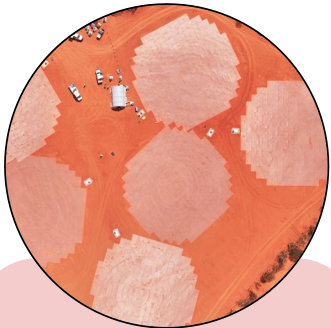
*Submitted papers by SWG category - Italy*



# SKAO milestones

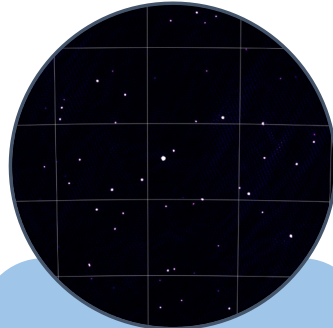
Community involvement starts

Credit: J. Heywood, SARAO



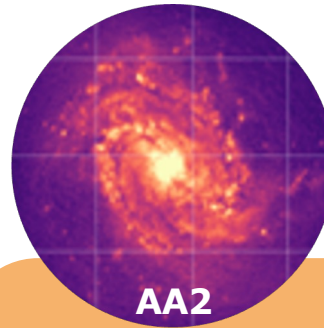
## Construction

- Building antennas, dishes, roads etc!
- Followed by Assembly, Integration and Verification

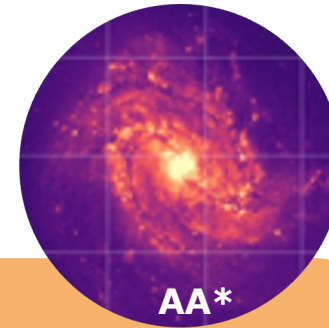


## Commissioning

- SKAO activity
- Includes system verification and science commissioning



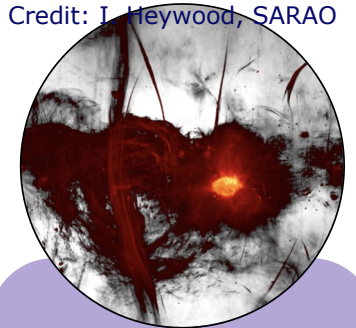
AA2



AA\*

## Science Verification

- A full dress rehearsal of the end-to-end system for every mode of operation
- Once modes and pipelines are working, the community can submit target ideas
- Data will be publicly available for scrutiny
- Build trust and fostering an early science return



## Cycle 0

- Shared-risk PI projects
- SRCNet resources ready for user
- Proprietary periods



Now

Now

First half 2027

First half 2029

2030



Now

Soon

First half 2029

First half 2031

2032





# SKA-Low Construction Progress

- AA0.5 (4 stations) achieved
- ~15,000 antennas assembled = 62 stations
  - 45,312 more antennas on site awaiting assembly
- Tracks, power & fibre: South, East, North arms 100% complete (excluding Kalli Station)
- Mesh: 100% complete south arm, 90% north, 100 (60) % east, core 71 (54) %



# SKA-Mid Dish Structure Progress



SKA001



SKA100



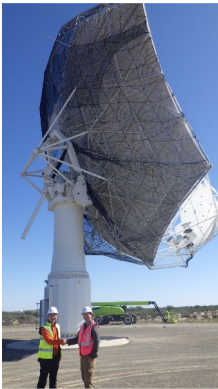
SKA036



SKA034

AA0.5 Fringes

AA0.5



SKA063



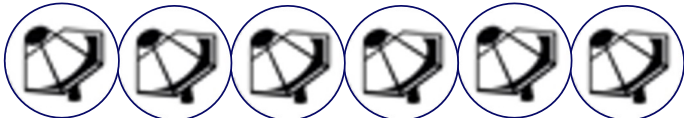
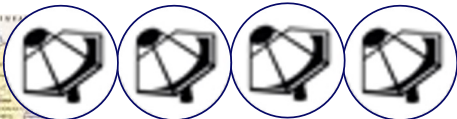
SKA031



SKA077



SKA046



AA1

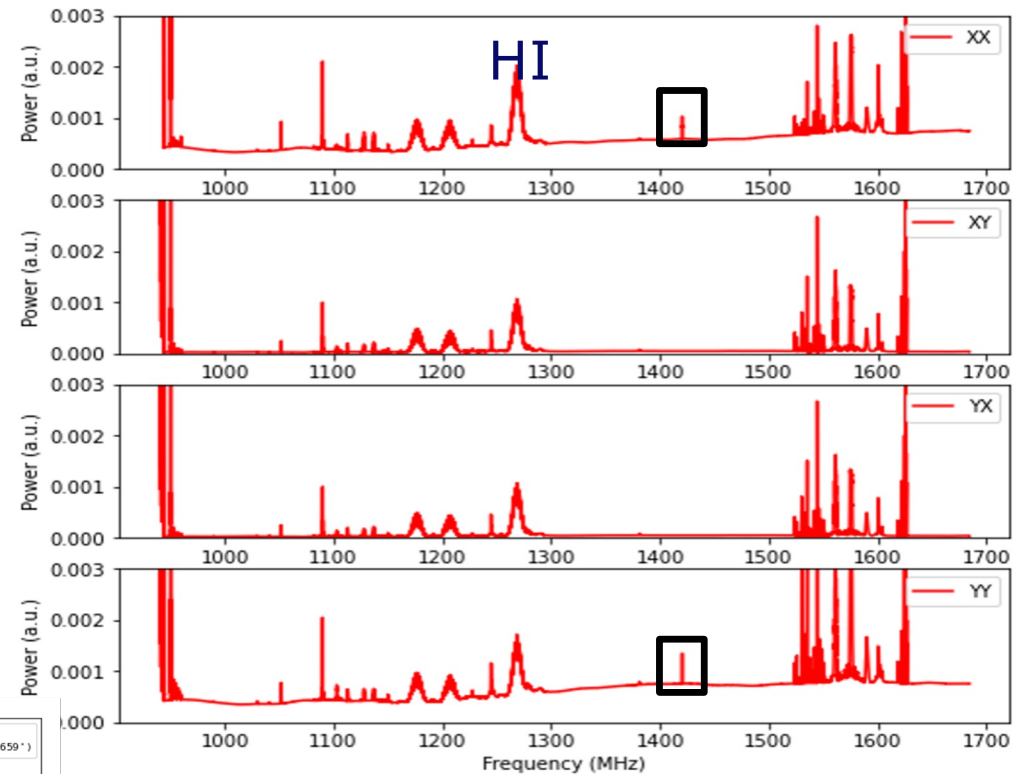
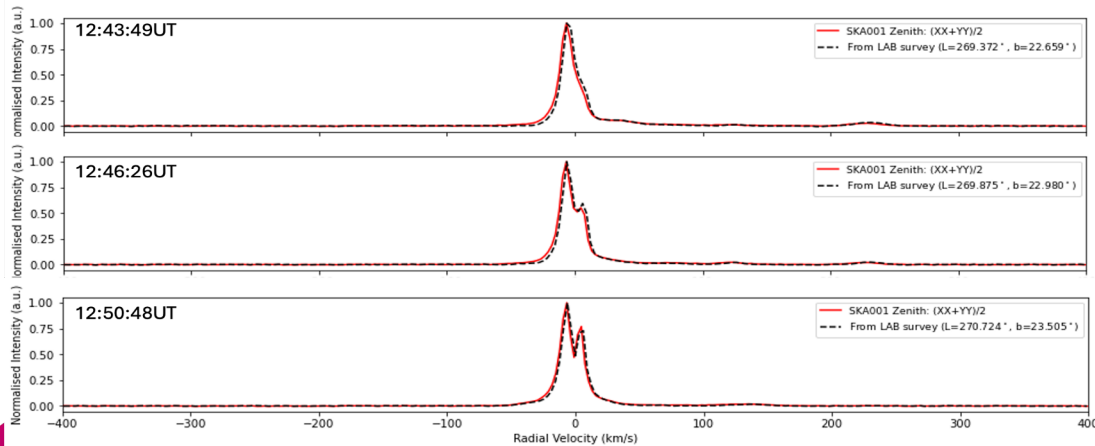
AA2 (Factory Acceptance Testing) - China





# SKA-Mid Full Signal Chain Tests

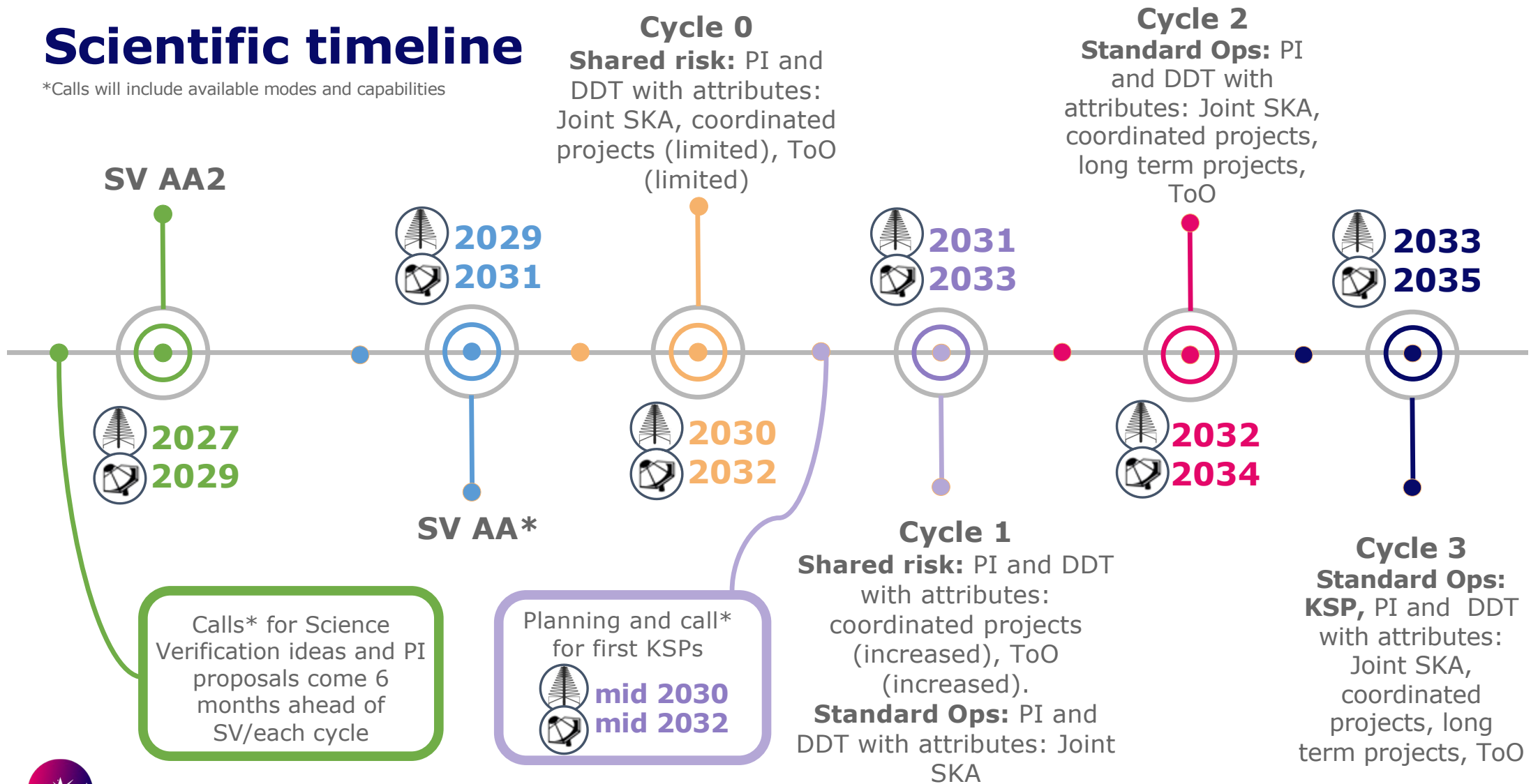
- Galactic HI seen in autocorrelations (XX, YY)
- Line profiles match LAB survey
- Band 2 SPF, digitiser, Talon correlator, software all work as a system



- Single-dish spectra
- Pointing at zenith
- Full signal chain

# Scientific timeline

\*Calls will include available modes and capabilities



# Delivery of modes and ODP capabilities



## Community availability of capabilities

Capabilities will be available to the community once:

**The end-to-end system is in place**

**They pass science commissioning**

**Roll through Science verification, shared risk and into standard operations**



## Delivery of modes & ODPs

Timeline that we are **confident is achievable** and allows the telescope to be commissioned in a sensible order

- Focused development adds confidence in our delivery
- **empowers the community** to build realistic and successful scientific plans.



This is our best understanding today, any evolution of the plan will be communicated as we progress





# Rollout of modes and ODPs

## community engagement and feedback

- Presented in Gorlitz
  - Talk
  - Printed posters in SWG parallel sessions
  - Bespoke discussions between SciOps and SWGs
- Follow up email to ALL SWG members and ALL participants of the meeting asking for feedback
- Available on our website
- Further asked SWG chairs if they had coordinated feedback



# AA2: Science verification

**Low: 2027** 68 stations

**Mid: 2029** 66 dishes

Calibrated, averaged and gridded visibilities, image cubes

Early implementation of continuum subtraction

**Single beam**, single pointing

Max 4k channels  
Max 4h observations

Calibrated, averaged and gridded visibilities, image cubes


Full BW (**150 MHz**/800 MHz)

**Single beam**, single pointing

Max 40 channels  
Max 4h observations

Single subarray

Sidereal and non-sidereal tracking

 Telescope mode

Observatory Data Products (ODPs) and their capabilities

 Continuum  
 Spectral  
 Beamformed  
 Transient

PST at least 1 beam, full processing

**8 PSS beams**

**1 VLBI beam**



# AA\*: Science verification

**Low: 2029** 257\* (307) stations

**Mid: 2031** 144 dishes

## Continuum subtraction

Up to 8 station beams  
At least 1 substation  
arrangement

Max 4h observations, up to 8  
images

Zoom resolution (up to 0.21  
kHz)

Calibrated, averaged and  
gridded visibilities, image  
cubes


Up to 16 station beams  
At least 1 substation  
arrangement

Full BW (300 MHz/5 GHz)

Max 8h observations, up to 16  
images

Single subarray

Sidereal and non-sidereal  
tracking

 Telescope mode

Observatory Data Products  
(ODPs) and their capabilities

 Continuum  
 Spectral  
 Beamformed  
 Transient

2 / 8 PST beams

8 / 200 PSS beams

1 / 4 VLBI beams



\* Restored subject to availability of funds in 2031 funding cycle

# AA\*: Cycle 0

**Low: 2030** 257\* (307) stations

**Mid: 2032** 144 dishes

Multiple pointings (no joint deconvolution)

Basic image-based mosaicing

Up to 24 station beams  
2 substation arrangement

Max 12h observations

Zoom resolution (up to 0.21 kHz)

Multiple pointings (no joint deconvolution)


Up to 24 station beams  
2 substation arrangements

Max 50h observations

Transient buffer triggered by PSS observations (in SV)

4 subarrays

Drift scanning (few mins)

 Telescope mode

Observatory Data Products (ODPs) and their capabilities

 Continuum  
 Spectral  
 Beamformed  
 Transient

8 / 16 PST beams

50 / 200 PSS beams

2 / 4 VLBI beams



\* Restored subject to availability of funds in 2031 funding cycle

# AA\*: Cycle 1

**Low: 2031** 257\* (307) stations

**Mid: 2033** 144 dishes

Multiple pointings

Joint deconvolution (in SV)

Multiple pointings

Limited joint deconvolution (in SV)

Transient buffer triggered by PSS observations (in SV)

Fast imaging/Transient buffer triggered by fast imaging (in SV)

16 subarrays

Drift scanning (few mins)

Telescope mode

Observatory Data Products (ODPs) and their capabilities

Continuum  
Spectral  
Beamformed  
Transient

8 / 16 PST beams

50\* (250) / 200\* (1125) PSS beams

4 / 4 VLBI beams



\* Restored subject to availability of funds in 2031 funding cycle



# AA\*: Cycle 2

**Low: 2032** 257\* (307) stations

**Mid: 2034** 144 dishes

Multiple pointings  
Joint deconvolution  
**Limited** / **Full** PLDP generation

Multiple pointings  
Joint deconvolution  
**Limited** / **Full** PLDP generation

Full drift scan (in SV)  
Wide-area scanning (in SV)


Transient buffer triggered by  
PSS observations

Fast imaging/Transient buffer  
triggered by fast imaging (in  
**shared risk**/ **SV**)

**8** / **16** PST beams

**50\*** (250)/ **200\*** (1125) PSS  
beams

**4** / **4** VLBI beams

 Telescope mode

Observatory Data Products  
(ODPs) and their capabilities

 Continuum  
 Spectral  
 Beamformed  
 Transient



\* Restored subject to availability of funds in 2031 funding cycle

# AA\*: Cycle 3

**Low: 2033** 257\* (307) stations

**Mid: 2035** 144 dishes

Multiple pointings  
Joint deconvolution  
Full PLDP generation  
Autocorrelaton data (shared risk)

Multiple pointings  
Full Joint deconvolution  
Full PLDP generation

Full drift scan (in SV)  
Wide-area scanning

Transient buffer triggered by  
PSS observations

Fast imaging/Transient buffer  
triggered by fast imaging

8 / 16 PST beams

50\* (250)/ 200\* (1125) PSS  
beams

4 / 4 VLBI beams

 Telescope mode

Observatory Data Products  
(ODPs) and their capabilities

 Continuum  
 Spectral  
 Beamformed  
 Transient



\* Restored subject to availability of funds in 2031 funding cycle

# Updates to the rollout?

- Rollout based on what we think we can deliver when, but:
- Continuing to investigate:
  - Concerns from the Cosmology SWG over the current timing of Wide Area Scanning (SV with this mode starting in 2034)
    - What pipelines are available within the community and;
    - Do we think current pipelines would significantly accelerate the delivery of SKAO pipelines
  - Concerns from the Transient SWG over the current timing of fast imaging (SV with this mode starting in 2032 for Low and 2033 for Mid)

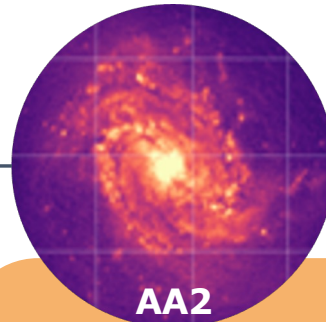


# SKAO milestones

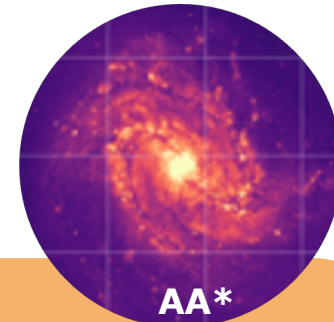


**Call for SV  
ideas ~ 6  
months  
before**

Community involvement starts



AA2

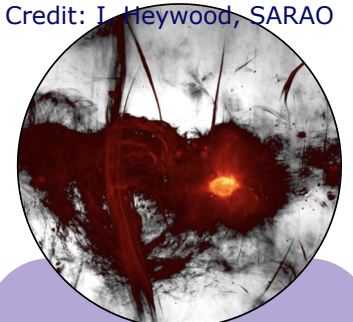


AA\*

## Science Verification

- A full dress rehearsal of the end-to-end system for every mode of operation
- Once modes and pipelines are working, the community can submit target ideas
- Data will be publicly available for scrutiny
- Build trust and fostering an early science return

Credit: J. Heywood, SARAO



## Cycle 0

- Shared-risk PI projects
- SRCNet resources ready for user
- proprietary periods



**Middle 2026**

**First half 2027**

**First half 2029**

**2030**



**Middle 2028**

**First half 2029**

**First half 2031**

**2032**



Thank you for your time...

We recognise and acknowledge the Indigenous peoples and cultures that have traditionally lived on the lands on which our facilities are located.

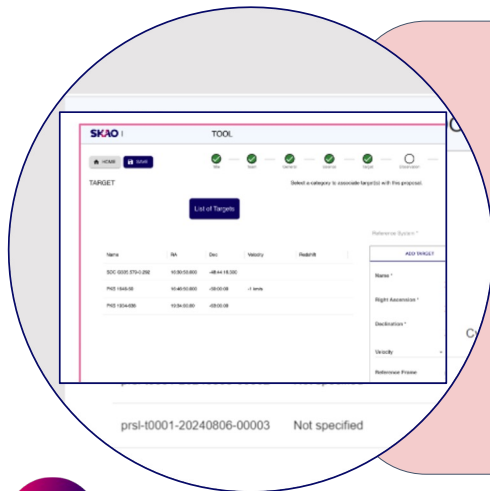
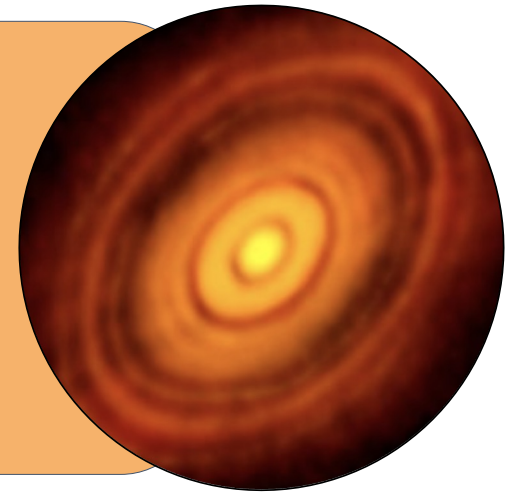




# Science verification

HL Tau Science Verification image from ALMA! Credit: ALMA (NRAO/ESO/NAOJ); C. Brogan, B. Saxton (NRAO/AUI/NSF)

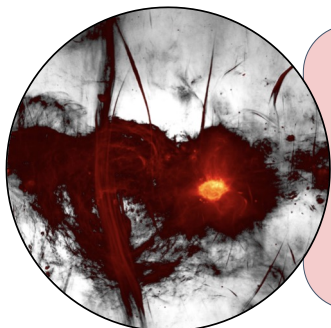
- First chance the community will have to access SKA data - you will be closely involved!
  - Provide the ideas and targets
  - Public data releases for scientific exploitation and fostering an early science return!
  - Validating data reduction workflows (allowing the SKAO to begin to build trust in our generated data products) and the user experience
  - Begin to understand the potential of an observing mode



- End-to-end test of the system
- Important milestone before we put out the first call for proposals
  - Do all our systems work? submission, telescopes, Correlator Beam Formers, observing mode, tools, data processing, data delivery, SRCNet analysis
  - Are the Observatory staff prepared?
- Allows us to build trust in our systems and pipelines, build momentum, demonstrate potential, science!



# What we will provide and support in Science Verification

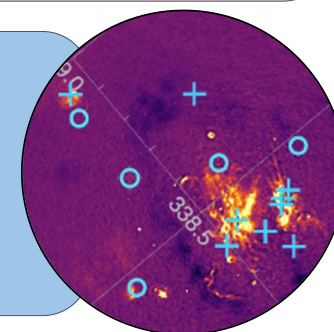


## Data products

- Fully processed Observatory Data Products (ODPs)
- Calibrated constituent data products. Calibrated, averaged visibilities (including calibration solutions), time domain products.
  - **Verification of pipelines** is an important step in SV
  - **Important for building trust**
- Verification periodically as modes/array mature/grow (even into Cycle 0)

- A description of the processing that has been done to arrive at the fully processed data product (e.g., workflow scripts).
- Quality assessment report.
- Versioning information for the software, relevant copies of telescope model data etc
- Local sky model

## Ancillary Products



## Support

- **A memo series** so that early career researchers, e.g., could receive technical credit, especially if a dataset doesn't result in publication
- **A helpdesk, community training and workshops.** Including documentation, videos, online drop in sessions, podcasts, community days.
- Limited SRCNet resources for analysis, visualisation, and Advanced Data Product creation.



# Programme Schedule & Risk Review (PSRR)

Covered in SKAO's  
webinar by Naomi  
McClure-Griffiths on  
18/11

- Started in early 2025 to fully review the schedule (phase 1) and risk/cost (phase 2) of the Construction programme
- Phase 1 completed before Görlitz SKA Science Meeting
  - Outcome was the ~2 yr schedule slip of SKA-Mid
- Phase 2 has been underway between July - November 2025
  - Resulted in the identification of extra cost risk means the project cannot deliver AA\* on budget
    - Risk identified from lessons from our delivery experience to date, cost escalation, scope gap, contractor failure, schedule delay
- Implemented extensive Observatory-wide cost-savings

To ensure costs do not exceed the budget up to 2031, we have had to make the difficult decision to defer some scientific capability



# Deferred capabilities

To help resolve the budget we will defer until post-2030 the delivery of 4 capabilities:

- 1) 50 stations in the core of SKA-Low. The initial delivery will be 257 stations, of which 149 are in the core.
- 2) PSS on SKA-Mid will be limited to 200 beams until *at least* Cycle 0 as advertised
- 3) PSS on SKA-Low will be limited to 50 beams in the early observing cycles
- 4) The SKA-Low CBF will match AA\* (it was previously over-scoped to match AA4)

**It is our clear intent to reinstate the capabilities as soon as financially viable starting from 2031, or sooner, to minimise the impact on science.**



# Planned Array Assemblies

