



LIFE =
Large Interferometer For Exoplanets

The LIFE Mission –
Characterizing other worlds and
searching for life

ETH zürich

PlanetS
National Centre of Competence in Research

ETH zürich | **SPACE**



CENTRE FOR
ORIGIN &
PREVALENCE OF
LIFE

Shaping the Italian contribution to HWO

July 11, 2025

Authors:

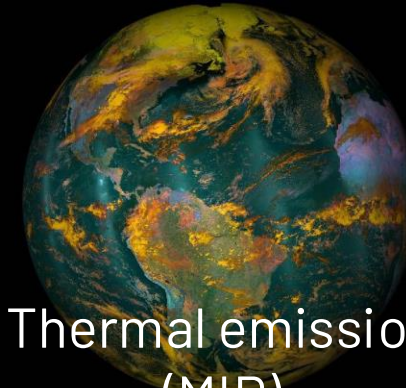
Sascha P. Quanz (PI)

for the
LIFE project

LIFE is unique among future facilities to directly detect terrestrial exoplanets and search for indications of life



Reflected light
(UV / Optical / NIR)

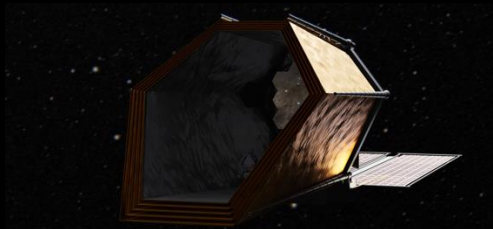


Thermal emission
(MIR)

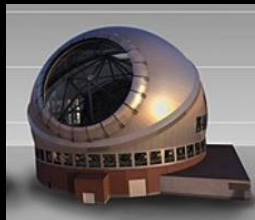
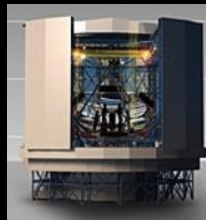
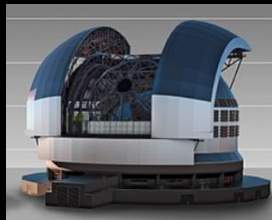


Solar-type
stars

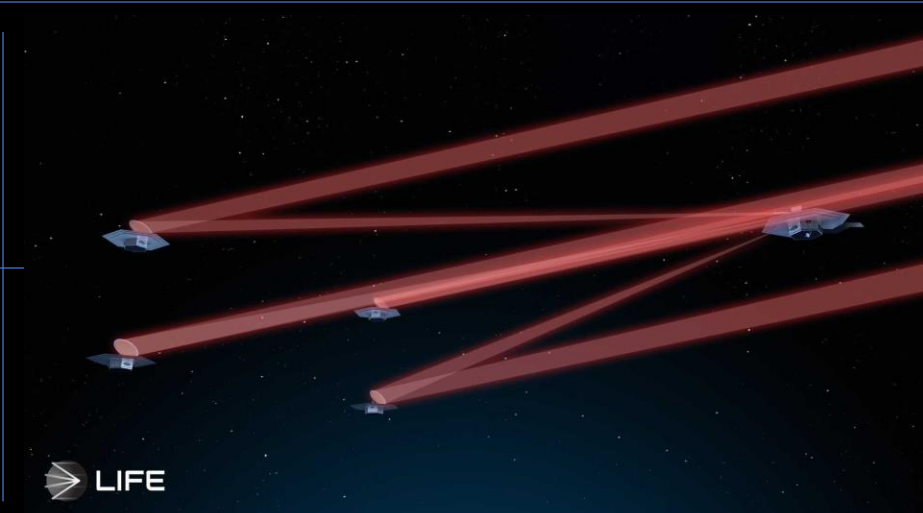
NASA's
HWO



ELTs



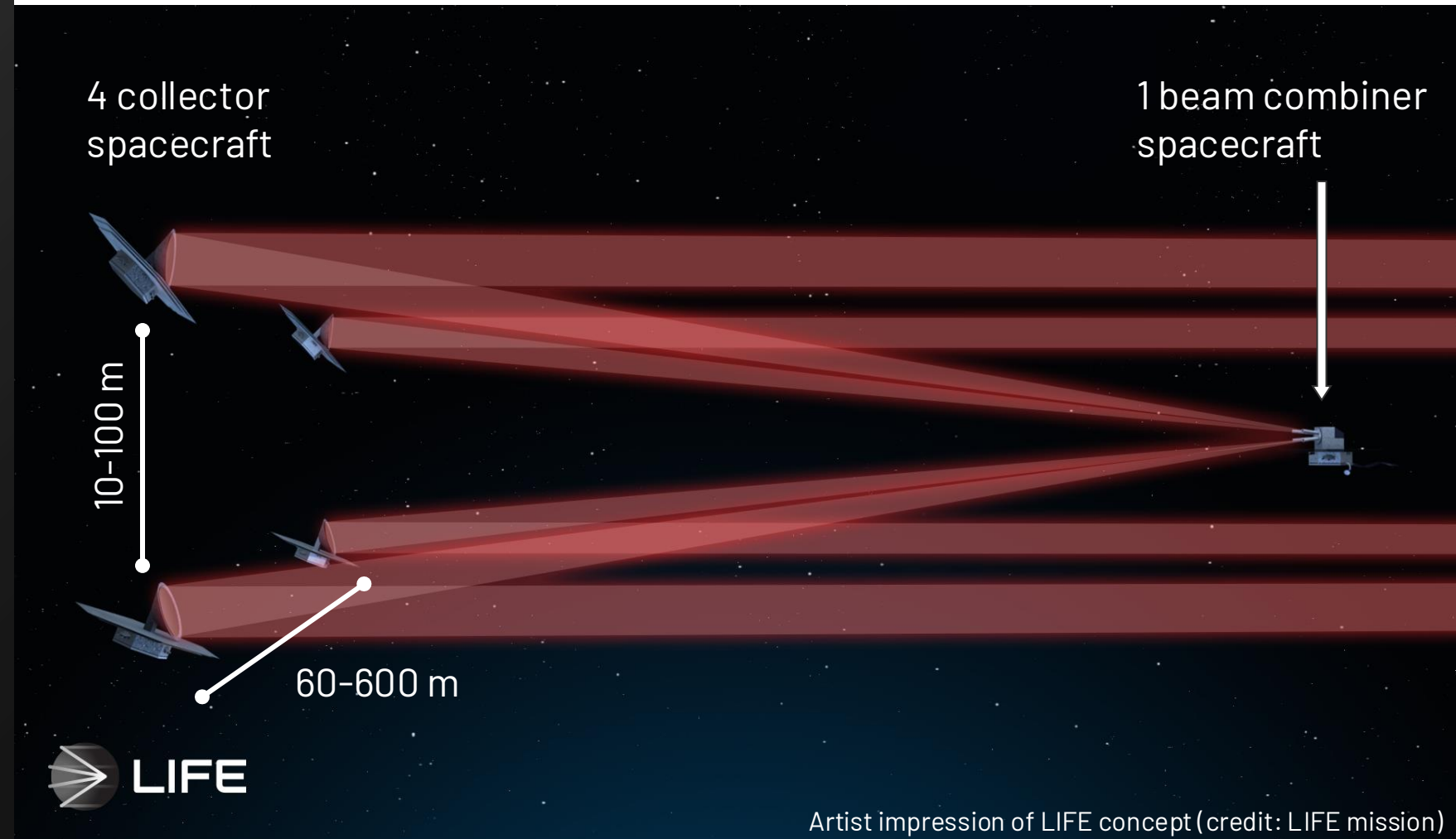
M stars



LIFE

The LIFE mission

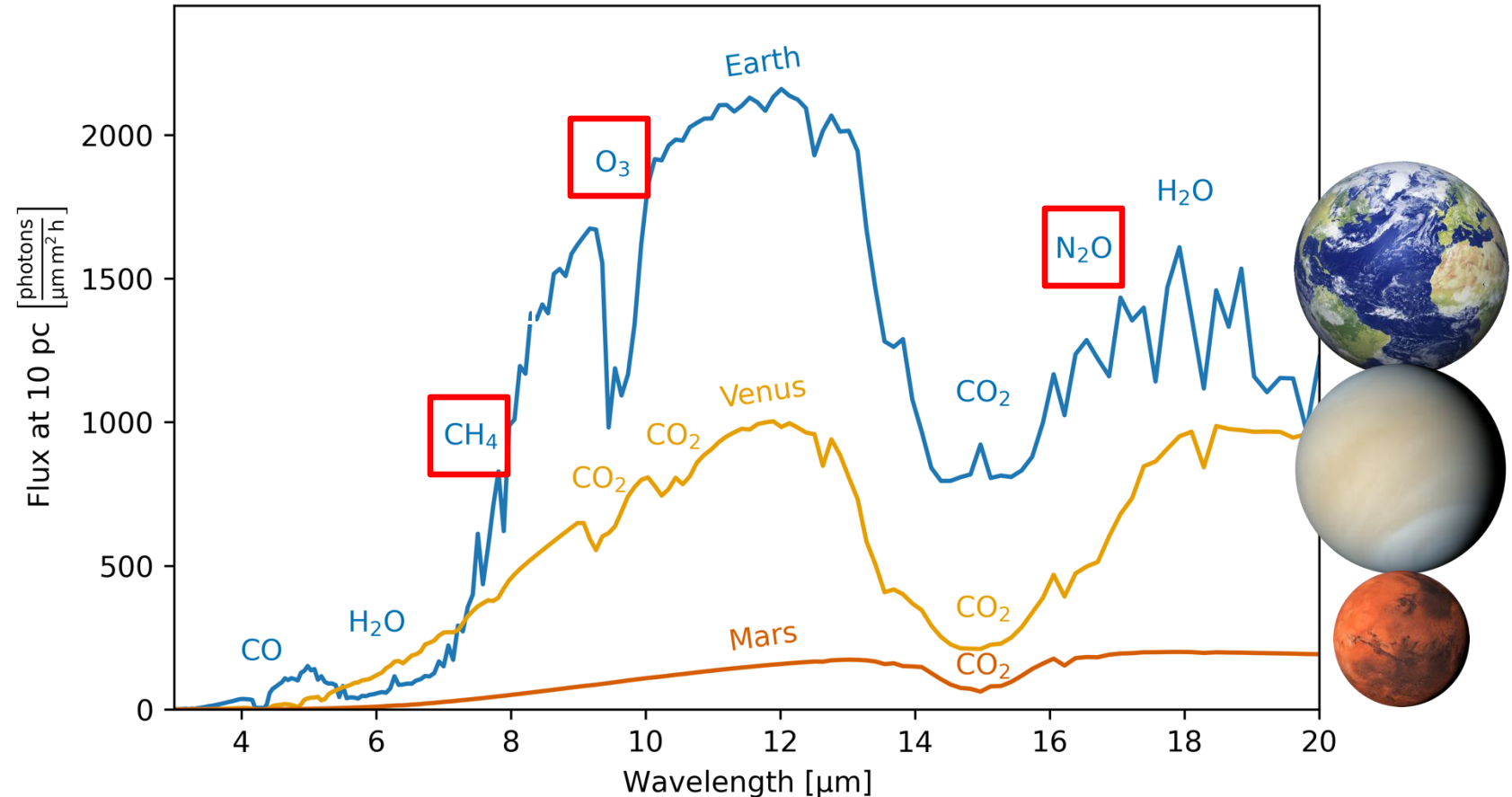
- LIFE is a space-based formation-flying nulling interferometer
- It consists of 4 collector spacecraft in a rectangular array and a central beam combiner spacecraft above the array
- The separation between the collectors can be freely adjusted to optimize the performance for each nearby star
- Like the James Webb Space Telescope, LIFE will orbit around Lagrange Point 2
- The nominal mission lifetime will be 5-6 years
- LIFE covers the mid-infrared wavelength range between $\sim 6\text{--}16\ \mu\text{m}$ (requirement) / $\sim 4\text{--}18.5\ \mu\text{m}$ (goal) with a spectral resolution of $R = \lambda/\delta\lambda \sim 100$



Investigating other worlds

- LIFE's wavelength range is chosen to cover the peak of the thermal emission of temperate terrestrial planets
- This wavelength range features absorption bands of major atmospheric constituents including molecules that are only present because biological activity such as ozone (O_3), methane (CH_4) and nitrous oxide (N_2O)

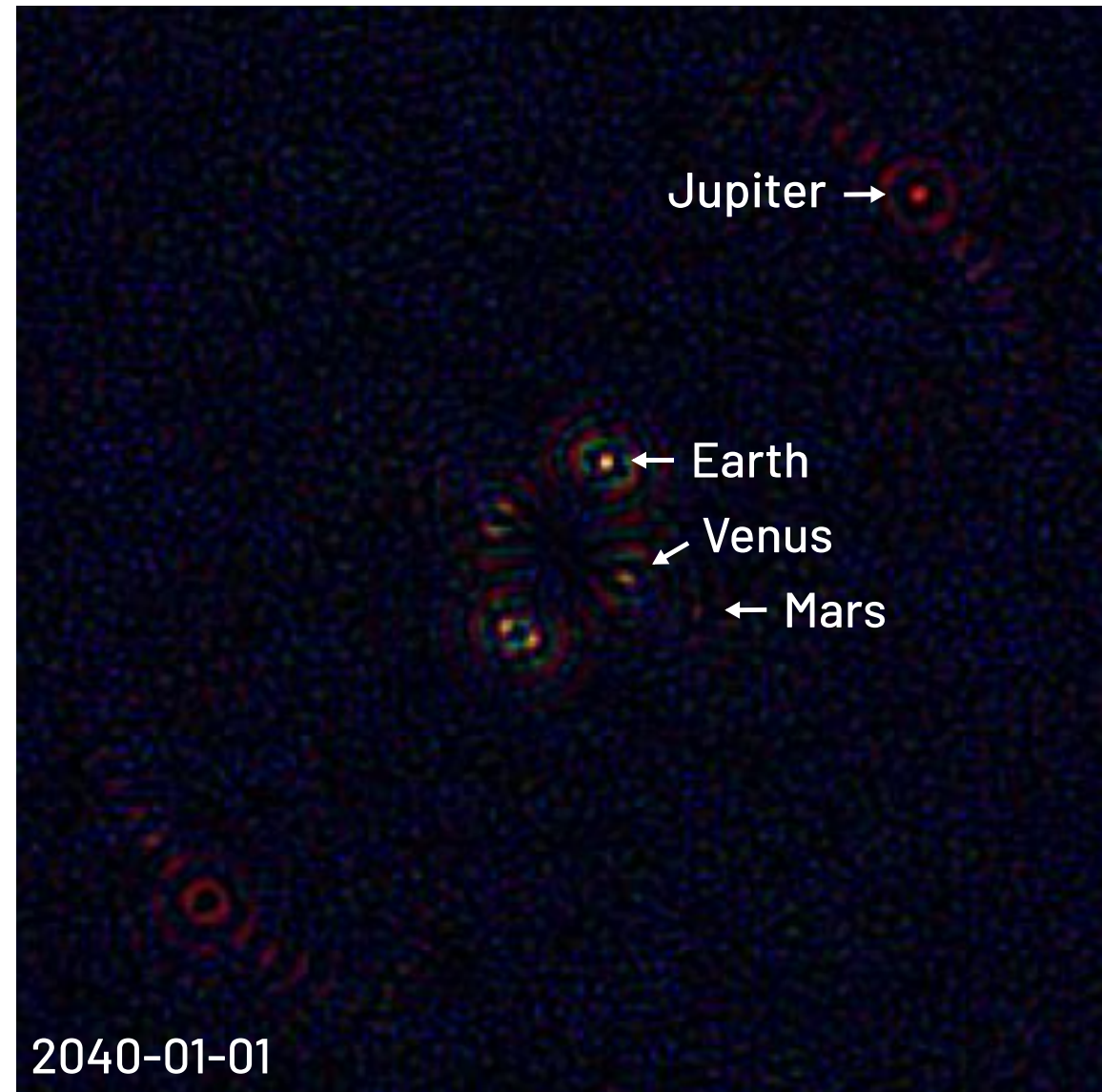
Emission spectra of terrestrial planets in our Solar System



Investigating other worlds

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Our Solar System as seen with LIFE from 10 pc distance



Exoplanet characterization: the mid-infrared advantage

In contrast to exoplanet observations in reflected light in the optical/near-infrared, LIFE will...



...directly constrain the **pressure-temperature structure** of exoplanet atmospheres



...access (multiple) atmospheric absorption bands of **major molecules** such as H_2O and CO_2 as well as collision induced absorption from N_2 and O_2



...search for numerous **atmospheric biosignatures** in the context of terrestrial exoplanets and gas dominated sub-Neptunes (e.g., O_3 and CH_4 , but also N_2O , PH_3 , NH_3 , and C_5H_8)



...constrain directly **the effective temperature** of exoplanets and provide access to their **radii**



...deliver a higher detection yield during search phase as it is **less affected by the orbital phase function** of the exoplanets' emission compared to reflected light missions



...immediately **start observing already known small, temperate exoplanets** around nearby M-stars

LIFE: today's most transformative mission to search for life beyond the Solar System

LIFE gives access to a broad range of possible biospheres and provides the necessary context for their interpretation

30 – 50 temperate, terrestrial exoplanets around Sun-like stars (F0-K5 stars)
15 – 25 temperate, terrestrial exoplanets around low-luminosity stars (K6-M3)
10-20 temperate, terrestrial exoplanets around very low-luminosity stars (M4-M9)

Earth-like biospheres

LIFE can detect biospheres on planets beyond the Solar System that are like Earth's in their composition and evolution

Non-Earth-like biospheres

LIFE can detect biospheres on exoplanets that differ significantly from Earth's in their stellar environment or their composition

Technological signatures

LIFE can detect imprints of technology in the atmospheres of exoplanets

Planetary diversity as necessary context

LIFE will detect a diverse sample of hundreds of exoplanets providing the large reference sample needed to infer the existence of biospheres

LIFE: today's most transformative mission to search for life beyond the Solar System

NOT EXHAUSTIVE

LIFE gives access to a broad range of possible biospheres and provides the necessary context for their interpretation



Earth-like biospheres

Konrad et al. 2022, 2024
Alej et al. 2022, 2024
Mettler et al. 2023



Non-Earth-like biospheres

Angerhausen et al. 2023, 2024
Leung et al. 2024



Technological signatures

Schwieterman et al. 2024

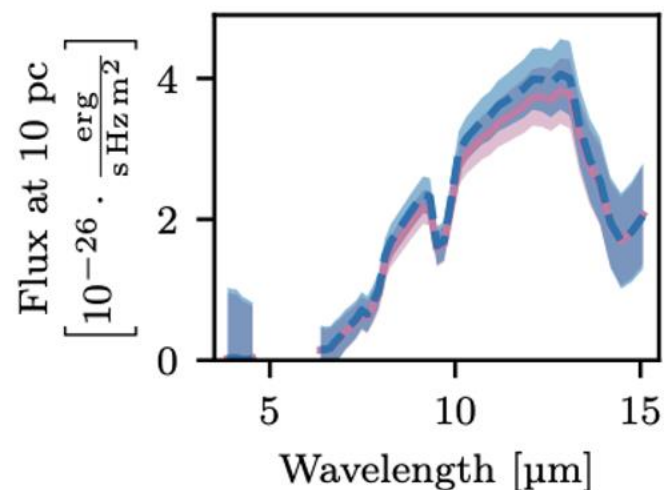
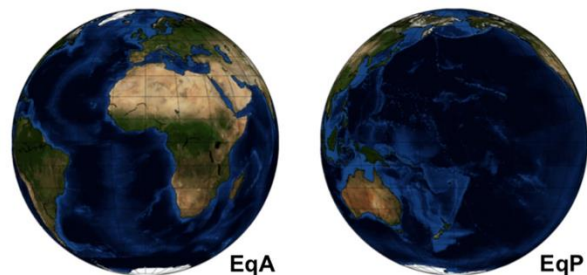
Planetary diversity as necessary context

Quanz et al. 2022; Kammerer et al. 2022; Carrión-González et al. 2023; Konrad et al. 2023; Hansen et al. 2025

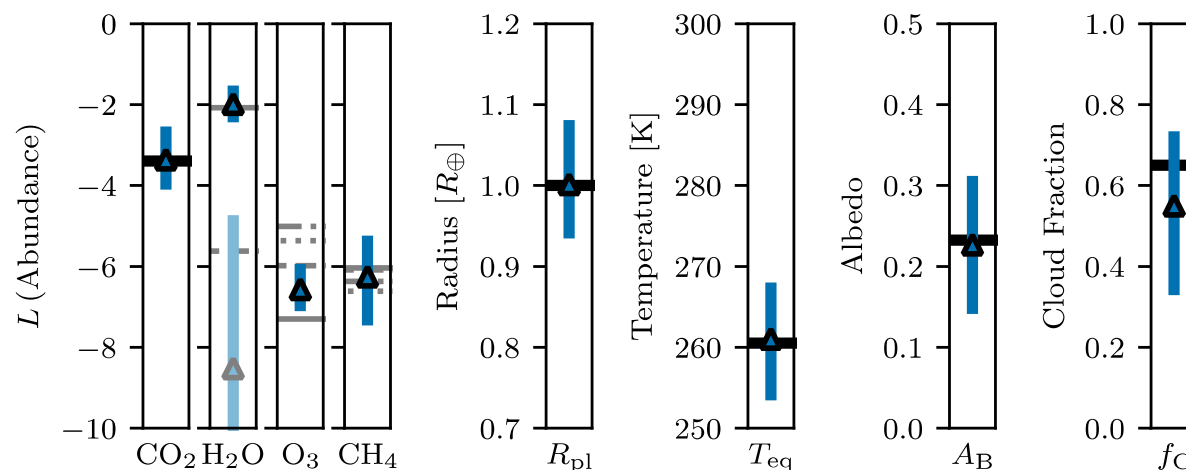
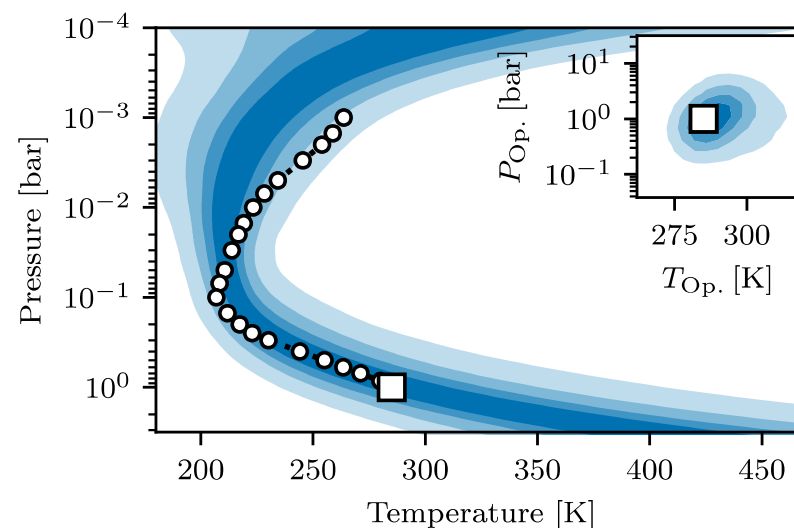
Atmospheric retrievals: LIFE observes “real” Earth

Atmospheric retrieval results based on simulated LIFE observations with SNR=10 and R=100 (~6–16 μm) for July equatorial view

Input data from Earth observing satellite

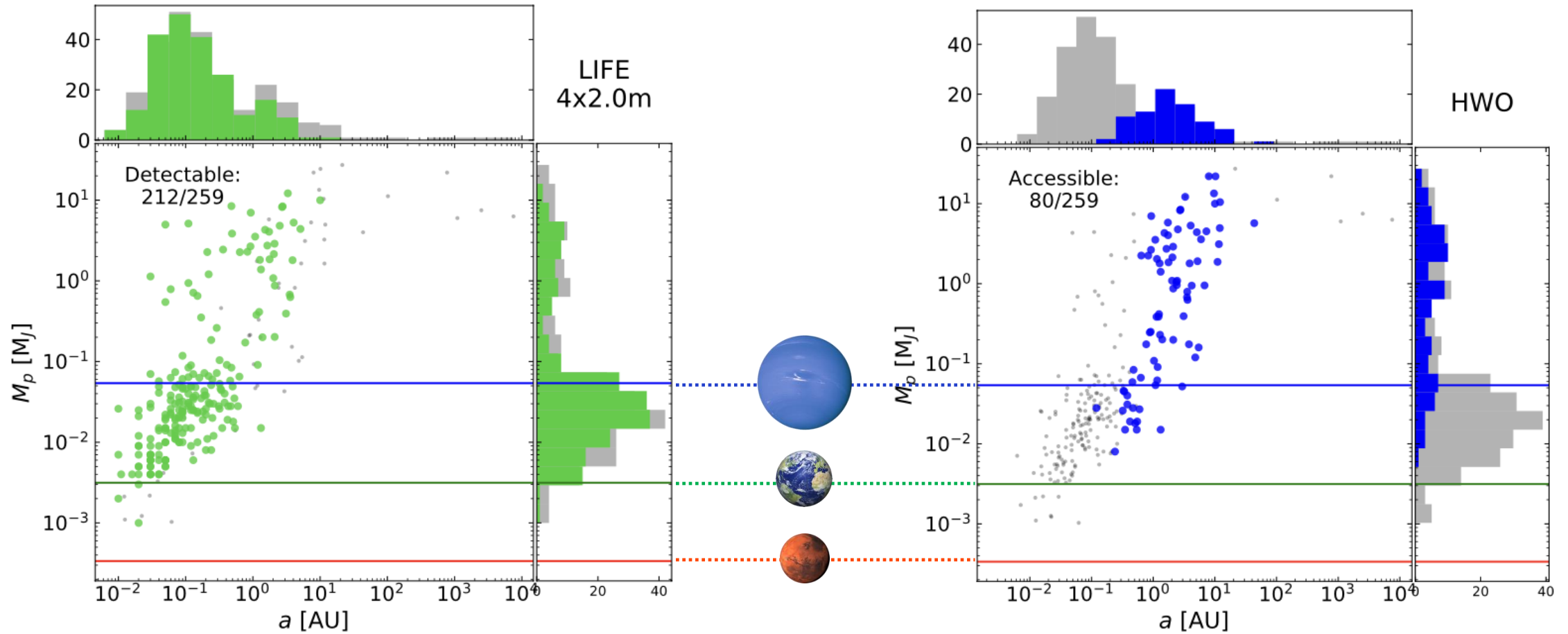


Atmospheric Retrieval



Characterizing known nearby planets from mission day 1

Detectability of known exoplanets within 20 pc

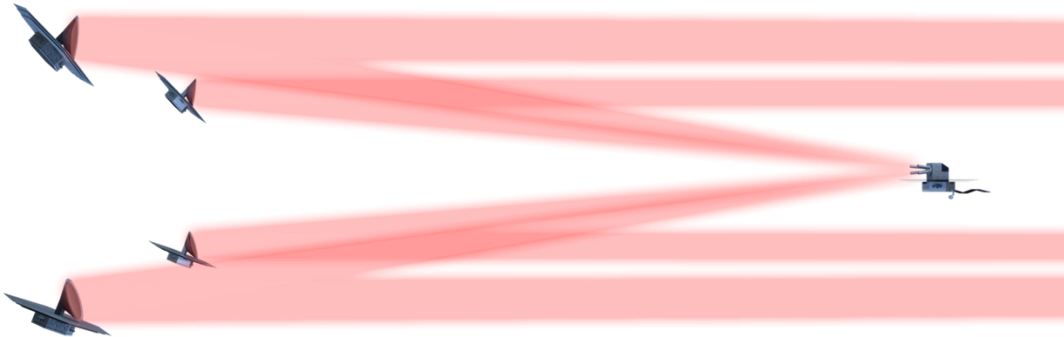


LIFE unleashed – beyond standard processes

We embrace scientific and technological challenges, question established mindsets, and pioneer new ways of collaborations

Goal

Implement today's most transformative space mission designed to search for life beyond the Solar System by 2040

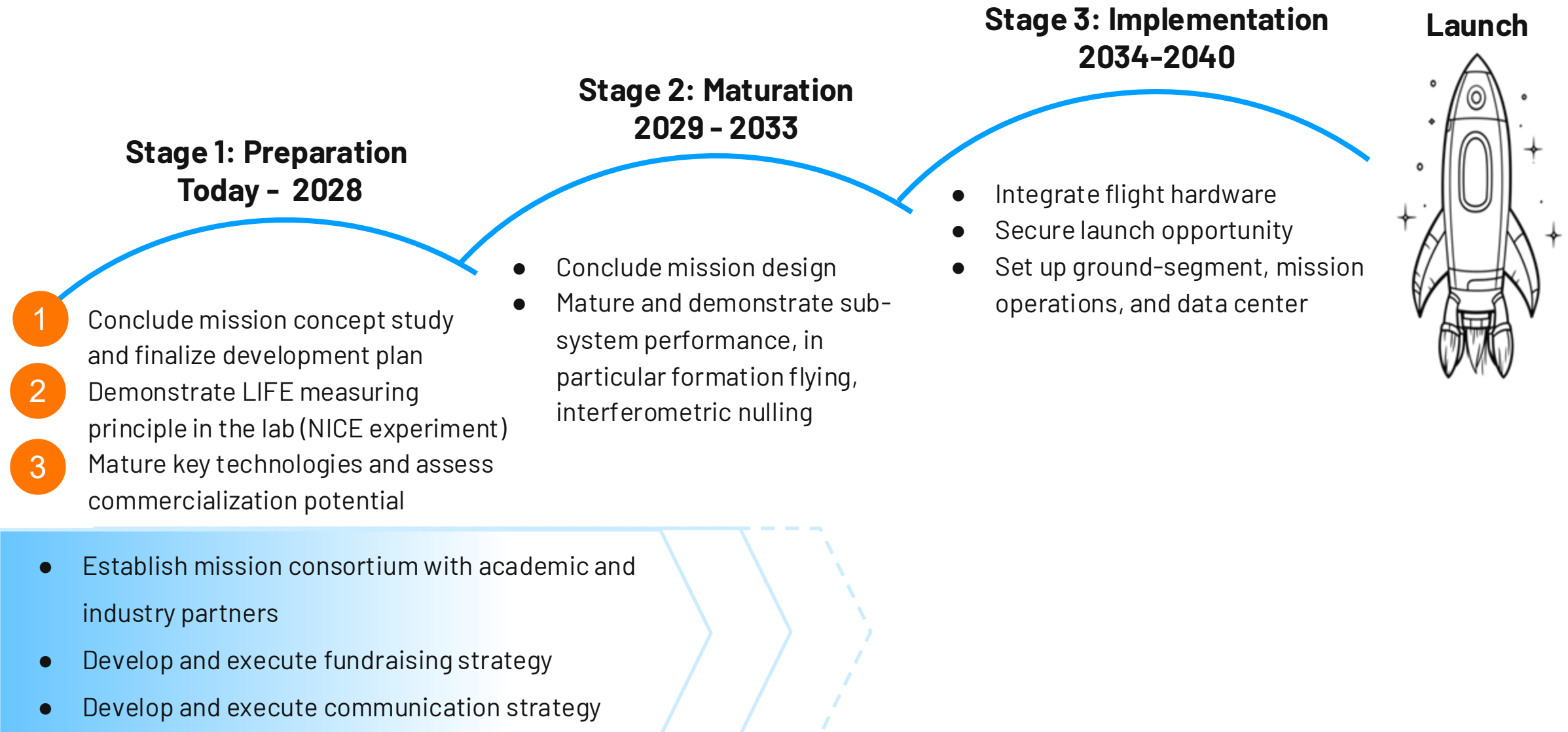


Approach

- Seek funding via donors, foundations, and sponsors
- Seek goal-oriented partnerships in industry and academia to
 - 1) Re-design the development and implementation process
 - 2) Develop and mature key technologies
 - 3) Leverage commercialization potential of technologies and processes (e.g., formation flying platforms, photonics, etc.)

Aiming at a launch in 2040 we consider 3 development and funding stages

INDICATIVE

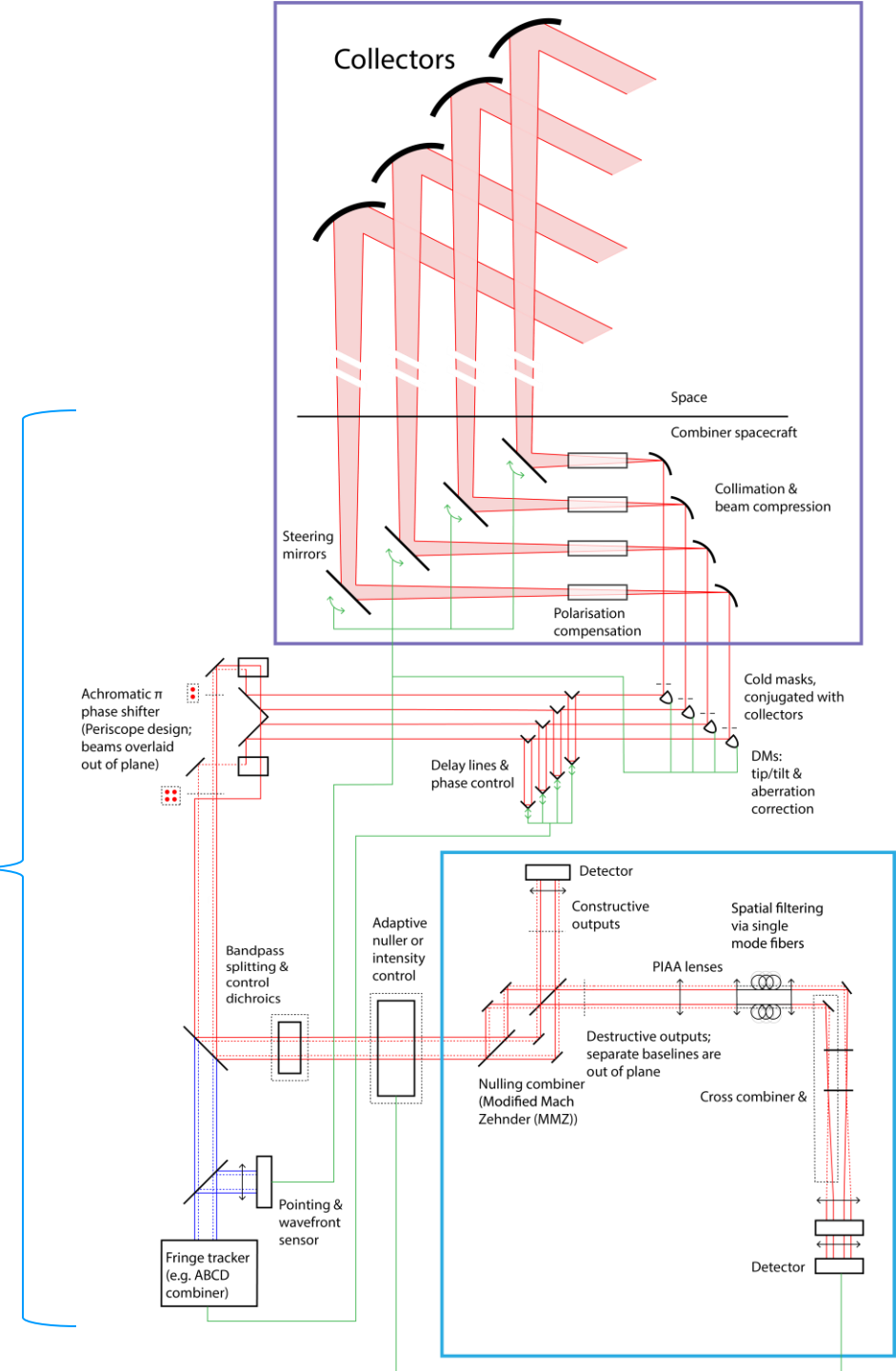
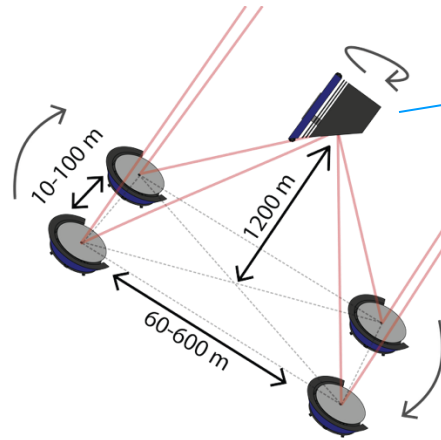


Stage 1: Preparation Today - end of 2028

- 1 Conclude mission concept study and development plan
- 2 Demonstrate LIFE measuring principle in the lab
- 3 Mature key technologies and investigate commercialization potential

Optical Concept

- Key building blocks identified
- Definition of major architecture trades
- Refinement of Mission Requirements
- Pre-studies launched



Stage 1: Preparation Today - end of 2028

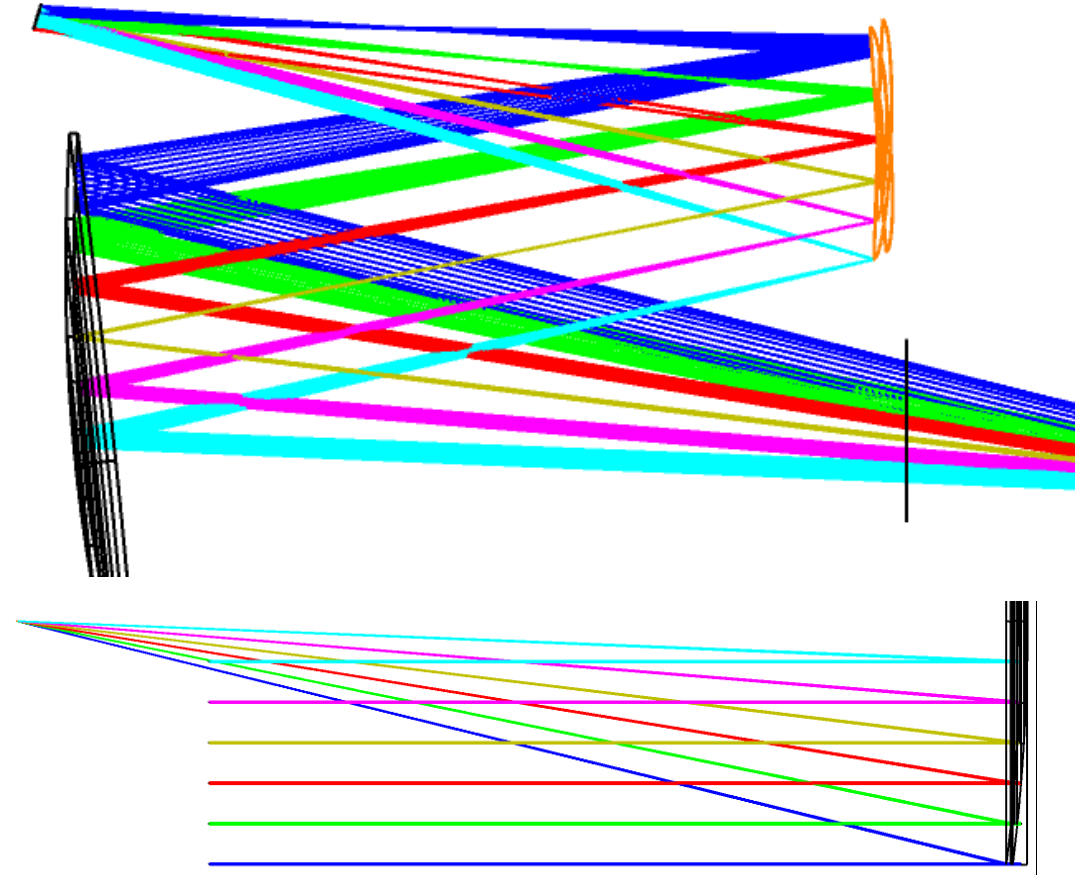
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Collector Optical Design

- Pre-study to understand aberration and pupil shape of spherical M1
- Investigation to correct the pupil distortions by using two freeform mirrors in the beam combiner optics
- Spherical aberrations and astigmatism resulting present in each individual baseline
- → Easily corrected with a Deformable Mirror
- Strong variations in pupil shape for each baseline
- → Much harder to correct



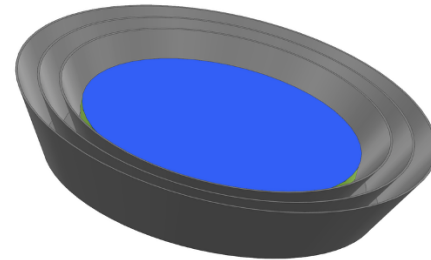
Stage 1: Preparation Today - end of 2028

beyond gravity

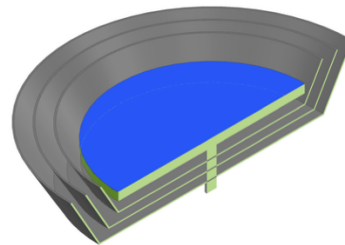
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Project Update: Collector Spacecraft Thermal Control Concept

Thermal Concept Design



CAD Model

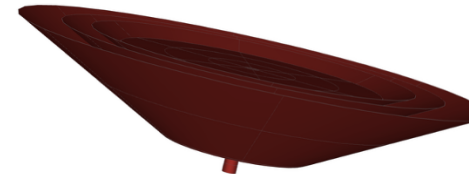
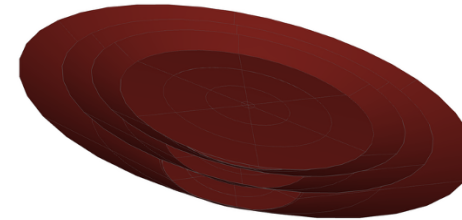


Cross-section view of CAD Model

- BG input: V-groove composition, geometry
- BG input: GFRP strut geometry



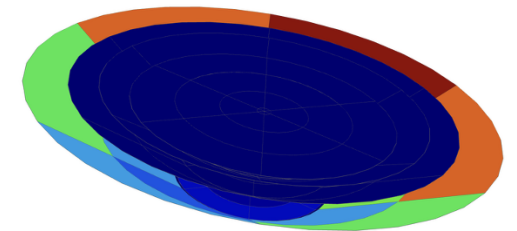
Geometric Simplification



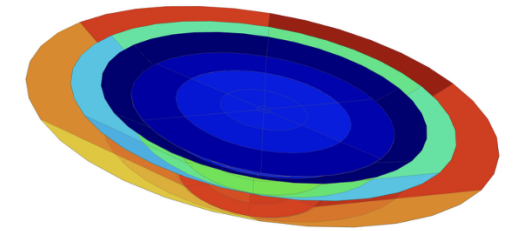
Simplified Model using Shell Elements



Thermal Analysis



Solar Flux



Temperature Distribution
(! first result → needs to be optimized)

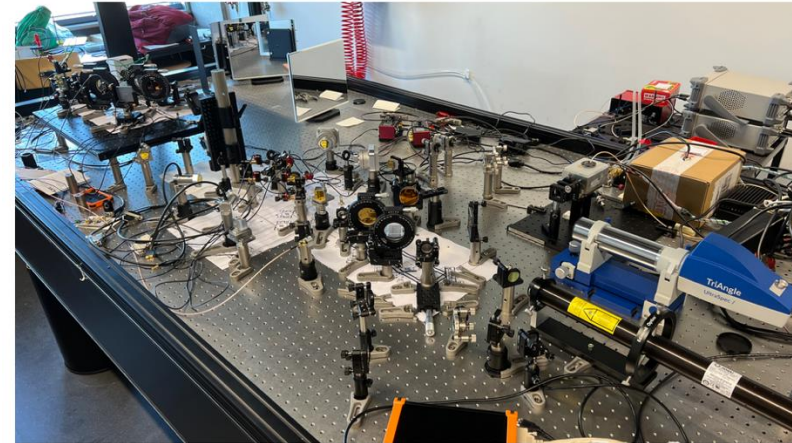
- BG input: material properties (thermal conductivity, emissivity, density, specific heat)

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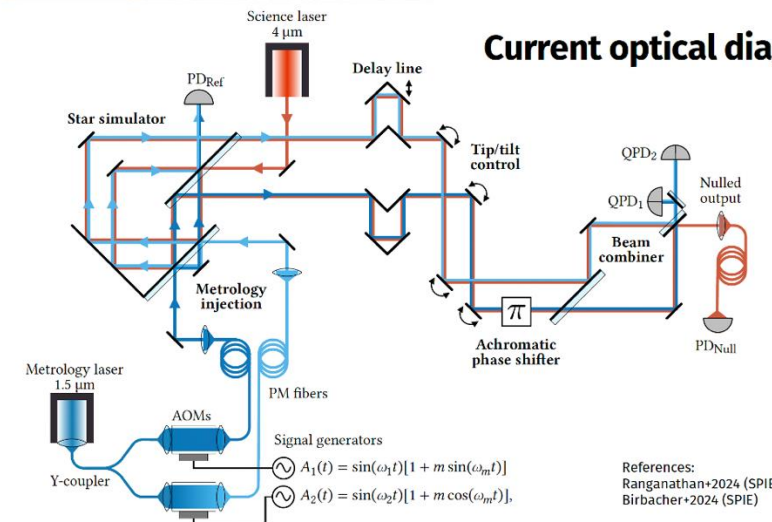
NICE objectives:

Demonstrate a broadband null with a null depth of 10^{-5} and stability of 10^{-8} while maintaining a high system throughput, and consequently a high level of sensitivity



Warm precursor setup of NICE

Simplified Layout



Current optical diagram

References:
Ranganathan+2024 (SPIE)
Birbacher+2024 (SPIE)

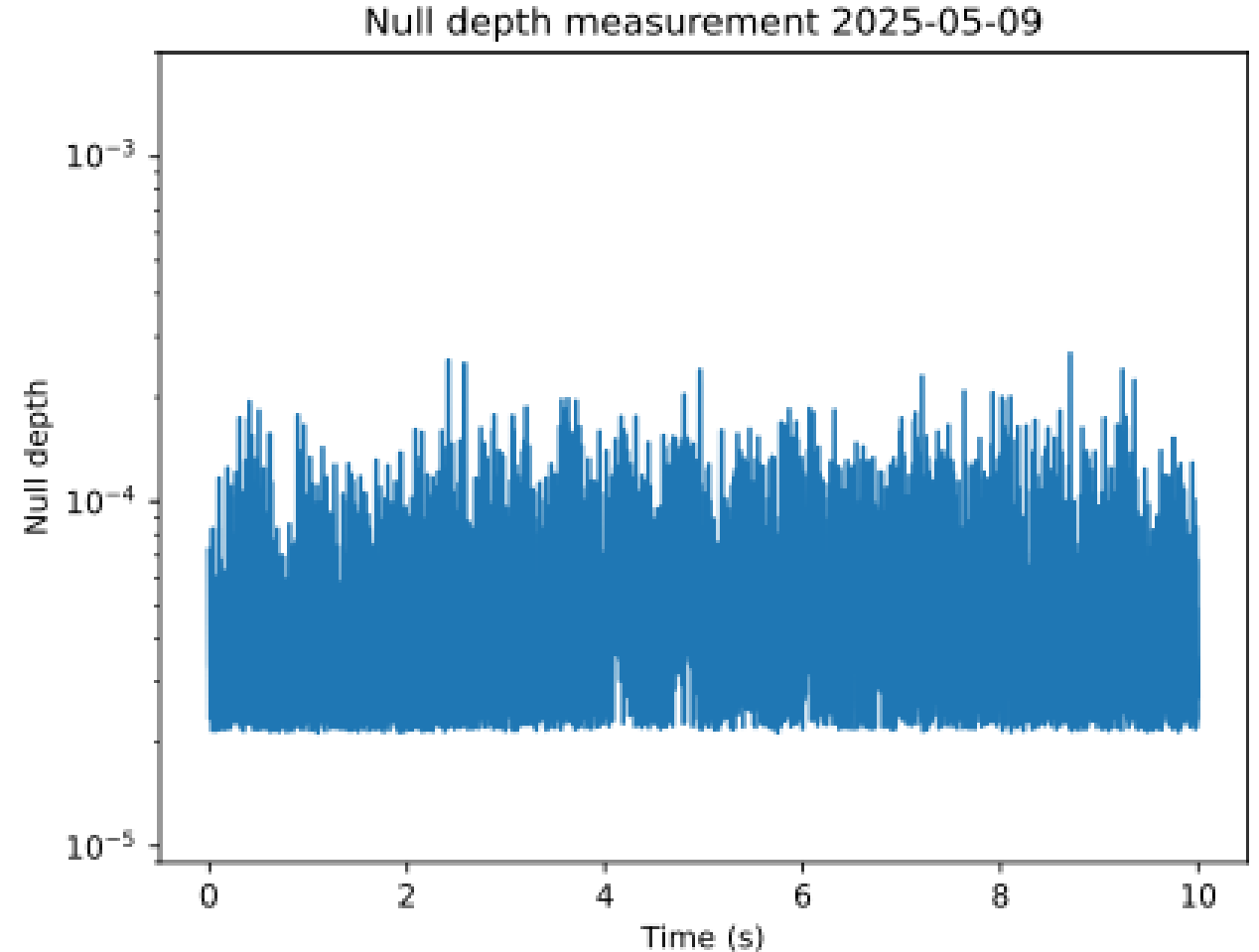
Credit: T. Birbacher, M. Ranganathan

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Status:

- Mean recorded null depth of $<5 \cdot 10^{-5}$, with a minimum around $2 \cdot 10^{-5}$ at 4 micron
- Fast OPD control loop allowing noise to be reduced at the sub nm level up to 1kHz
- Bench upgrade in progress to enhance stability and allow for broadband operation.



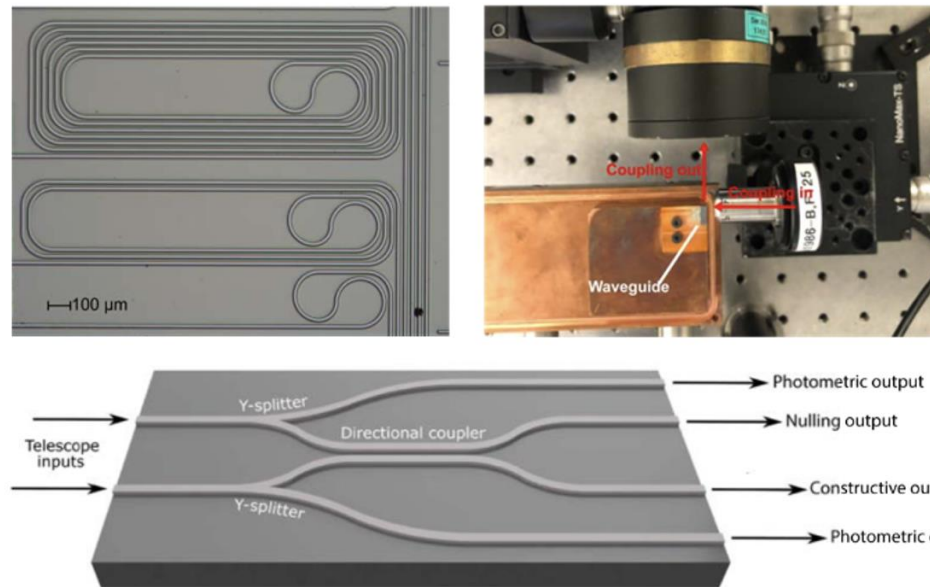
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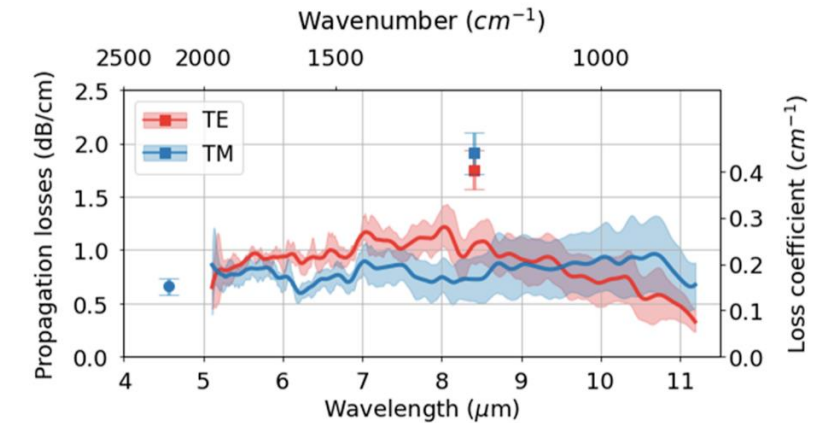
Development of InGaAs waveguides with InP cladding and Fe doping

MPIS4LIFE Collaboration with Institute for Quantum Electronics (Prof. J. Faist, Prof. R. Grange)

MIR waveguides / integrated optics



Measured transmission losses (Montesinos, 2024)



Stage 1: Preparation Today - end of 2028

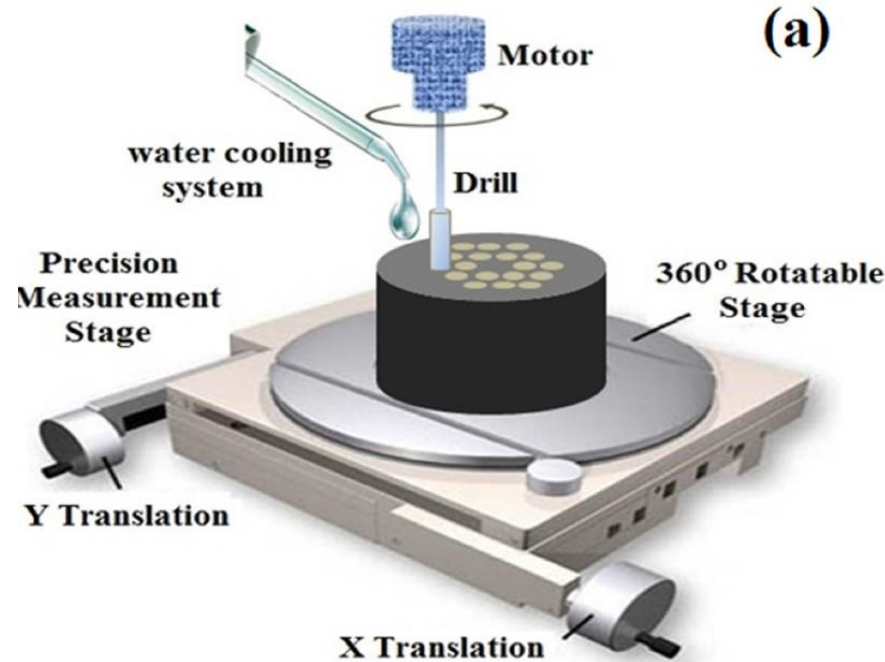
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Spatial Filters - Phase Induced Amplitude Apodization in combination with Photonic Crystal Fibres

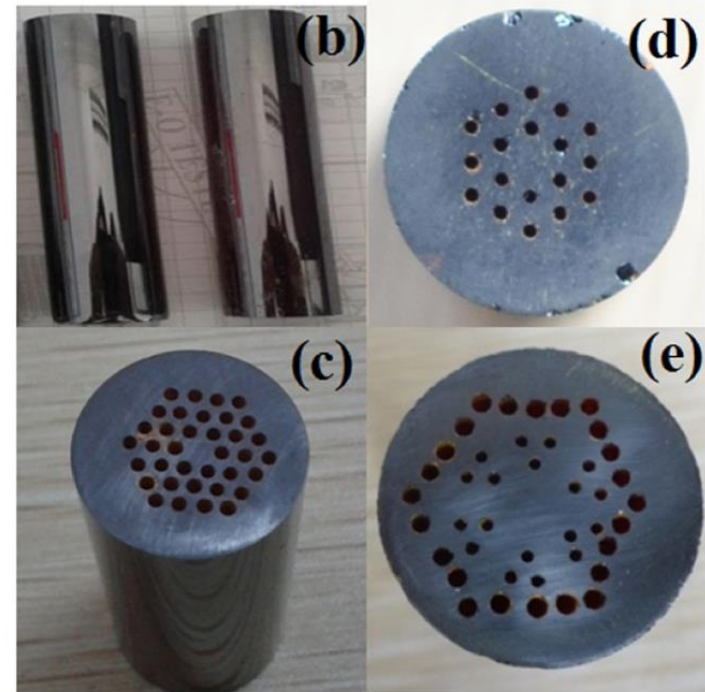
M. Ireland, ANU, Australia



Australian
National
University



(a)



(b)

(d)

(c)

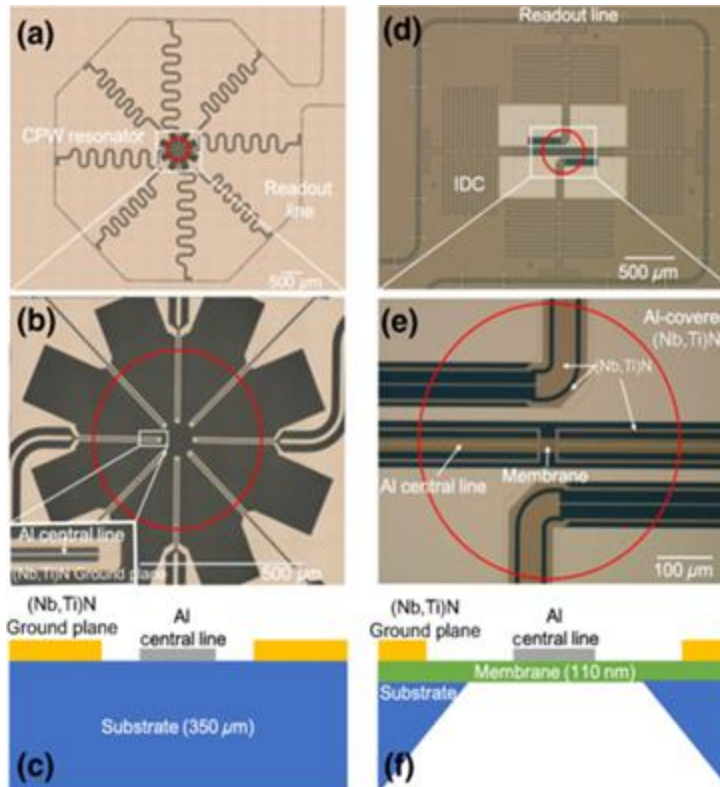
(e)

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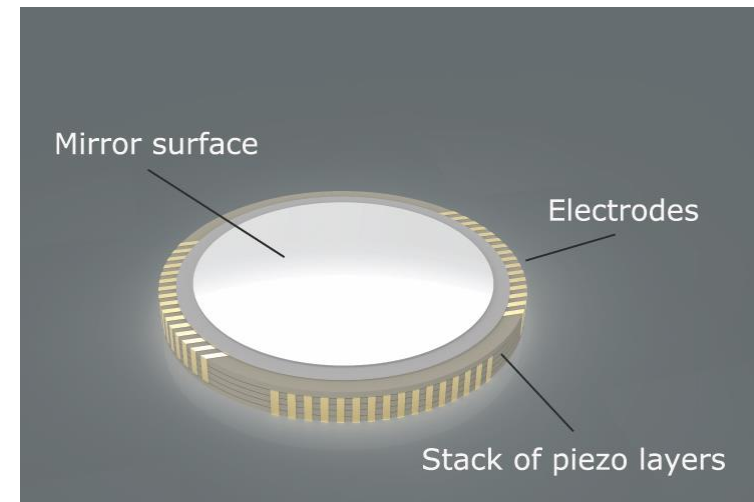
MKID detectors

P. de Visser (SRON)



Cryogenic DMs

Robert Huisman (SRON)



SRON

Join and engage in LIFE-related activities

Bottom-up engagement in LIFE Science community

Dutch/Belgium LIFE Days September 2024



Discussions for Italian LIFE Days are ongoing



ELEONORA ALEI

Science Team Lead
NASA GSFC



TIM LICHTENBERG

Science Team Lead
University of Groningen

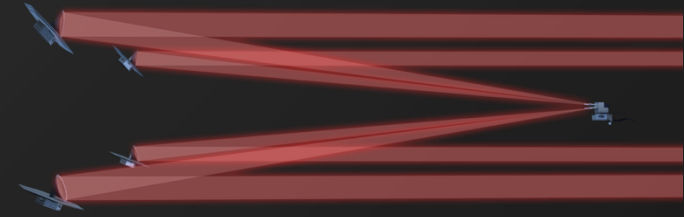
Monthly LIFE Science
community calls

Abstract submission open!
Deadline: August 31



Information / contact:
www.life-space-mission.com





The LIFE mission

*We inspire and unite humankind in the way we perceive ourselves and our home planet Earth.
We encourage current and future generations to pursue bold visions.*

*We embrace scientific and technological challenges, question established mindsets, and
pioneer new ways of collaborations.*

*We develop and implement today's most transformative space mission to search for life beyond
the Solar System and investigate the diversity of other worlds.*



Backup