

The Euclid MERge Processing Function

Erik Romelli, INAF – OATs
on behalf on the OU-MER team

7° Meeting Nazionale della collaborazione Euclid
Bologna, 30-06-2025

OU MER team



WP5200 OU - MERa – OATS

- Erik Romelli (**Lead**)
- Samuele Galeotta (**Dev-Lead**)
- Marius D. Lepinzan

WP5300 OU – MERb – OARoma

- Marco Castellano
- Emiliano Merlin
- Fernando Caro
- Thomas Gasparetto
- Adriano Fontana
- Paola di Mauro



- Martin Kuemmel (**Lead**)
- Daniel Hernández Lang
- Javier Gracia Carpio



- Herve Dole (**Co-Lead**)
- Elie Subrié
- Loic Morin
- Mermiam Ezziati
- Remi Cabanac



- Marc Huertas Company
- Hui Kong
- Tyann Dumerchat



- Mike Welmsley

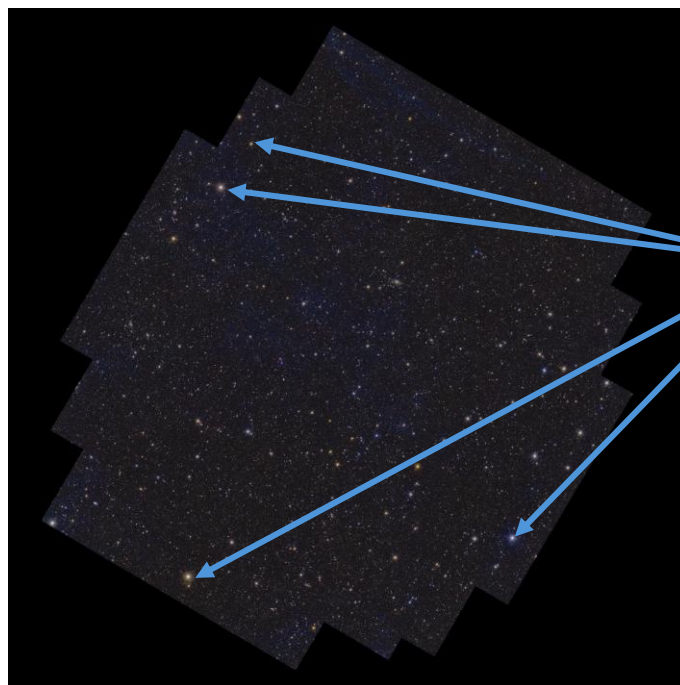


- Yuedong Feng

OU MER definition and summary

MER is the **cataloguing unit** of the *Euclid* SGS

⇒ Responsible for the creation of the **official source list**



Euclid unique ID

Coordinates (R.A., Dec)

Fluxes (I_E , Y_E , J_E , H_E , g , i , r , u , z)

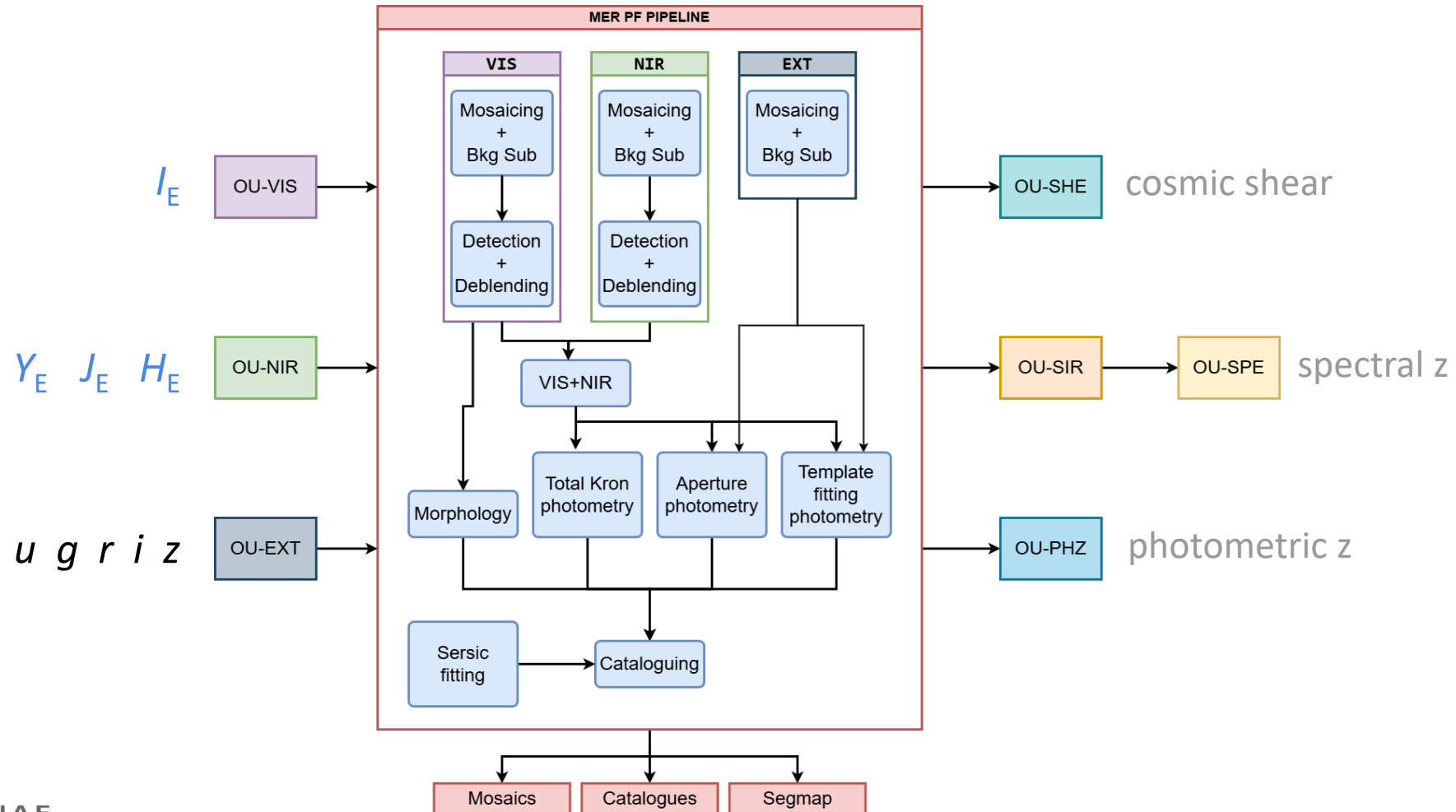
Morphology

Flags

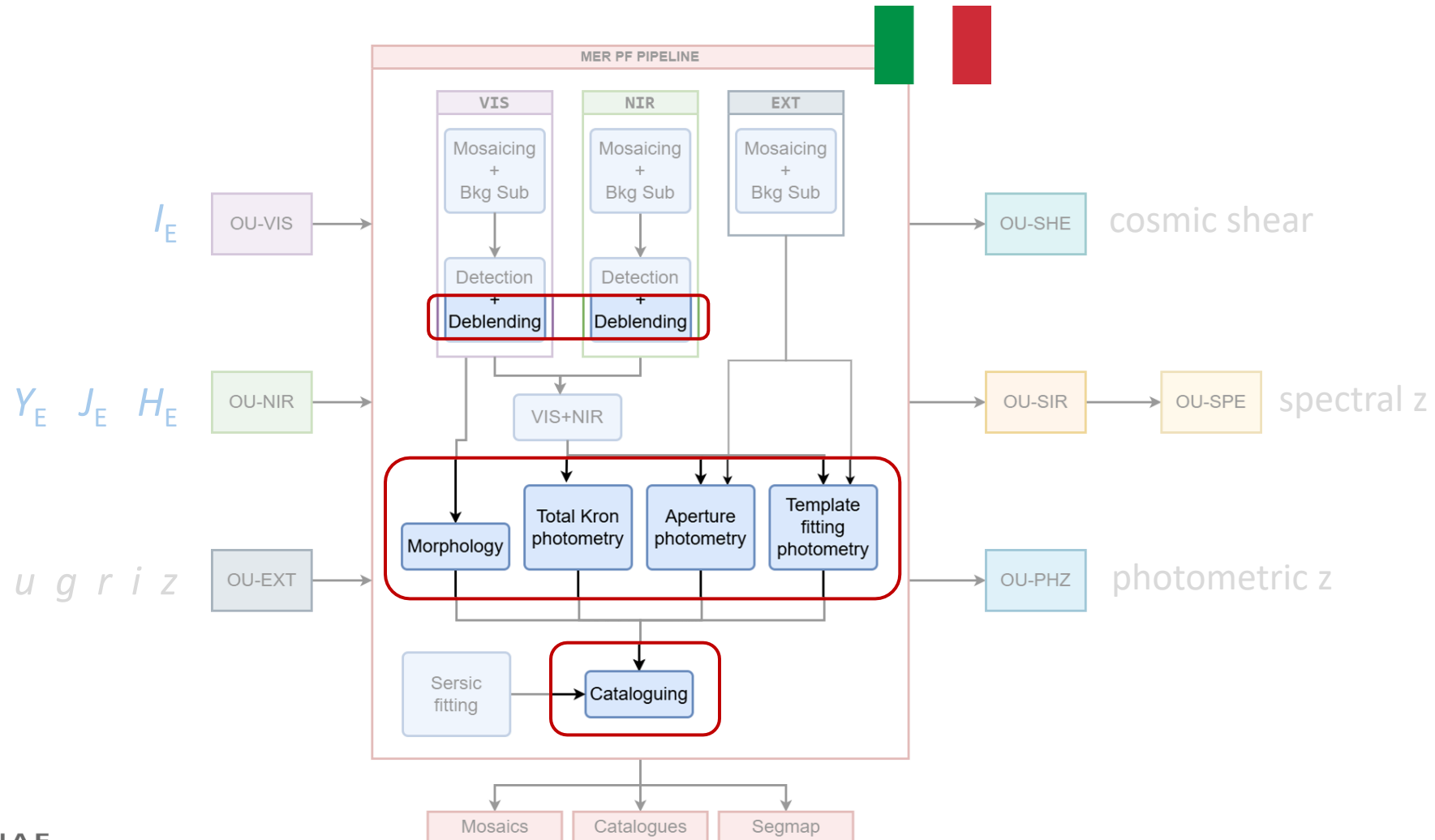
Credits: Esa/Euclid/Euclid Consortium/Nasa

Image processing by J.-C. Cuillandre, E. Bertin, G. Anselmi

Pipeline overview



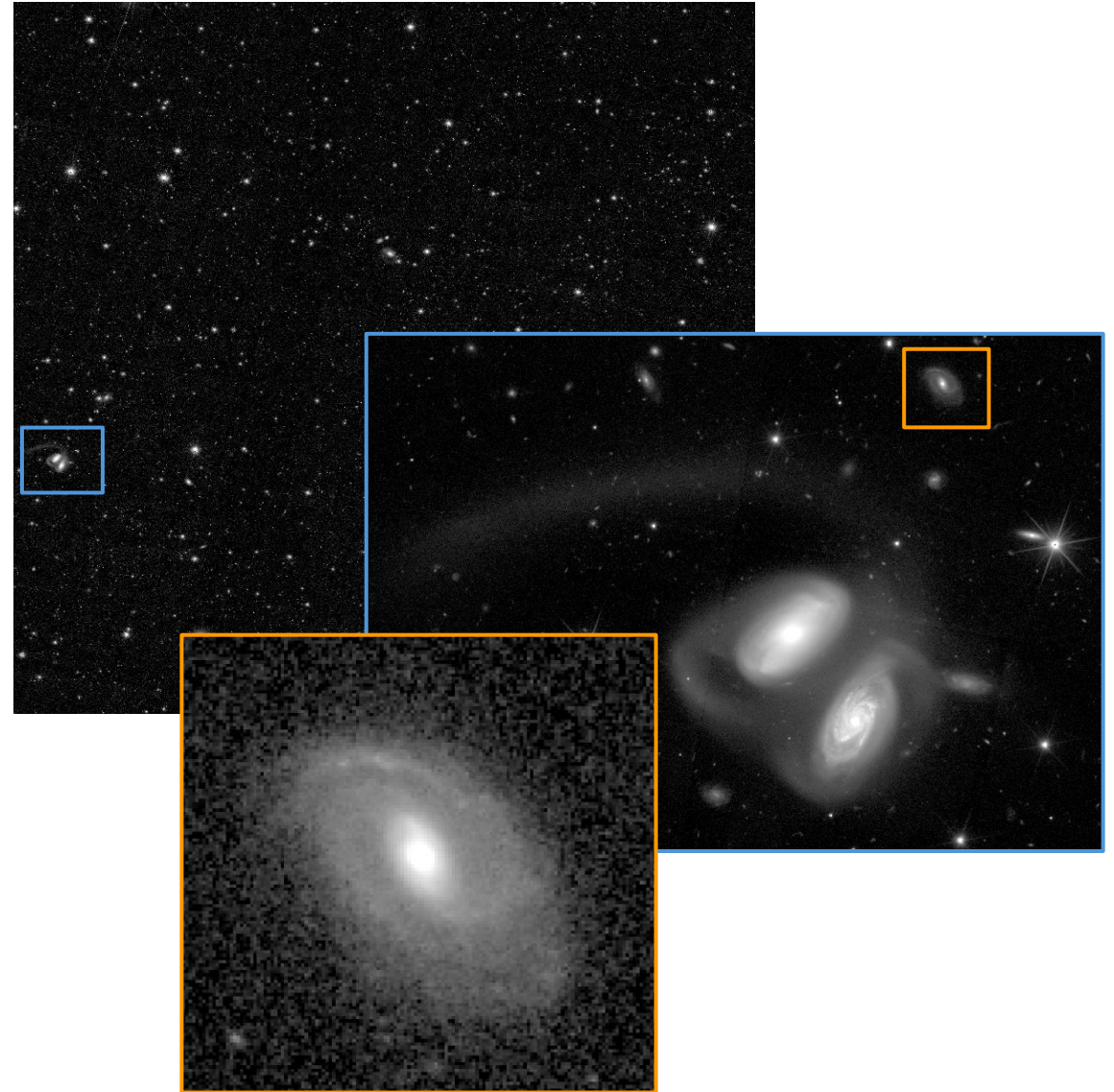
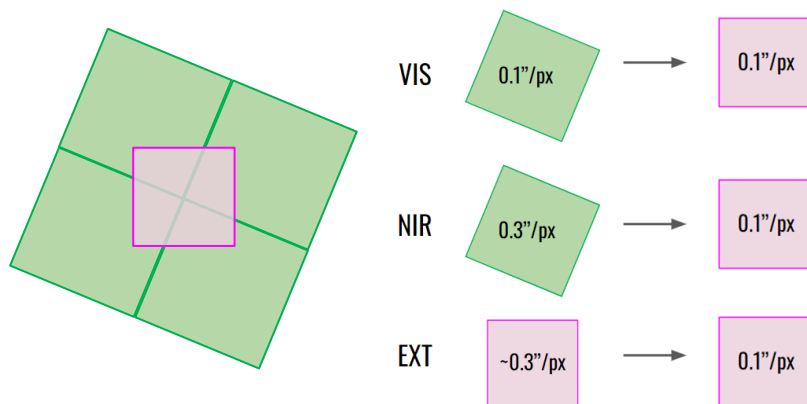
Pipeline overview



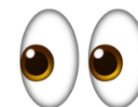
Products

Mosaics

- One product for each band
- Common pixel scale
- Common WCS



Catalogues



Hands-on session – Wed 2 @ 16:00
Overview of the OU-MER Data Model

- One product for each tile
- ~120k sources
- 1000+ columns
- Complete description: [Data Product Description Document](#)

Photometry

- Kron total flux
- PSF-matched aperture photometry
- Template fitting photometry
- Sersic model fitting photometry

Morphology

- Non parametric (Concentration, Asymmetry, Smoothness)
- Sersic Model Fitting morphology
- Machine Learning morphology

Photometry Cookbook

Photometry

- OBJECT_ID
- RIGHT_ASCENSION
- DECLINATION
- DET_QUALITY_FLAG
- FLUX_DETECTION_TOTAL
- FLUX_<band>_<n>FWHM_APER
- FLUX_<band>_TEMPLFIT
- FLUX_VIS_TO_<band>_TEMPLFIT
- FLUX_<band>_SERSIC

■ Colors

$$H-g = -2.5 \log_{10} \left(\frac{\text{FLUX_H_2FWHM_APER}}{\text{FLUX_EXT_G_DECAM_2FWHM_APER}} \right)$$

■ Total Fluxes

$$F_{H,\text{tot}} = \text{FLUX_DETECTION_TOTAL} \left(\frac{\text{FLUX_H_2FWHM_APER}}{\text{FLUX_VIS_2FWHM_APER}} \right)$$

■ Magnitudes

$$m_{\text{AB}} = -2.5 \log_{10} (F_{\mu\text{Jy}}) + 23.9$$

Examples using Aperture photometry

Processing status

SGS processing milestones

Quick Release 1 (March 2025)

[Euclid Collaboration: Romelli et al](#)

Q1 data come with some *caveats*:

[MER Release Notes](#)

v11.0.1

355 tiles - 63.1 deg²

Data Release 1 (Late 2026)

No major updates w.r.t Q1

Bug fixes and performance improvements

WIDE + DEEP processing

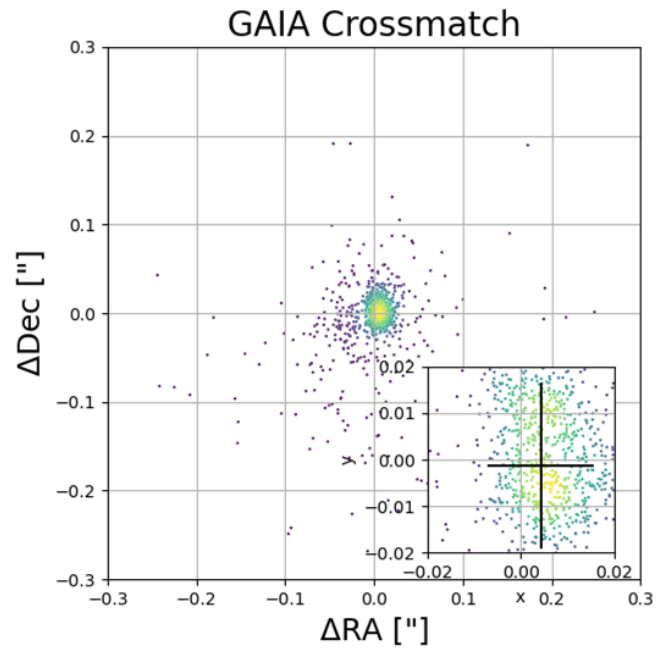
v11.3.1

~10k tiles - ~1900 deg²

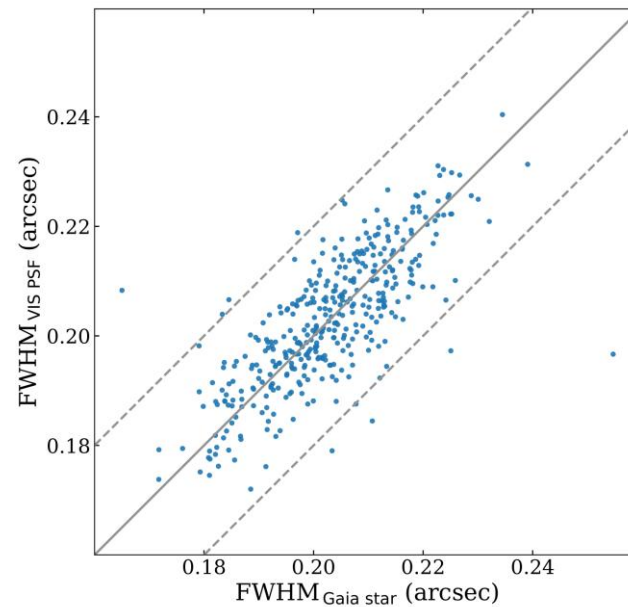
Gaia Validation

Validation of one MER tile

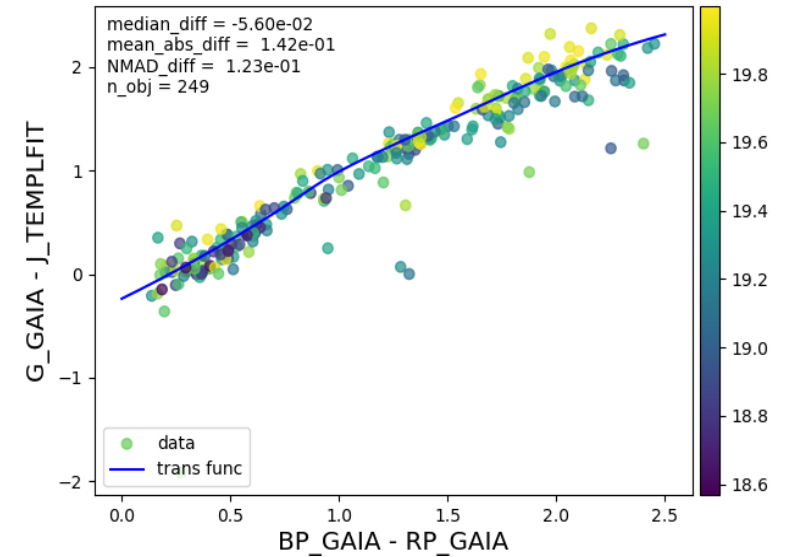
Astrometry



PSFs



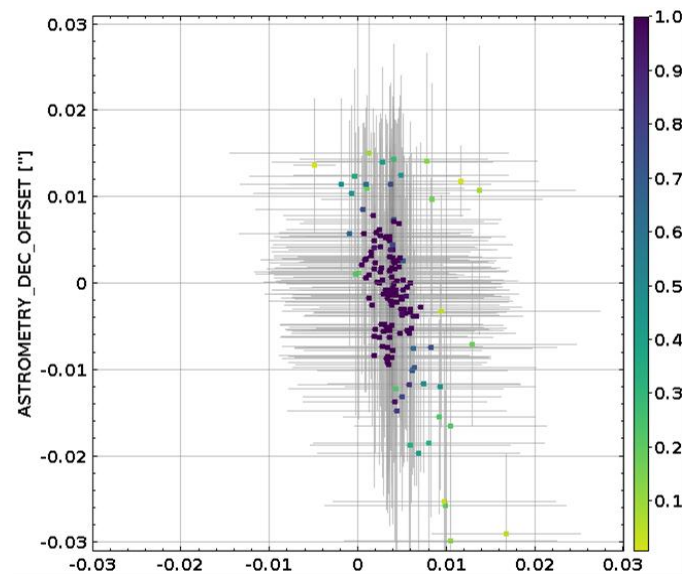
Photometry



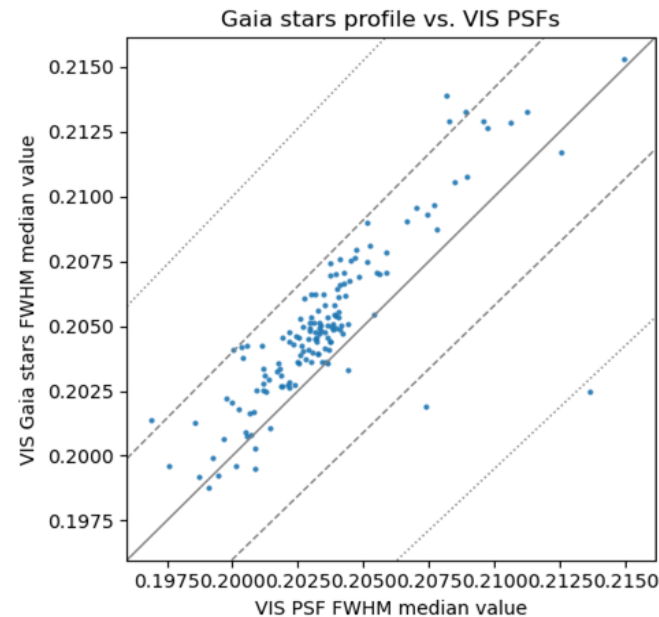
Gaia Validation

Validation of a Skyspatch (collection of tiles) – EDF-S from Q1

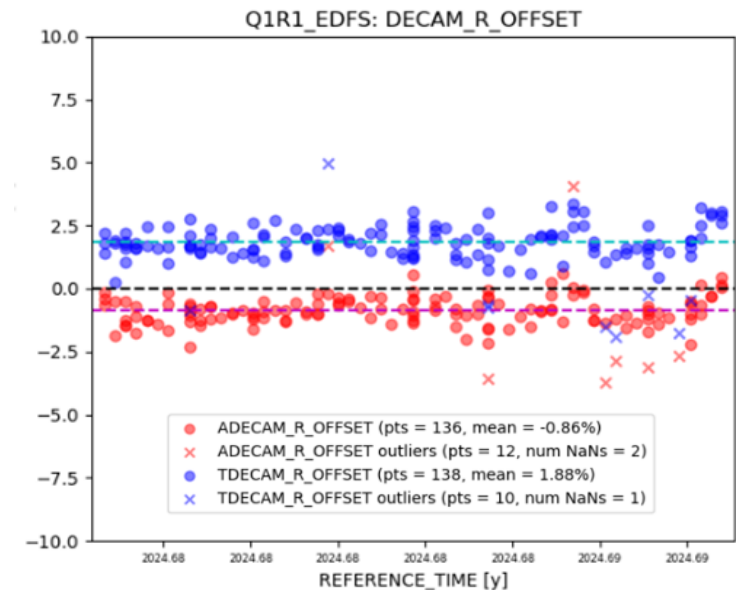
Astrometry



PSFs



Photometry



Towards DR2

DR2 planned updates and new features

Photometry and flagging

Validation of extended objects

Improve/update of
POINT_LIKE_PROB and
SPURIOUS_PROB

Improve bright star masking

Morphology

Improve (if possible) CAS
measurements

Disk+Sersic fitting

Pipeline

Re-design of the DEEP processing

RMS with and w/o Poisson noise
in all bands

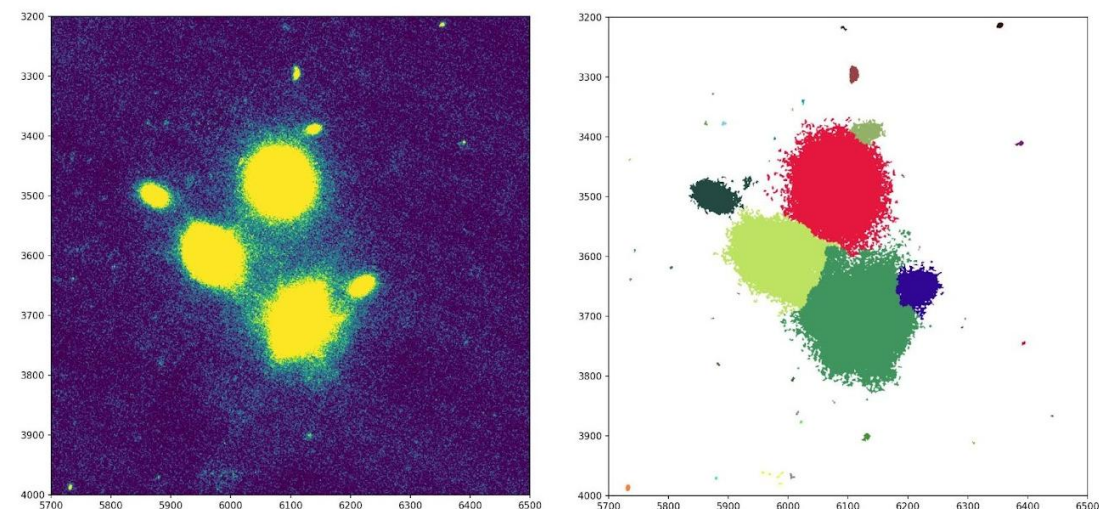
New deblending algorithm

Parent Redmine issue: [#31676](#)

Deblending with ADP

- Advanced Density Peak (ADP) identifies cluster centers, borders, and performs hierarchical merging in the density probability (3D clustering)
- Natural extension to structured grids by treating neighboring pixels as nearest neighbor (2D deblending)
- Tests on-going on synthetic segmentation maps and Q1 data

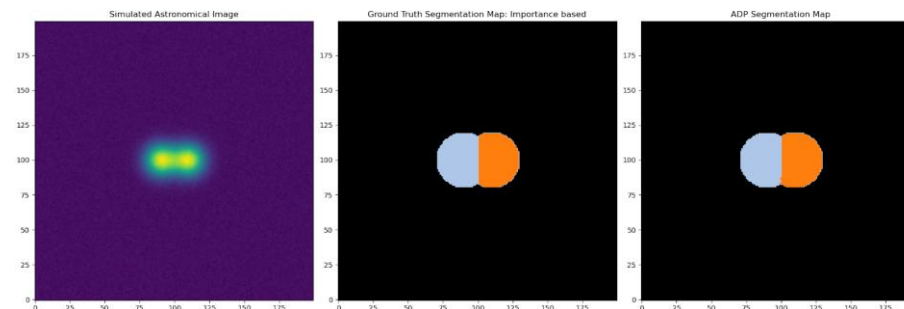
Courtesy of M.D. Lepinzan & F. Tomba



Simulated image

Sim. segmap

ADP segmap



Q&A