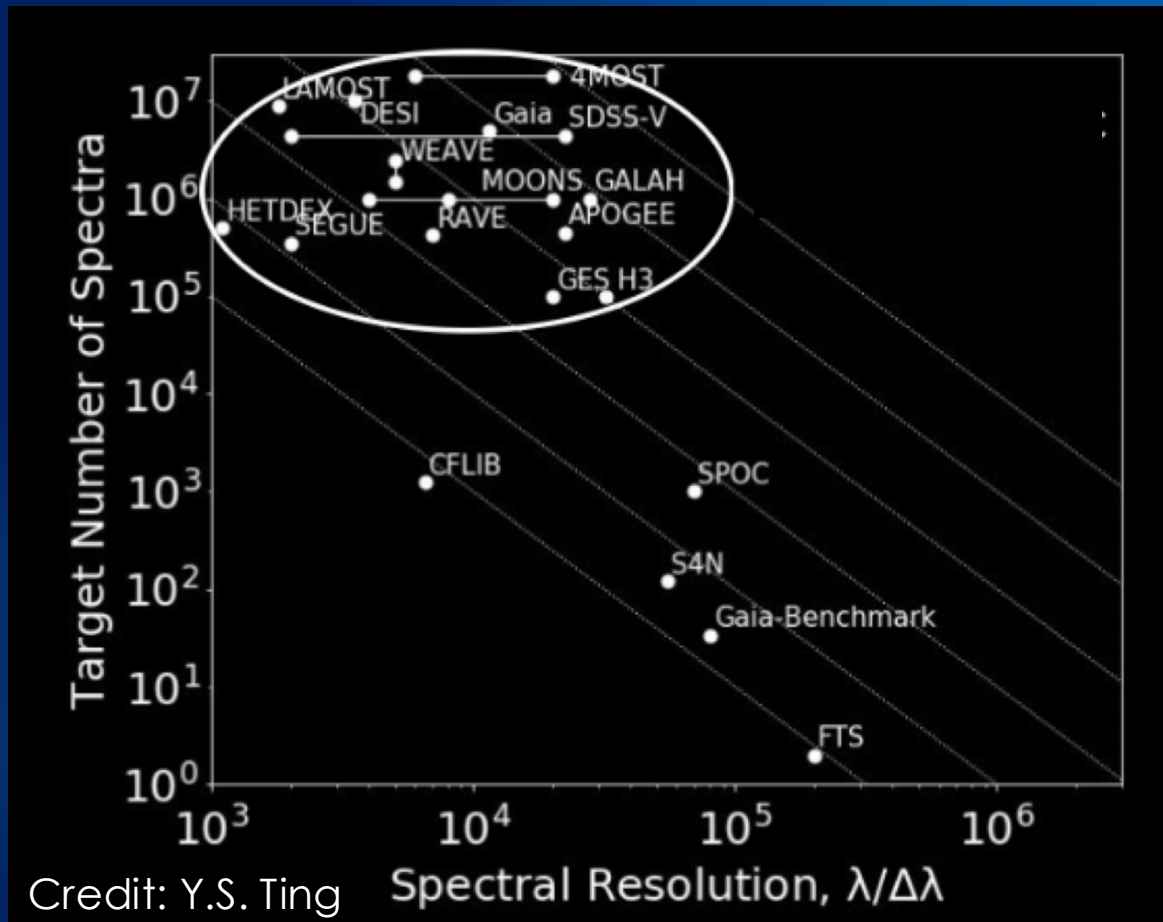


The evolving role of AI in astronomical spectroscopy and future telescope

G. SACCO, N. CANDEBAT, F. BELFIORE, M. GINOLFI, S. ZIBETTI

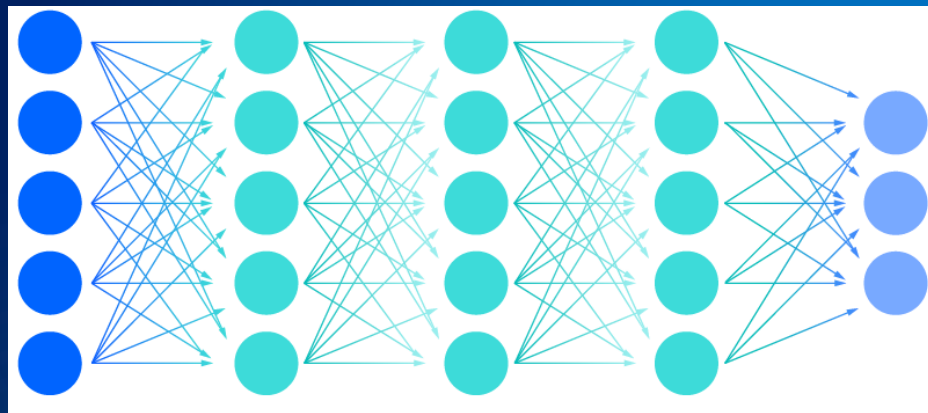
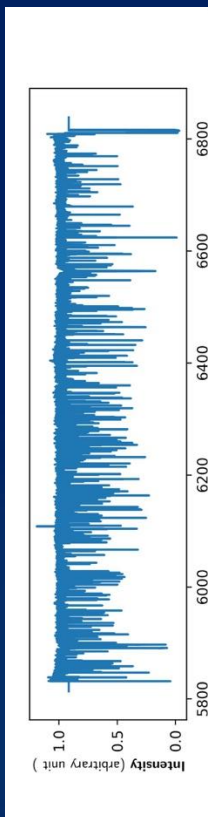
The need of AI for astronomical spectroscopy



Large amount of data
+
Complex physical models and
multiple parameters
+
Effects of atmosphere and
instruments

Traditional methods
for spectral analysis
are not adequate

First steps of AI in astronomical spectroscopy



[e.g. Ness+2015, Ting+2019,
Keown+2019, Ambrosch+2023]

Teff,
Logg,
[Fe/H],
etc..

Strengths

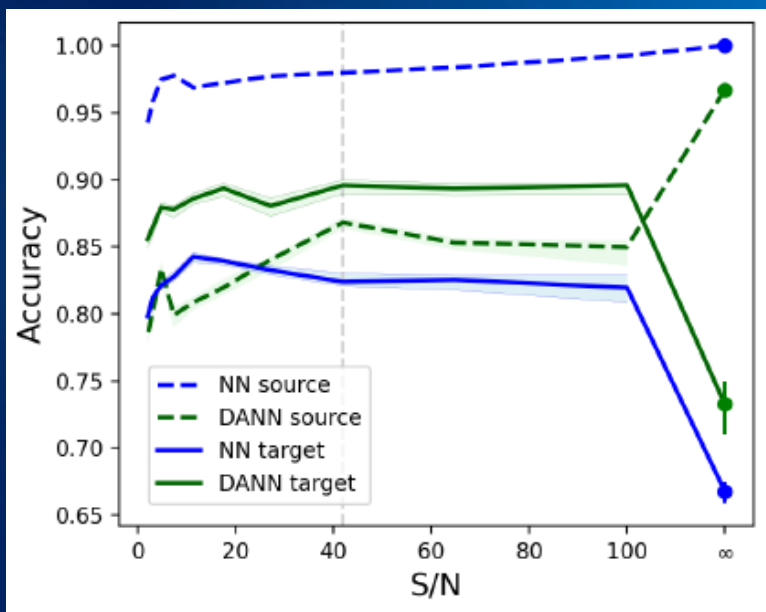
- Much faster
- Better with low res spectra
- Better with Low SNR spectra

Weaknesses

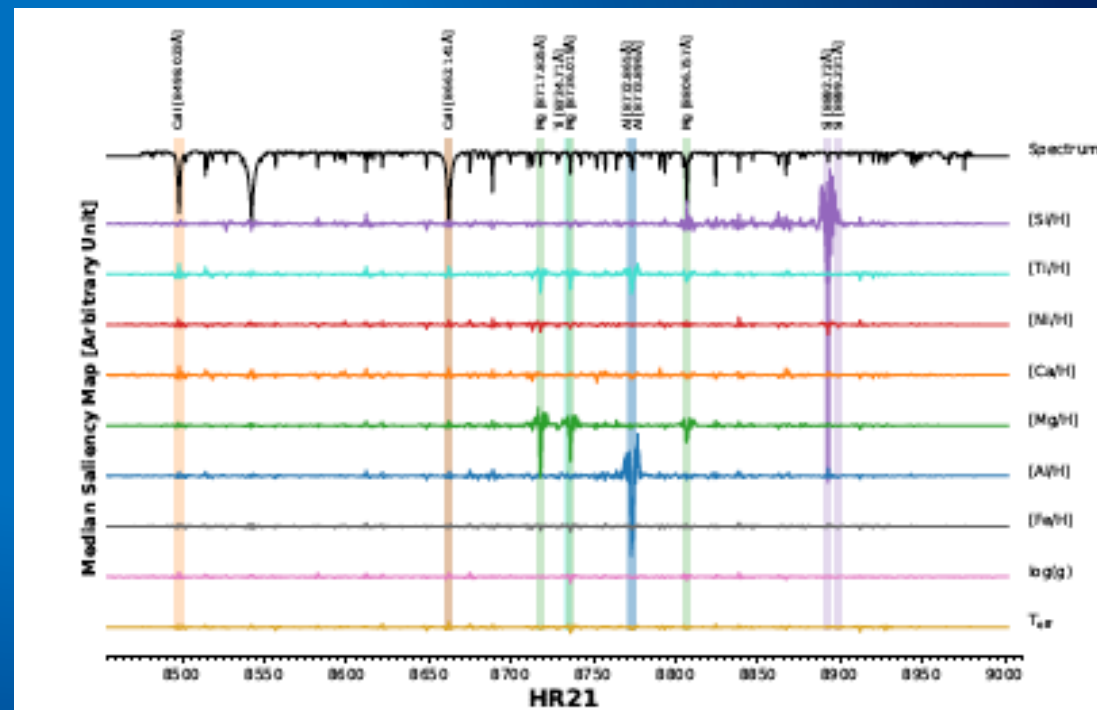
- Error estimates
- Interpretability
- Results depend on the training sample

Scientific AI for astronomical spectroscopy

Domain adaptation to use models for training
(E.g. O'Brian+2021, Belfiore+2025)

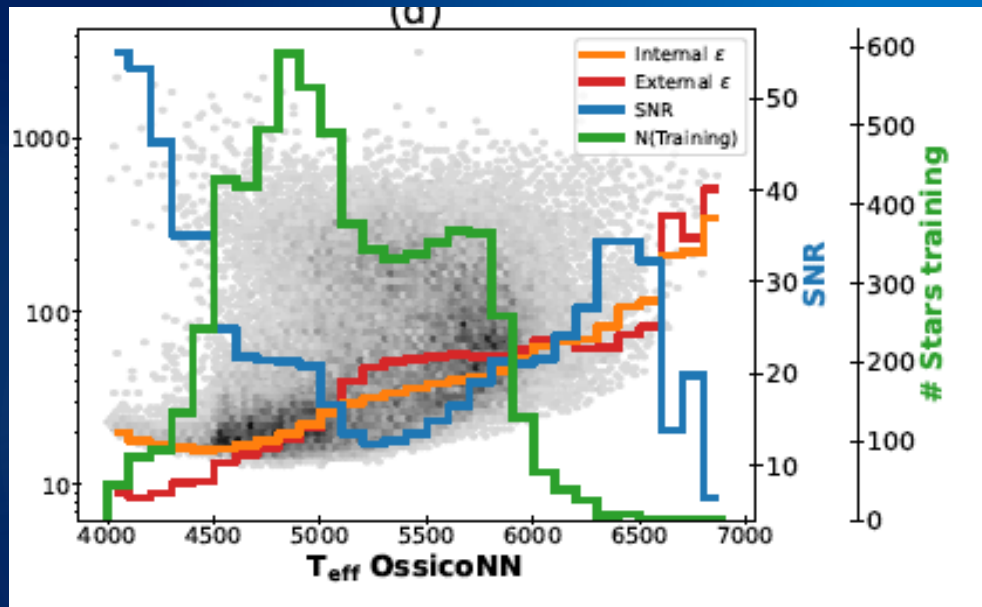


New tools for interpretability
(e. g., Ambrosch+2023, Candebat, Sacco+2024)



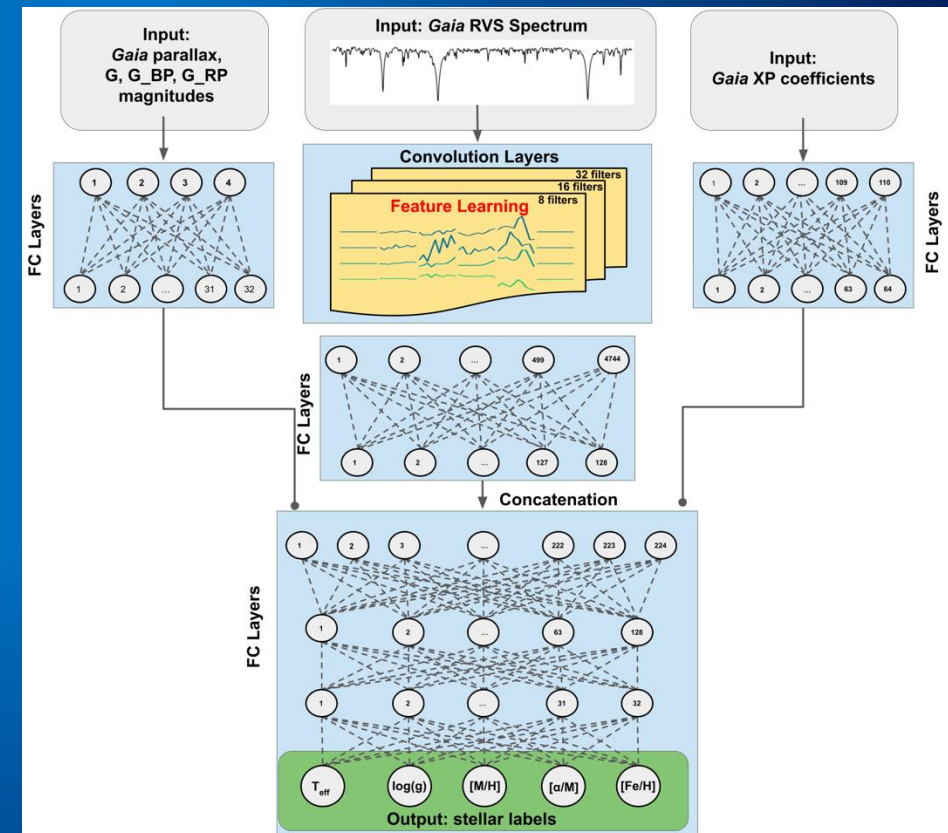
Scientific AI for astronomical spectroscopy

Normalizing flows to calculate
posterior distribution
[e.g. Kang+2022, 2023, Iglesias-
Navarro+2024 Candebat+ 2024]

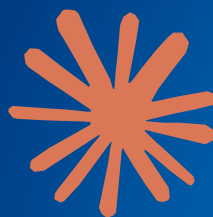
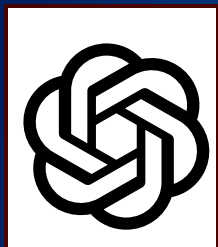


Candebat+2024

Combination of photometry and
spectroscopy
[Guiglion+2024, Candebat+, in prep]



AI 2.0: foundation models



Specific and small dataset

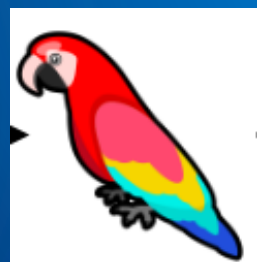


Very large Dataset



LLM Architecture

Pretraining



Foundation Model

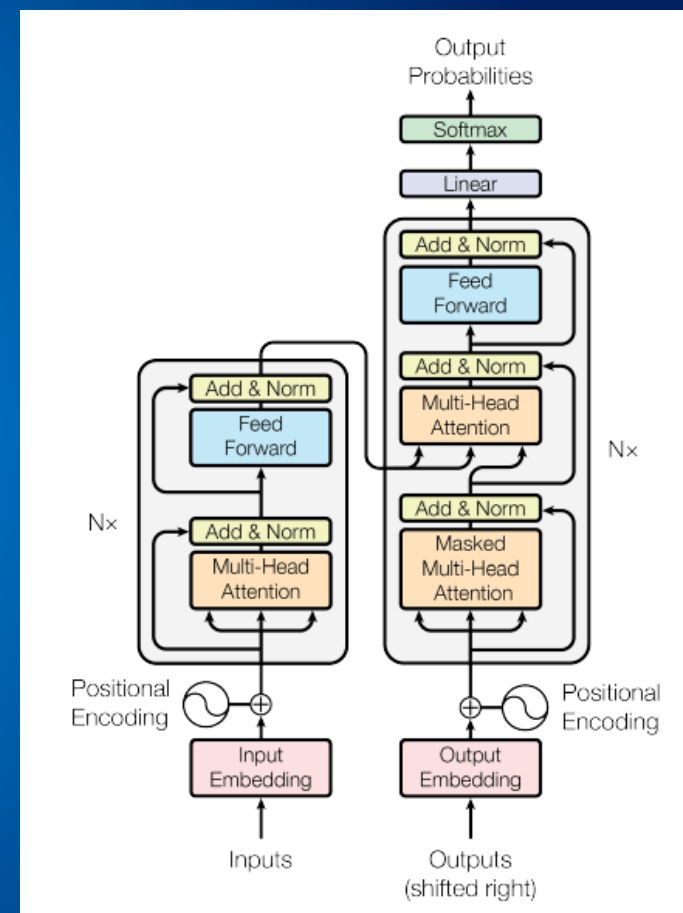


Finetuning



Finetuned LLM

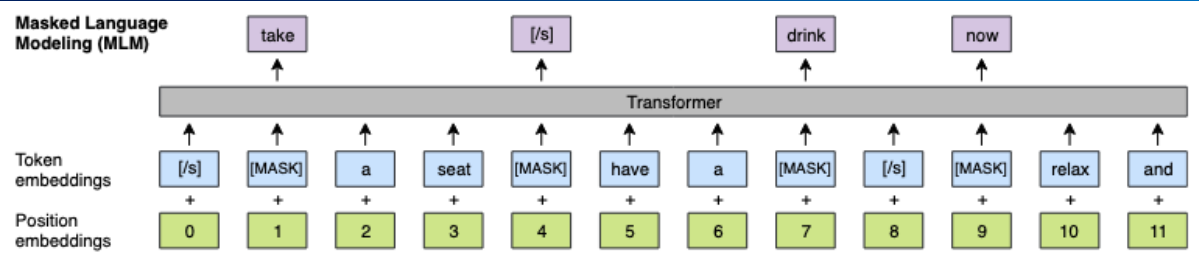
Attention is all you need



Vaswani+2017

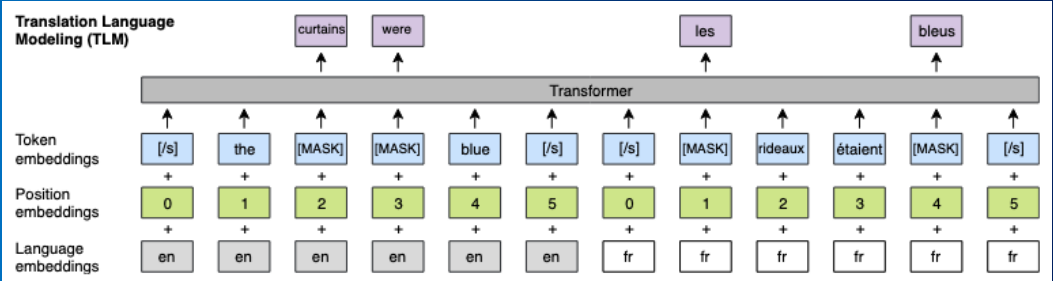
AI 2.0: foundation models

Self-Supervised Learning pre-training

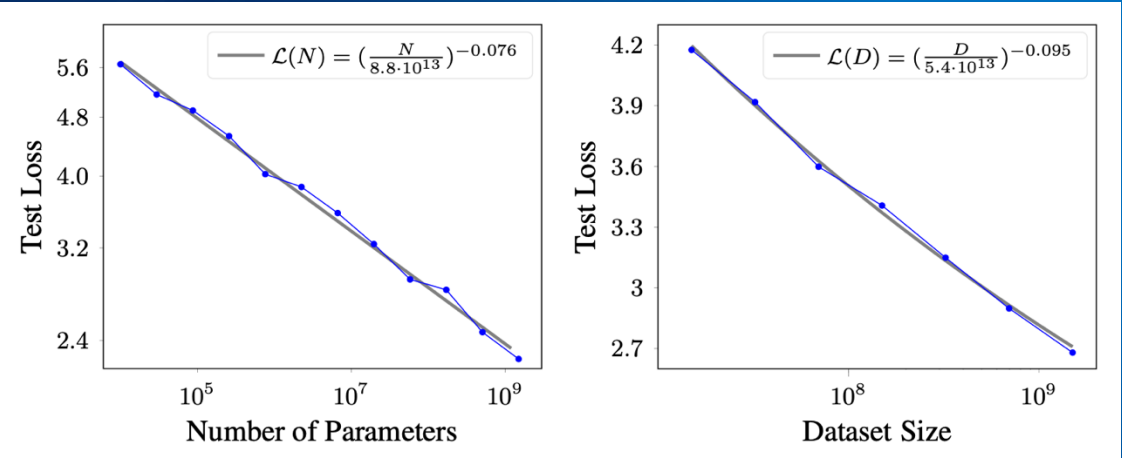


Lample & Conneau 2019

Cross-Domain Generalization

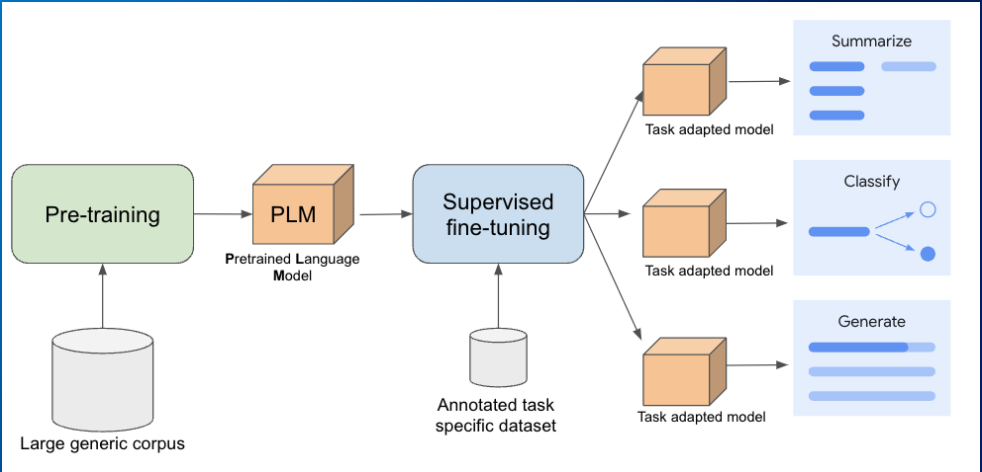


Scaling Law



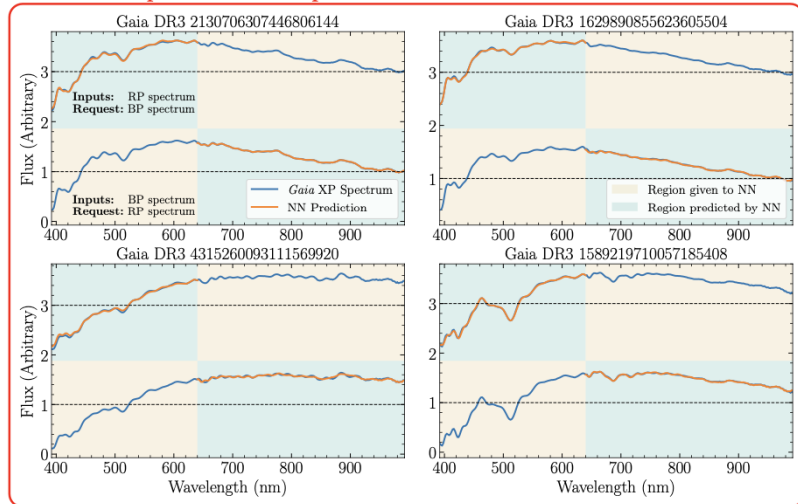
Xio&Zhu 2025

Versatility

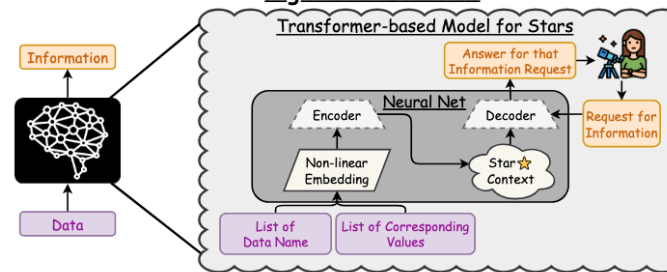


Foundation models in astronomy

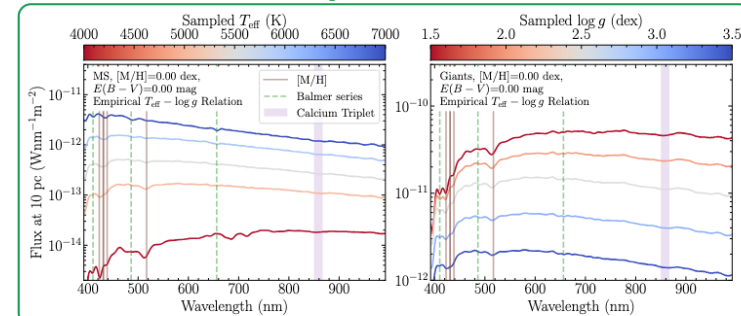
Task: Stellar Spectra to Stellar Spectra



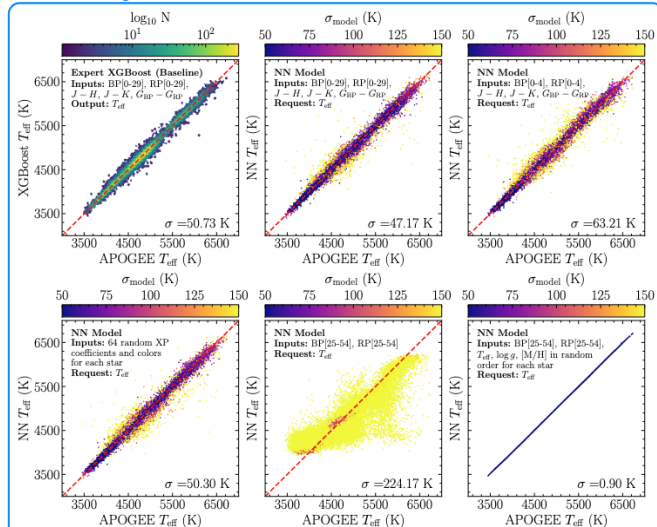
High-level Overview



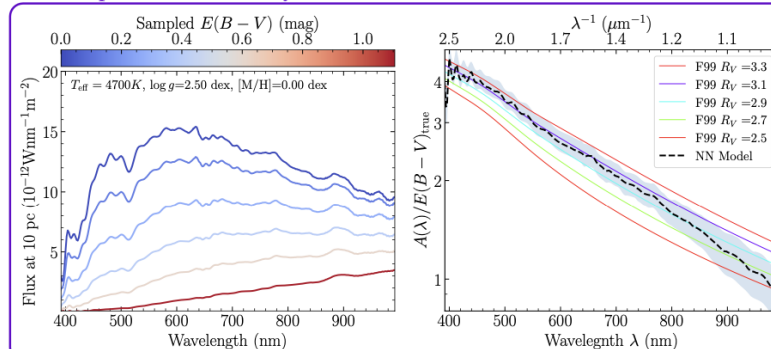
Task: Stellar Parameters to Stellar Spectra



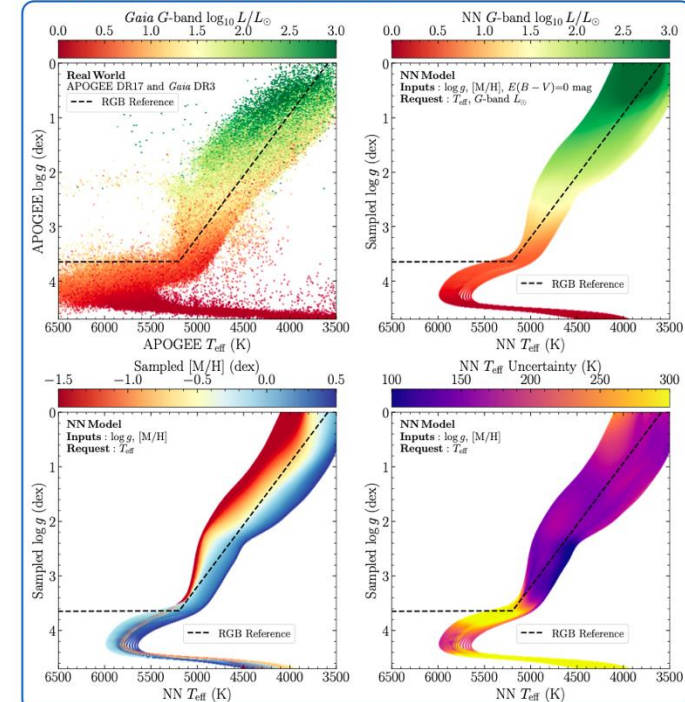
Task: Stellar Spectra to Stellar Parameters



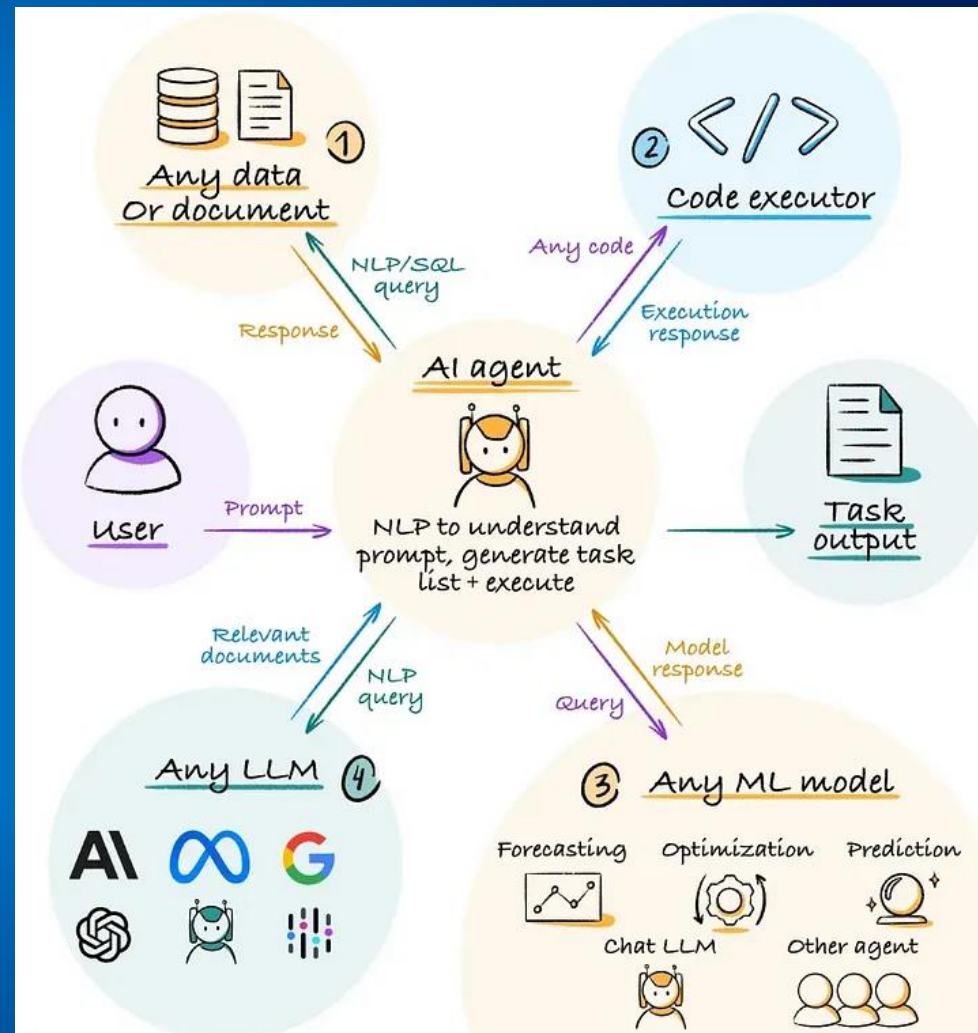
Task: Empirical Law Recovery



Task: Stellar Parameters to Stellar Parameters

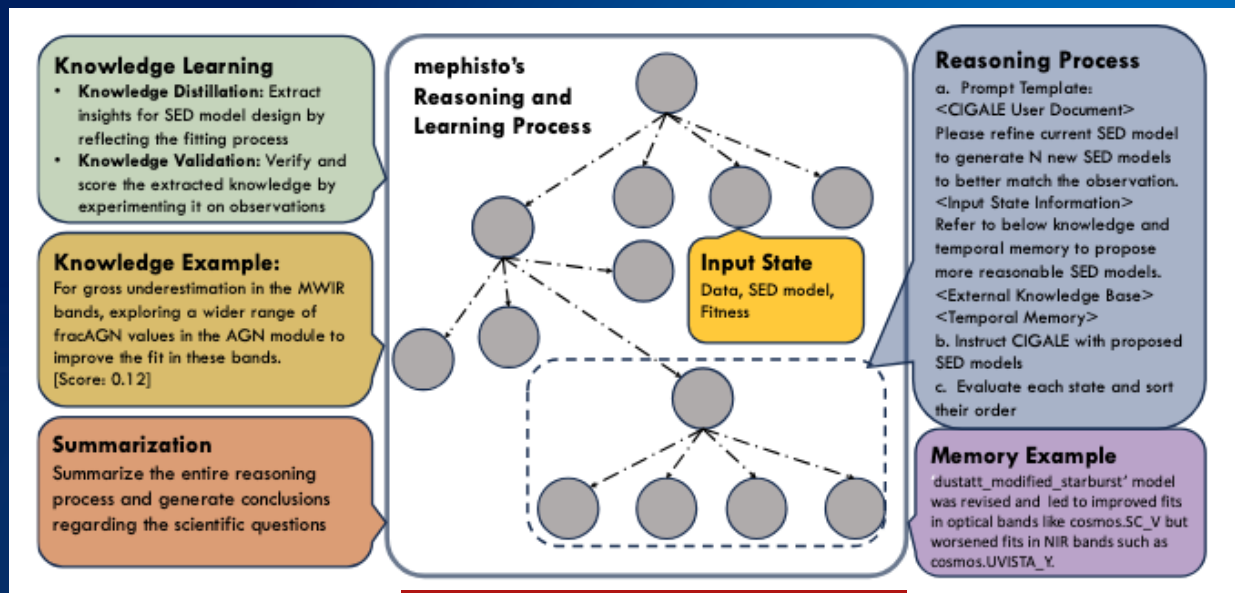


AI 2.0: AI agents



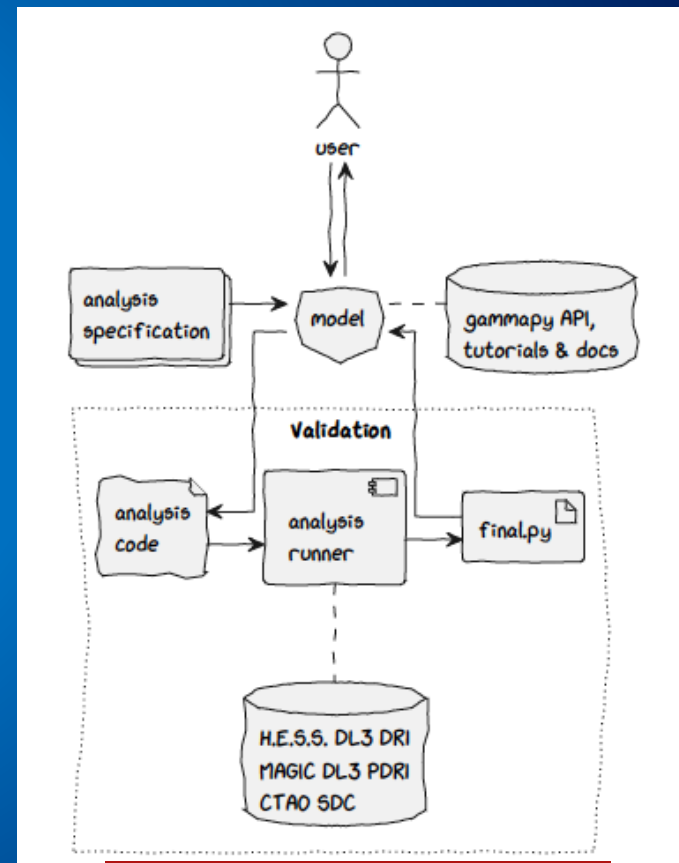
AI agents in Astronomy

An AI agent able to use CIGALE to fit SED



Sun+2024

AI agent for Gamma ray astronomy



Kostunin+2024

AI and spectroscopy towards 2040+

Deploy current AI tools
for the next generation
of IFU and MOS

Develop new tools for
data reduction, survey
strategy, scheduling

Foundation models and
Agents for Astronomy?

**An INAF data analysis hub for the next
ESO facility**