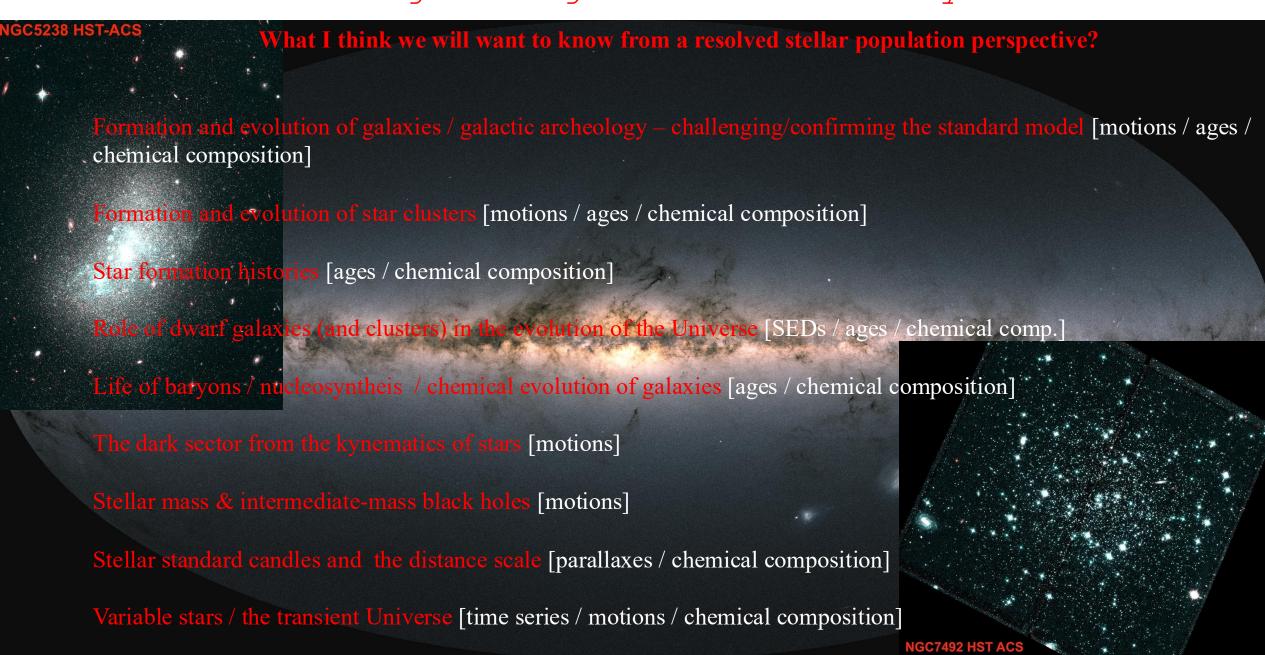
Michele Bellazzini (INAF - OAS Bologna)

Brainstorming!
Trying to be representative of a community...



Horizon: next large ESO facility after ELT

An effort worth trying but ... hard to tell the future



A complete census of the Local Universe (faint LF vs Λ CDM) Finding out all the galaxies and star clusters in the LG:

GaiaNIR

Deep wide-field imaging and photometry

All-sky astrometry + photometry

Galactic dynamics and archeology Absolute proper motions over the Galaxy and its immediate surroundings; geometric distances to as many stars and standard candles as possible:





Not suitable for ESO '40

What we miss:
A northern version of LSST
An all-sky imaging survey from space in the optical-NIR



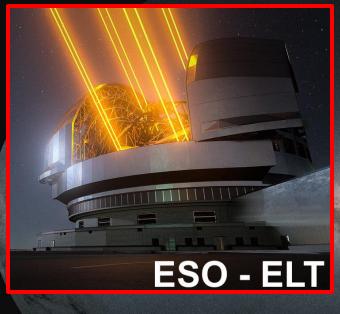
What we miss:

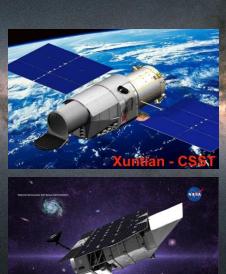
An effective successor of Gaia really able to cope with crowding to go much deeper in the densest regions of the MW and to pierce the veil of dust

Not suitable for ESO '40

Resolving distant galaxies into stars, extremely dense/crowded environments High precision kinematics of systems on small scales

High spatial resolution: photometry, spectroscopy and astrometry





THE NANCY GRACE ROMAN
SPACE TELESCOPE



What we miss:
A further step ahead.
Efficient AO in the optical?
Interferometry?

Core business of the current upcoming ESO facility

Future ELT instrumentation + space facilities

Tracing HI distribution and kinematics also very helpful for many goals in this and in the previous slide: FAST, Chords, SKA – we should be reasonably safe on this side

We can co ahead in the exercise of going through my list of wishes:

Variable stars – transient Universe → LSST

Intermediate-mass black holes → ELT

We can do it after my main conclusion if you like...

A classic holy grail for resolved stellar populations is: AGE

...but, unfortunately, we cannot build an age-meter. We need magnitudes, distances and chemical compositions to be obtained with different instruments/facilities to be combined together

Or we need asteroseismology (to measure stellar mass) but this is not easy for faint/distant stars and it is done better from space.

It does not seem an appropriate driver for this specific project at this stage

So my main conclusion is that we need spectra to be acquired over a wide field and/or with integral field at an industrial pace

- To match Gaia and Gaia-NIR catalogues with RV and chemical composition
- To follow-up discoveries from imaging surveys and future imaging facilities
- To patrol the Galaxy and the Local Group down to much fainter magnitudes wrt, e.g, WEAVE and 4most
- To sample velocity fields in stellar systems on scales from arcmin to degrees: the most interesting range for tracing dark matter
- To go in full survey mode, allowing new discoveries
- To have a very flexible, multi-purpose facility

The actual set-up by which which we get industrial scale wide-field and/or integral field spectroscopy is not critical: large FoV, high multiplexing, integral field capabilities, spectral resolution are critical

A spectral resolution of about 6-8000 would be enough to achieve many of the main goals but, obviously, high resolution opens extremely valuable windows. Vision and perspective call for both. **However, a system**working at low-res with higher efficiency + advanced machine learning+AI may be highly competitive in the '40s

What can drive radically different choices:

- The emergency of a fundamental scientific issue or a drastical change in priorities requiring a specific experiment / instrumental set up

- A breakthrough technologic advancement enabling science that is currently out of reach or unthinkable



Thank you!