

Survey of Surveys: Evaluating Methods for Homogenizing Stellar Parameters Across Spectroscopic Surveys

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Motivation

Challenges: Heterogeneity in data from different spectroscopic surveys.

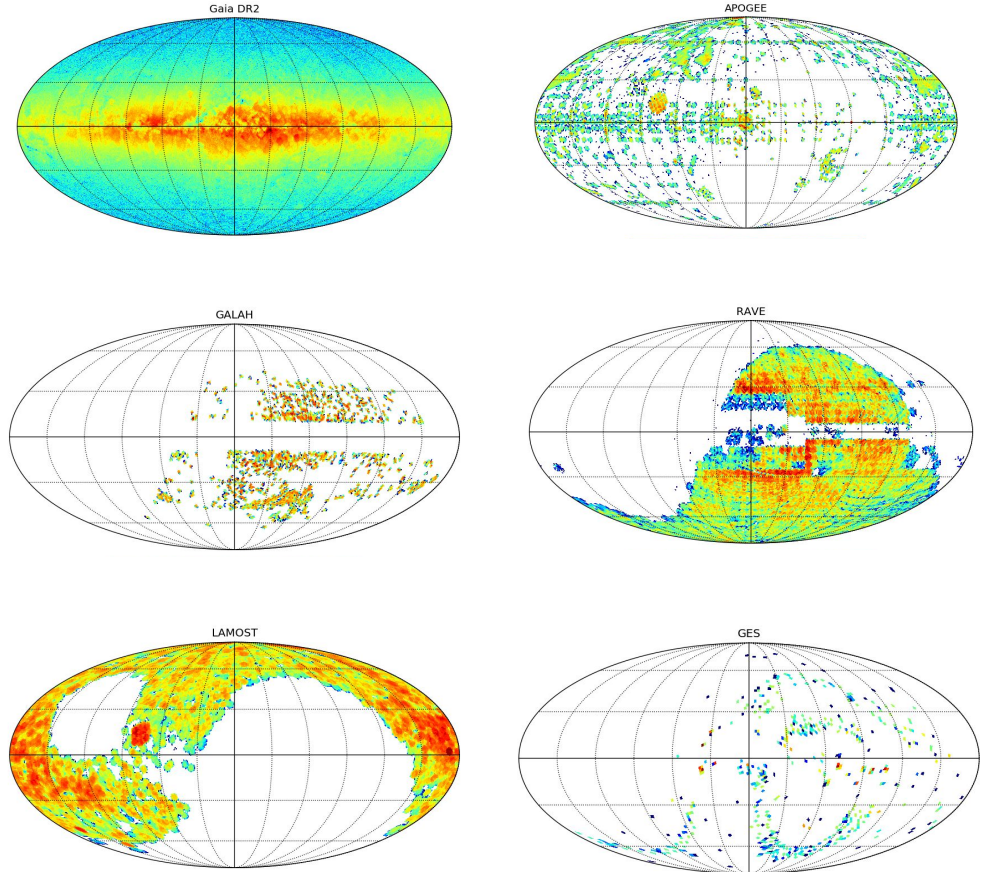
Objective: Create a unified catalogue for stellar parameters, ensuring consistency across datasets.

First results: Survey of Surveys (SoS) DR 1/2

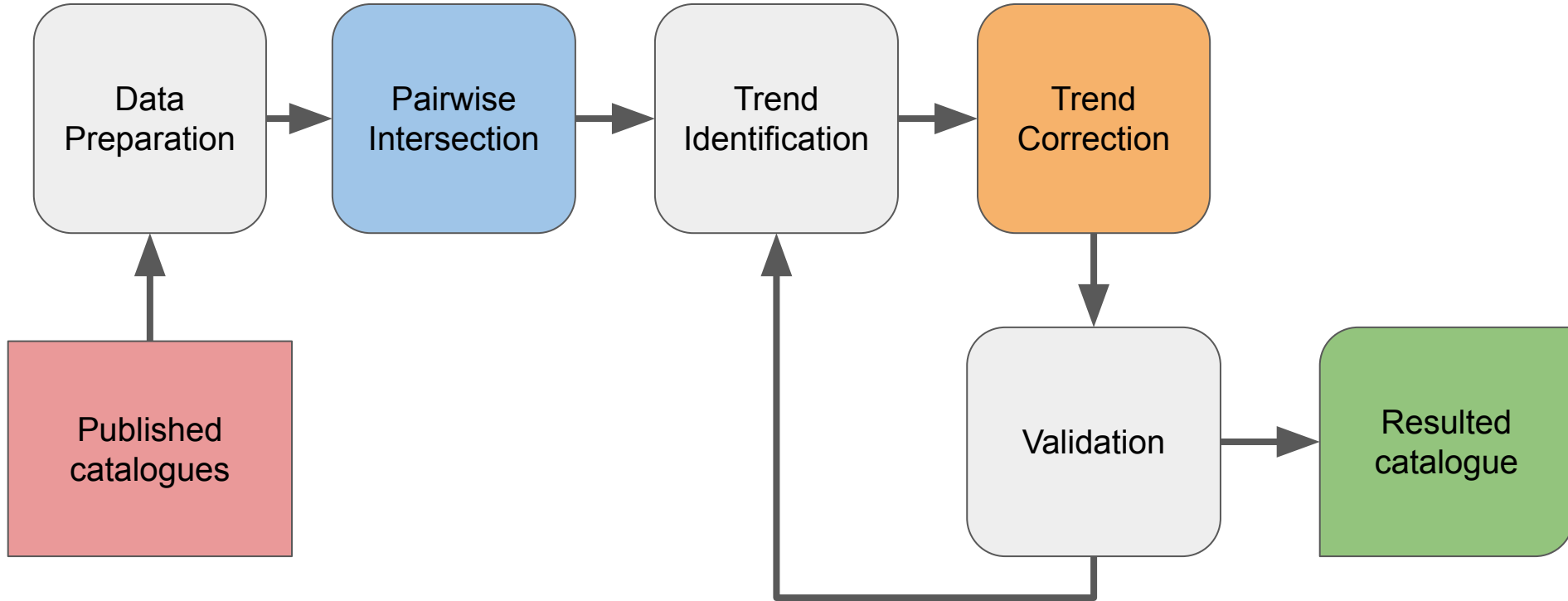
Tsantaki et al., 2022 - Radial velocities.

Turchi et al. (in preparation) - Homogenized spectroscopic stellar parameters + 19M stellar parameters from photometry with ML.

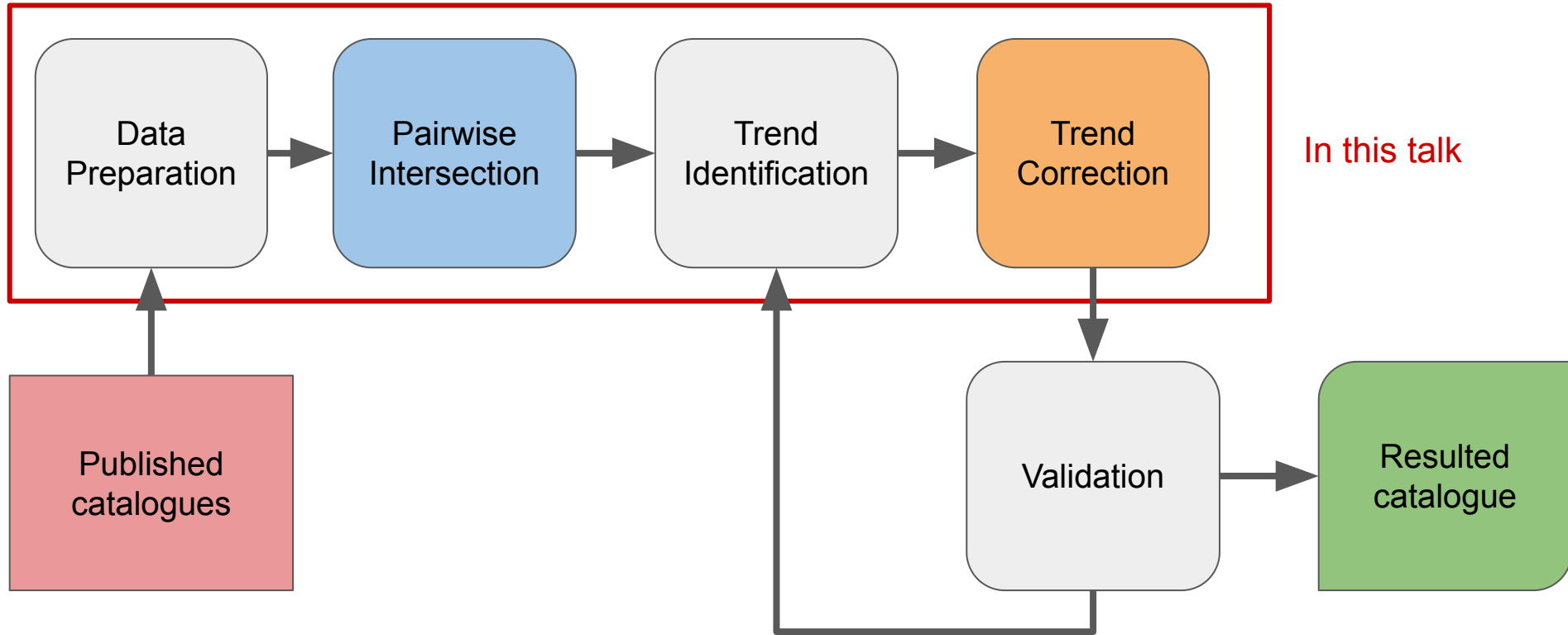
Now: Recent data releases provide improved opportunities for cross-survey calibration.



Strategy Overview



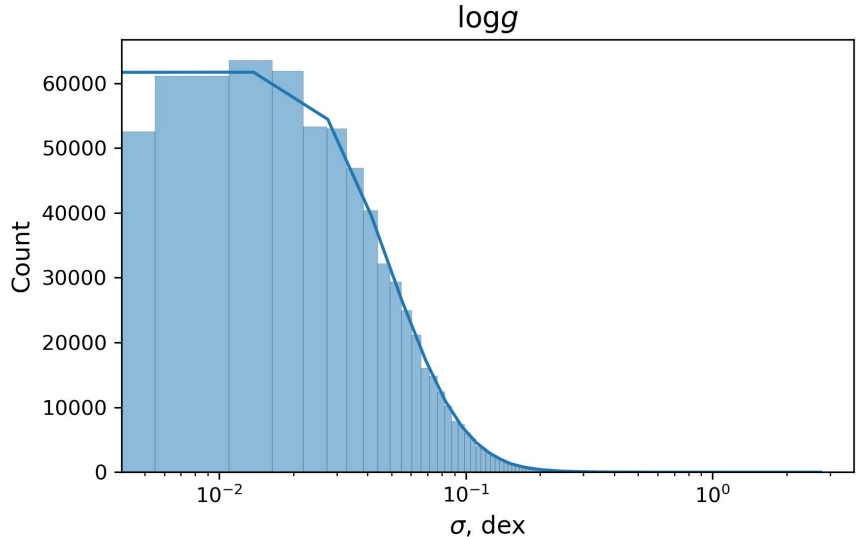
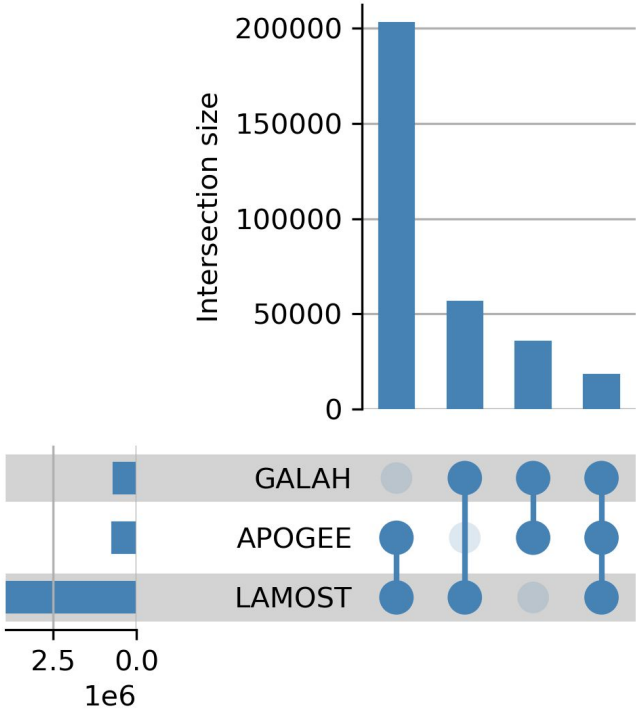
Strategy Overview



Data Selection and Preprocessing

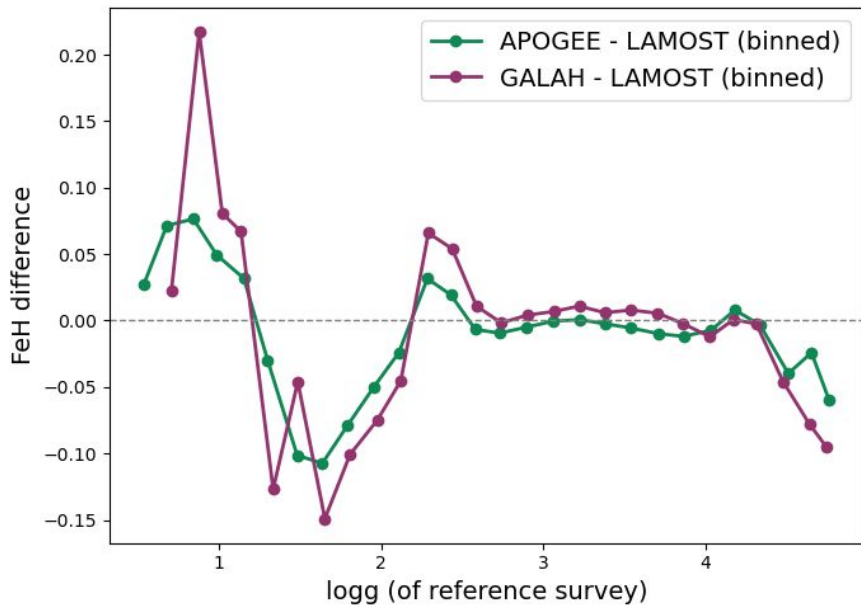
Surveys: APOGEE, GALAH, LAMOST

Filters: quality flags, duplicates with large parameter spread.

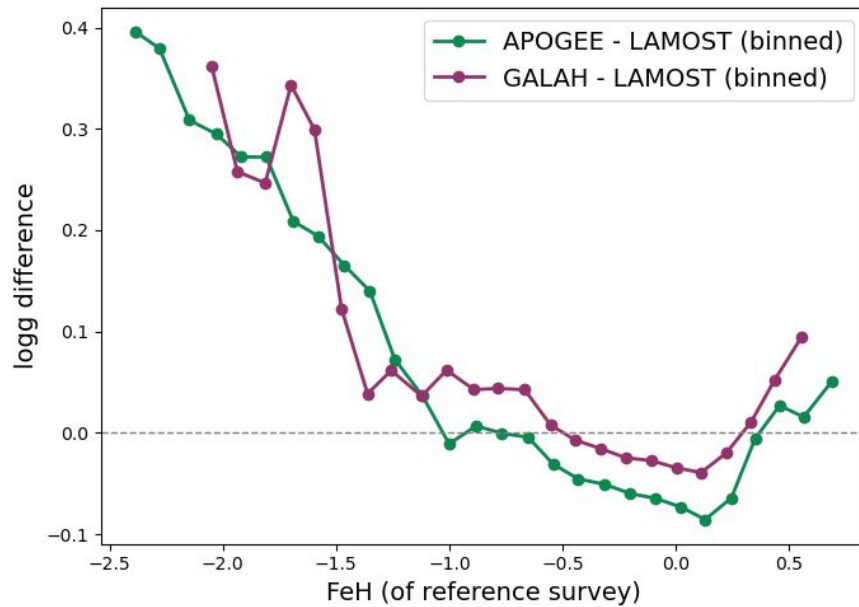




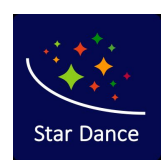
Trend Detection



On y-axis: the difference between LAMOST and APOGEE/GALAH [Fe/H] measurements.

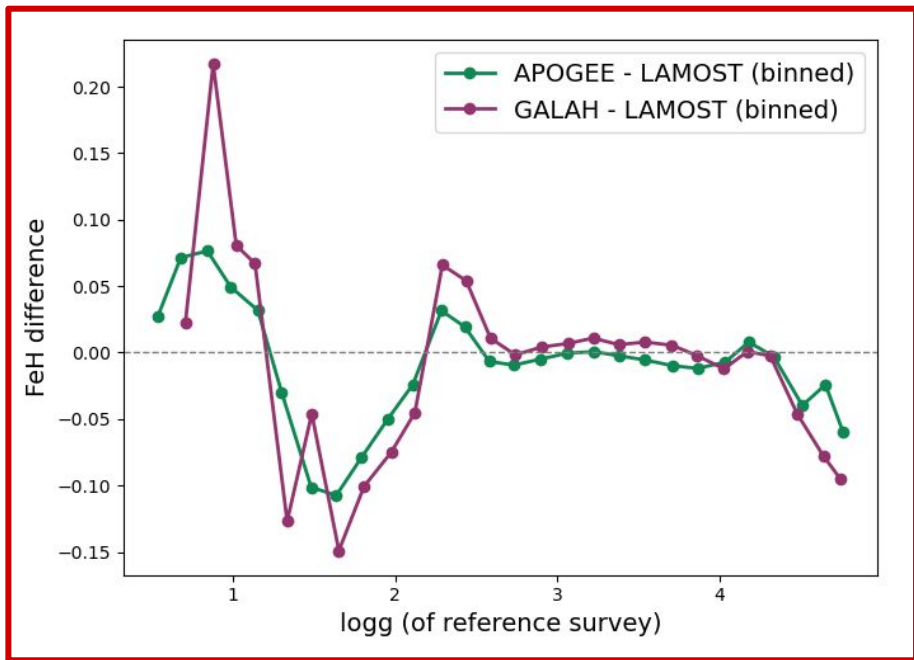


On y-axis: the difference between LAMOST and APOGEE/GALAH **logg** measurements.

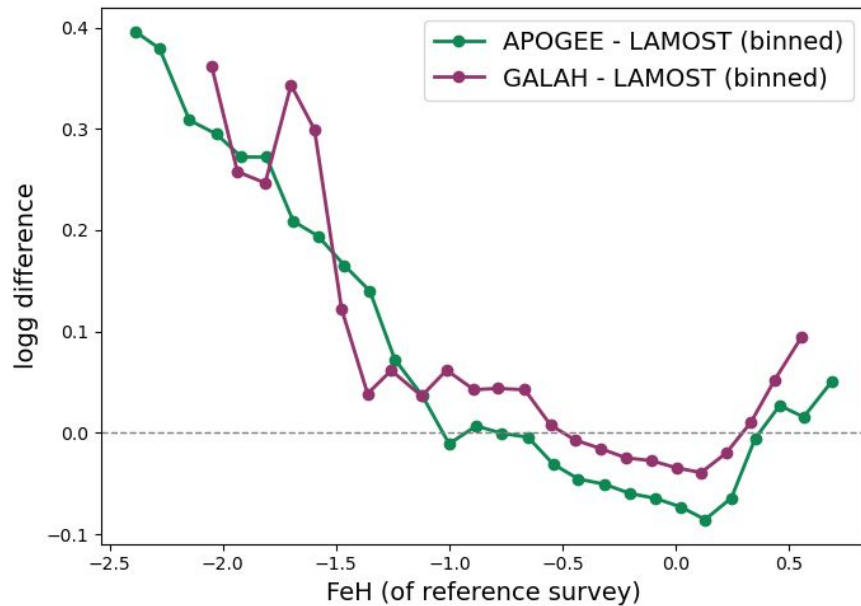


Trend Detection

In this talk



On y-axis: the difference between LAMOST and APOGEE/GALAH [Fe/H] measurements.

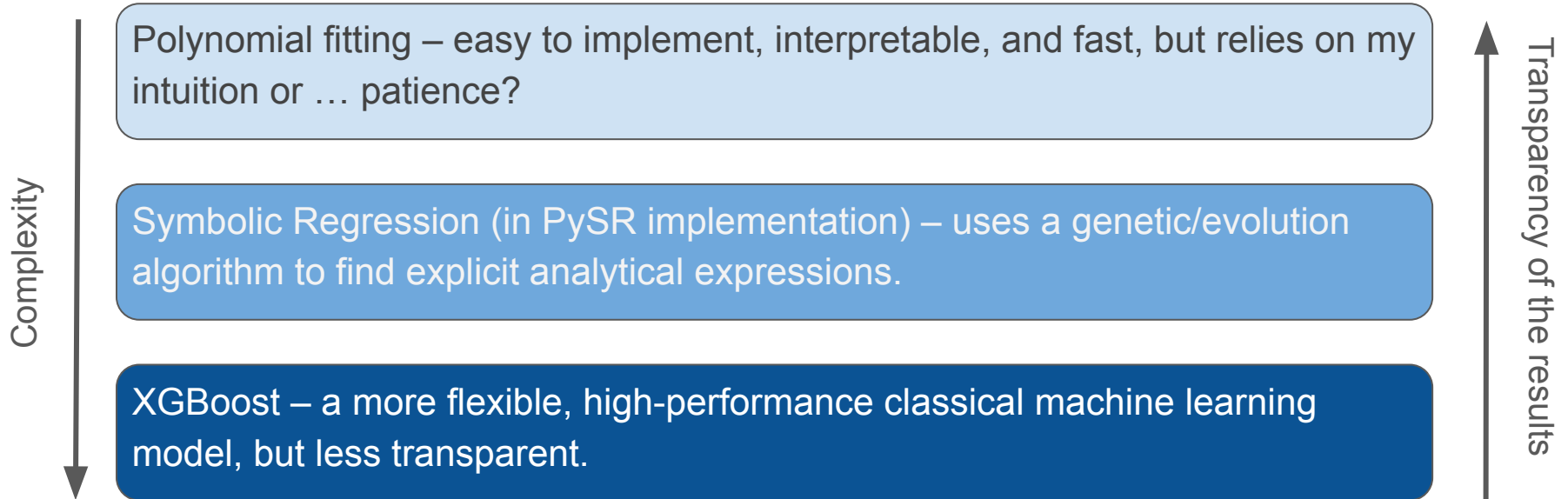


Inputs to the models are normalized effective temperature ($T_{\text{eff}}/5040 \text{ K}$), surface gravity ($\log g$), and metallicity ($[\text{Fe}/\text{H}]$).

We use the union of GALAH and APOGEE as the reference because LAMOST shows a consistent metallicity trend when compared to both surveys.



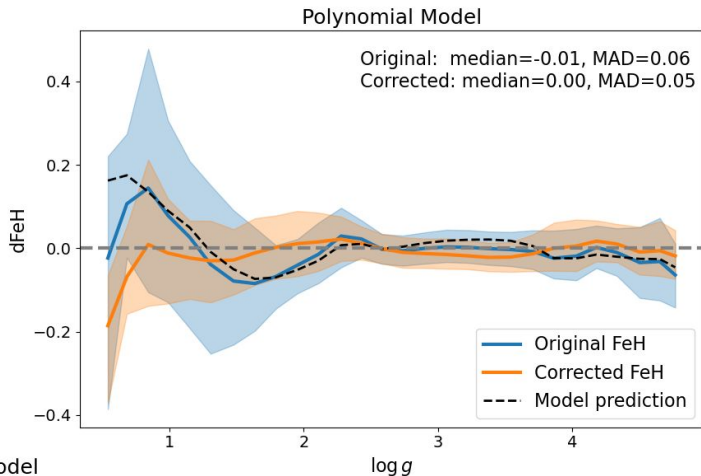
Method Comparison: Goals & Expectations



The goal: to reduce systematic trends in $[\text{Fe}/\text{H}]$ across the parameter space **without introducing artifacts**, ensuring smooth, consistent calibration.

Test case of FeH: Polynomial Fitting (4th order)

x-axis: $\log g$ →

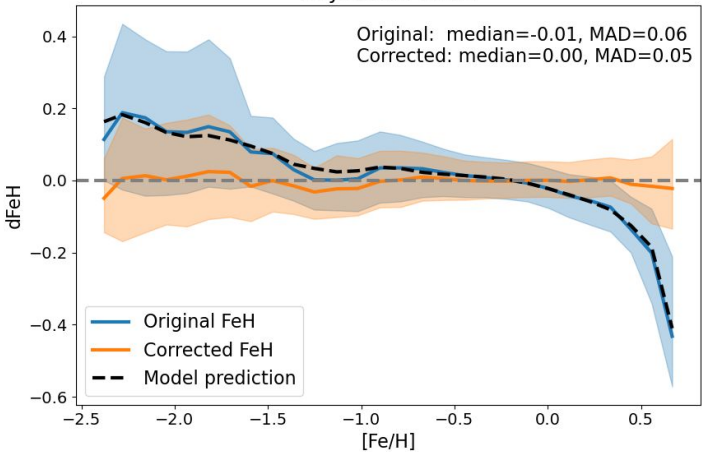


The 4th-order polynomial with crossed terms effectively eliminates the trend dependent on $[\text{Fe}/\text{H}]$.

Residual offsets remain at low $\log g$ (< 1 dex) and high temperatures (> 7000 K).

Binned plots show the 16th-84th percentile range as shaded areas.

Polynomial Model



← x-axis: $[\text{Fe}/\text{H}]$

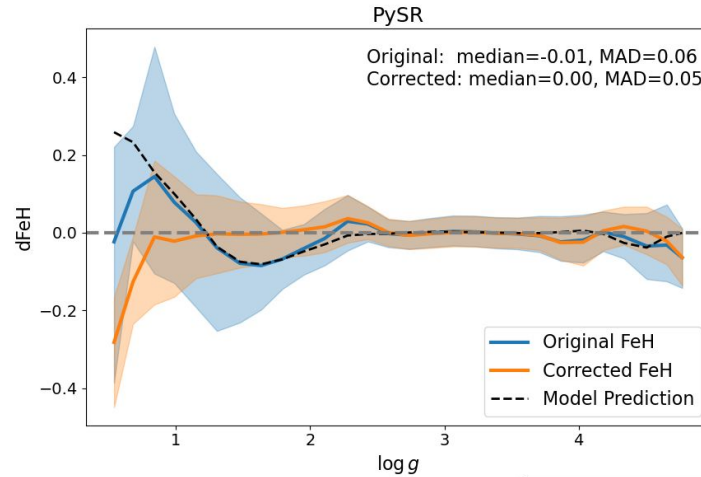
Test case of FeH: Symbolic Regression (PySR)

The PySR symbolic regression method shows similar performance to the 4th-order polynomial.

It also struggles with extreme values, especially at low $\log g$ (< 1 dex) and high T_{eff} (> 7000 K). Moreover, it fails at the edges of metallicity ($[\text{Fe}/\text{H}] < -2$ dex or $[\text{Fe}/\text{H}] > 0.2$ dex).

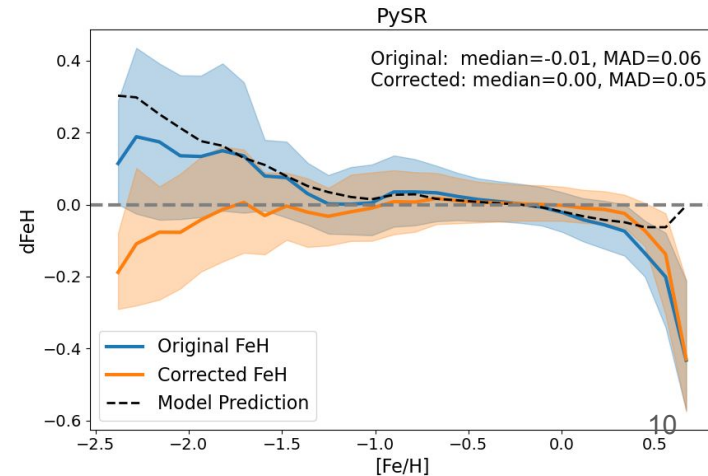
While it reduces the trend, it does not significantly improve calibration in these regions.

Or it needs significantly more computational time to converge.

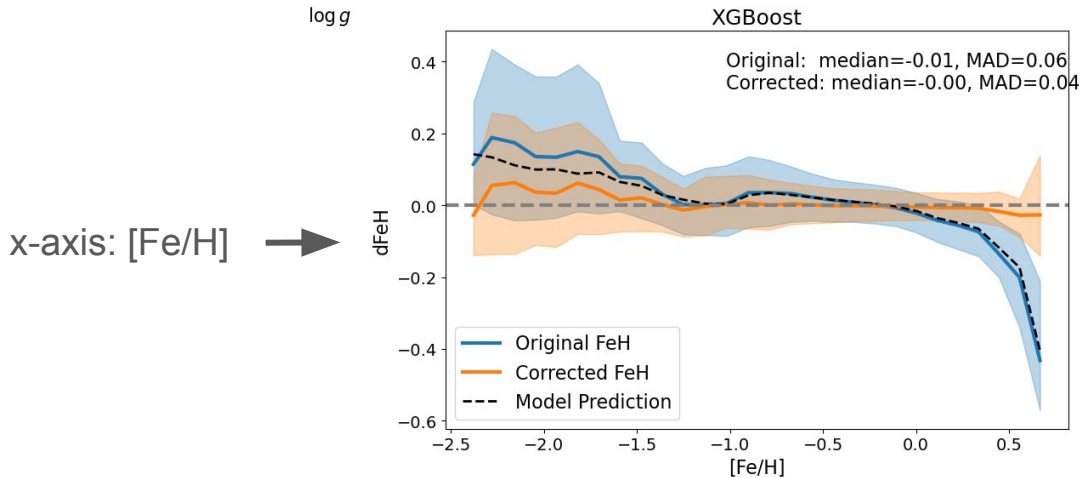
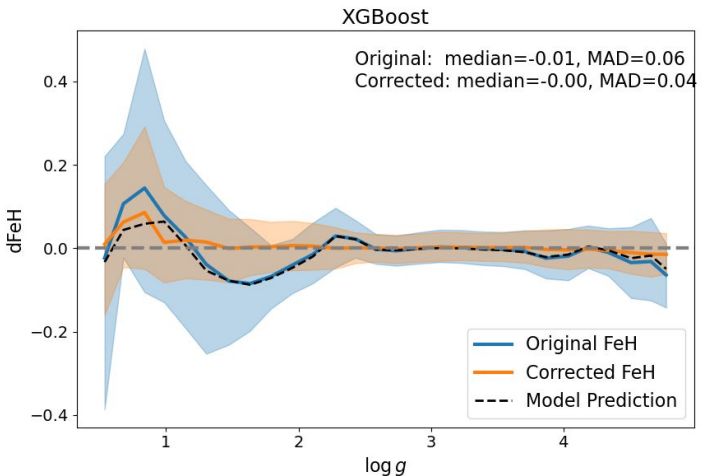


← x-axis: log g

x-axis: [Fe/H] →



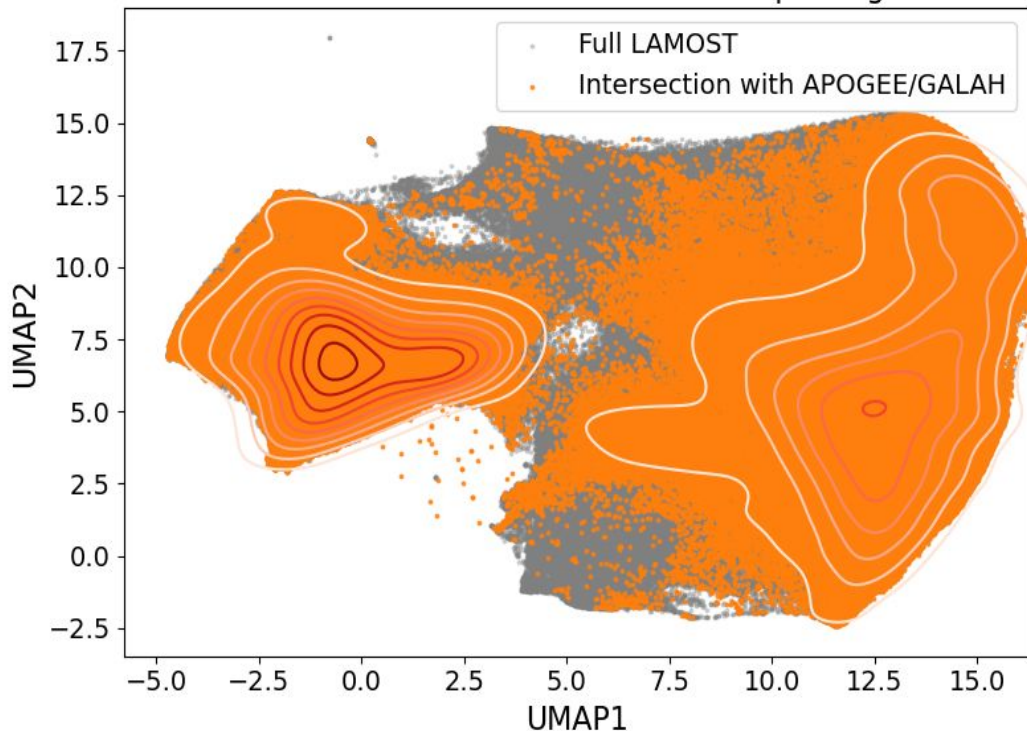
Test case of FeH: XGBoost



XGBoost provides superior performance in correcting trends across the parameter space, including in regions with extreme [Fe/H], low logg (< 1), and high Teff (> 7000 K).

While the quartiles are higher at the edges, the median is well-calibrated across all parameters, showing consistent improvements in overall performance.

UMAP (Uniform Manifold Approximation and Projection): Coverage Exploration



UMAP exploration used the following parameters:

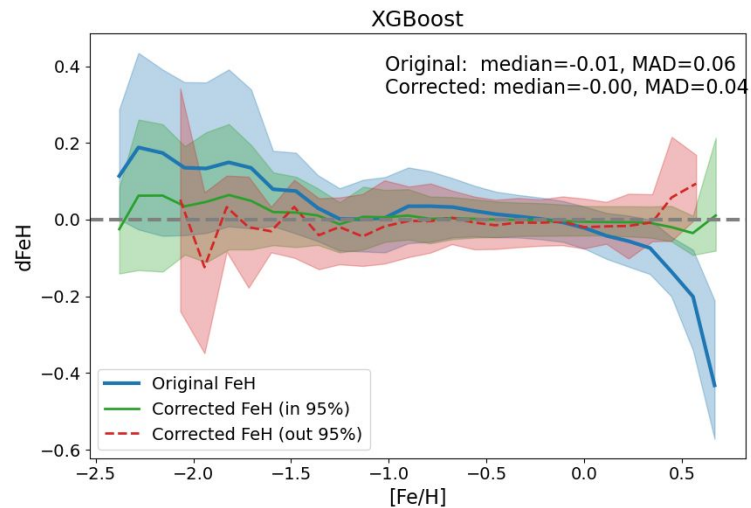
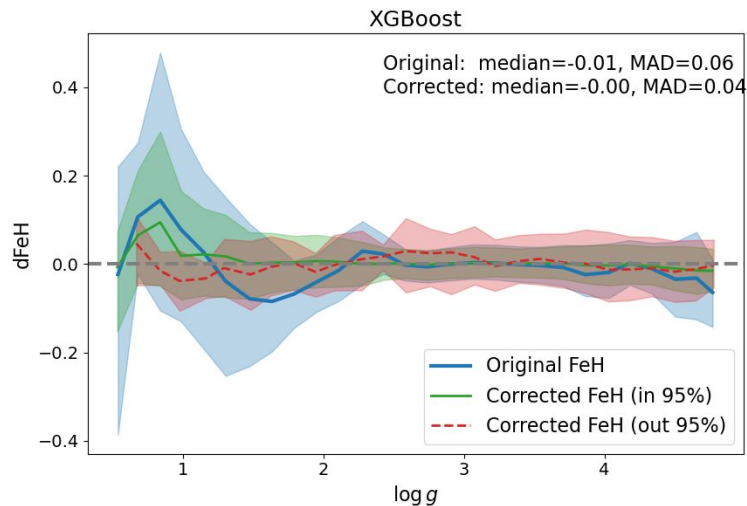
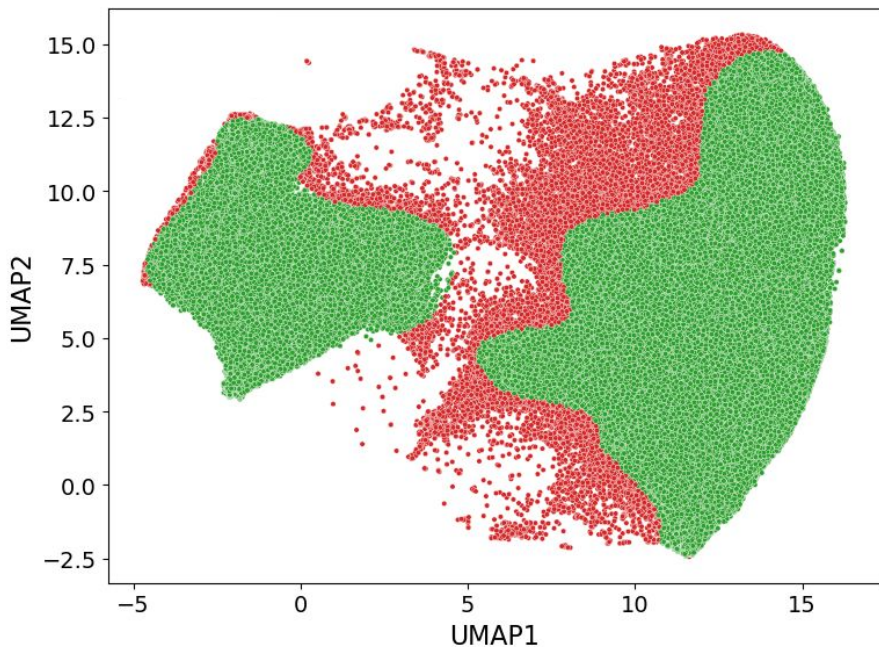
- 8 parameters: T_{eff} , $\log g$, $[\text{Fe}/\text{H}]$, SNR, radial velocity, Gaia magnitude, distance and the estimation of extinction

Key finding:

- Only part of the parameter space is well-covered by the LAMOST-GALAH/APOGEE intersection

Generalization Challenges

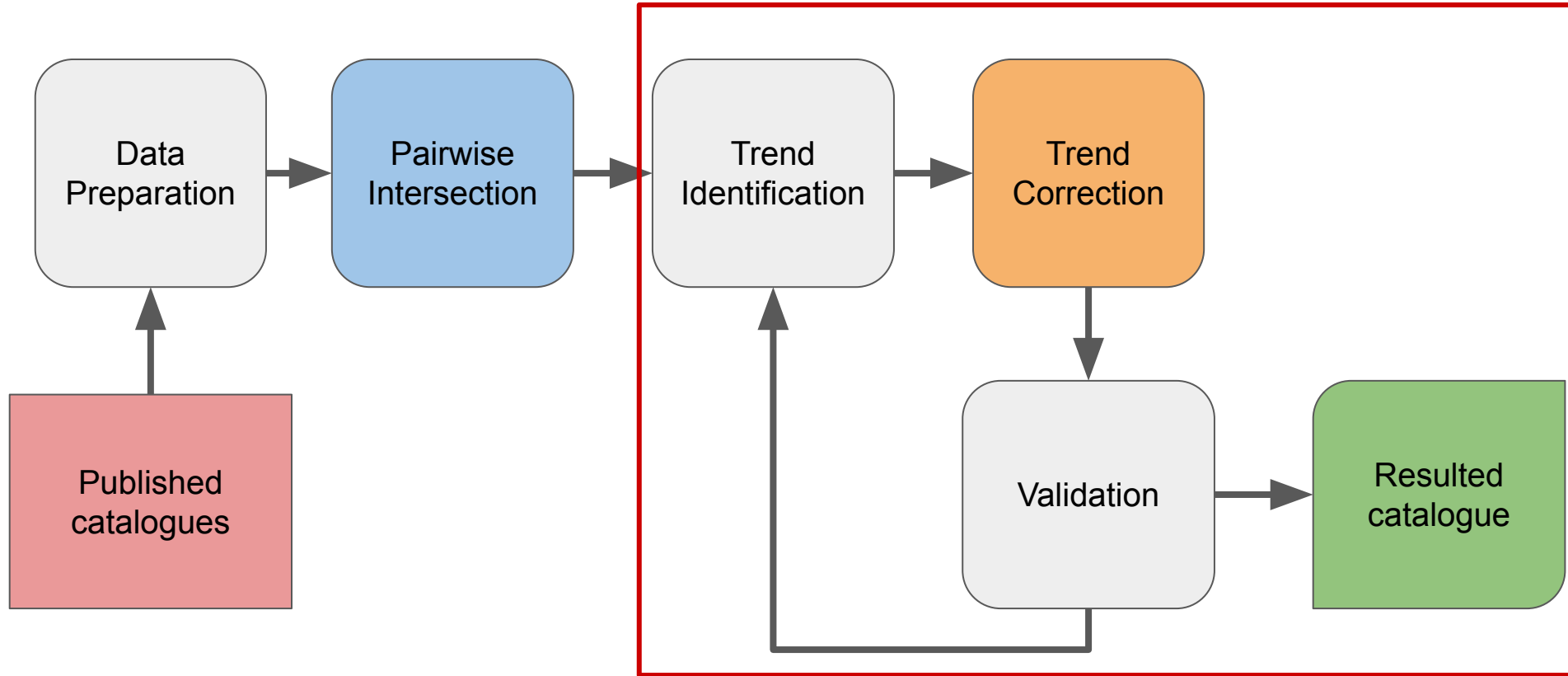
We divide intersection in two groups: most populated in the parameter space 95% (in green) and least populated 5% (in red):



Summary

- LAMOST metallicities show systematic trends when compared with high-resolution surveys (GALAH/APOGEE)
- Polynomial correction (4th order with cross-terms): Removes most trends in $[Fe/H]$, but struggles at low $\log g$ and high T_{eff} .
- Symbolic Regression (PySR): Interpretable formulas, but less effective at edges of the parameter space. Also requires a lot of time to find the right formula.
- XGBoost: Best correction performance across full parameter range, effective in low-density / edge regions.
- UMAP exploration reveals that $\sim 33\%$ of LAMOST stars lie outside well-calibrated intersection. Generalization to full dataset needs further validation.

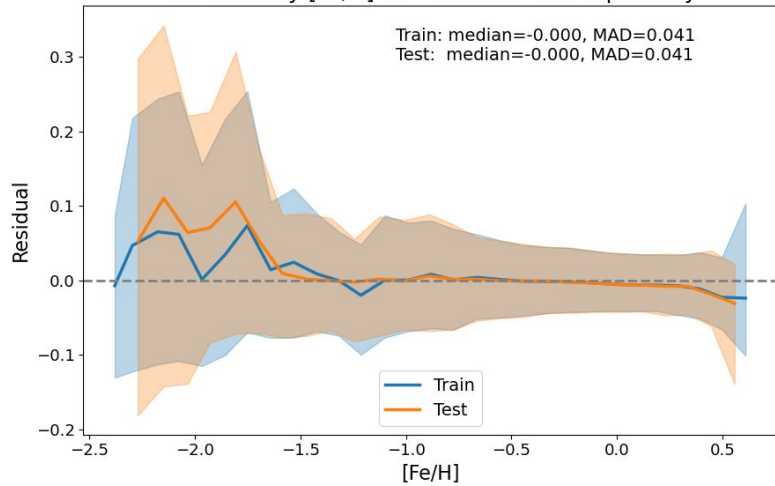
Future Work



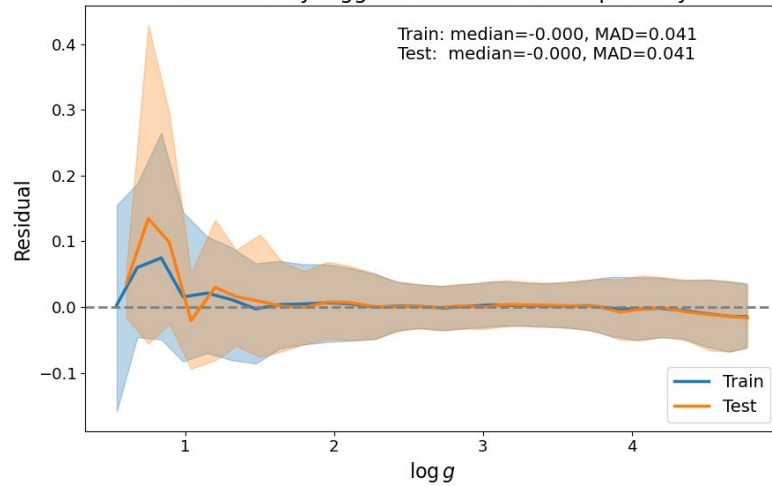
Thank you!
Questions?



Residuals by [Fe/H] for Train and Test Separately



Residuals by log g for Train and Test Separately



Residuals by Teff for Train and Test Separately

