

Vision-Language Models for Radio Astronomy

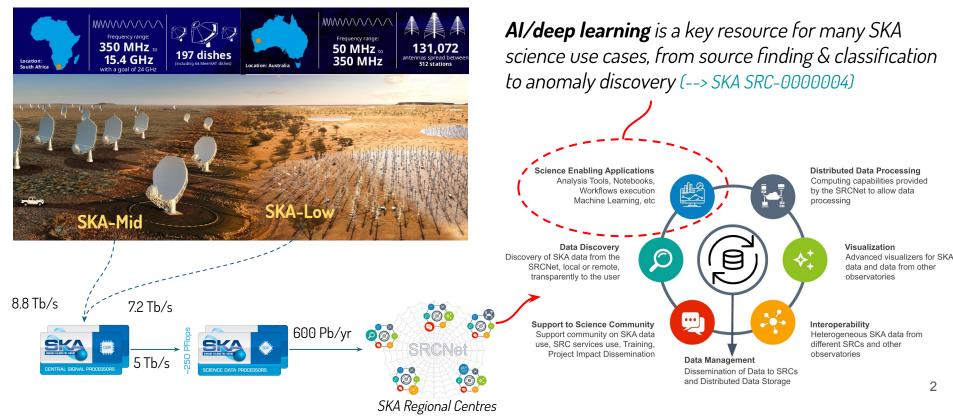




The Square Kilometer Array



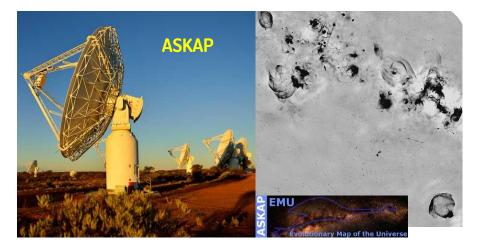
SKA will generate massive radio data volumes
 ✓ Automated extraction of science-ready data is a major challenge





SKA precursors surveys and upgrade operations are ongoing ...

- ✓ Anticipating SKA challenges, serving as analysis test benches for SKA
- ✓ Boosting new developments in data processing software



ASKAP EMU Survey

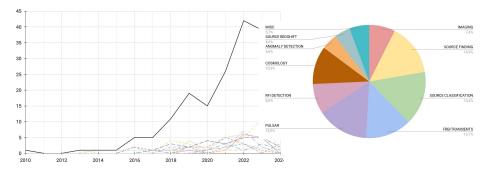
- Early Science (SCORPIO field, 912 MHz) (2018-2019)
- *Pilot 1* (SCORPIO field, 1243 MHz) (2019)
- Pilot 2 (SCORPIO & other GP fields, 943 MHz) (2021)
- EMU main survey (943 MHz) started



SARAO MeerKAT Galactic Plane Survey (SMGPS)

- Surveyed area: lbl<1.5°, 2°<l<60°, 252°<l<358°
- Frequency: 886 1678 MHz
- Theoretical LAS @ 1.284 GHz: 27 arcmin

《 ML in Radio Astronomy



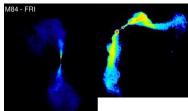


Detection/Segmentation e.g. to extract sources from maps or to segment source regions



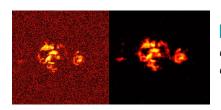
Regression/Inference

e.g. to estimate source physical parameters (redshift, flux, ...) or model parameters



Classification

e.g. to classify source type or morphology, or to find sub-groups



Denoising

e.g. to clean data from background or instrumentation noise patterns

Anomaly Detection

e.g. to search for astrophysical objects with peculiar morphology or to identify outliers/anomalies in the data





Representation Learning

e.g. to extract features/parameters from the data, view/discover hidden patterns

Data Generation

e.g. to generate simulated/synthetic data, map inpainting, etc

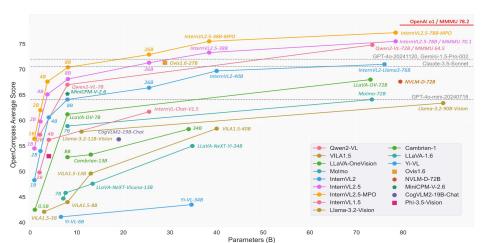
\bigcirc | Vision Language Models (VLMs)

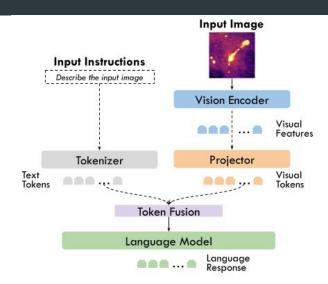
VLMs used for various tasks on image + text inputs

- ✓ Image Captioning
- ✓ Visual Question Answering (VQA)
- ✓ Text-to-Image
- ✓ Image Search/Retrieval

■ Why exploring VLMs for astronomy?

 Instruct and perform visual tasks (data quality assessment, source detection/classification, anomaly/similarity search, etc) without coding through a text-based interface and examples





Leaderboards of open/commercial VLMs (March 2025)

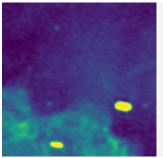
Rank 🔺	Method 🔺	Param (B)	Avg Score	Rank* (UB)	Model	Arena Score
1	InternVL 2.5-78B- MPO	78	80.3	1	ChatGPT-40-latest (2025-01-29)	1279
				1	GPT-4.5-Preview	1269
2	InternVL 2.5-38B-	38B- 38	78.3	1	Gemini-2.0-Flash-Thinking-Exp-01-21	1278
	MPO			3	Gemini-2.0-Pro-Exp-02-05	1241
3	Qwen2.5- VL-72B	73.4	78.1	4	01-2024-12-17	1232
4	0vis2- 34B	34.9	77.5	4	Claude 3.7 Sonnet	1223
5	InternVL 2.5-26B- MPO	26	76.4	4	Gemini.2.0.Flash.001	1227
				5	Gemini-1.5-Pro-002	1222

🗠 | Radio Benchmarks

We have set up 6 benchmarks on radio source detection/classification

B1 - Extended/diffuse radio source detection

- Multi-label multi-class classification
- ✓ Dataset: ~5,700 ASKAP/MeerKAT labelled images with extended (33%), diffuse (4%) and large-scale diffuse (7%) radio sources



Sample image {EXTENDED, DIFFUSE-LARGE}

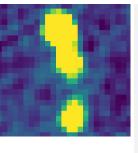
Context: ...

Question: Which of these morphological classes of radio sources do you see in the image?

EXTENDED DIFFUSE DIFFUSE-LARGE NONE

B2 - Source morphology classification

- ✓ Single-label multi-class classification
- ✓ Dataset: ~3,800 VLA source-zoomed images of 6 morphological classes (~600 per class): 1C-1P, 1C-2P, 1C-3P, 2C-2P, 2C-3P, 3C-3P



Context: ...

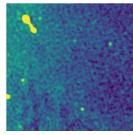
Question: Which of these morphological classes of radio sources do you see in the image?

1C-1P 1C-2C 1C-3P 2C-2P 2C-3P 3C-3P

Radio Benchmarks

B3 - Extended Radio Galaxy Detection

- ✓ Binary classification
- ✓ Dataset: ~5,700 ASKAP/MeerKAT images with extended radio galaxies (~18%)



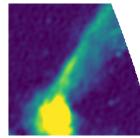
Context: ...

Question: Do you see any likely radio galaxy with an extended morphology in the image?

YES NO

B5 - Image peculiarity classification

- ✓ Single-label multi-class classification
- ✓ Dataset: ~5,700 ASKAP/MeerKAT images labelled as ORDINARY (~63%), COMPLEX (~35%), PECULIAR (~3%)



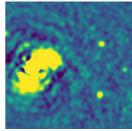
Context: ...

Question: Can you identify which peculiarity class the presented image belongs to?

ORDINARY COMPLEX PECULIAR

B4 - Artefact Detection

- ✓ Binary classification
- Dataset: ~5,700 ASKAP/MeerKAT images with imaging artefacts (~7%)



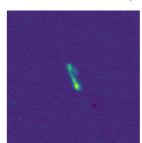
Context: ...

Question: Do you see any imaging artefact with a ring pattern around bright sources in the image?

YES NO

B6 - Radio Galaxy Morphology Classification

- Binary classification
- ✓ Dataset: ~832 VLA images centred on FR-I (~48%) and FR-II (~52%) radio galaxies

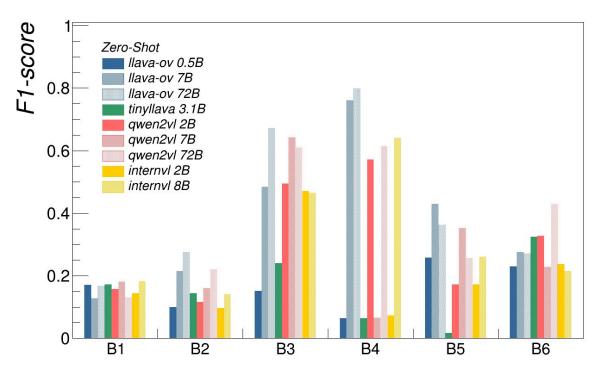


Context: ...

Question: Which of these morphological classes of radio galaxy do you see in the image?

Zero-Shot Results

Evaluating open VLMs of various sizes on radio benchmarks



- ✓ Overall, poor performance across all benchmarks
- Larger models performing better, as expected
- ✓ Best results on B3 (galaxy detection)
 & B4 (artefact detection)
- LLaVA slightly outperforming other models on most benchmarks

need models specialized on radio tasks

ⓒ | The *radio-llava* model

Fine-tuning LLaVA-OneVision 7B on radio data

- ✓ Vision encoder frozen
- ✓ LLM (qwen2) & projector free
- ✓ Full vs LORA fine-tuning tested
- ✓ Shallow (1 epoch) vs deeper (3 epochs) fine-tuning
- ✓ Standard vs alternative hyperparameter choices

Two training datasets created

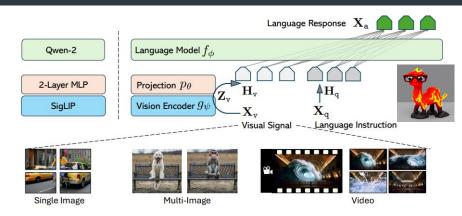
Q&A dataset

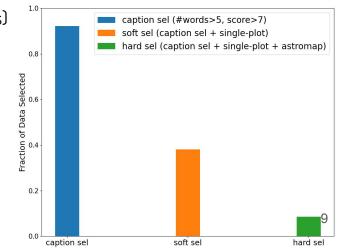
- ✓ assembled from *fine* & *coarse-grained* annotated radio datasets
- ✓ 59k radio images (source-zoomed & wide-field, ASKAP/MeerKAT + others)
- \checkmark ~1.5 M user-assistant conversations generated from label information
 - *Q: Can you describe the image content?*

Q: Can you provide the bounding box of sources with class X in the image? Q: Do you see source/pattern X in the image?

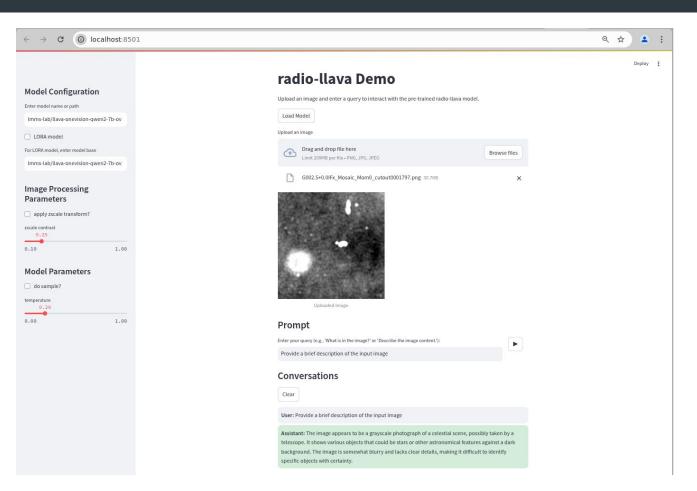
Caption dataset

- ✓ assembled from arXiv papers (2000–2025) on radioastronomical topics
- ✓ ~38k figure-caption pairs after quality selection (*soft sel*)
 - *Q: Can you describe the image*





(한) | The radio-llava model

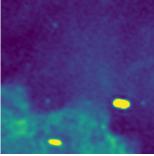


For more details: Riggi et al, 2025, https://arxiv.org/abs/2503.23859

📠 | Fine-tuning Results (B1)

B1 - Extended/diffuse radio source detection

- ✓ Multi-label multi-class classification
- ✓ Dataset: ~5,700 ASKAP/MeerKAT labelled images with extended (33%), diffuse (4%) and large-scale diffuse (7%)



Context: ...

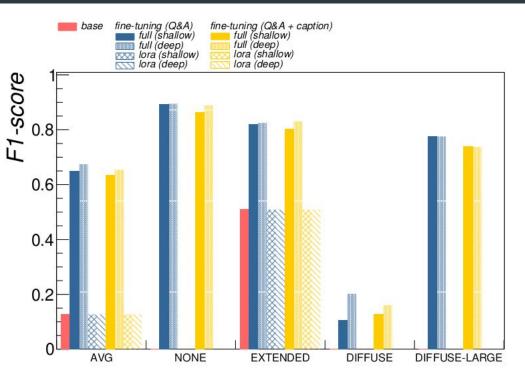
Question: Which of these morphological classes of radio sources do you see in the image?

EXTENDED DIFFUSE DIFFUSE-LARGE NONE

Sample image {EXTENDED, DIFFUSE-LARGE}



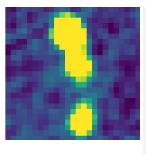
- ✓ +40% wrt base model
- ✓ Caption data slightly degrade
- Marginal improvements with deeper fine-tuning
- ✓ Diffuse sources still largely missed
- ✓ LORA fine-tuning not effective
- ✓ Suboptimal wrt vision-only fine-tuning (F1~80%)



📠 | Fine-tuning Results (B2)

B2 - Source morphology classification

- ✓ Single-label multi-class classification
- ✓ Dataset: ~3,800 VLA source-zoomed images of 6 morphological classes (~600 per class): 1C-1P, 1C-2P, 1C-3P, 2C-2P, 2C-3P, 3C-3P



Context: ...

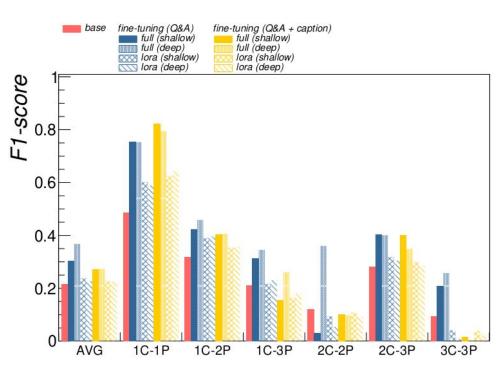
Question: Which of these morphological classes of radio sources do you see in the image?

1C-1P 1C-2C 1C-3P 2C-2P

2C-3P 3C-3P



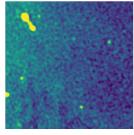
- ✓ +10% wrt base model
- ✓ Caption data slightly degrade
- ✓ Marginal improvements with deeper fine-tuning
 - Poor classification overall
- LORA fine-tuning not effective



📠 | Fine-tuning Results (B3)

B3 - Extended Radio Galaxy Detection

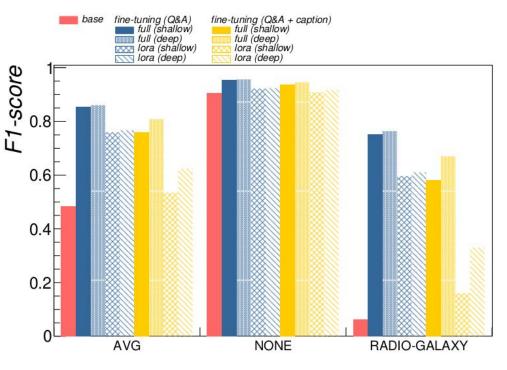
- ✓ Binary classification
- ✓ Dataset: ~5,700 ASKAP/MeerKAT images with extended radio galaxies (~18%)



Context: ...

Question: Do you see any likely radio galaxy with an extended morphology in the image?

YES NO



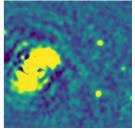


- ✓ +30% wrt base model
- ✓ Caption data slightly degrade
- ✓ Negligible improvements with deeper fine-tuning

I Fine-tuning Results (B4)

B4 - Artefact Detection

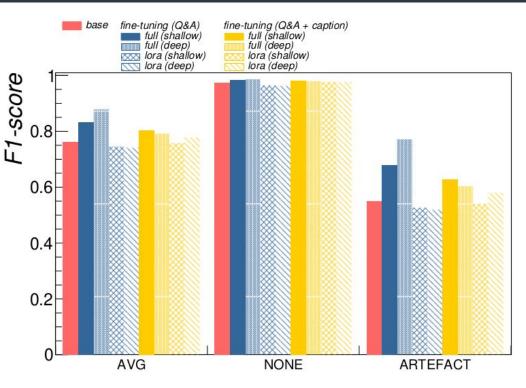
- ✓ Binary classification
- ✓ Dataset: ~5,700 ASKAP/MeerKAT images with imaging artefacts (~7%)



Context: ...

Question: Do you see any imaging artefact with a ring pattern around bright sources in the image?

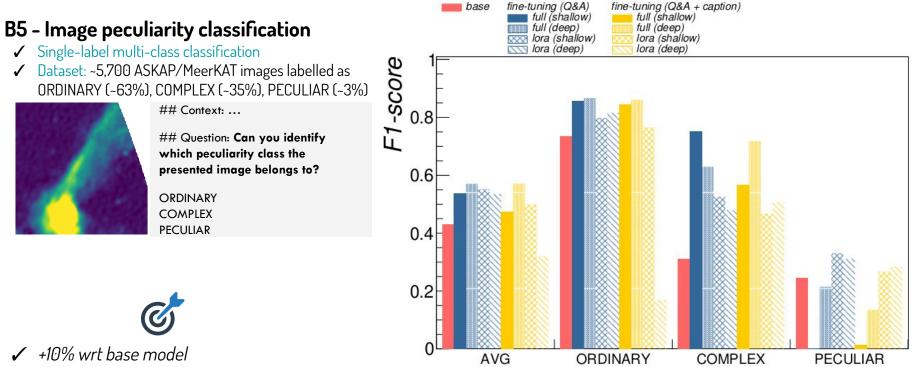
YES NO





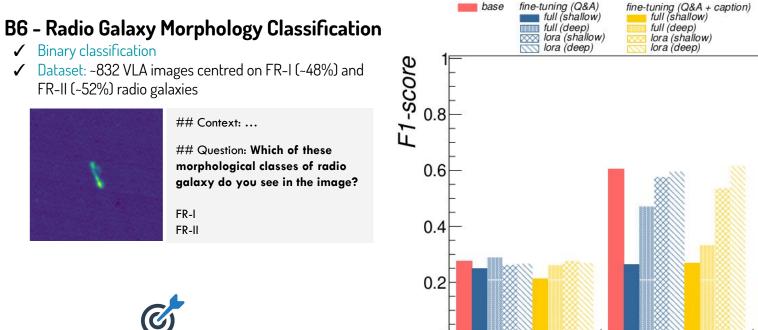
- ✓ +10% wrt base model
- Caption data slightly degrade
- ✓ Marginal improvements with deeper fine-tuning
- LORA fine-tuning not effective

📠 | Fine-tuning Results (B5)



- Caption data slightly degrade
- Peculiar frames largely missed

Fine-tuning Results (B6)



n

AVG

FR-I

 \checkmark

- No improvements wrt base model /
- Poor classification overall 1

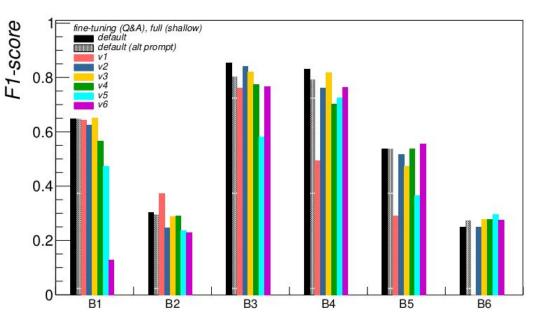
FR-II

Fine-tuning Results (B1-B6)

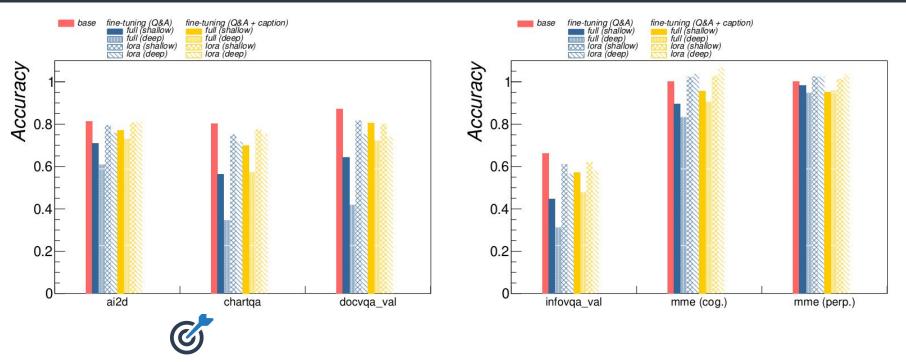
No significant improvement with alternative model hyperparameters

- ✓ Alternative learning rates or learning schedules
- ✓ Alternative batch sizes
- ✓ Alternative LORA parameters

Model variant	Parameters		
v1	$lr=5 \times 10^{-5}$		
v2	$lr=5 \times 10^{-6}$		
v3	warmup_ratio=0.01		
v4	cosine_with_min_lr (lr=5 $\times10^{-6}$)		
v5	effective_batch_size=256		
v6	LORA rank=128, alpha=256		



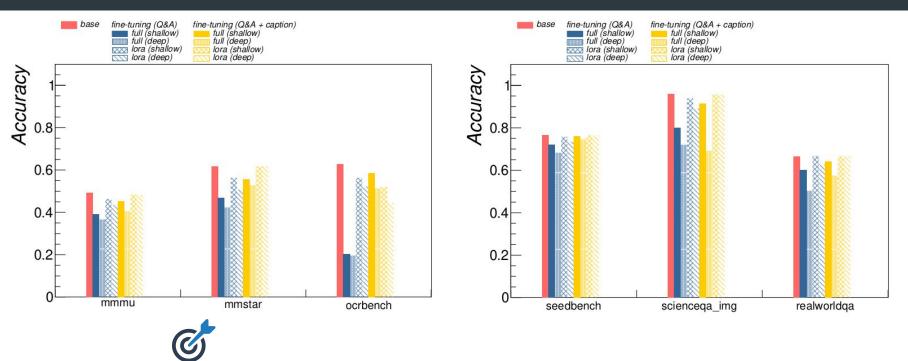
Fine-tuning Results (Multi-Modal Benchmarks)



Catastrophic forgetting of previously learned multi-modal tasks

- ✓ 20-30% accuracy drop
- ✓ getting worse in deeper training runs
- ✓ caption data & LORA fine-tuning mitigating

Fine-tuning Results (Multi-Modal Benchmarks)



Catastrophic forgetting of previously learned multi-modal tasks

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- ✓ getting worse in deeper training runs
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Solution Summary

VLMs offering an intuitive text-based interface for running radio source analysis

First attempt to evaluate VLMs on radio data

- ✓ Full fine-tuning improves mostly B1 and B3 by >30% F1-score
- ✓ LoRA gains are smaller but avoid catastrophic forgetting
- ✓ Best results still behind vision-only task-customized models
- ✓ Caption data improve instruction following & generalization to non-radio benchmarks
- ✓ Minor gains from prompt or hyperparameter changes

Challenges & limitations

- Multimodal misalignment & training data quality/diversity are key bottlenecks to performance
- ✓ Strategies to mitigate catastrophic forgetting on non-radio tasks



Q | Large Language Models (LLMs)

LLMs increasingly used for various tasks

- Drafting & analyzing documents
- Generating & modifying software codes
- Inspecting/analyzing/formatting data

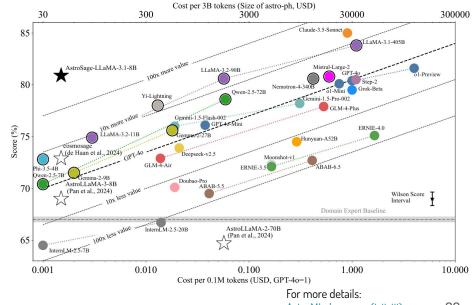
■ Why developing LLMs for astronomy?

- Existing LLMs less effective on astro tasks or having prohibitive inference costs for large-scale analysis
- Data privacy reasons

Ongoing projects focusing on small LLMs

Astrollama, cosmosage, astroLLM, astroBERT, AstroSage, ...

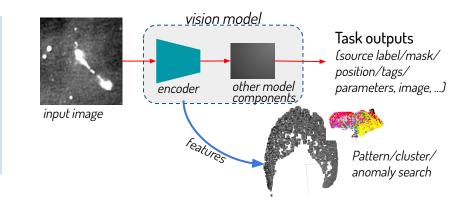
- Trained on entire ArXiV astroph paper corpus (+ other resources) via instruction fine-tuning or RAG
- Catastrophic forgetting is a major challenge



\mathbb{Q} | Vision Models (VMs)

VMs used for various tasks on visual data

- Source classification
- Source segmentation
- ✓ Anomaly search
- Data generation
- Similarity search



VM including vision encoders (CNN/ViT-based)

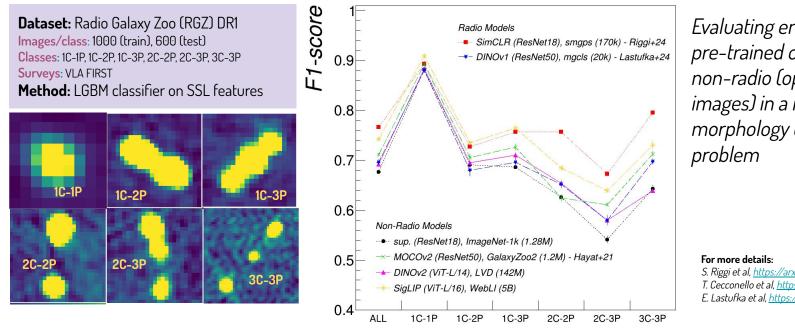
- Extracting features from visual inputs
- Pre-trained on large data samples (often in a self-supervised way)
- ✓ Model & data representation learnt can be used for data inspection and analysis and re-used for different tasks

■ Why developing VMs for astronomy?

- Extract high-quality features from astronomical data
- Improve supervised analysis in the low-label regime

Various foundational models developed

- ✓ Radio SSL models provide best performance on radio data/tasks
- ✓ Marginal improvements in some tasks compared to ViT-pretrained models or fully-supervised models
- ✓ Ongoing efforts: scale-up to millions of images, improve downstream task datasets



Evaluating encoders pre-trained on radio vs non-radio (optical, natural images) in a radio source morphology classification problem

For more details: S. Riggi et al, <u>https://arxiv.org/abs/2411.08519</u> (2024) T. Cecconello et al, <u>https://arxiv.org/abs/2411.14078</u> (2024) E. Lastufka et al, <u>https://arxiv.org/pdf/2409.11175</u> (2024)