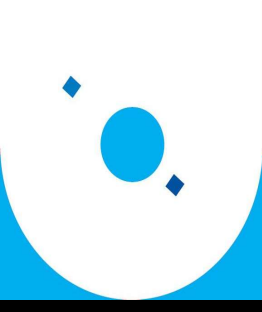


WST

the Wide-field
Spectroscopic
Telescope

Surveying the Universe in the 2040's and beyond

Italian Workshop
in memory of Bianca Garilli

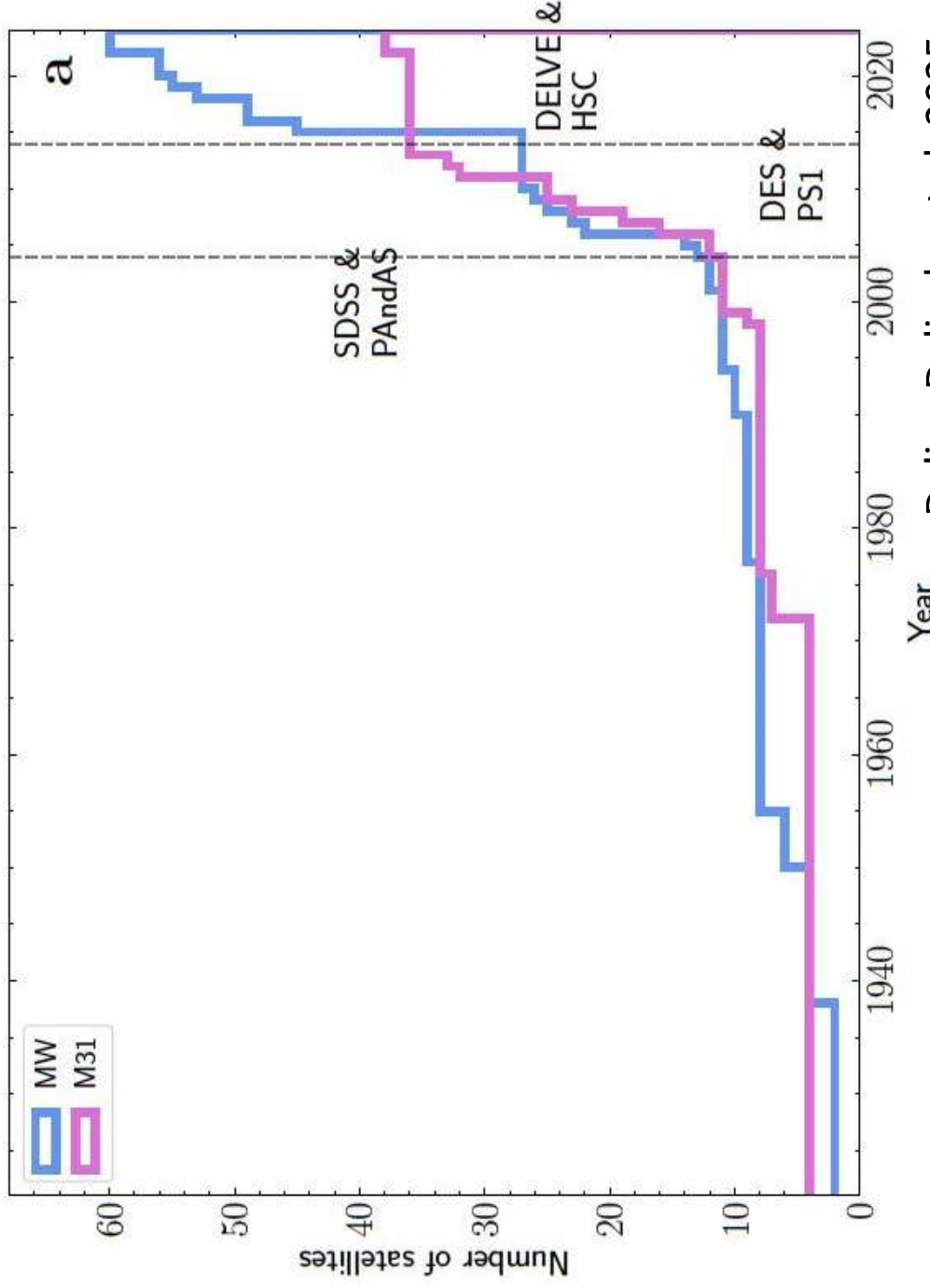
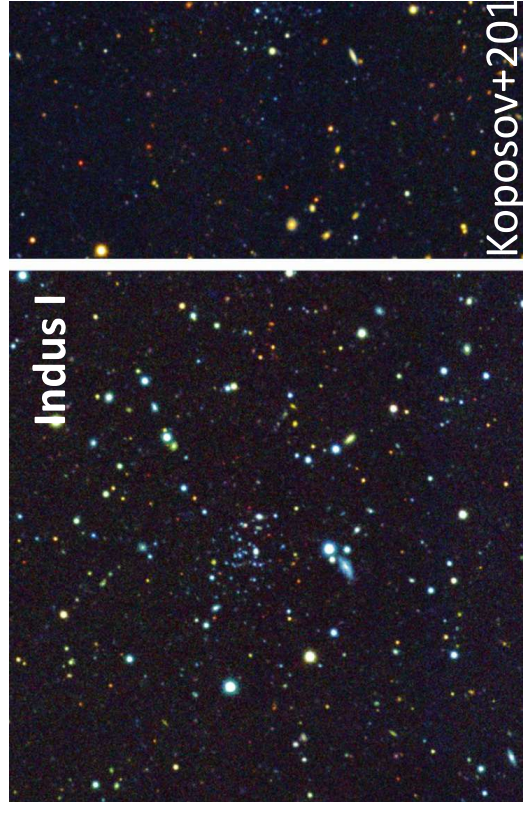
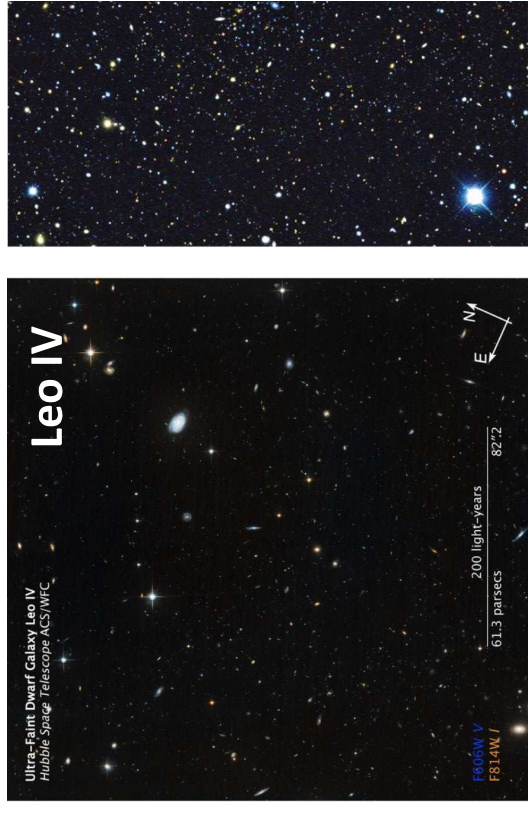


Unveiling the Secrets of Ultra-Faint Dwarf Galaxies w WST

Massimiliano Gatto (INAF – Osservatorio Astronomico di Capodimonte)

The Ultra-Faint Dwarf Galaxies

ULTRA-FAINT DWARF GALAXY

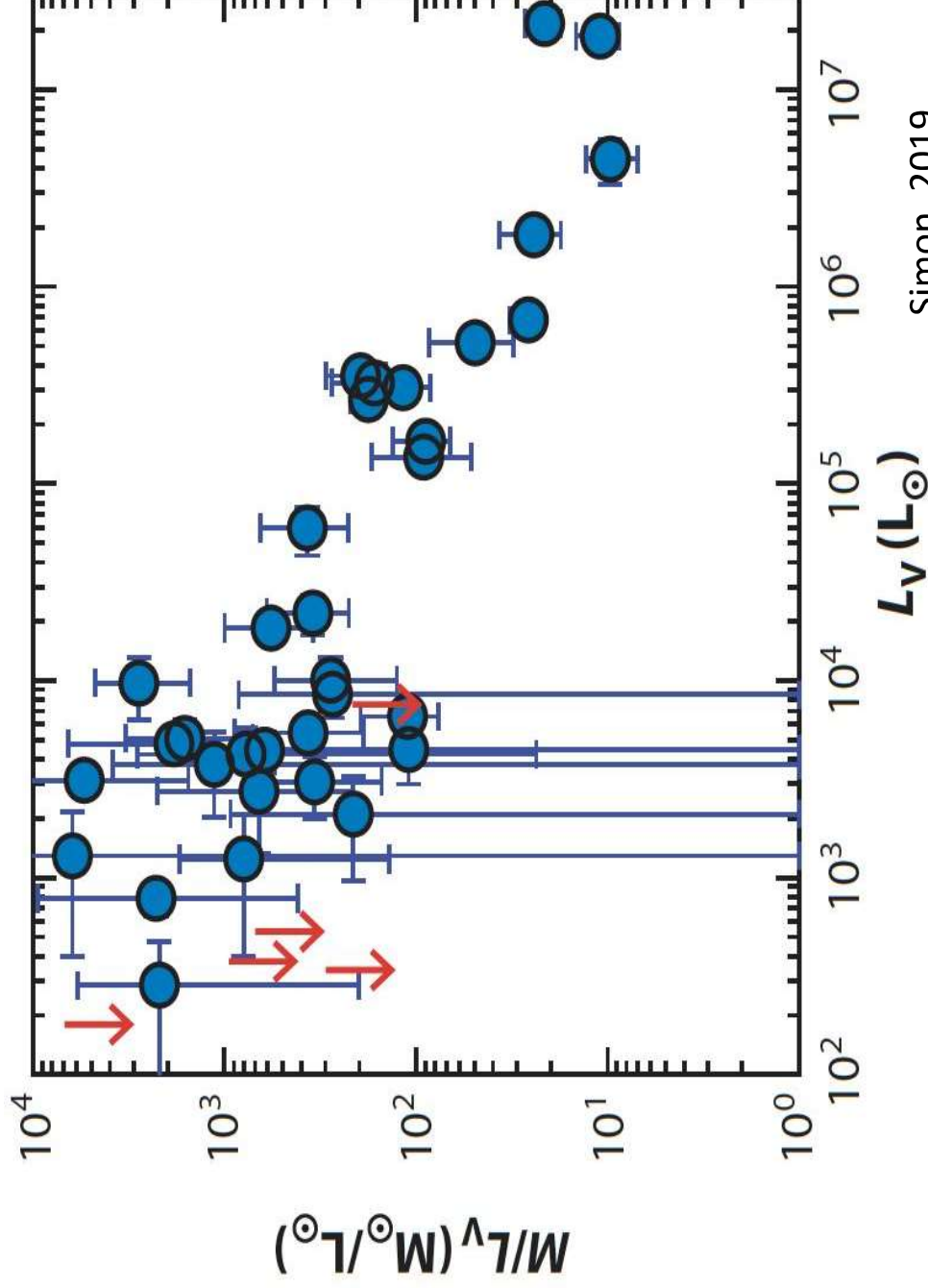


Doliva-Dolinsky et al. 2025

Koposov+2011

The UFDs as cosmic probes

The most dark matter-dominated systems known



- Nature of the Dark Matter
- Core - Cuspy problem

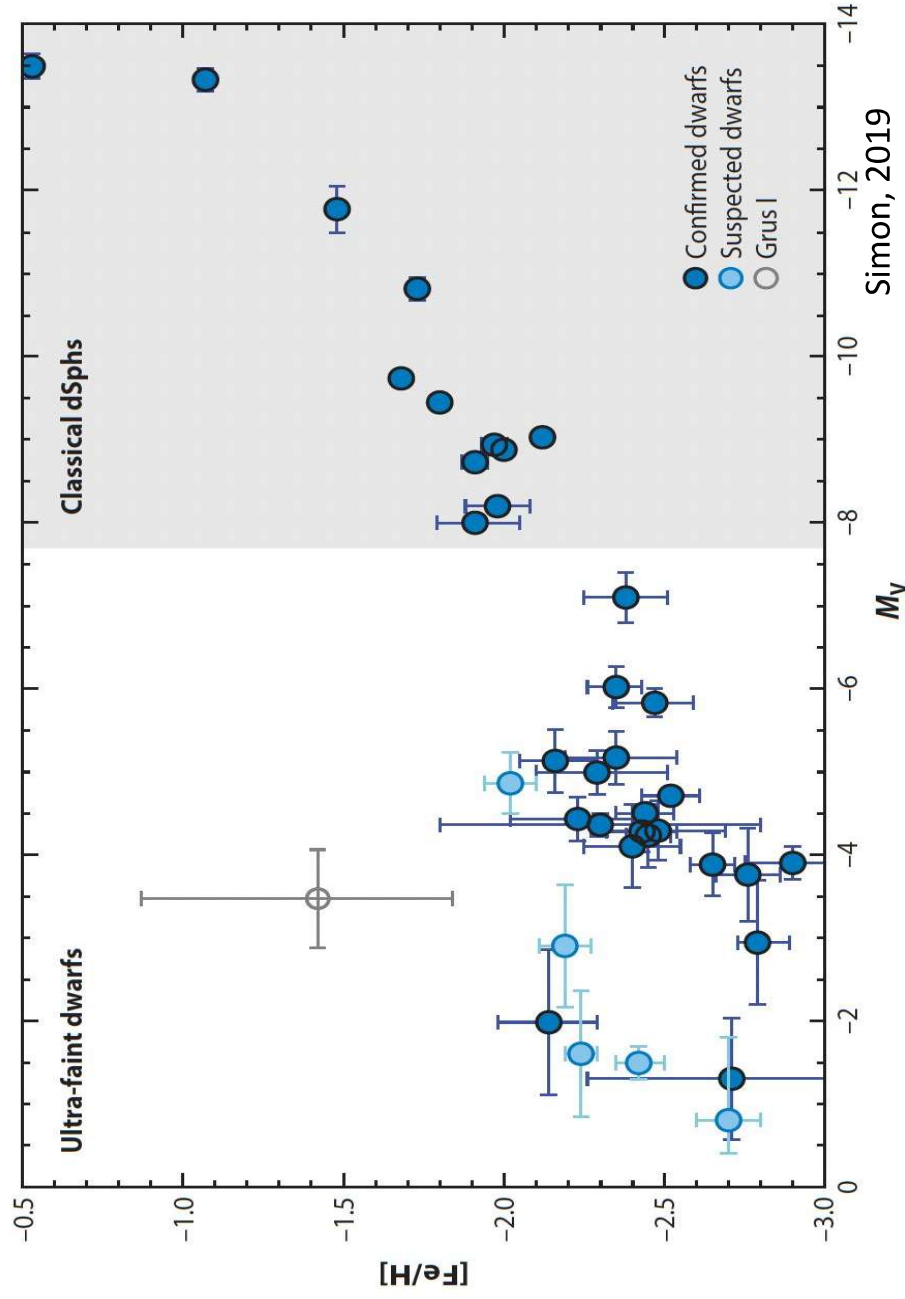
Velocity dispersions of about 3-8 km/s
(Simon, 2019)

CAVEATS:

- Few Bright Stars
- Outliers
- Binaries

The UFDs as proxy of early chemical evolution

The most metal-poor galaxies known



Simon, 2019

□ Nucleosynthesis

□ Population III

Nearly half of known MW UFDs lack metallicity information

CAVEATS:

- Few Bright Stars
- Heterogenous sample
- Lack of chemical abundances

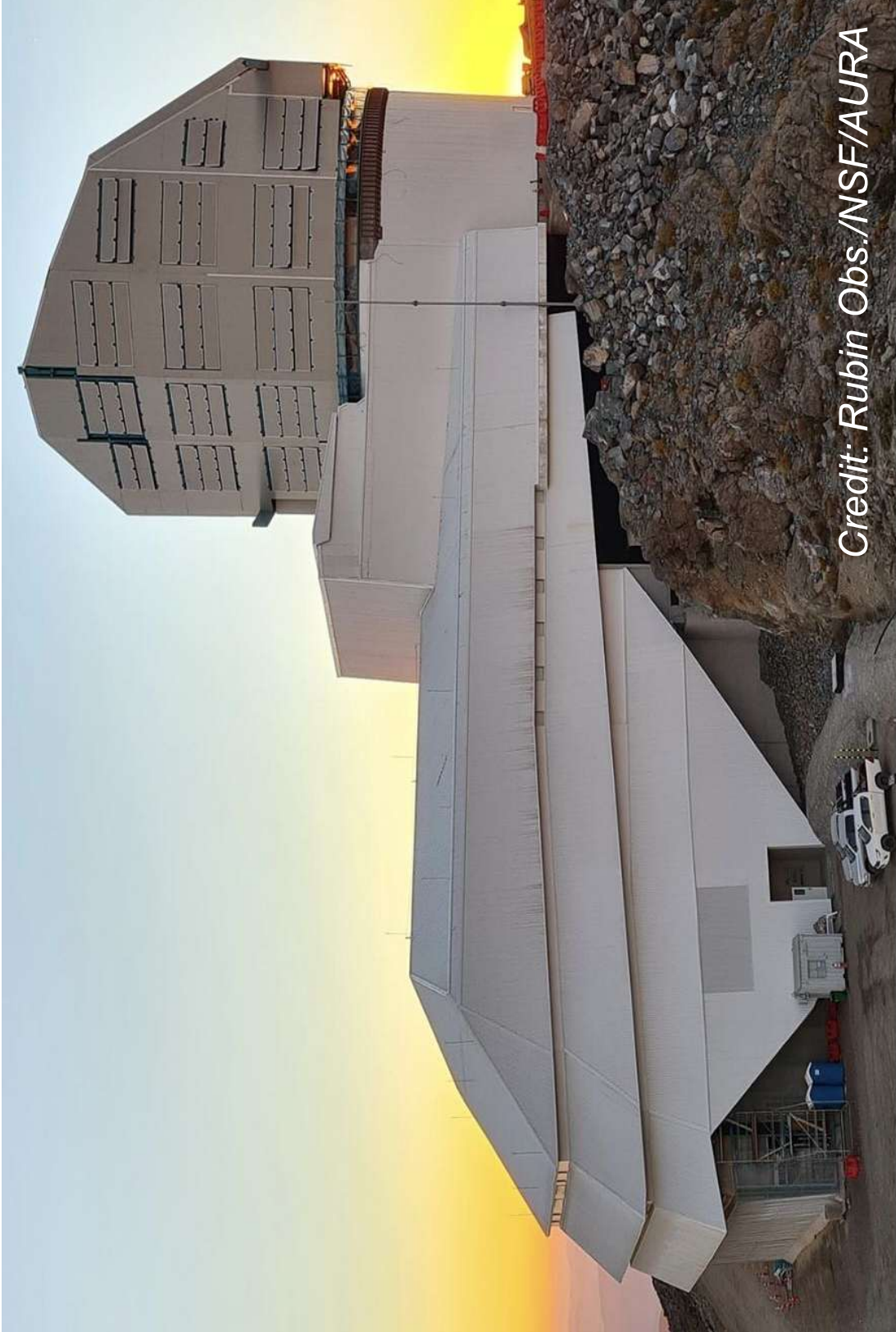
The UFDs in the next decade

Vera Rubin Observatory - LSST

- Six filters (ugrizy)
 - Repeated images.
 - 9.6 deg² field of view
-
- $r \sim 24.5$ with a single exposure
 - $r \sim 27.5$ after 10 years

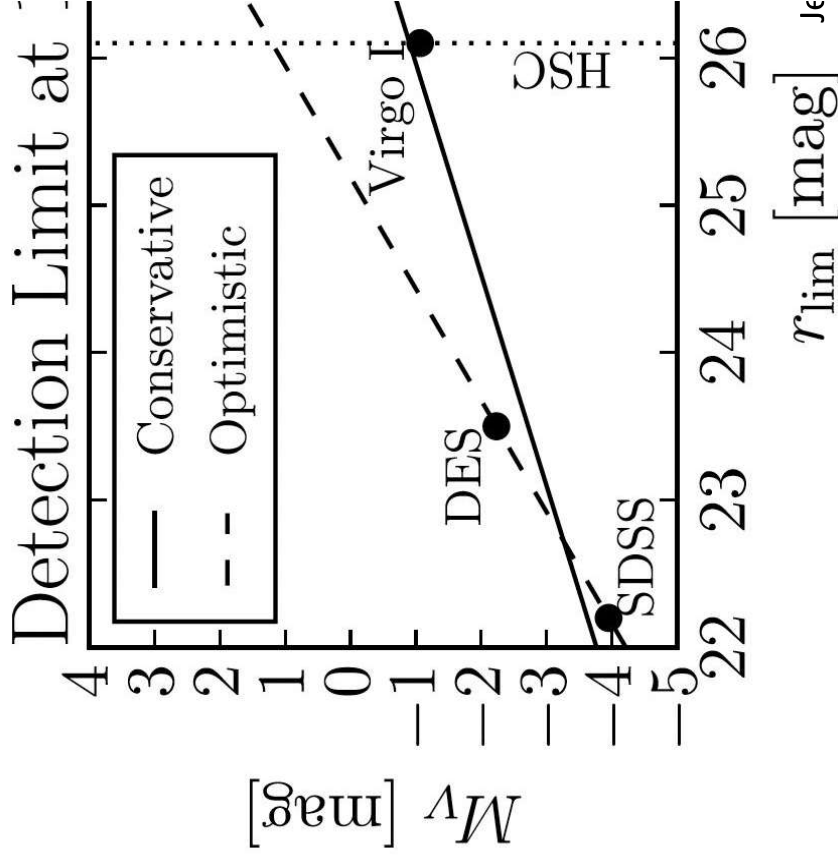
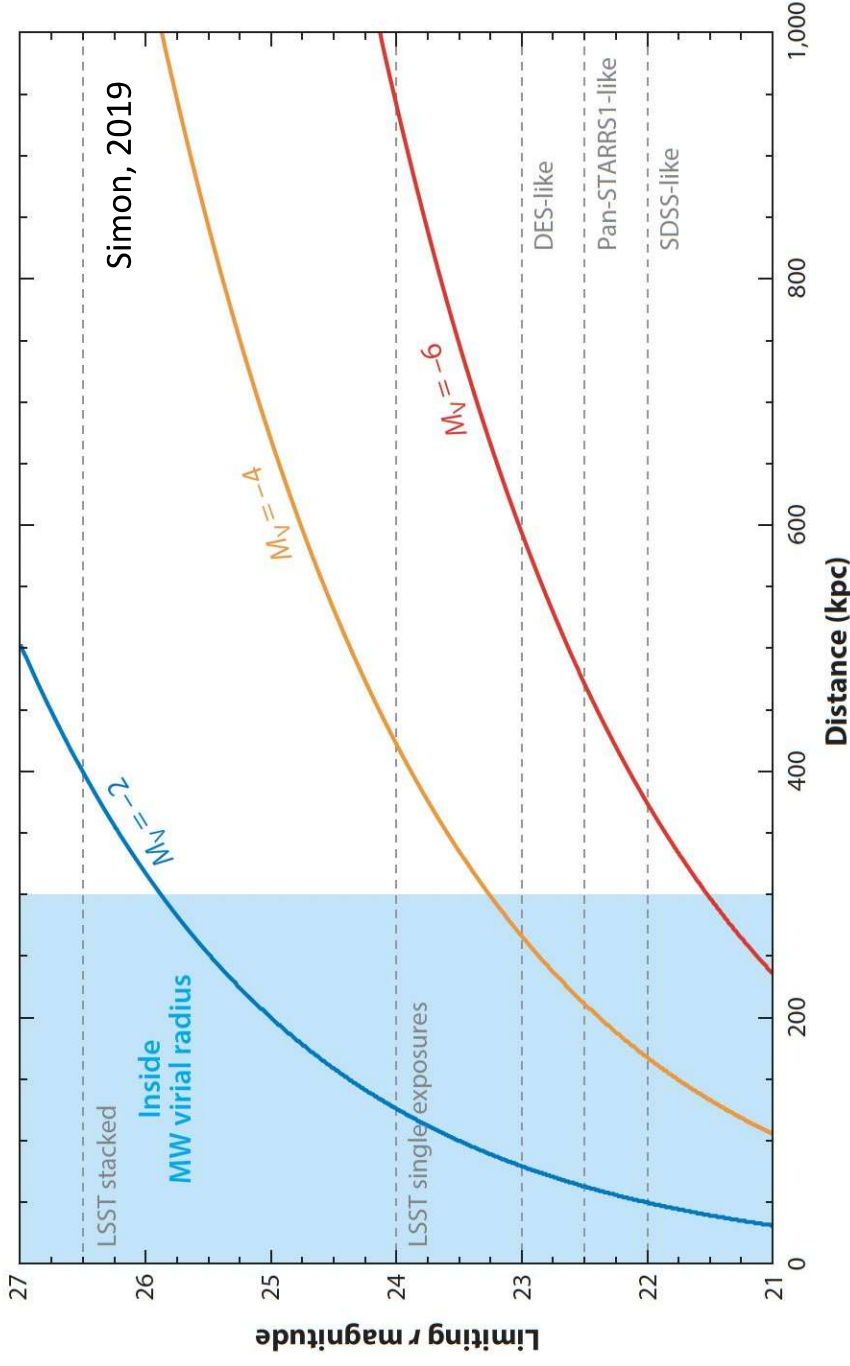


LSST will resolve
sequence turnoff stars
300 kpc



Credit: Rubin Obs./NSF/AURA

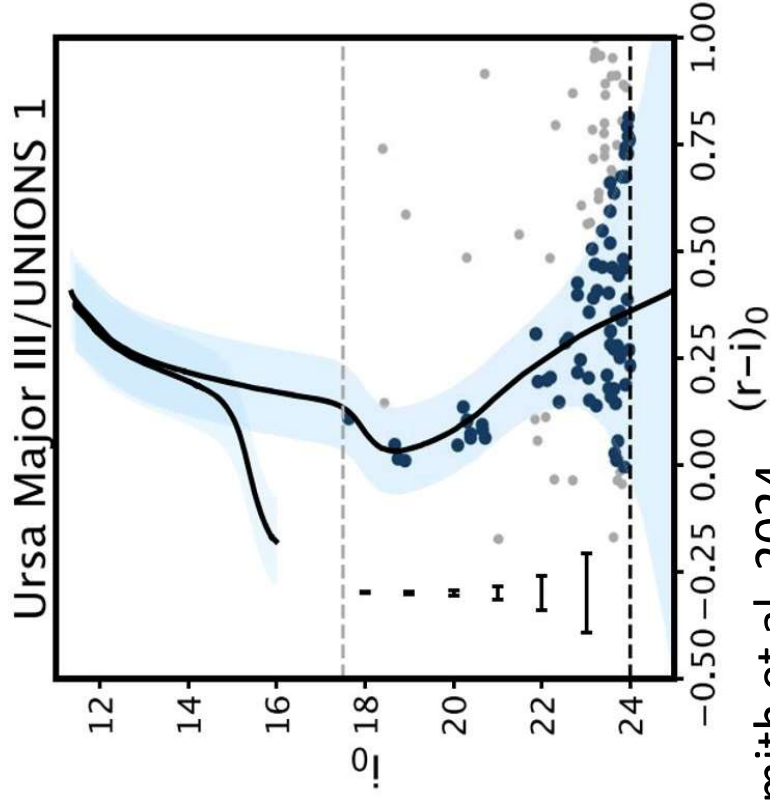
The UFDs in the next decade with the LSST



$$M_* < 100 M_{\odot}$$

Ursa Major III/UNIONS 1

- ❖ Spectroscopically confirmed with Keck II/DEIMOS
- ❖ Absolute Magnitude of $M_V = +2.2$ mag
- ❖ Stellar Mass of $M_* = 16 M_\odot$
- ❖ Darkest UFD ever discovered
- ❖ Heliocentric distance ≈ 10 kpc



Smith et al. 2024

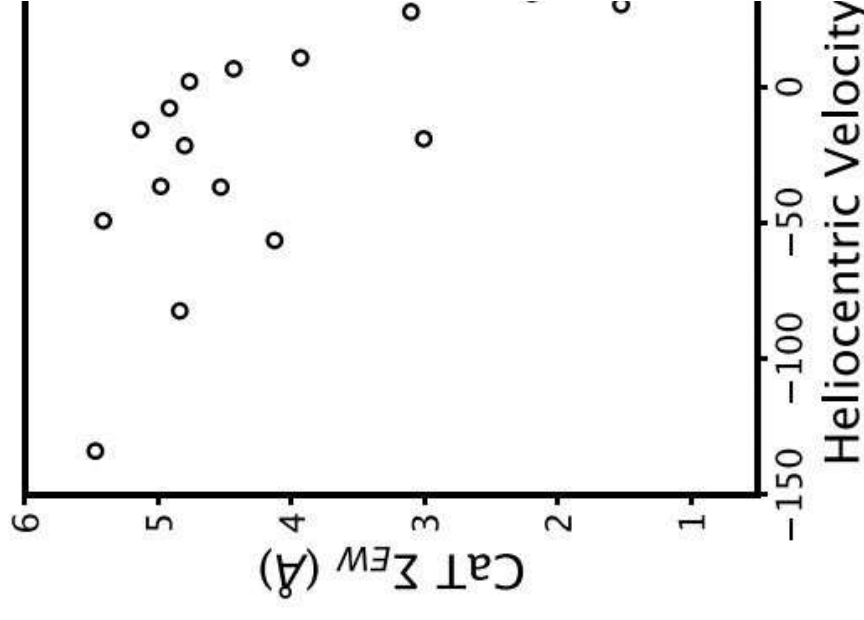
❖ Spectroscopically confirmed with Keck II/DEIMOS

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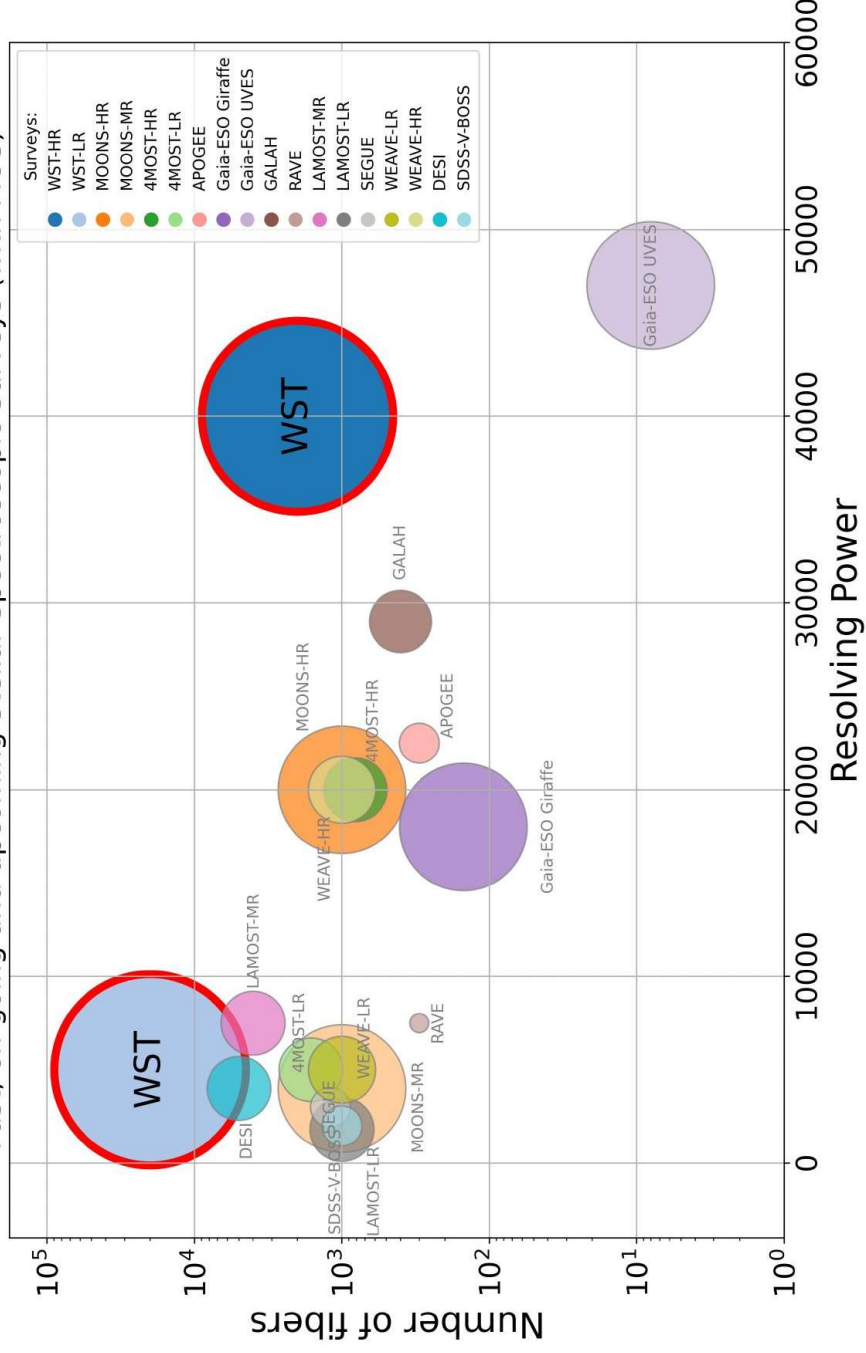
Expected hundreds of UFDs similar to UM

(e.g., Tollerud+2008, Hargis+2014, Jethwa+2018, Newton+2018, Sim

Important implications for the minimum halo threshold able to trigger the formation of s

The Wide-Field Spectroscopic Telescope

Past, on-going and upcoming Stellar Spectroscopic Surveys (with MOS)



Facility/Instrument	First light	Aperture (M1 in m)	Field of View (sq. deg)	Etendue	Multiplexing	Wavelength coverage (μm)
TMT/WFOS	2030+	30	0.007	4.9	80	0.30-1.0
VLT/FLAMES-GIRAFFE	2002	8.2	0.14	7.4	132	0.37-0.95
VLT/MOONS	2024	8.2	0.14	7.4	1000	0.65-1.8 3 windows
ELT/MOSAIC	2030+	39	0.009	10.7	200 100	0.39-1.80 windows
WHT/WEAVE	2022	4	3.14	39.5	1000	0.37-1.00 3 windows
GMT/MANIFEST	2030+	25.4	0.09	45.5	250	0.35-0.95
VISTA/4MOST	2024	4	4.2	52.8	1600 800	0.37-0.95 3 windows
Subaru/PFS	2024	8.2	1.25	65.9	2400	0.38-1.26
Mayall/DESI	2019	4	7.1	89.2	5000	0.36-0.98
MSE-QM	2030+	11.25	1.5	149	20000	0.36-1.8 0.36-0.95 3 windows
MUST	2030+	6.5	5	166	10000	0.36-1.0
MegaMapper	2030+	6.5	7.1	235	20000	0.36-0.98
WST	2030+	12	3.1	350	20000 2000	0.37-0.97 windows

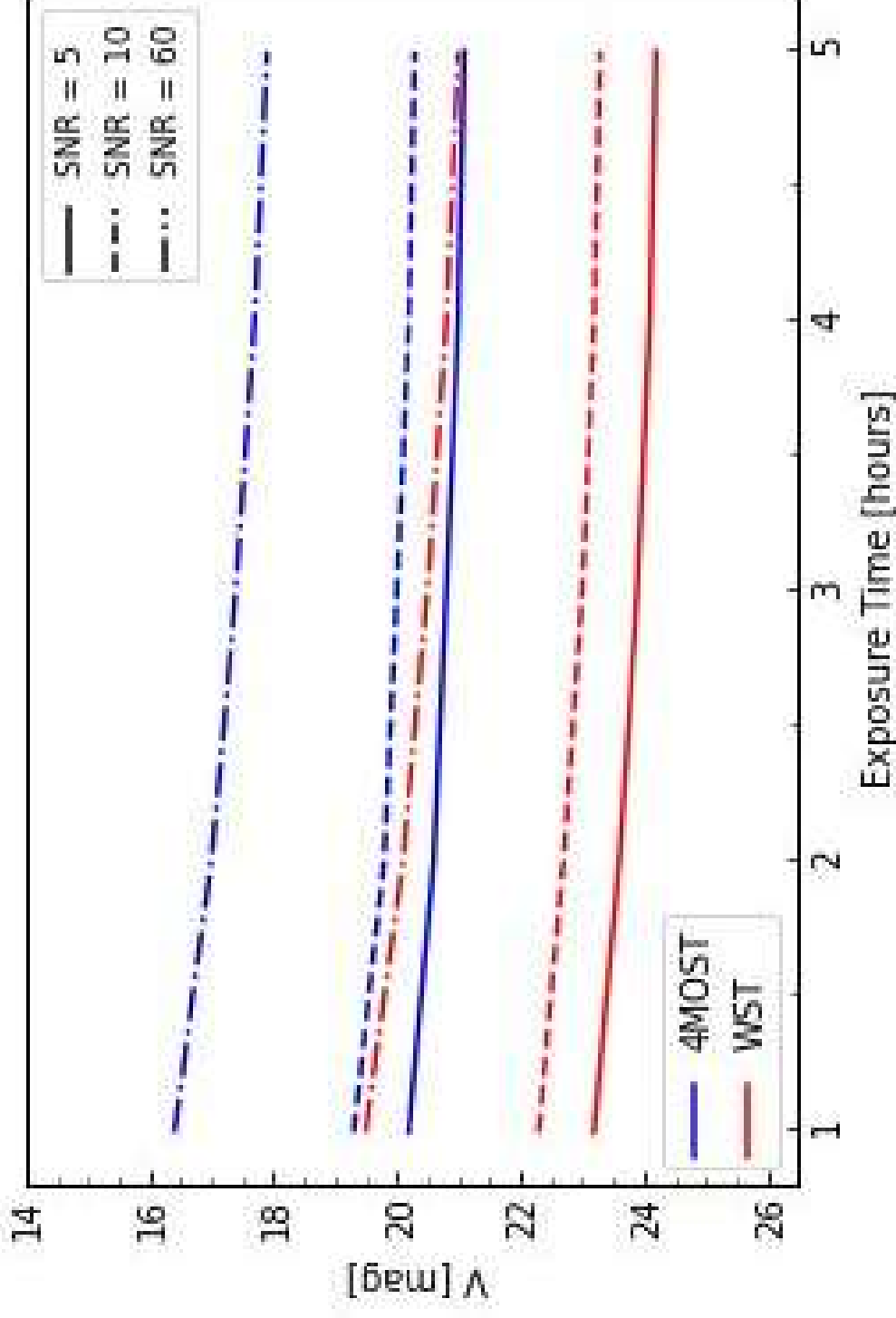
The WST White paper

WST Capabilities

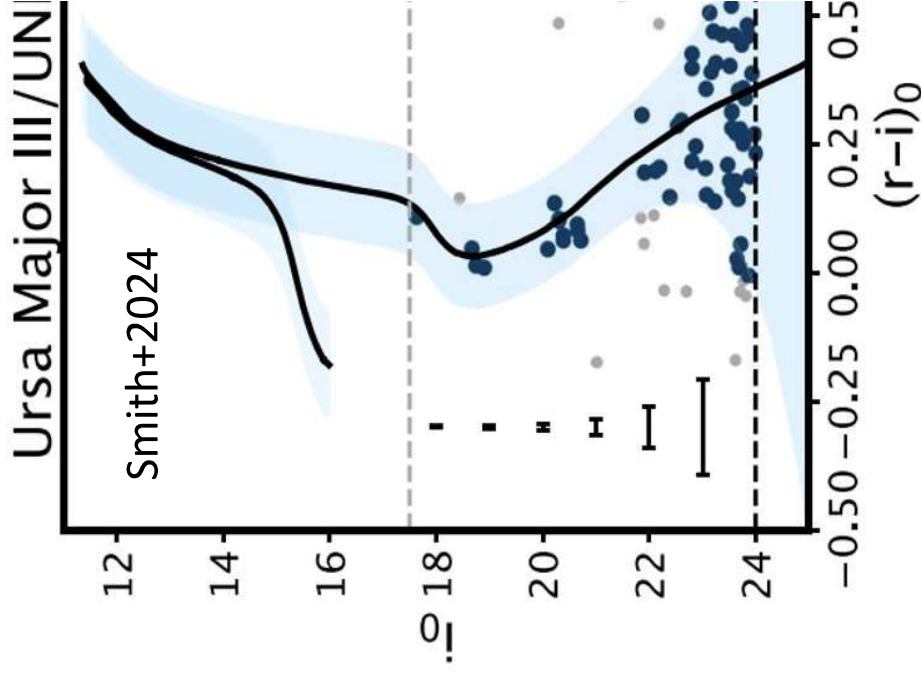
<https://github.com/RolandBacon/pyetc>

LOW RESOLUTION MOS

Main sequence star; Dark; Seeing = 0.8"; Airmass = 1.0"



- $M_V = +2.2$ mag
- No post main-sequence



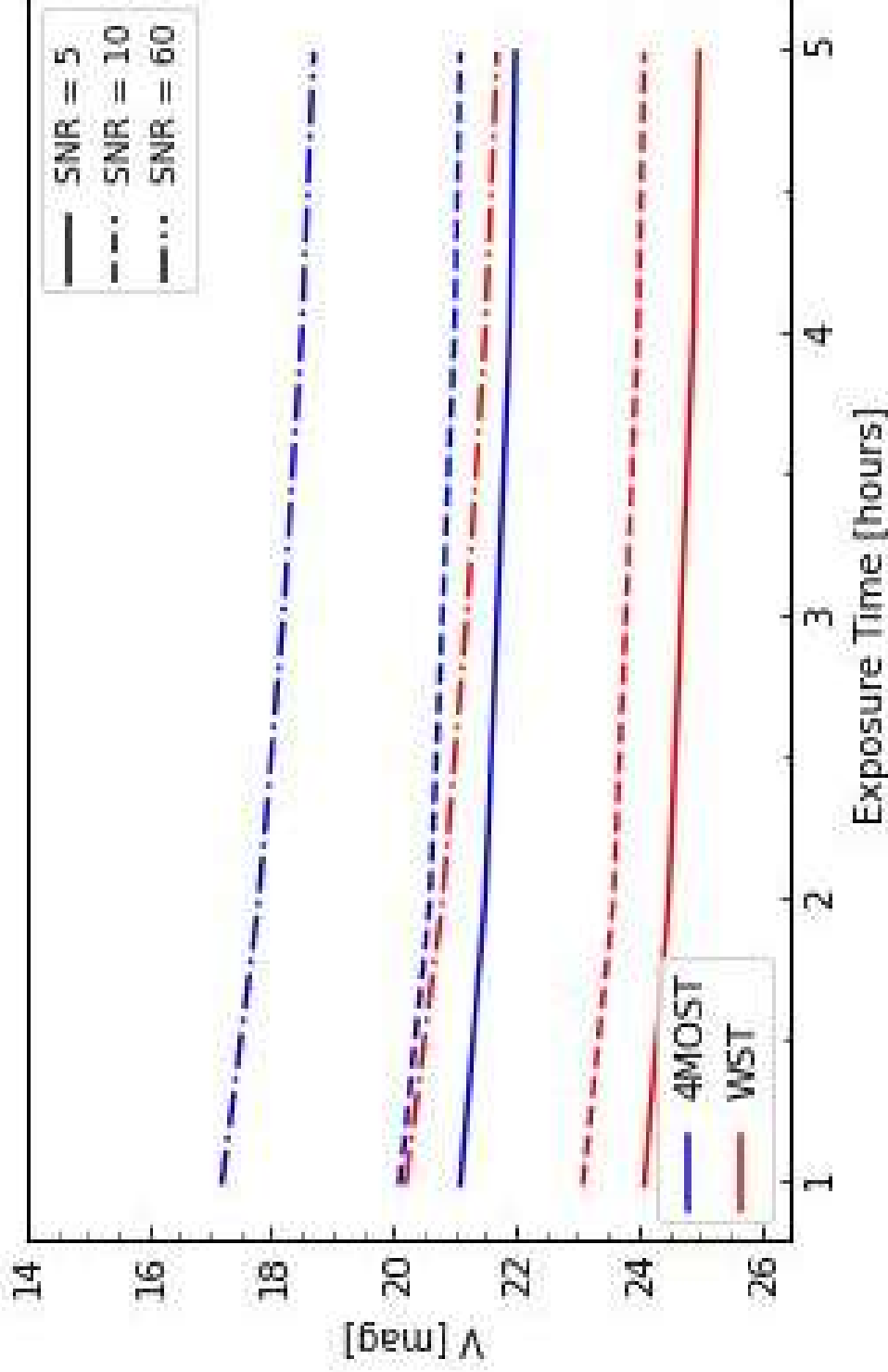
- ❖ Radial Velocities and $[\text{Fe}/\text{H}]$ u
- ❖ Velocity Dispersion up to 80
- ❖ Chemical abundances (Ca-Mg)

WST Capabilities

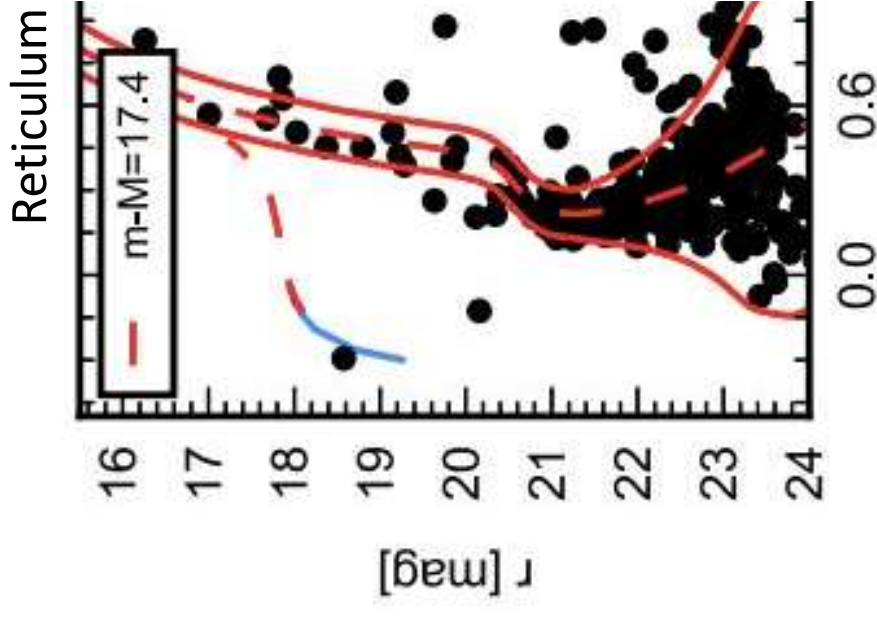
<https://github.com/RolandBacon/pyetc>

LOW RESOLUTION MOS

Red giant branch star; Dark; Seeing = 0.8"; Airmass = 1.0"



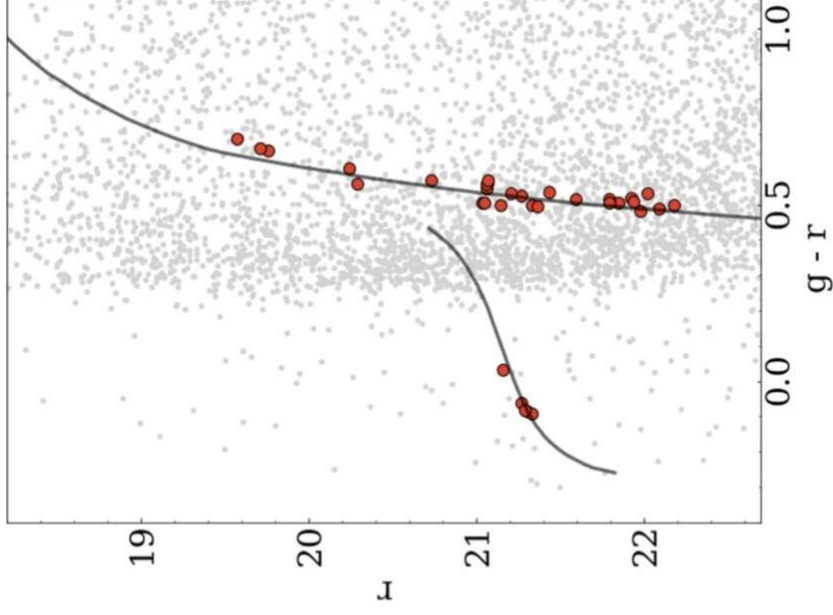
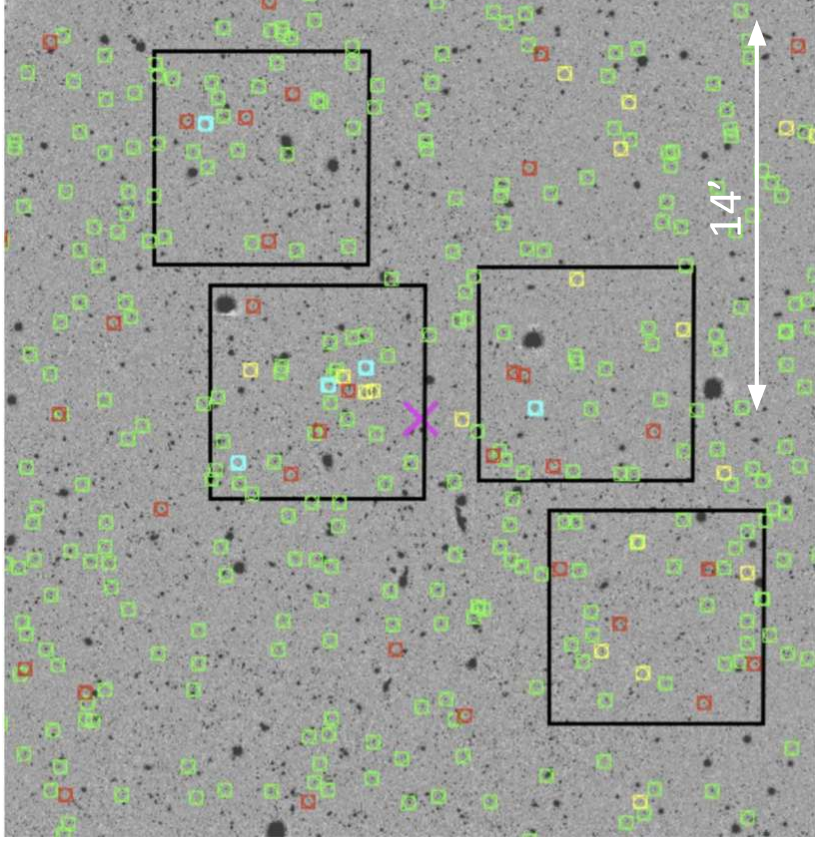
- $M_V = -2.7$ mag
- Few RGB Stars



- ❖ Radial Velocities and $[Fe/H]$ up to 100 kpc
- ❖ Velocity Dispersion up to 300 kpc
- ❖ Chemical abundances (Ca-Mg-N) up to 100 kpc
- ❖ High Resolution up to ≈ 80 kpc

WST Capabilities

Sextans II



- ❖ 150 kpc MSTO at $r \cong 24'$
- ❖ $M_V = -3.9$ mag
- ❖ $r_h \cong 200$ pc \rightarrow 5 arcmin

FORS2@VLT Follo

- ❖ 13 hours
- ❖ 32 stars down to $r \cong 24'$
- ❖ Sampled only up to $R \cong 10'$
- ❖ Outskirts unexplored

WST

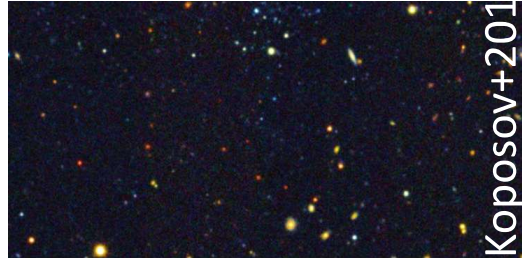
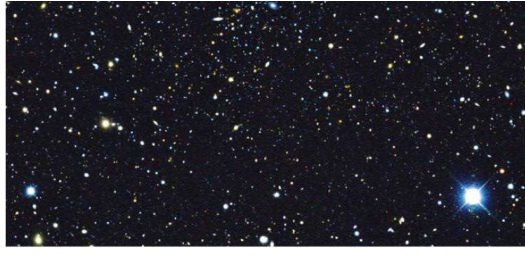
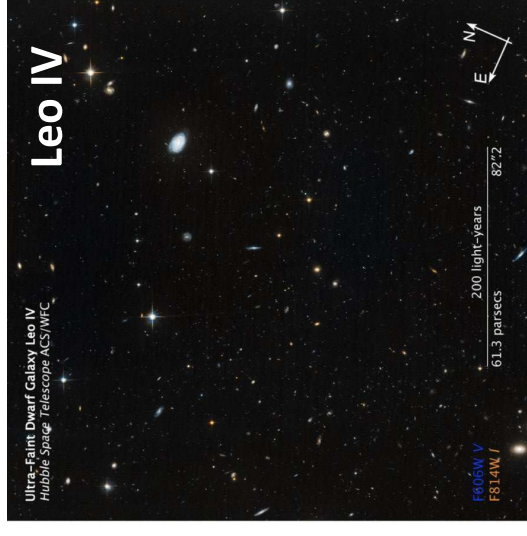
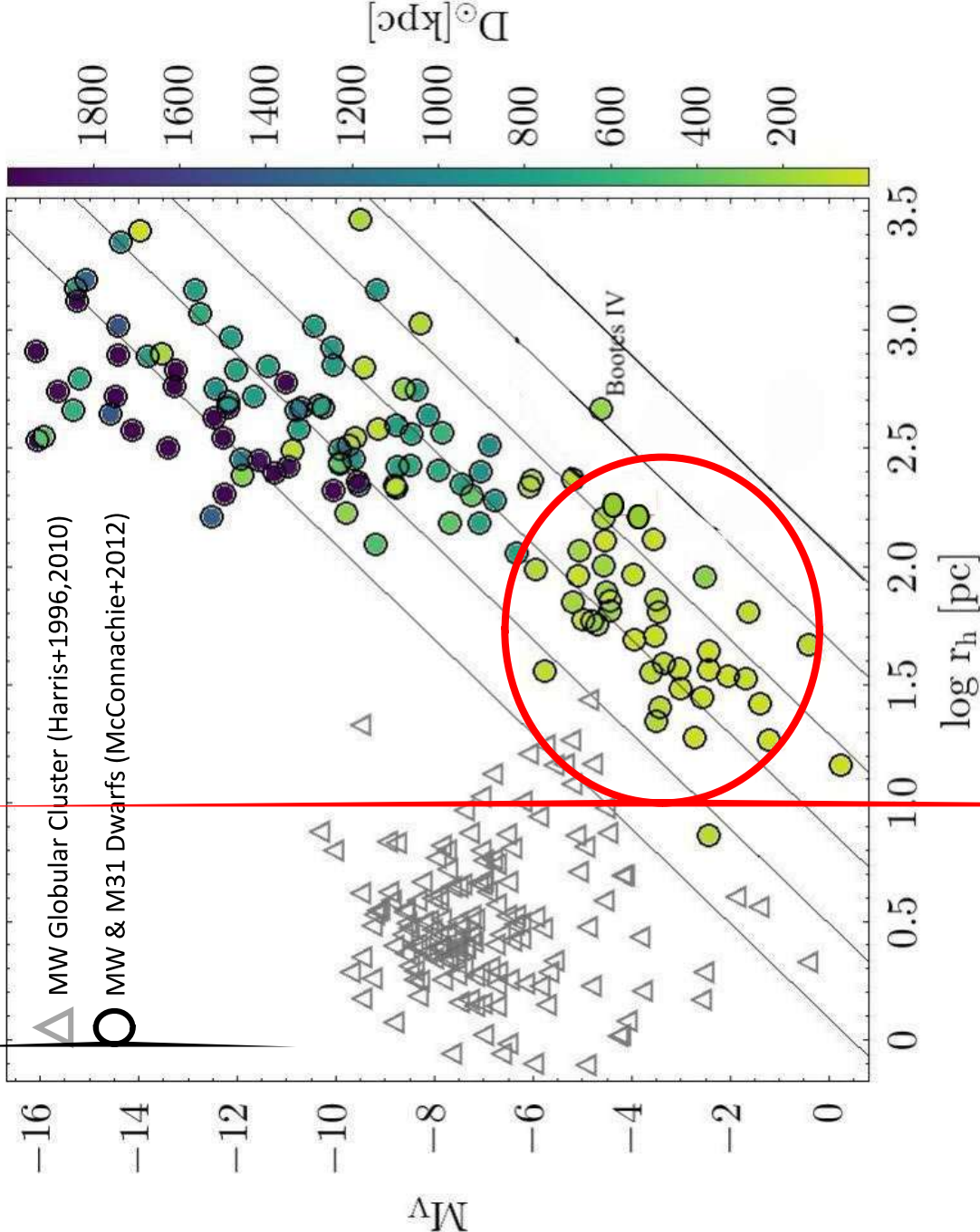
- 1 hour \rightarrow RV and $[\text{Fe}/\text{H}]$ for all stars down to \cong MSTO
- 1 single pointing \rightarrow Complete catalog of spectroscopic confirm
- Possibilities to unveil tidal streams
- Probing the presence of metallicity gradients

Summary

- ✓ Spectroscopic counterpart of the LSST
- ✓ Constrain dark matter content in UFDs
- ✓ Orbital properties of Milky Way Satellites
- ✓ Determine the metallicity distribution functions
- ✓ Chemical abundances of a significant fraction of stars
- ✓ Unveil tidal structures in UFDs
- ✓ Quantify the fraction of stripped stars as a function of galactocentric radius

The Ultra-Faint Dwarf Galaxies

ULTRA-FAINT DWARF GALAXY

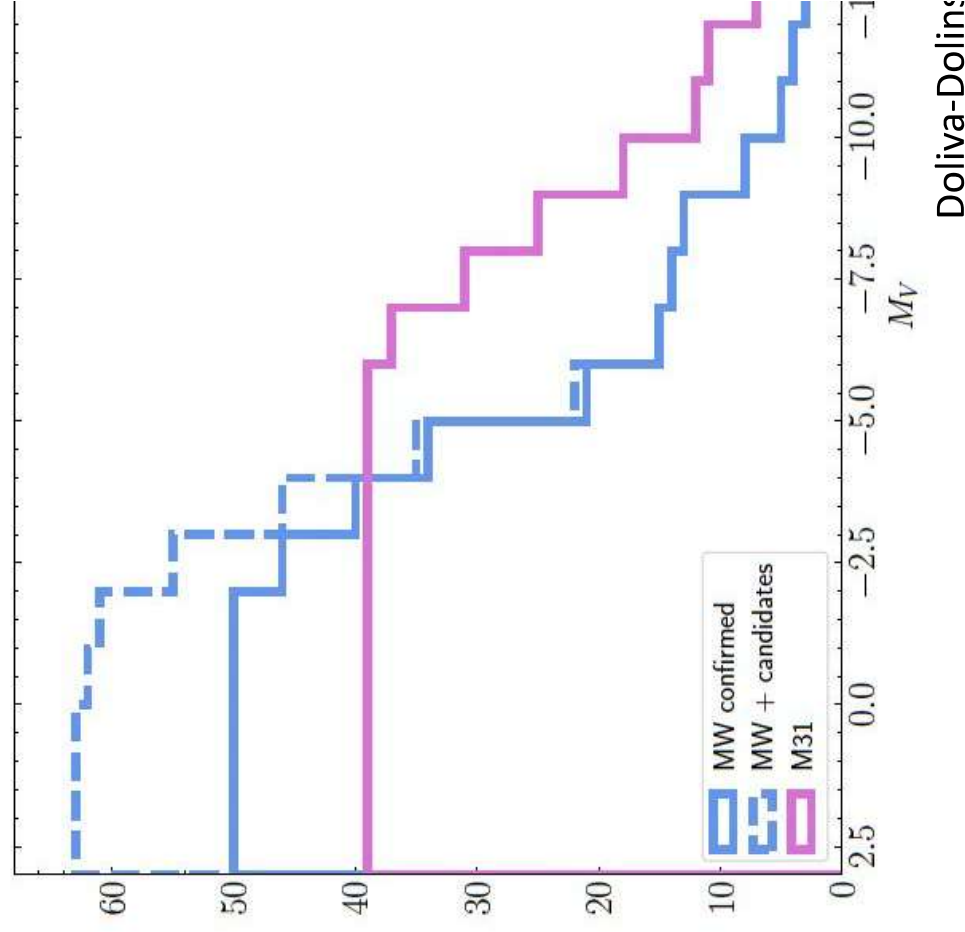


Koposov+2011

Confirming a UFD is challenging especially at faint magnitudes

Metallicity
Spread

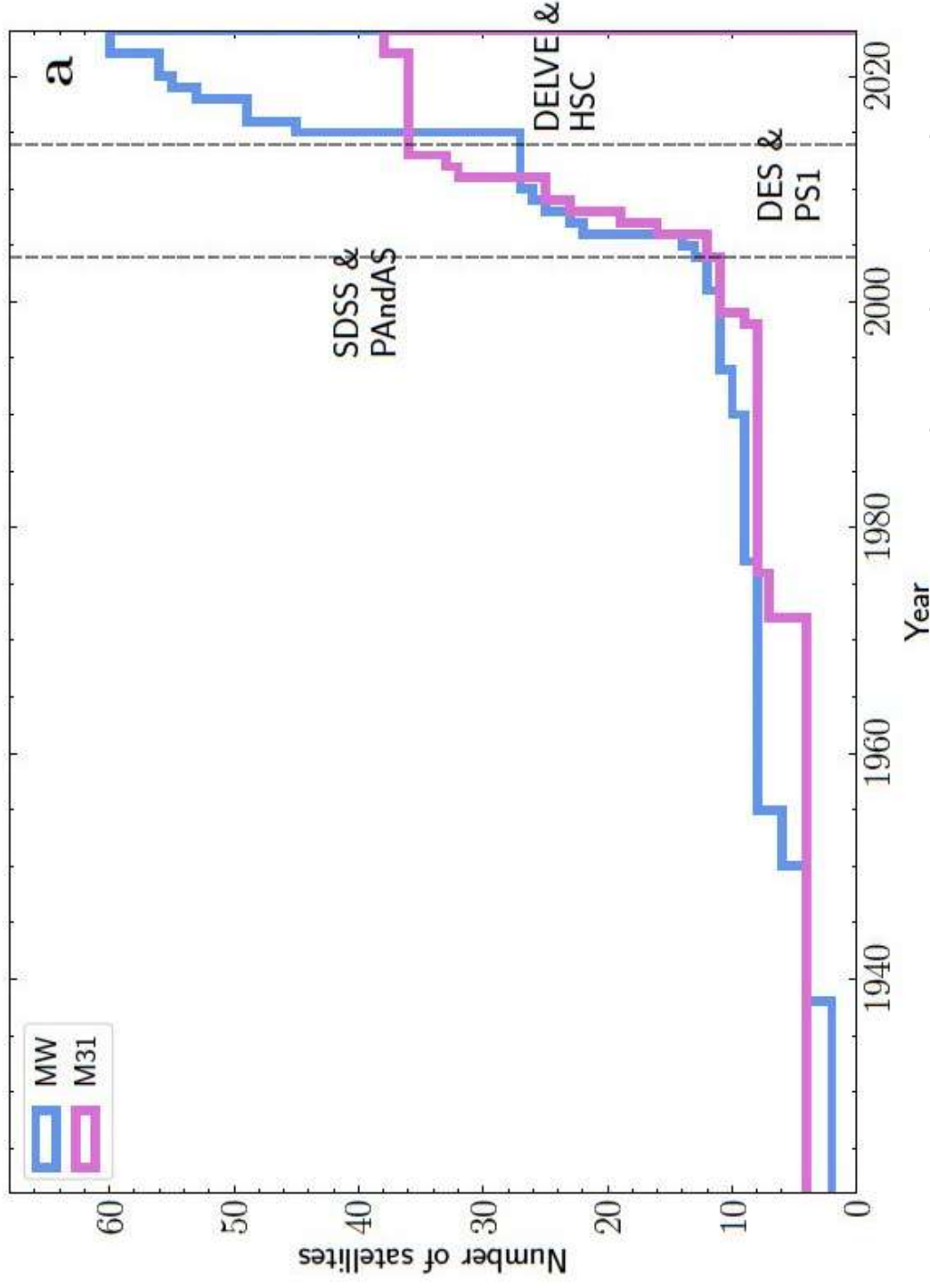
Presence of Dark
Matter



**SPECTROSCOPIC
FOLLOW-UP**

Doliva-Dolins

The census of Milky Way satellites



- About 60 MW satellites (confirmed + candidates)
- The advent of the era enables to sample faint-end of the Luminosity Function

Doliva-Dolinsky et al. 2025