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## The WST science drivers: transformational science in the 2040s

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WST: a facility to answer a wide range of cutting-edge scientific questions that cannot be addressed with current or planned facilities



Astrophysics > Instrumentation and Methods for Astrophysics

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#### The Wide-field Spectroscopic Telescope (WST) Science White Paper

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The Wide-field Spectroscopic Telescope (WST) is proposed as a new facility dedicated to the efficient delivery of spectroscopic surveys. This white paper summarises the initial concept as well as the corresponding science cases. WST will feature simultaneous operation of a large field-of-view (3 sq. degree), a high multiplex (20,000) multi-object spectrograph (MOS) and a giant 3x3 sq. arcmin integral field spectrograph (IFS). In scientific capability these requirements place WST far ahead of existing and planned facilities. Given the current investment in deep imaging surveys and noting the diagnostic power of spectroscopy, WST will fill a crucial gap in astronomical capability and work synergistically with future ground and space-based facilities. This white paper shows that WST can address outstanding scientific questions in the areas of cosmology; galaxy assembly, evolution, and enrichment, including our own Milky Way; origin of stars and planets; time domain and multi-messenger astrophysics. WST's uniquely rich dataset will deliver unforeseen discoveries in many of these areas. The WST Science Team (already including more than 500 scientists worldwide) is open to the all astronomical community. To register in the WST Science Team please visit this https URL

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## Time-domain: exciting science on all scales





Transient astronomy (stellar flares, SNe, TDEs, GRBs, etc.)



Gravitational Wave Source Characterization, Calibration sources



Supermassive BHs (AGN, QSOs)



## Transients post LSST/Roman era

In the 2040s LSST will have discovered hundreds of thousands of transients. Our understanding of transientrelated phenomena will be transformed. We will probably have even more questions than today.

Build time-domain capabilities (e.g., Operational model, fast data processing) in WST to enable new science

- Fiber-level ToO: large scale follow-up of transients with WST will be key to get a full picture of their nature.
- **Spectroscopic variability**: WST will issue alerts enabling the detection of new classes of transiets.



## Detection and characterization of the EM of GW events from the next generation of detectors

Einstein Telescope will detect 10<sup>5</sup> binary neutron stars mergers per year, also prior of the mergers during the in spiral phase.

WST spectroscopic identification and characterization of the EM counterparts will be crucial for kilonovae population studies and study of the extreme physics during the merger phase.





Credit to Sofia Bisero (GEPI)



#### **Galactic science cases**

Three science themes:

- Origins of elements
- Origin of the Milky Way system
- Origin of stars and planets

#### Science in the 2040s: Building upon Gaia+4MOST+MOONS+WEAVE



# (X. Luri & the DPAC-CU2)

#### **Origins of elements**

- Neutron capture elements (r-process, i-process, other yet unknown channels?)
- Type Ia supernovae: how many sub-types? How many of them important as nucleosynthetic sources?
- HR needed to probe diverse nucleosynthetic channels
- Not only reaching fainter targets, but adding new key and precise information





## The formation history of the MW and its satellites

Detailed abundances for studies of the chemical enrichment history of all populations/structures

Whole sky survey (HR+LR) -Full chemical information and RVs up to G~16-17 mag - A few element abundances and RVs up to G~18.5 mag - At larger distances: metallicities and RVs



Credit to Laura Magrini (INAF)



## The small-scale matter cycle



To access the relevant physics regulating the consumption of gas, the production of dust and metals and star-formation one needs to reach "cloud-scales" (e.g. average distance between HII regions)

Credit Francesco Belfiore (INAF)

across a range of

density).

100 pc resolution



## How galaxy evolution is driven by gas flow as a function of environment

How the Cosmic Web impact the evolution of galaxies?

Reconstruct the 3-D density distribution with an SDSS-fidelity from 6 million 1<z<4 galaxies over 200 deg<sup>2</sup>

The topology of the dark matter over large volumes from Lyα forest seen in absorption line spectra of background galaxies



MOS: IGM tomography in absorption and LSS in emission IFS: z=2-7 LAEs, small scale clustering, resolved CGM & IGM emission in Lyα



#### **Opening-up high redshift cosmology**



WST will conduct sub-percent galaxy clustering measurements beyond redshift of z~2

Sensitive constraints will be placed on :

- The level of non-gaussianity
- The sum of neutrino mass
- Inflationary model
  - Modified gravity





5

2

3

z

1





#### **Probing fundamental physics**



## Get involved in Wst



WST Special Session (SS7) at the EAS meeting in Cork (Ireland) June 23<sup>rd</sup>, 2025

### New scientific ideas are welcome!

Preparing a winning science case for the ESO Expanding Horizons proposal (due June 2027)





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## Thanks!