PRECURSOR ACTIVITY PRECEDING INTERACTING SUPERNOVAE

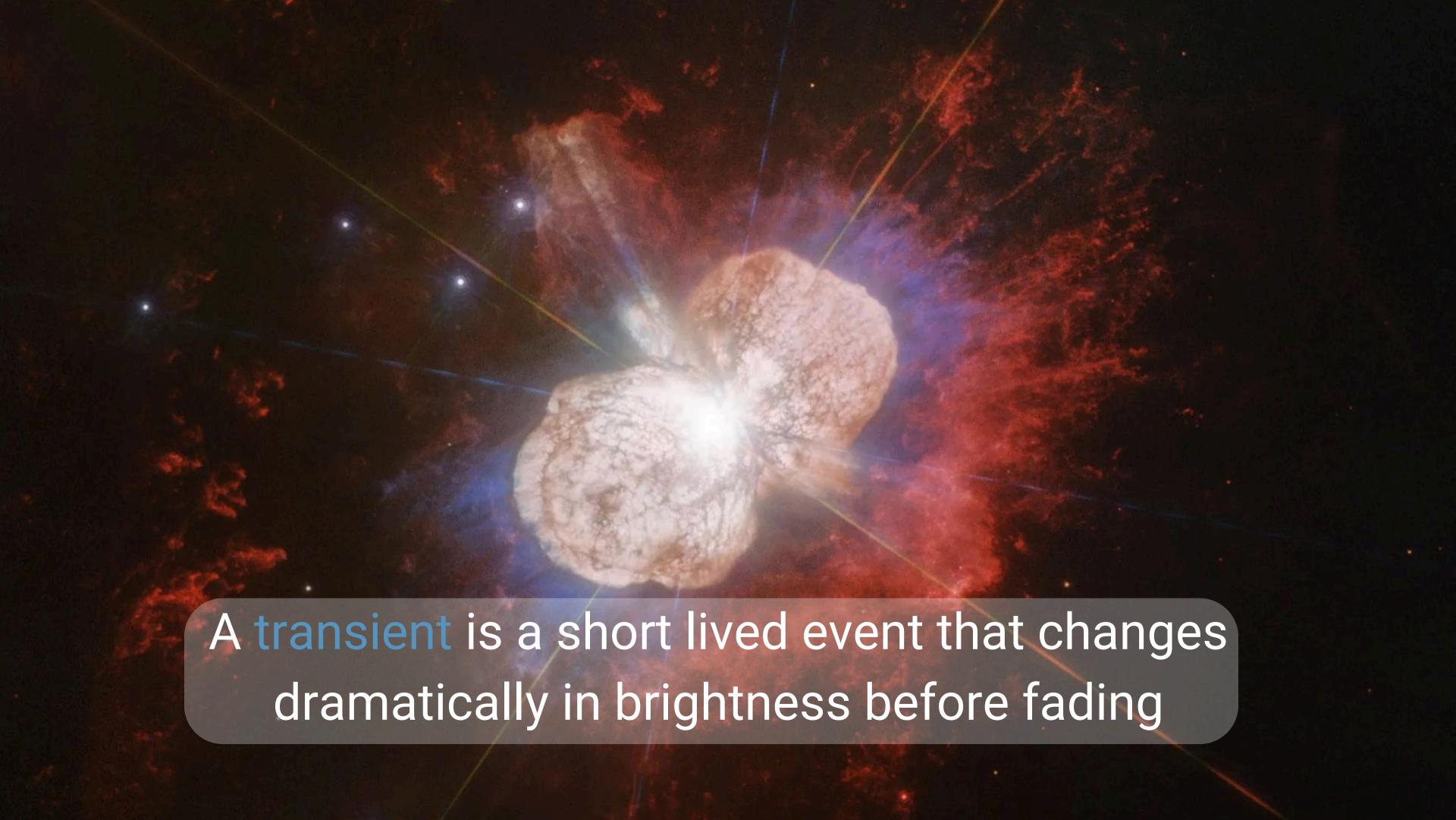
Seán Brennan

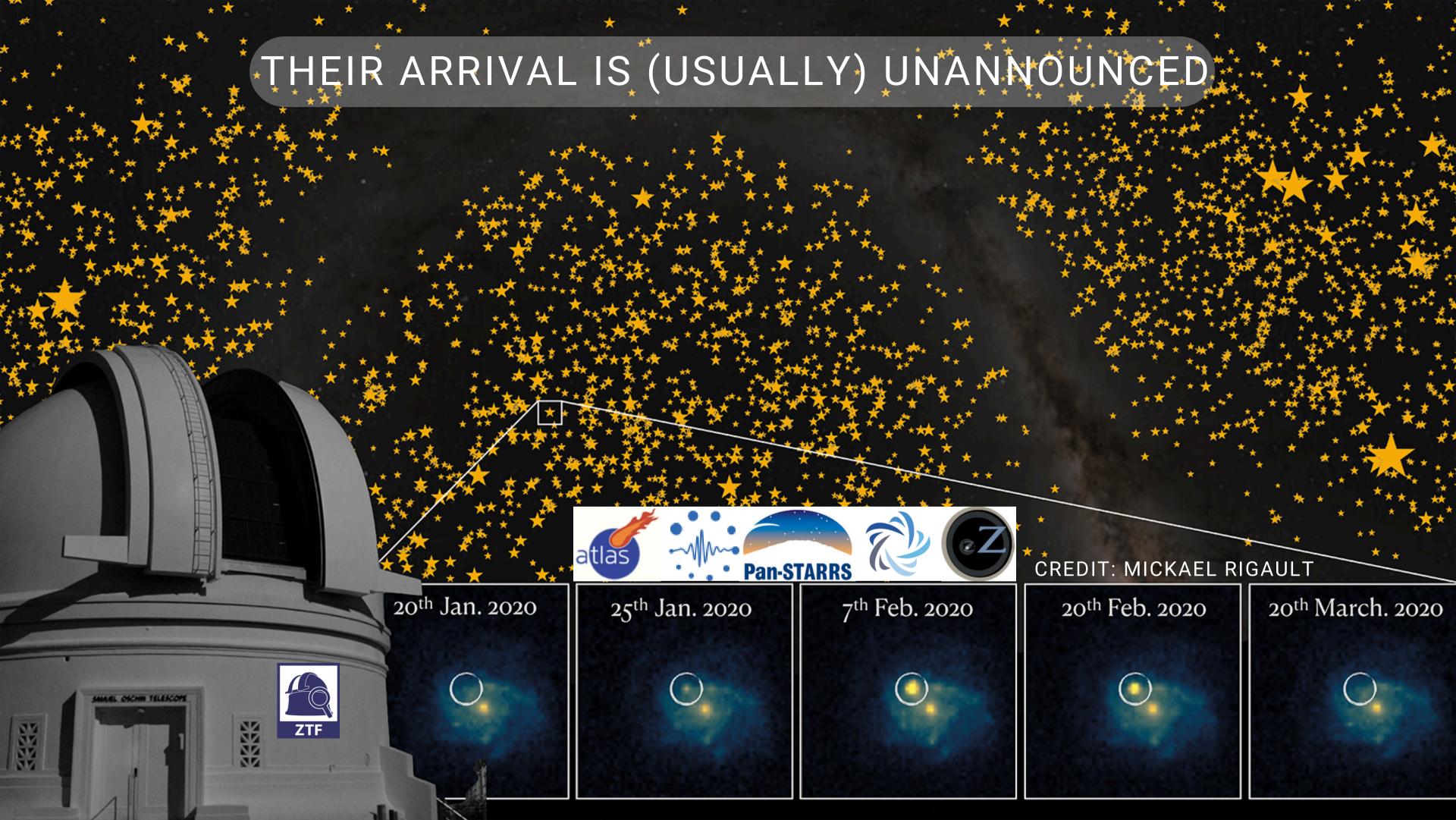
An Extraordinary Journey Into The Transient Sky Padua | Italy









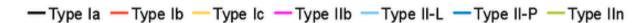


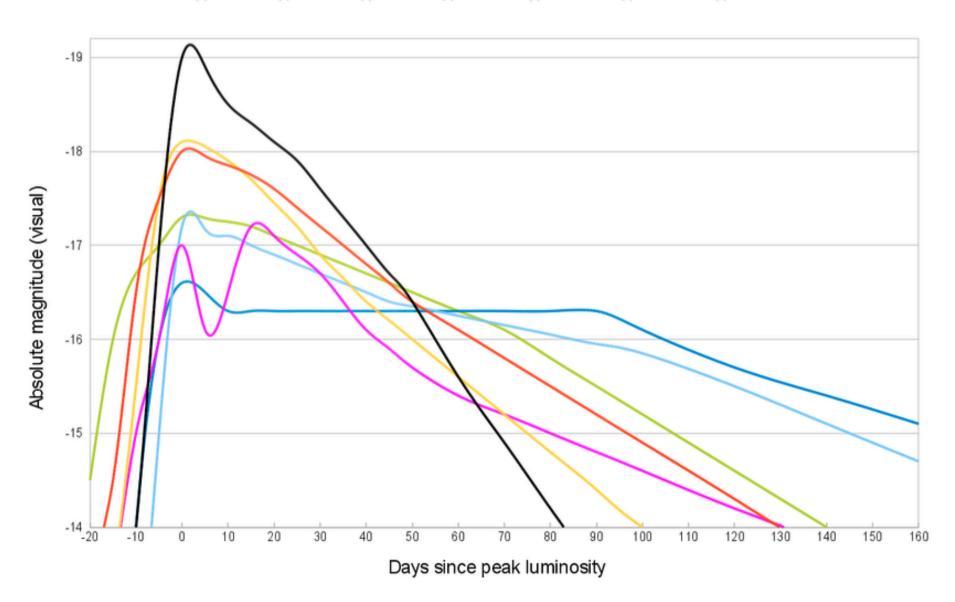


PROGENITORS OF CORE-COLLAPSE SUPERNOVAE

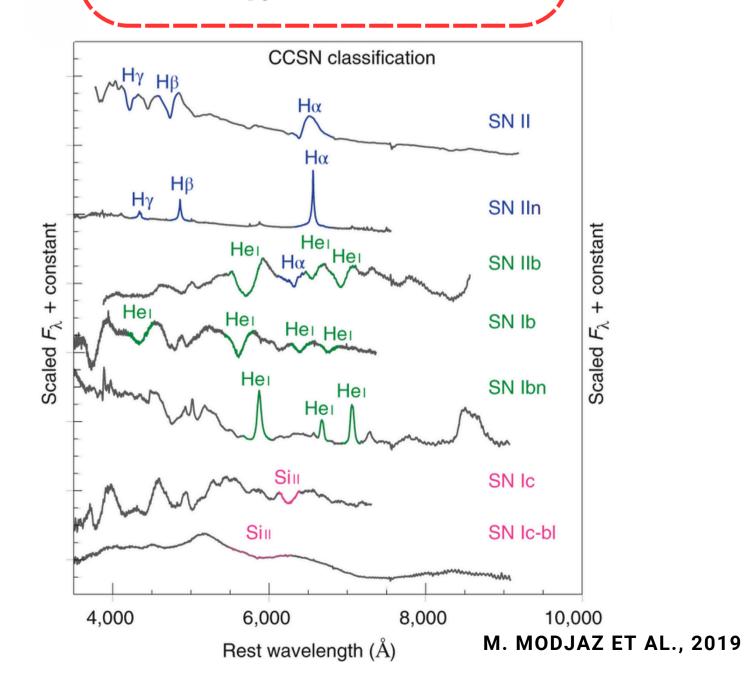
GAINING INSIGHT A POSTERIORI

How the explosion's *brightness* changes over time



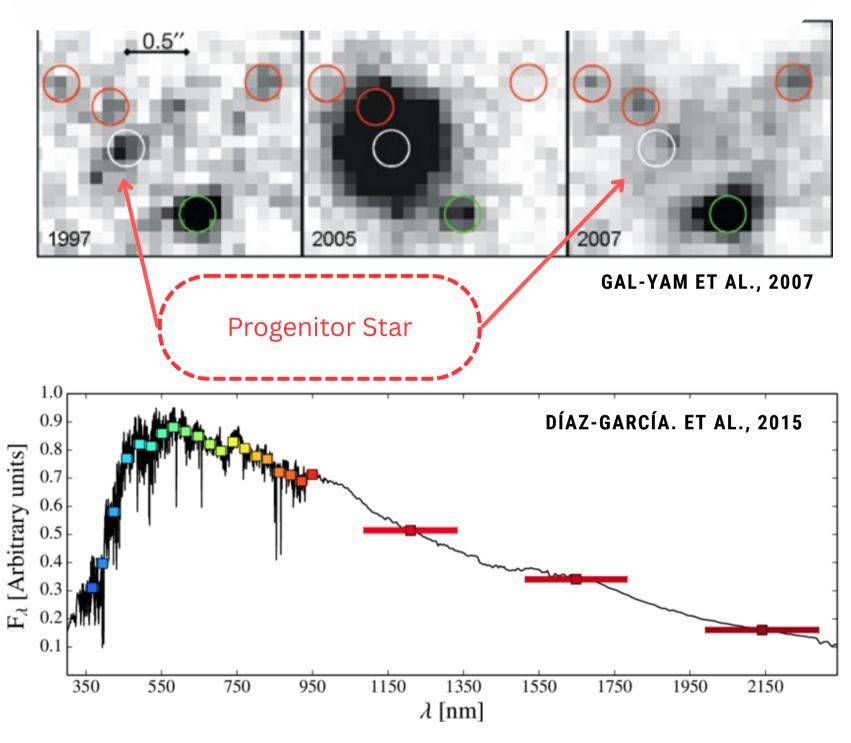


At what wavelengths the energy comes out

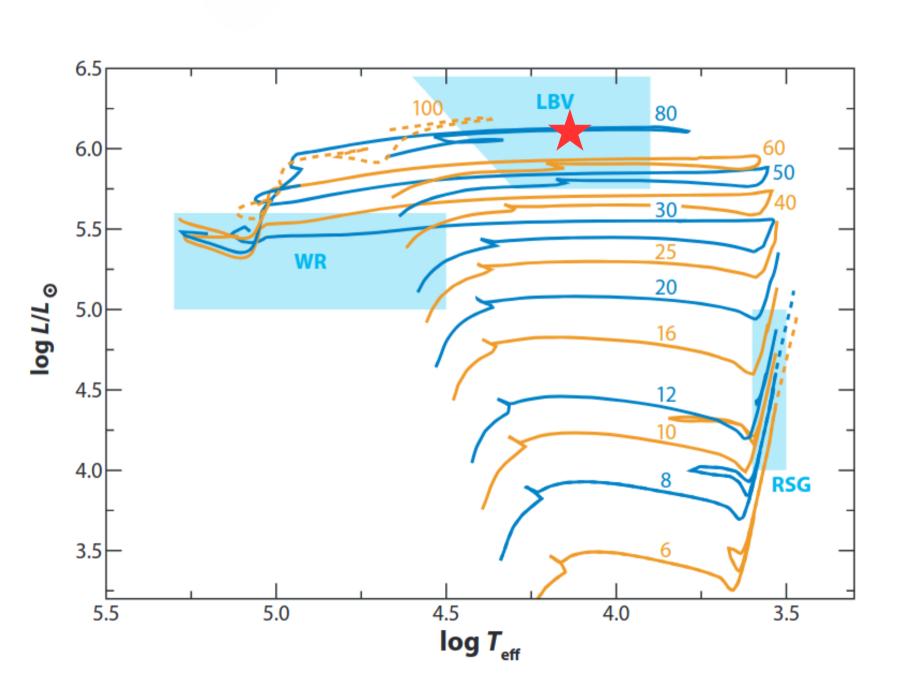


PROGENITORS OF CORE-COLLAPSE SUPERNOVAE

PRE- CORE-COLLAPSE DETECTIONS



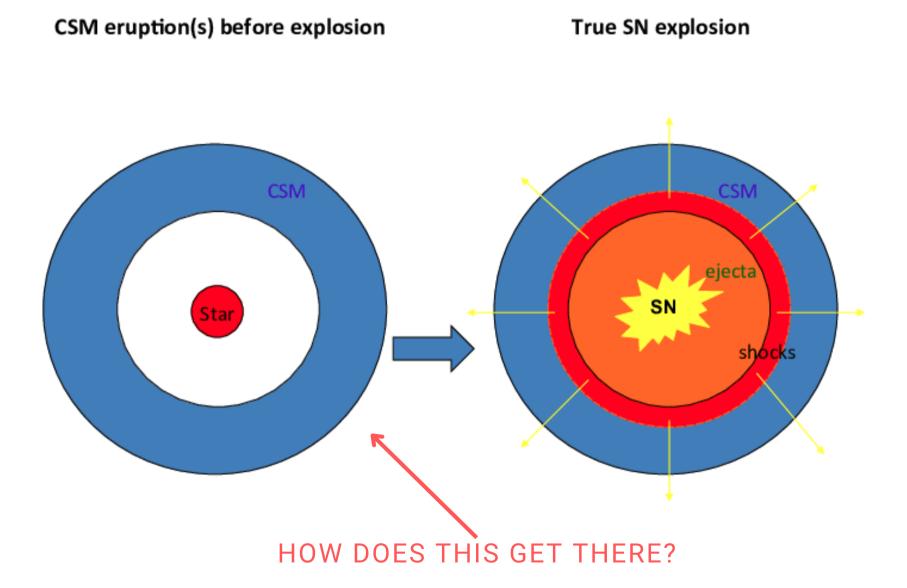
 $MODEL(OBSERVATIONS) \rightarrow L_{BOL}, T_{EFF}$



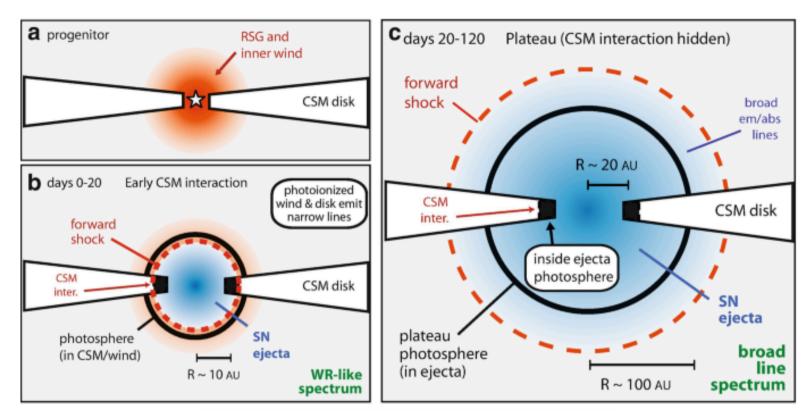
SMARTT. ET AL.,2009, ELDRIDGE & TOUT 2004

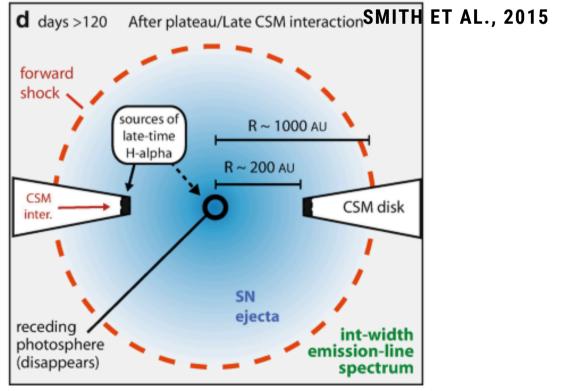
INTERACTING SUPERNOVAE

A) A SIMPLIFIED PICTURE



B) MORE-REALISTIC PICTURE (ALTHOUGH NOT WELL EXPLORED)





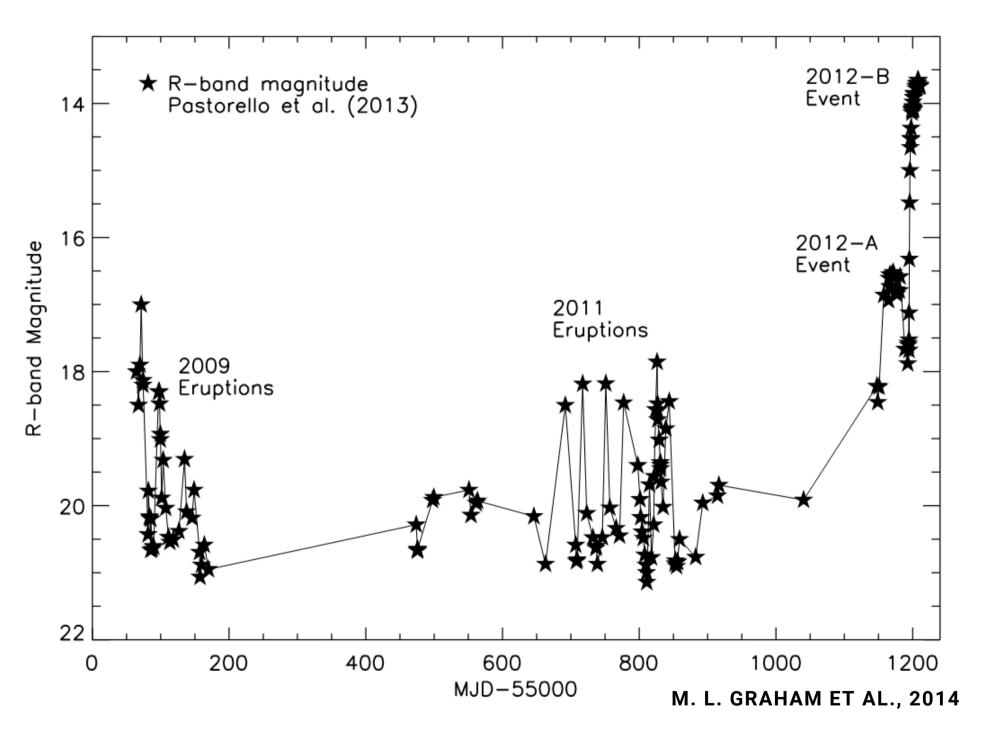
OVER 2/3 OF TYPE II PROGENITORS NEED $M_{\rm CSM} \ge 10^{-2.5} M_{\rm SUN}$ WITHIN $R_{\rm CSM} \sim 10^{15}$ CM

SN 2009IP-LIKE TRANSIENTS

YEARS LONG ``PRECURSOR ACTIVITY''
FOLLOWED BY TWO EXPLOSIONS —
THE LATTER BEING A GENUINE SN*

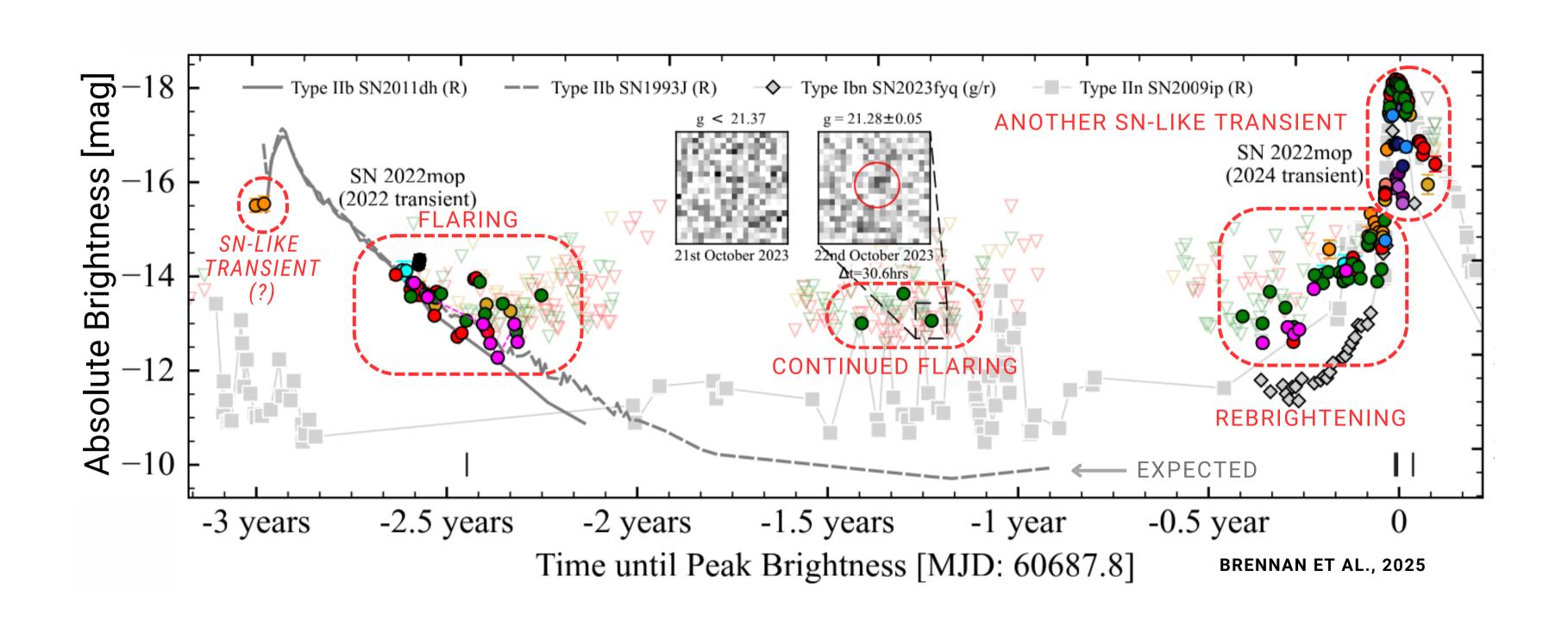


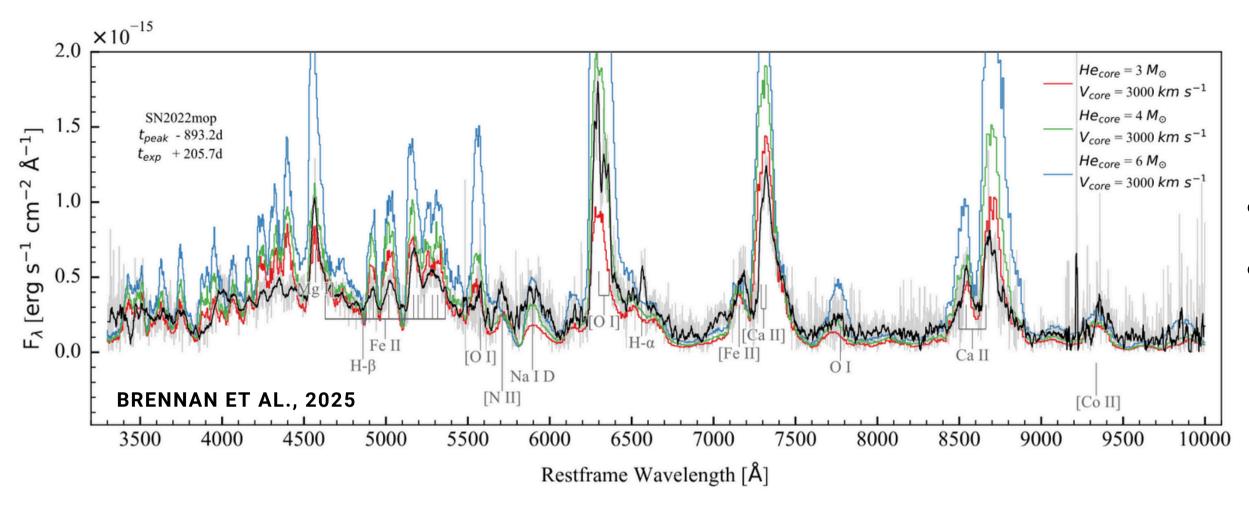




Papers on SN 2009ip: Smith et al., 2010; Foley et al., 2011; Fraser et al., 2013; Mauerhan et al., 2013; Ofek et al., 2013; Ouyed et al., 2013; Pastorello et al., 2013; Potashov et al., 2013; Prieto et al., 2013; Smith et al., 2013; Soker et al., 2013; Tsebrenko et al., 2013; Graham et al., 2014; Levesque et al., 2014; Margutti et al., 2014; Mauerhan et al., 2014; Smith et al., 2014; Fraser et al., 2015; Martin et al., 2015; Moriya et al., 2015; Smith et al., 2016; Graham et al., 2017; Pastorello et al., 2018; Chugai et al., 2022; Fransson et al., 2022; Pessi et al., 2022; Smith et al., 2022; Smith et al., 2022; Pessi et al., 2023.

PRECURSOR ACTIVITY PRECEDING INTERACTING SUPERNOVAE I: BRIDGING THE GAP WITH SN 2022MOP



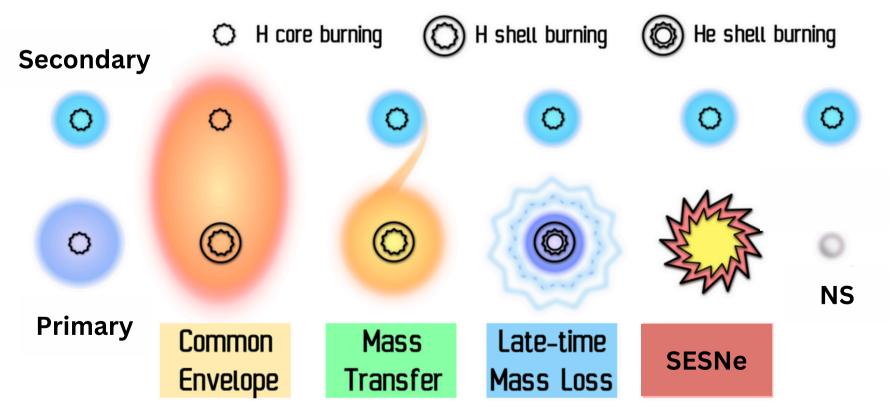


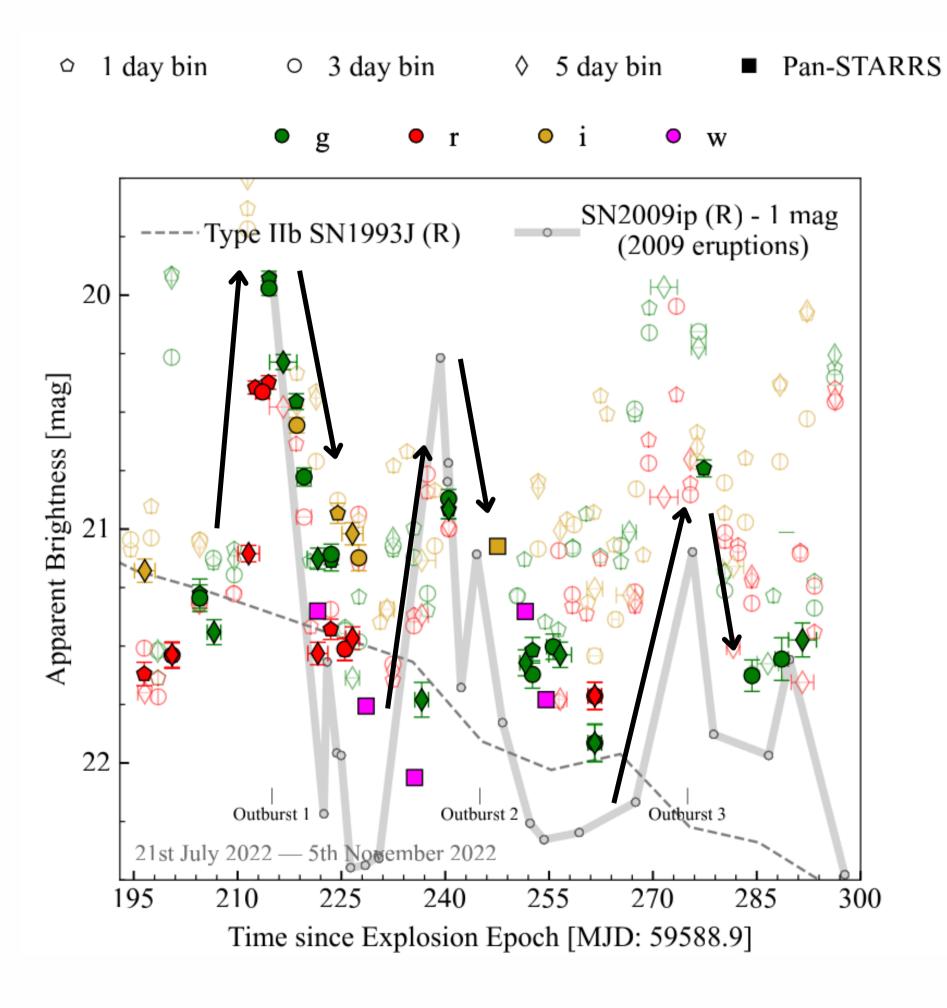
NEBULAR SPECTRUM

- [O I], MG I], AND [CA II]
- LATE TIME SPECTRUM OF A SUPERNOVA

(POSSIBLE) FORMATION PATHWAY

• NEUTRON STAR CONTINUES TO INTERACTION WITH COMPANION (?)



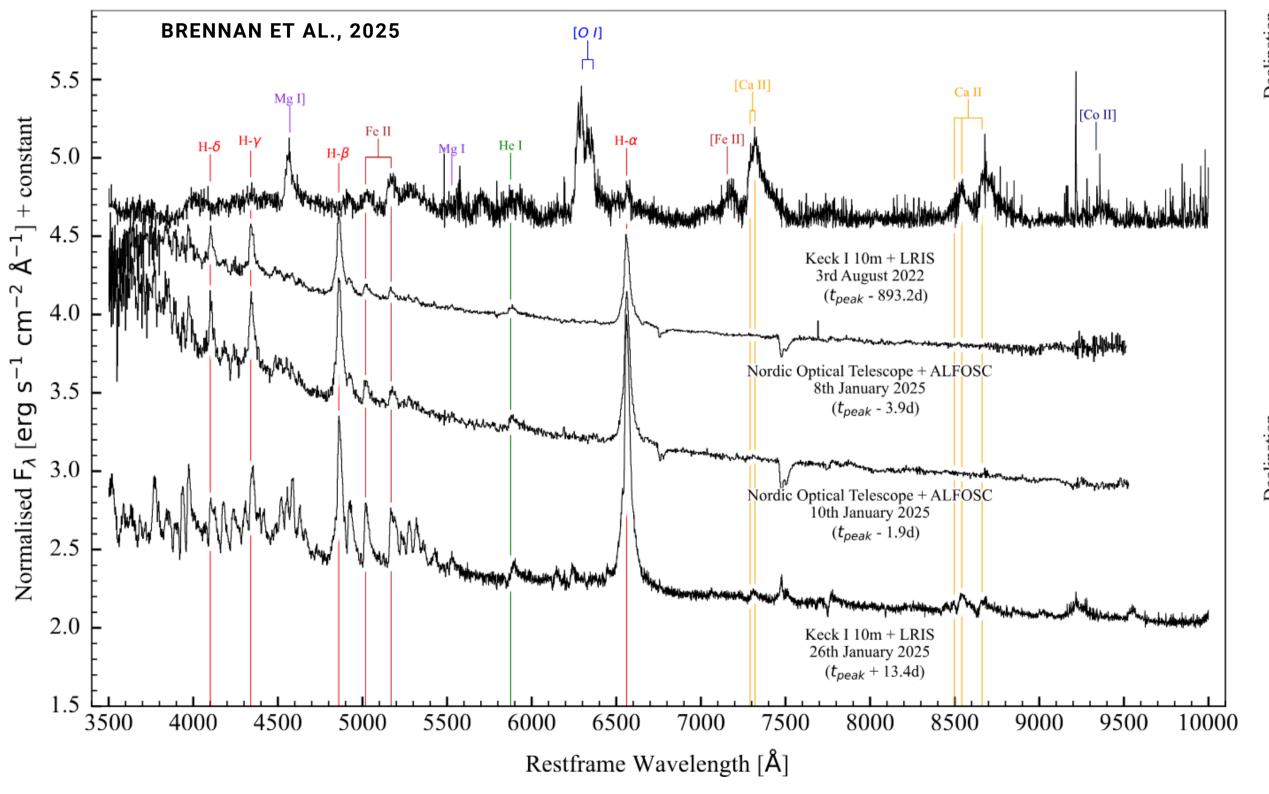


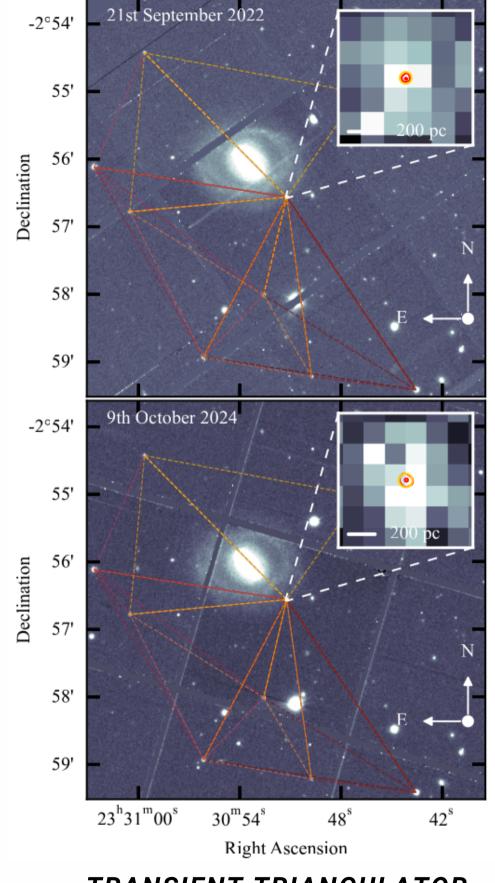
- ~200 DAYS POST EXPLOSION
- FLARING SEEN IN LATE TIME LIGHTCURVE - NOT RADIOACTIVE DECAY
- SIMILAR TO 2009 ERUPTIONS OF SN 2009IP
- ~30 DAY PERIOD



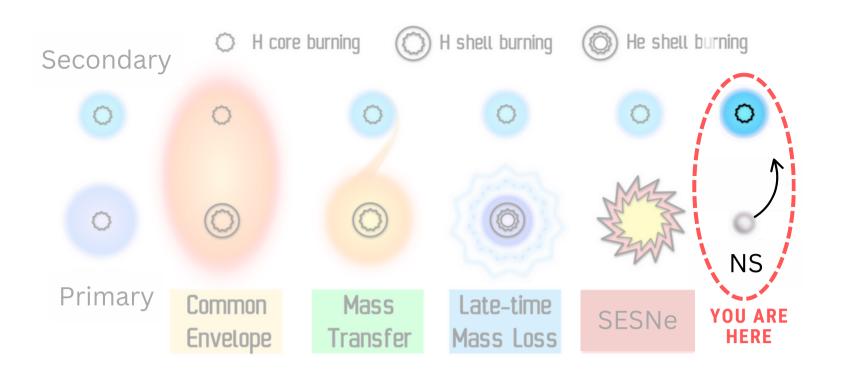
Image credit: Instituto de Astrofísica de Canarias.

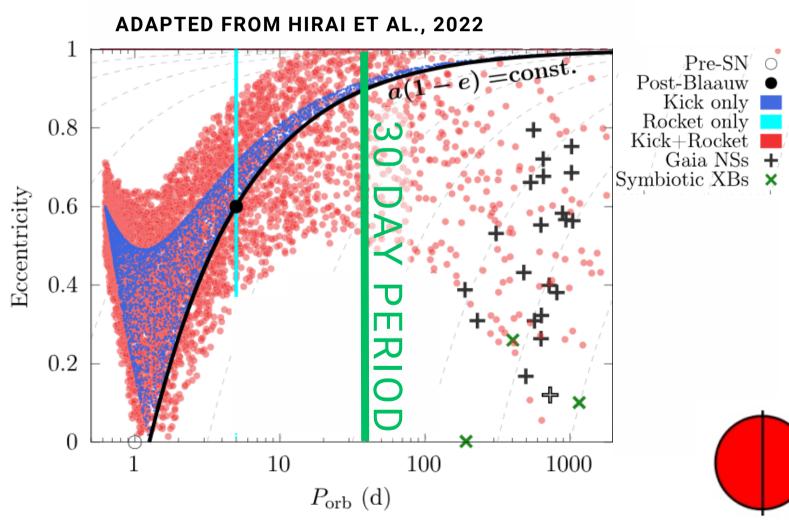
A HYDROGEN POOR SESN BECOMES A HYDROGEN RICH SN 3 YEARS LATER



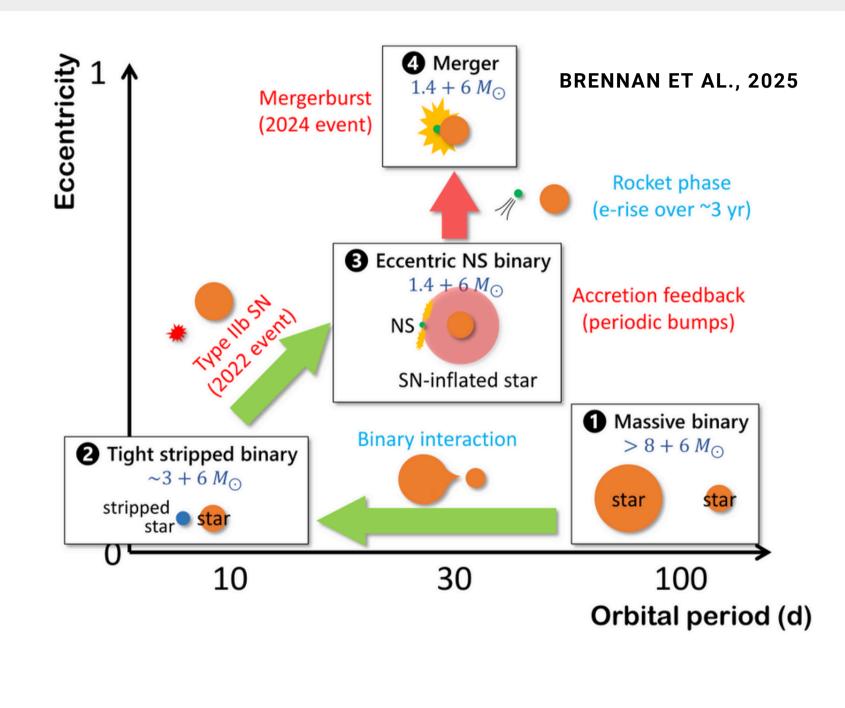


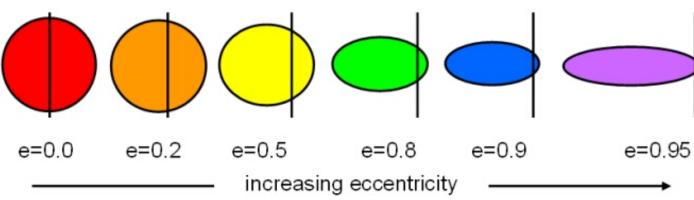
TRANSIENT TRIANGULATOR - (SOON) AVAILABLE ON GITHUB





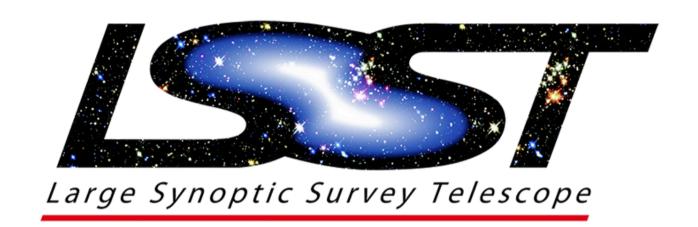
NEWLY FORMED NEUTRON STAR INTERACTS WITH ITS COMPANION: MERGERBURST





PRESTO - PRE SUPERNOVA TARGETED OBSERVATIONS

GOAL: IDENTIFY PRE SN ACTIVITY AND FOLLOWUP



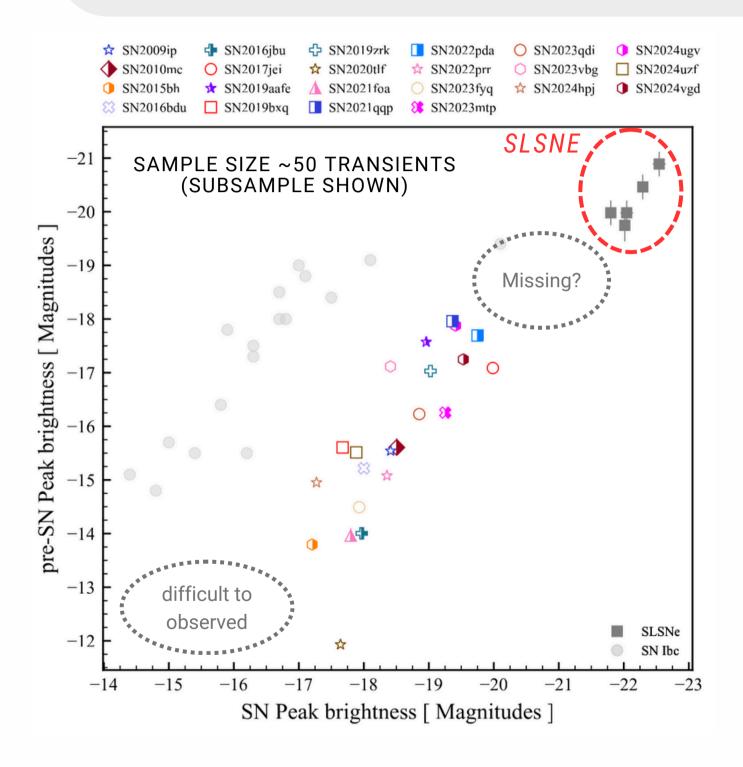


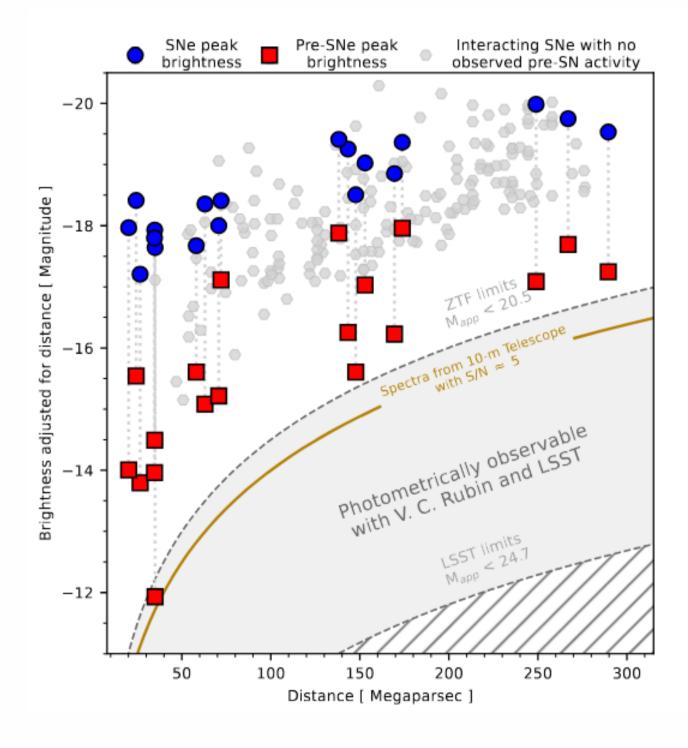
• IDENTIFY THIS ACTIVITY BEFORE CORE-COLLAPSE OCCURS

- HIGH CADENCE OBSERVATIONS
 - PERIODICITY AND IN-SPIRAL (~CHIRP SIGNAL)
- EXPLORE PANCHROMATICALLY
 - IR (DUST AND NEBULAR FEATURES)
 - UV XRAY (INTERACTION)

PRESTO - PRE SUPERNOVA TARGETED OBSERVATIONS

PRELIMINARY WORK



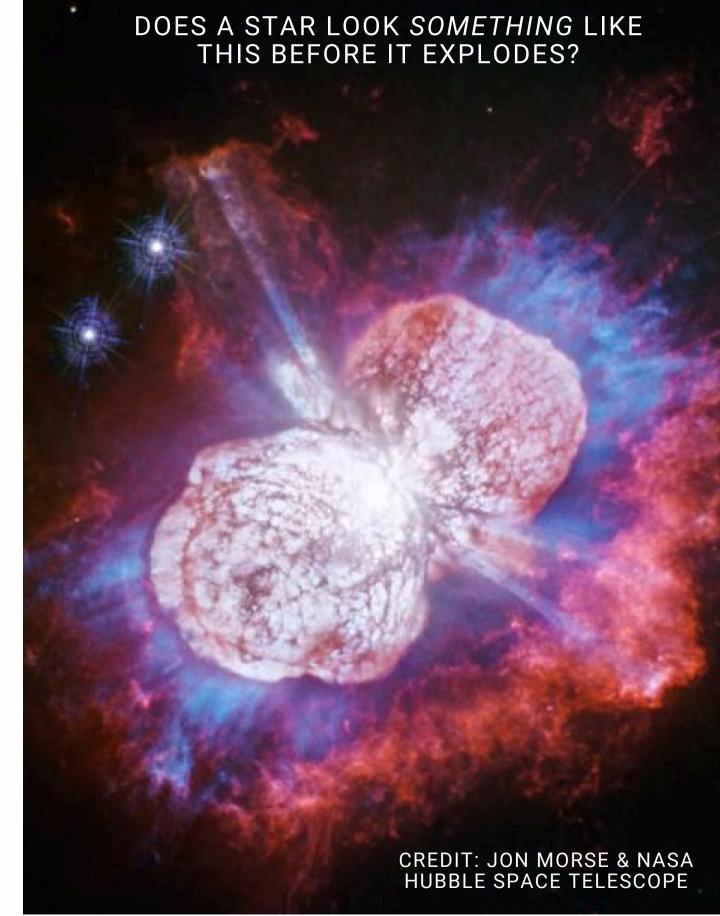


BRIGHTER ACTIVITY YIELDS BRIGHTER SUPERNOVAE

NEW MAGNITUDE SPACE TO EXPLORE WITH RUBIN

TAKE HOME POINTS

- IT IS POSSIBLE TO IDENTIFY A STAR BEFORE IT EXPLODES
- NO HIGH CADENCE FOLLOW UP (YET!)
- SELECTION CRITERIA (IN PROGRESS) FOR RUBIN ERA









Seán Brennan sean.brennan@astro.su.se

OPEN QUESTIONS

- MERGER BURST EXPLOSION PATHWAY (?)
- CAN A MERGER TRIGGER CORE-COLLAPSE?
 - ∘ IS THIS 'CORE-COLLAPSE'
 - SN V.S. SN IMPOSTOR
- PROGENITOR-SUPERNOVA CONNECTION
 - EFFECTS OF CSM
 - ASSUMPTION OF HYDROSTATIC-EQUILIBRIUM







Seán Brennan sean.brennan@astro.su.se

