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PIANO NAZIONALE
DI RIPRESA E RESILIENZA



Centro Nazionale di Ricerca in HPC,
Big Data and Quantum Computing

Mercury-Arxes and OPAL: an update

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Spoke 3 II Te

! WP1-WP2 monthly meeting !

Sp, Bologna Dec 17 -19, 2024

Parallelizing Mercury-Arxes: scientific rationale

- Observations allow to retrieve the chemical composition of exoplanet atmospheres
- For gas giants, atmospheric composition is a proxy for the composition of the whole body
- From the composition it is possible to reconstruct the formation and evolution of planets
- Investigating these processes in a cosmic context is fundamental to understand these processes

Pipeline of codes to self-consistently model how a planet looks like from its history:

- Chemical composition of the protoplanetary disk
- Planetary formation & migration
- Atmospheric chemistry
- Synthetic spectra

Parameter space exploration is required!

Parallelizing Mercury-Arxes: starting point and objectives

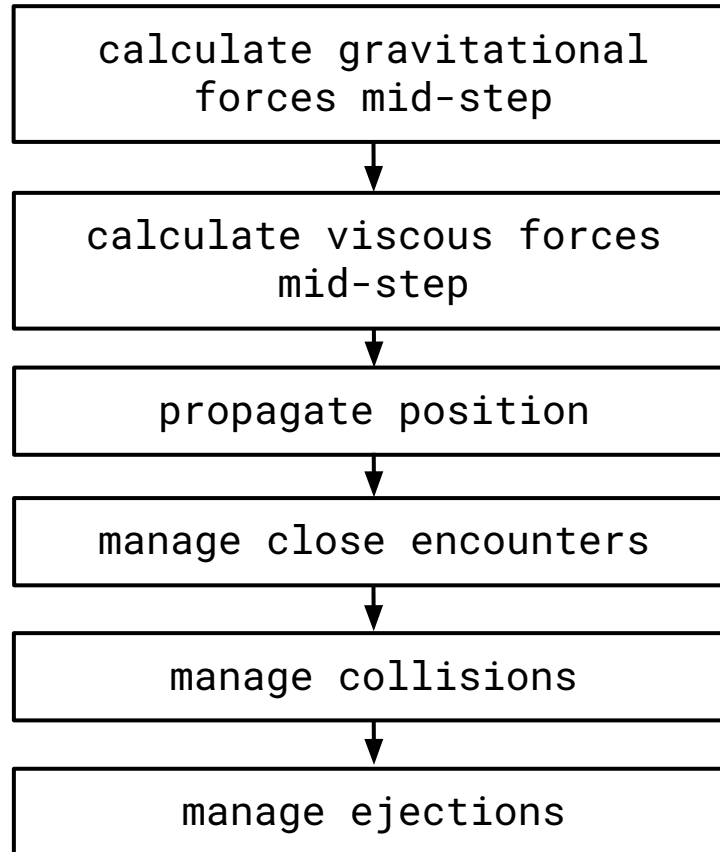
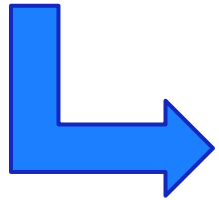
- Mercury (Chambers+99): multipurpose n-body integrator
- Mercury-Arxes (Turrini+21): branched out version tailored to study planetary formation:
 - based on Mercury 6
 - viscous interaction with a time-evolving disk
 - management of merging trees
 - hybrid symplectic integrator to conserve energy and momentum over ~Myr timescales
- Current typical simulation: <10 gravity-source planets, >1000 non gravity-source particles

GOAL: have all the particles as gravity-sources

- especially important at the beginning of the simulation, when most of the mass is in particles

Parallelizing Mercury-Arxes: methodology

Mercury-Arxes main loop:

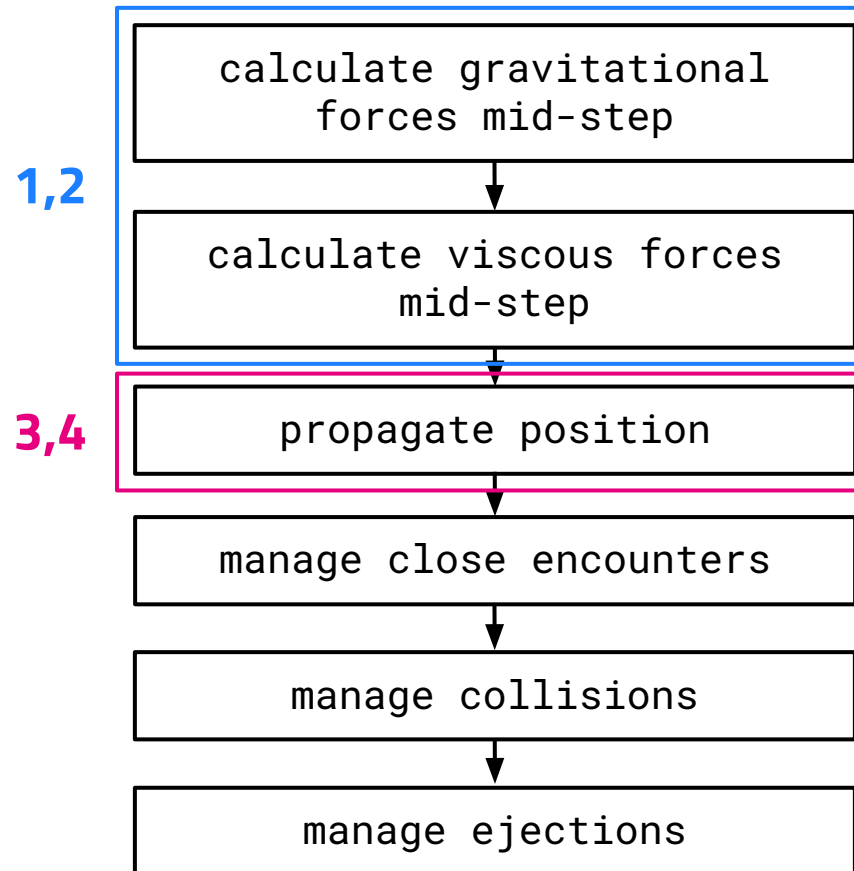


To achieve our goal, high level of parallelism is needed. In practice we:

- **parallelize the main calculations on GPU using OpenACC**
- **parallelize the rest of the code on CPU using OpenMP**

Why? Minimal need for code refactoring with respect to the potential gain

Parallelizing Mercury-Arxes: current status

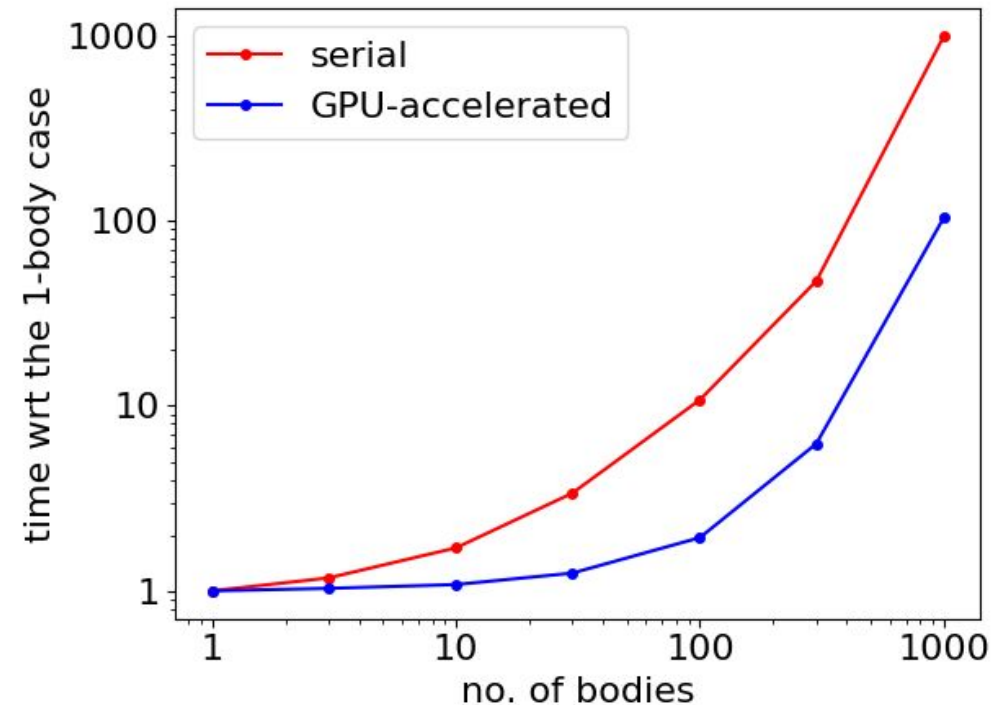


1. merged gravitational and non-gravitational forces in one routine
2. ported force calculation on GPU
3. translated the position propagation routines from C to Fortran
4. ported position propagation on GPU

(3) needed because of issues in the interaction between OpenACC and the external C routines - solution suggested by G. Lacopo (Univ. of Trieste, INAF-OATs)

Parallelizing Mercury-Arxes: main results

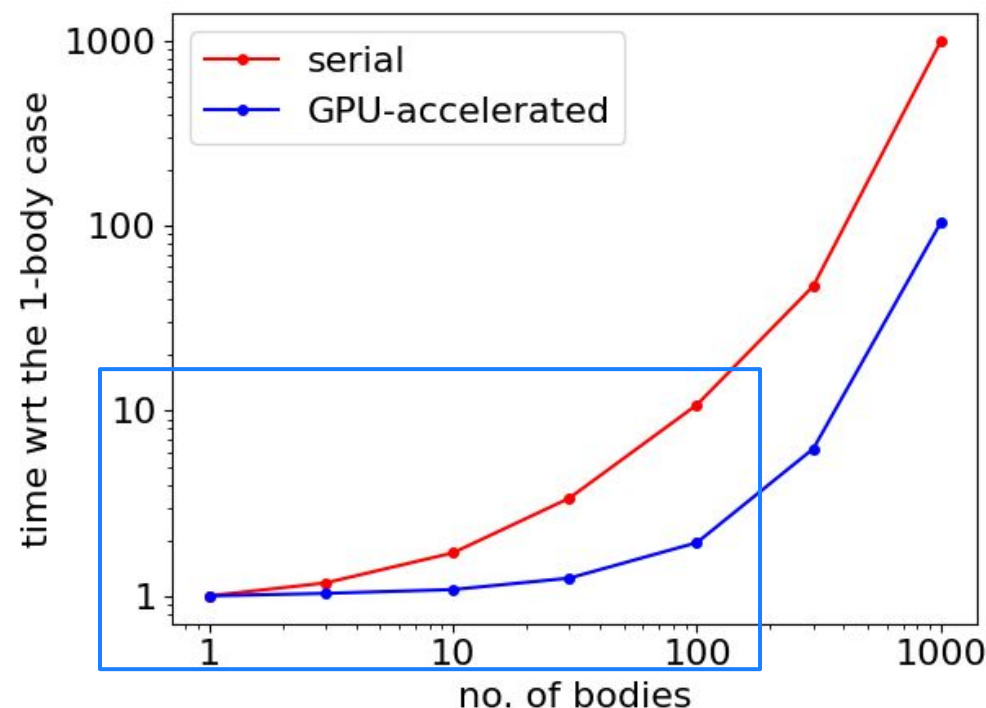
- marginal speedup from (1) and (3): few % runtime decrease
- large speedup from (2) and (4) despite:
 - inefficient data management
 - inefficient execution of the CPU part due to nvfortran
- substantially improved serviceability of the code



Parallelizing Mercury-Arxes: main results

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Fewer bodies → fewer (non-parallel) close encounters → larger advantage of the GPU-accelerated version

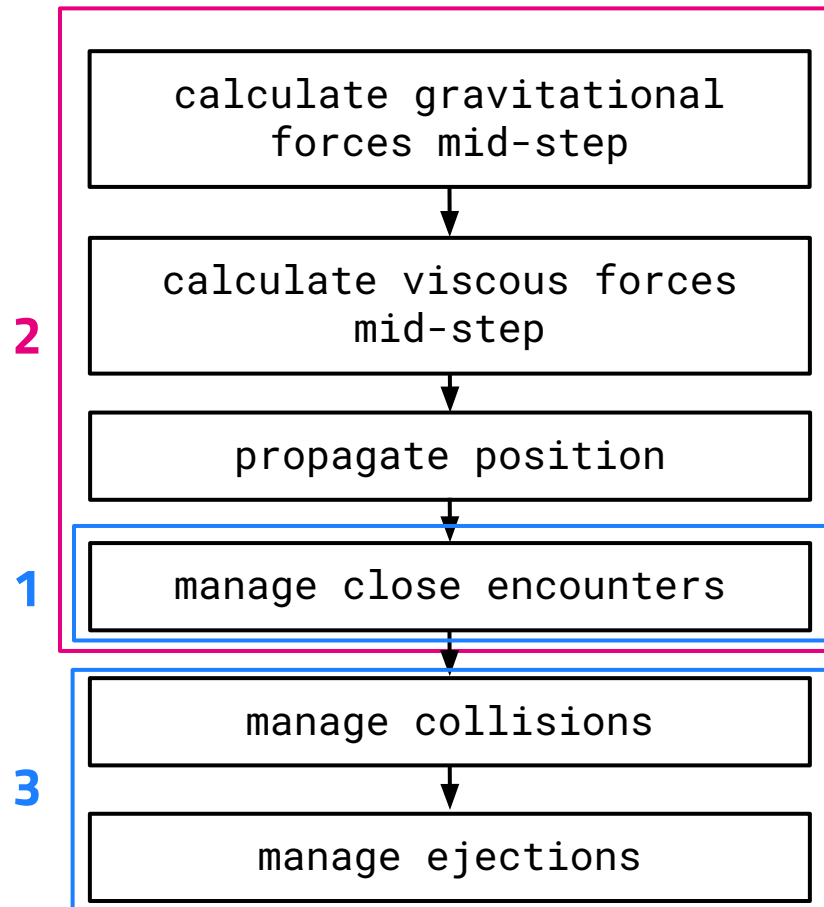


Parallelizing Mercury-Arxes: KPI & milestones

- Small delay with respect to the M9 target related to Mercury-Arxes of OPAL, mostly because of the unforeseen need for code refactoring
- Assessed progression on this target: 75%

ID	TARGET	KPI	MILESTONE
4	Assessment of the comparative performances of the OpenMP and OpenACC releases of Mercury-Arxes (Task A)	Report on the HPC assessment on Mercury-Arxes	M9

Mercury-Arxes: final steps



1. **try to parallelize close encounter management - trickier than the rest of the loop!**
2. **define a data region in order to cut down transfer times btw. host and device**
3. **parallelize collision and ejection management via OpenMP (minor)**

ENDPOINT: comparative performance assessment and deployment of the fully parallelized version of the code

What is OPAL:



A **PNRR Key Science Project** with two goals:

- produce a library of detailed synthetic atmospheric models of giant planets for the Ariel space mission (Tinetti+2018, 2020; Turrini+2018, 2022)
- port the Arxys planet formation suite of codes to HPC

Takes advantage of allocated time in:

Leonardo (4 million core hours)

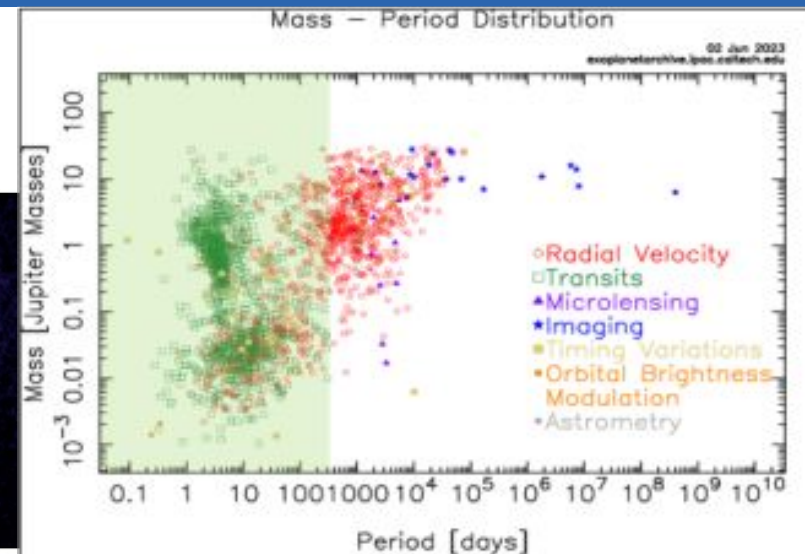
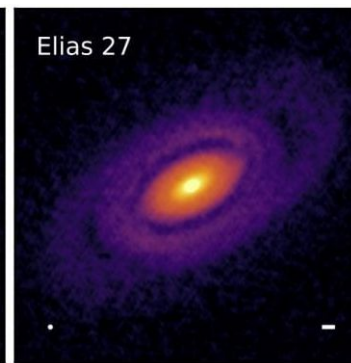
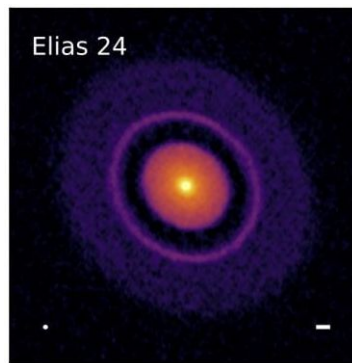
Pleiadi (2.5 million core hours)

as well as the dedicated cluster **GENESIS+**.

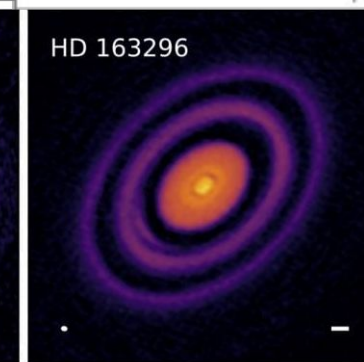
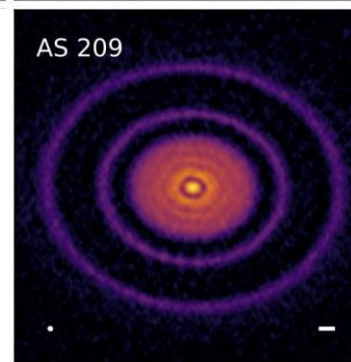
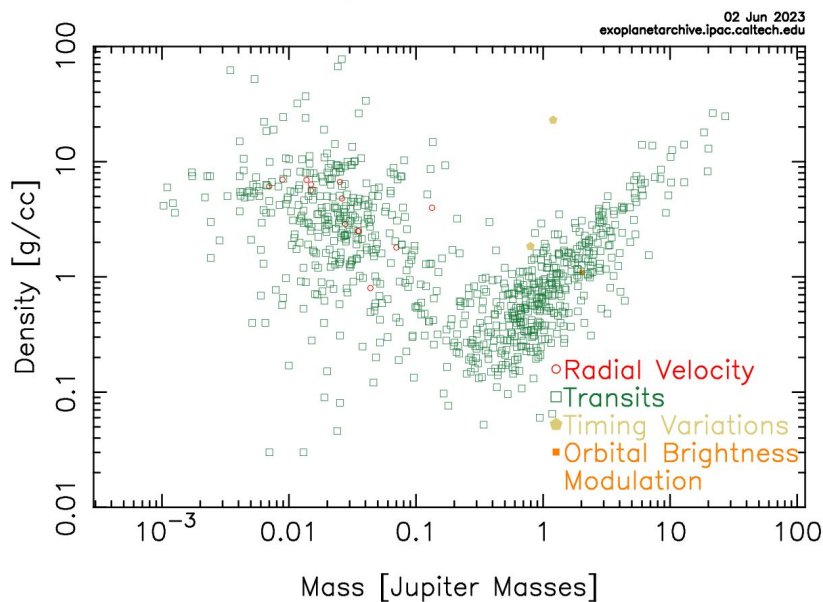
OPAL is a coordinated effort between the ARIEL, INAF and the PNRR

Why we need it:

- New observational data highlight our limited understanding of where planets form in disks



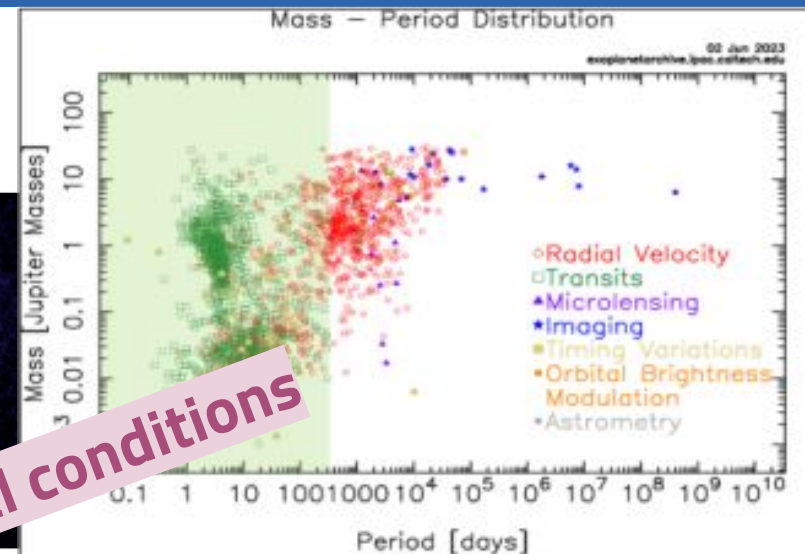
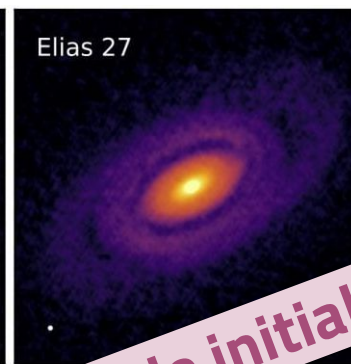
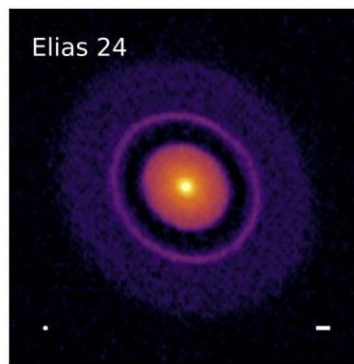
Density - Mass Distribution



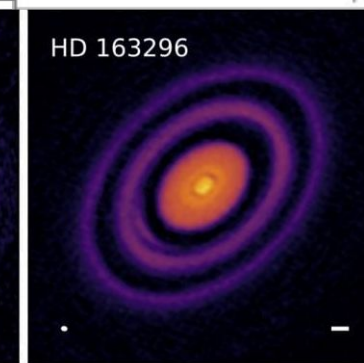
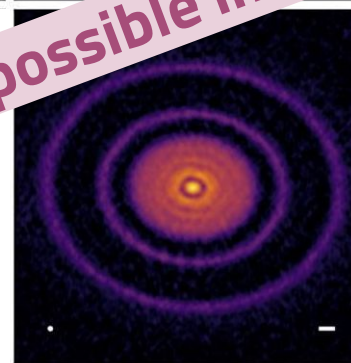
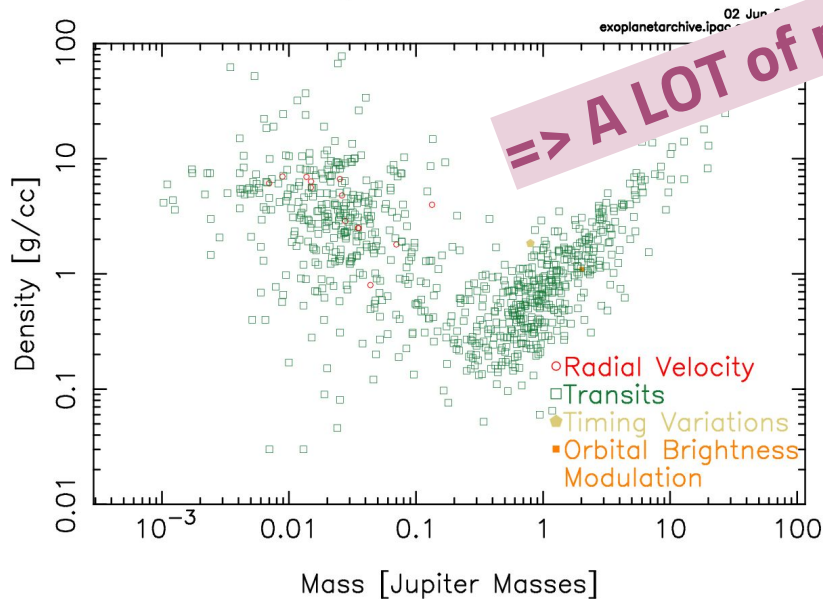
- Planets form and evolve under a wider range of conditions than previously thought.

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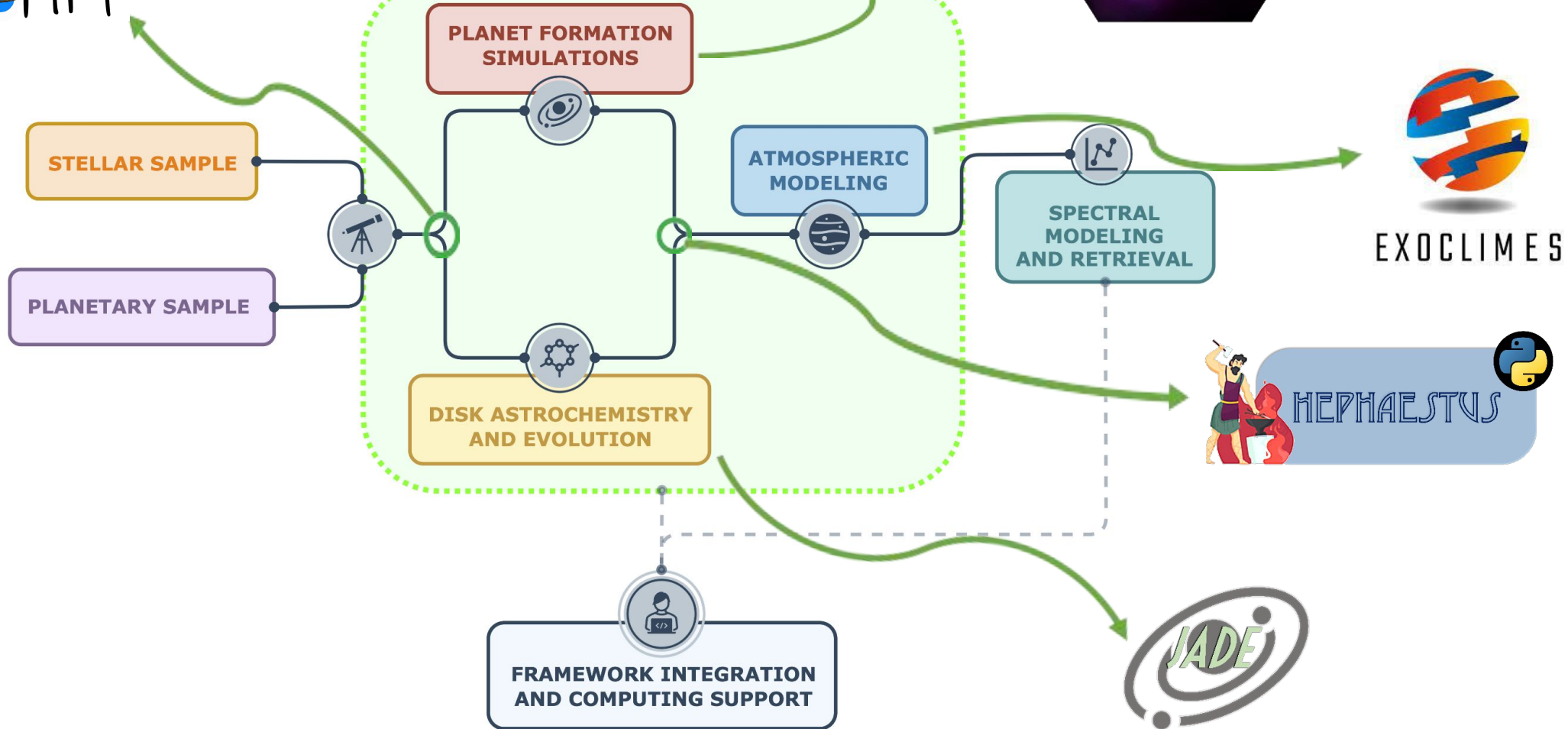
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- Planets form and evolve under a wider range of conditions than previously thought.

OPAL: Modelling Complexity

GOMIT



OPAL: Modelling Complexity



JADE (Pacetti et al. in prep.) is the multi-language (Python + Fortran) data-parallel code to jointly simulate the **physical and chemical evolution of protoplanetary disks** of the *Arxes* suite.

Mercury-Arxes (Turrini+2019,2021) is the parallel n-body code of the *Arxes* suite incorporating physical libraries to simulate **planet formation in protoplanetary disks**



HEPHAESTUS (Turrini+2021; Pacetti+2022) is the Python compositional post-processing tool of the *Arxes* suite that resolves planetary compositions across 20+ elements.

FastChem and **Vulcan** are atmospheric modelling codes from the University of Bern's ExoClimes suite convert Hephaestus' elemental compositions into atmospheric molecular compositions (e.g. Fonte+2023, Simonetti et al., under review).



EXOCLIMES

OPAL: Modelling Complexity

An extra word on Vulcan:

We are expanding the chemical network of Vulcan to also include the sedimentation and transport of the condensation of iron!

(and soon also the silicates!)

to jointly
uite.

MERCURY
ArXes

processing
ments.

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EXOCLIMES

GroMiT: The new kid in town!

GroMiT, the Growth and Migration Track code, tracks how a forming planet accretes pebbles and gas as well as how it migrates depending on the local environment of its native circumstellar disk (e.g. Johansen+2019, Ida+2016, Tanaka+2020).

Born as a plug-in for Mercury-ArXes to make it self-consistent it found plenty of applications also as a stand-alone code in exploring the formation histories of **observed planets (GAPS)**.

→ Single Planet Mode
(Polychroni+2023)

It comes in two flavours!

→ Population Synthesis Mode
(Polychroni+ in prep)

GroMiT

