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PIANO NAZIONALE  
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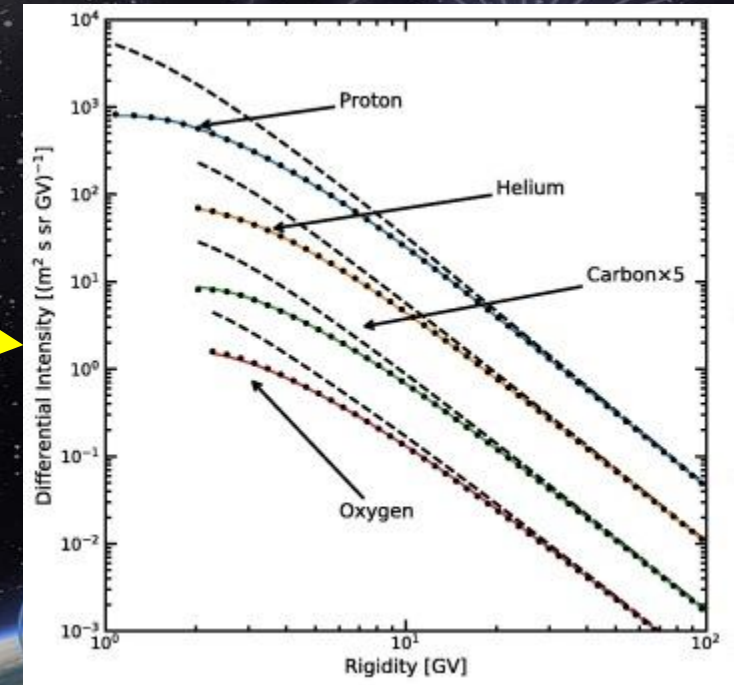
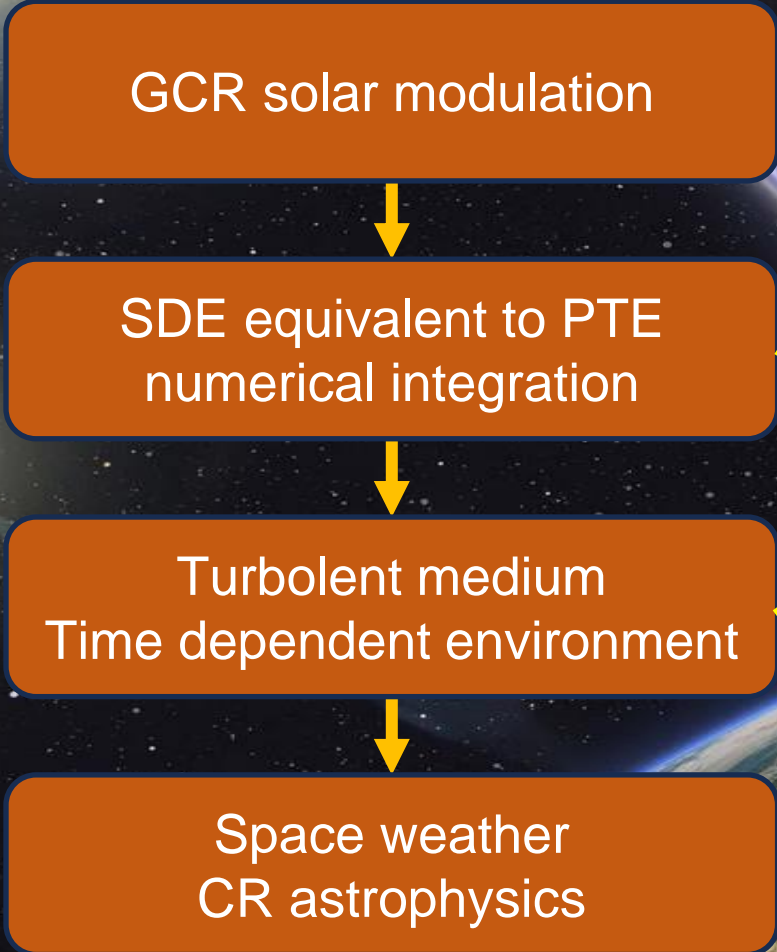


Centro Nazionale di Ricerca in HPC,  
Big Data and Quantum Computing

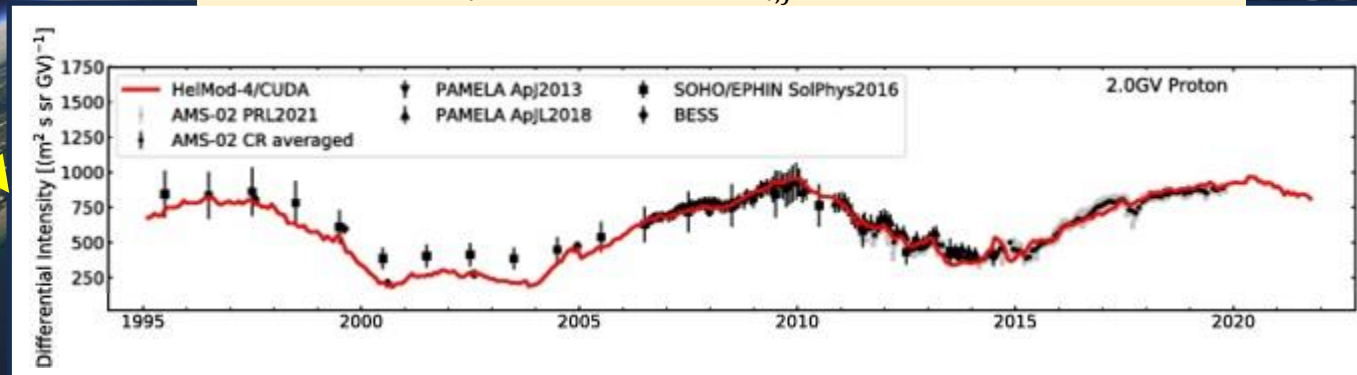
*COSMICA: a novel parallel GPU code for  
Cosmic Rays propagation in heliosphere*  
*Giovanni Cavallotto (INFN MiB), Stefano Della Torre (INFN MiB)*

Spoke 3 II Technical Workshop, Bologna Dec 17 -19, 2024

# Scientific Rationale

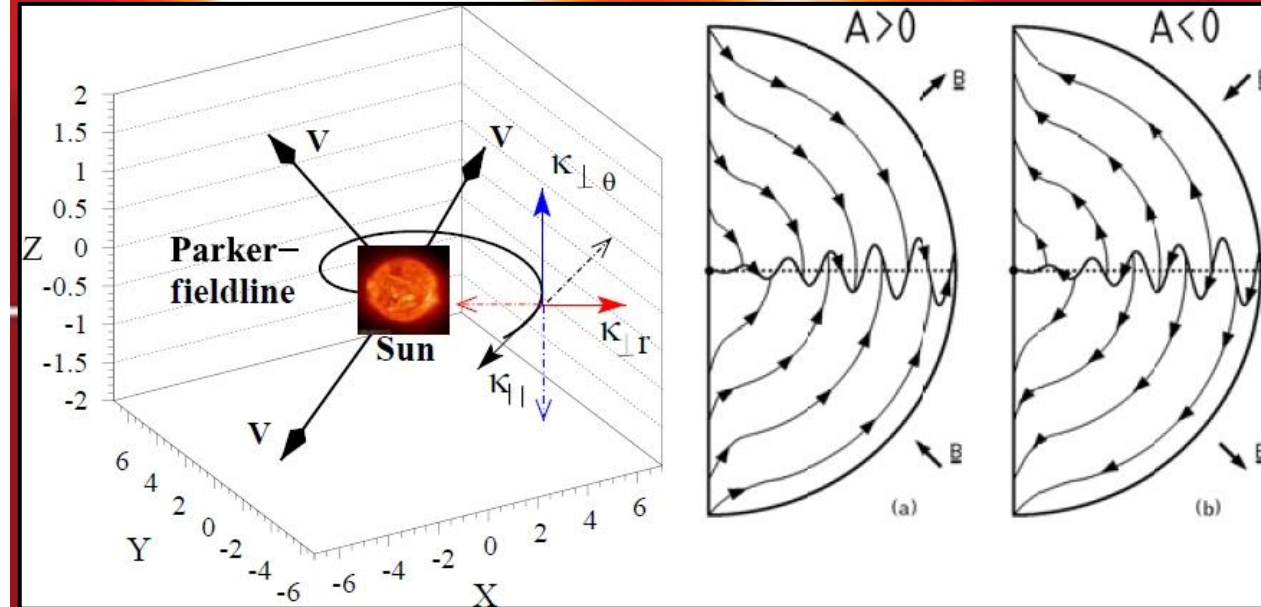
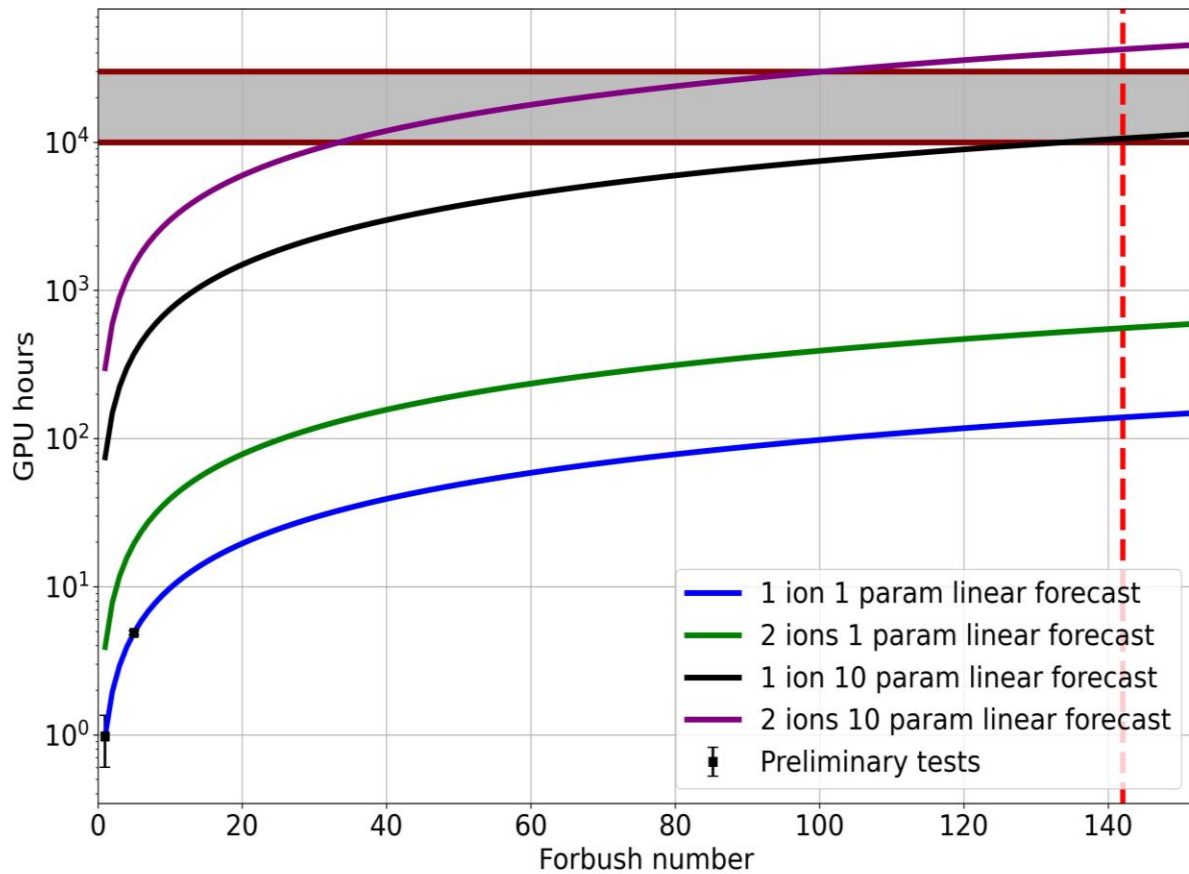


$$SDE: \frac{\partial U}{\partial s} = \sum_i A_i(s, y) \frac{\partial U}{\partial y_i} + \frac{1}{2} \sum_{i,j} C_{ij}(s, y) \frac{\partial^2 U}{\partial y_i \partial y_j} - L U + S$$

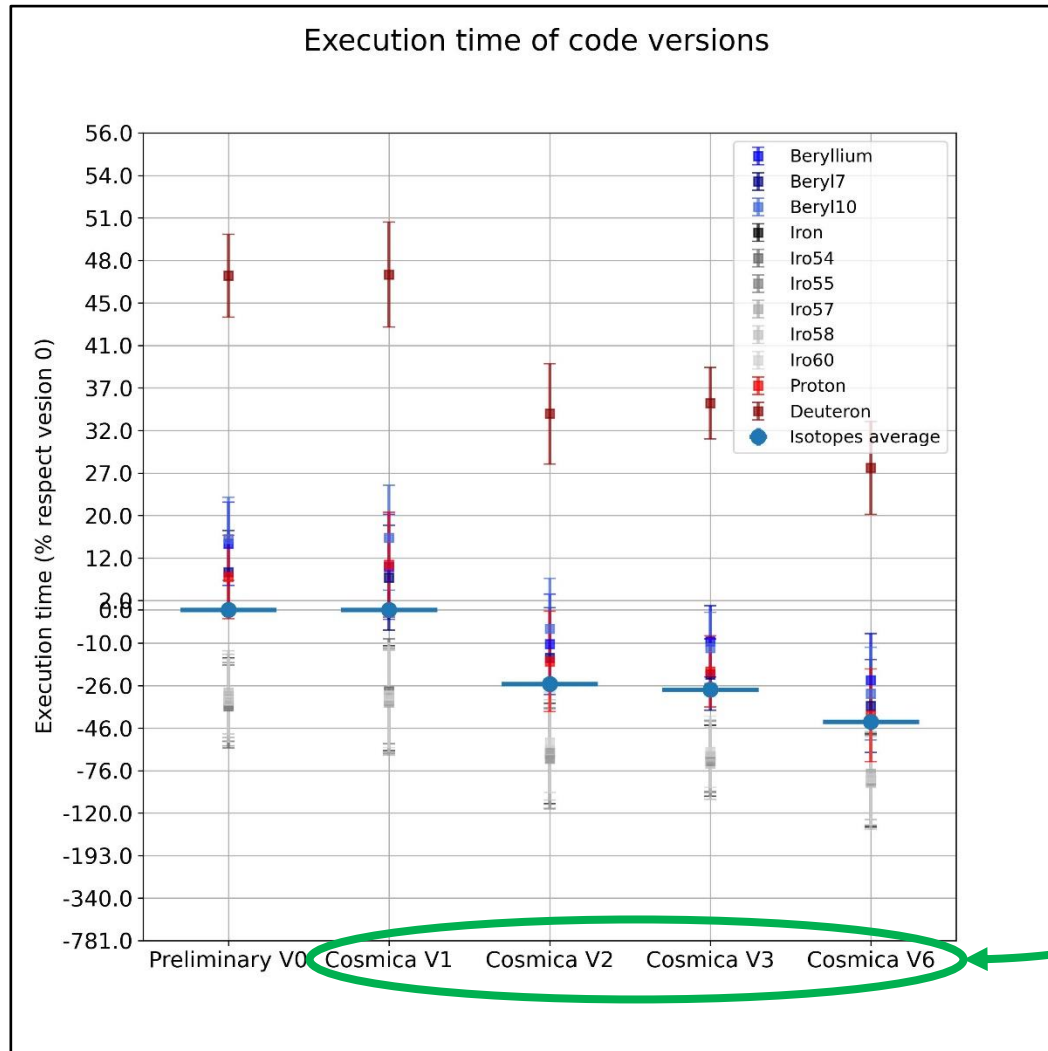


# Technical Objectives

Execution times for Forbush



# Methodologies and Solutions

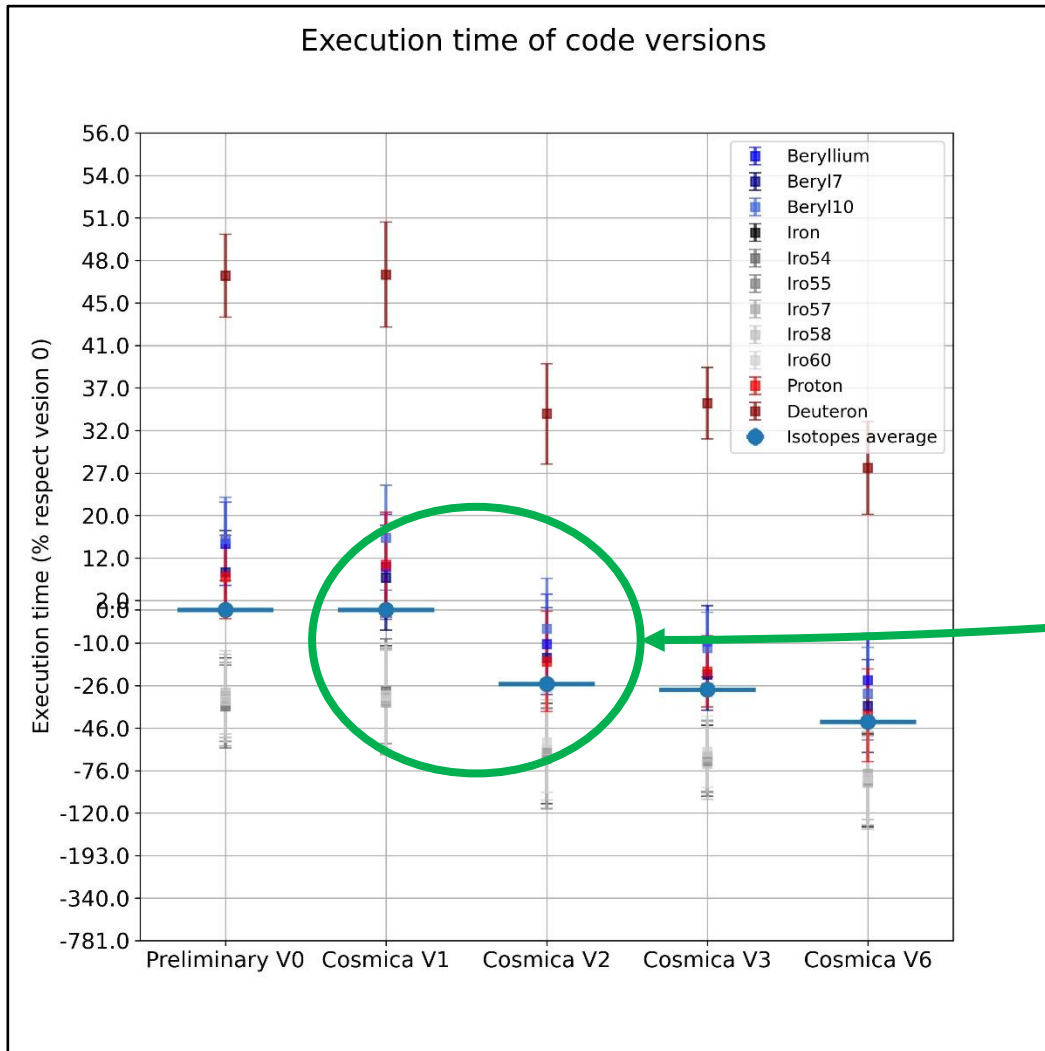


- Particle propagation independence
- SIMD parallelization on GPU  
CPU-GPU mixed CUDA code

- Propagation kernel bottlenecks profiling (Nsight)
- Optimization of stochastic computations, WpB and passage to rigidity

Available at the Spoke3 GitHub  
(open source)

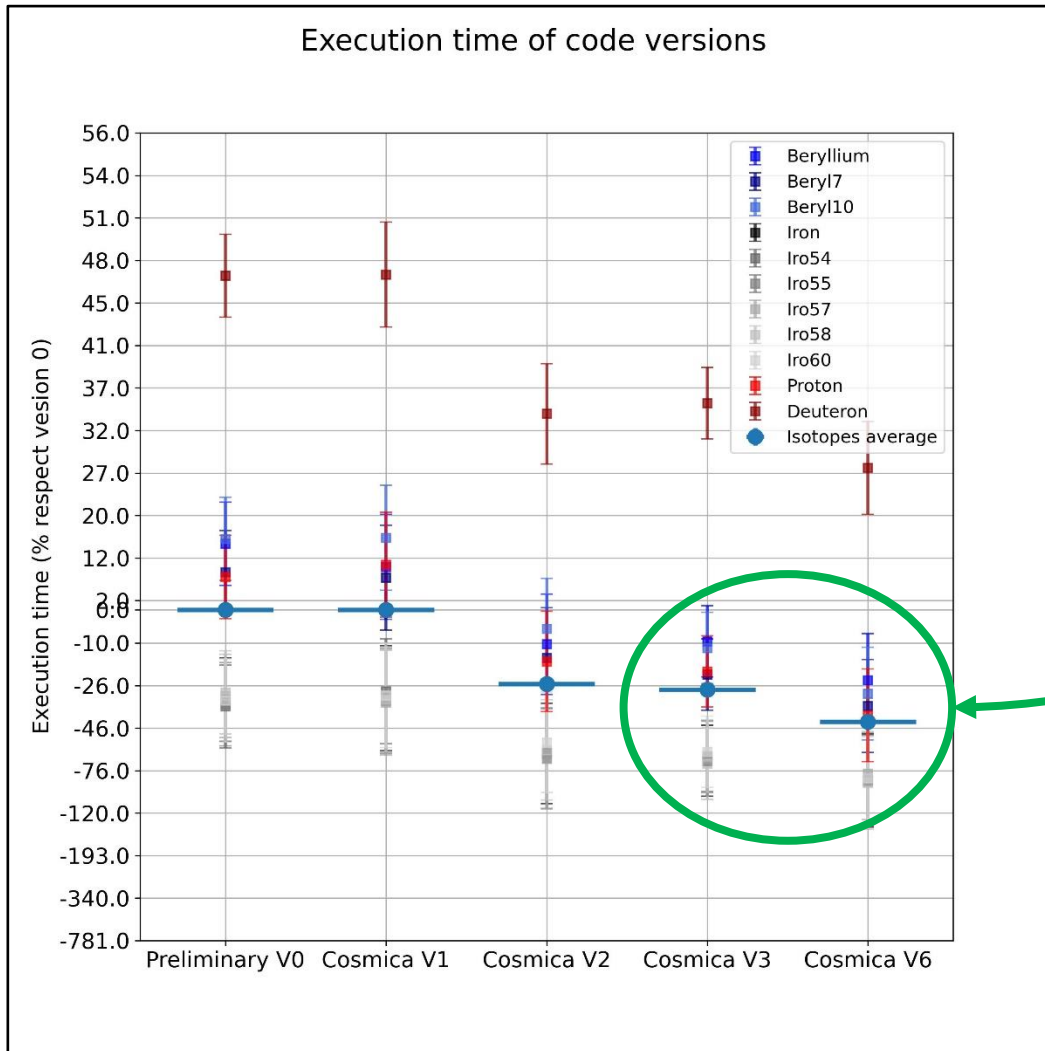
## Main Results



**V1 → V2 = 26% speedup:**

- Customized compilation flags → less registers
- Best WpB on A30 & A40 NVIDIA boards

## Main Results



**V3 → V6 = 19% speedup:**

- Switch from energy to momentum-rigidity PTE

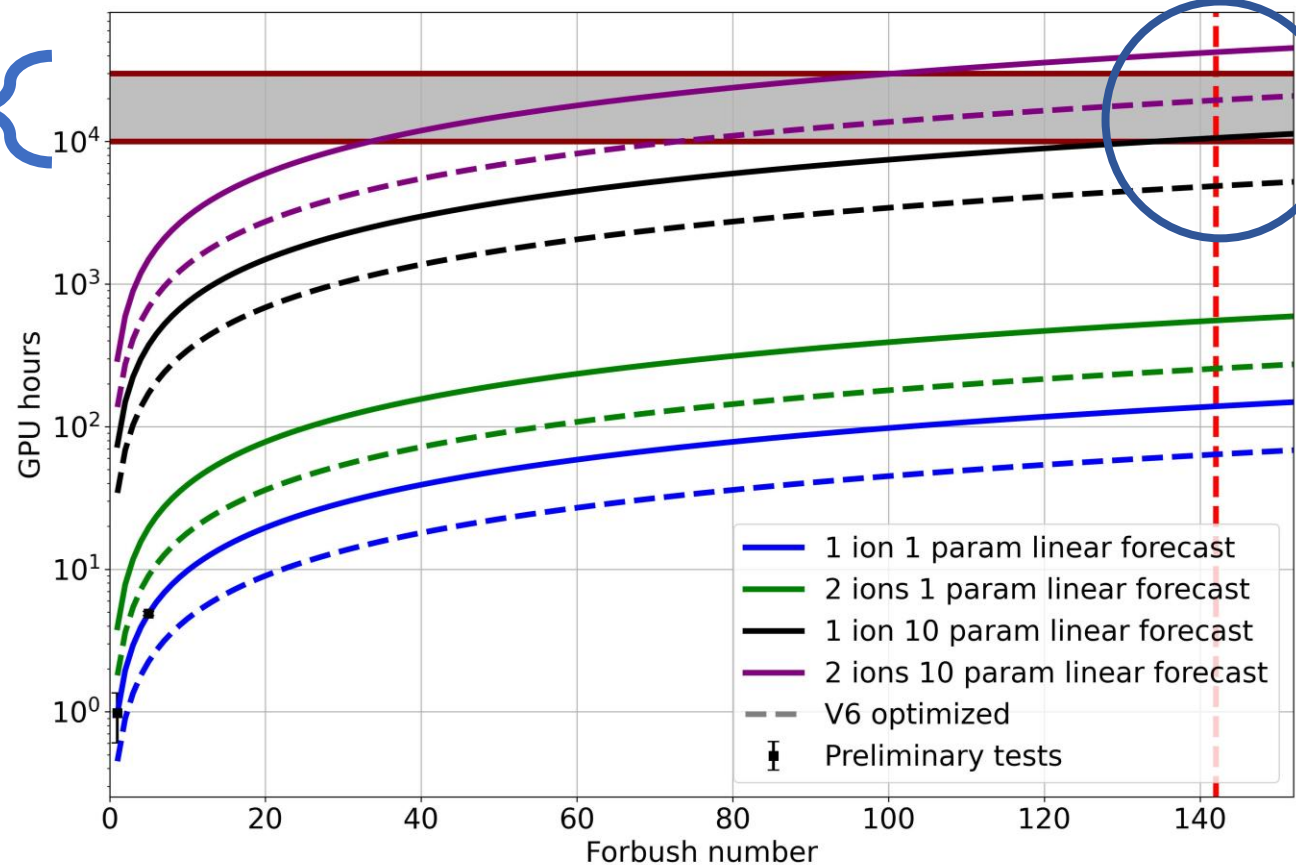


**V0 → V6 = 46%  
compressive speedup**

# Large scale real case

Execution times for Forbush

Minimal & maximum resources request at CINECA



- H not optimized **inside**
- H optimized **below**
- H and He not optimized **over**
- H and He optimized **just inside**

# Cosmica 1D model

Updated on the Spoke3 GitHub  
(open source)

## Cosmica 1D model

Here there are the codes and building scripts of the Cosmica 1D model of Cosmic Rays (CR) propagation in the heliosphere. These are the simplified version of the Cosmica code, which is 2D in modelling and 3D in propagation. The main algorithm is mantaine, but the propagation and implementation is reduced to its essential 1D components. This version of the model can be taken as toy model to understand the algorithm and perfor some test or start to develop a different physical propagation model.

**Tutorial COSMICA 1D**

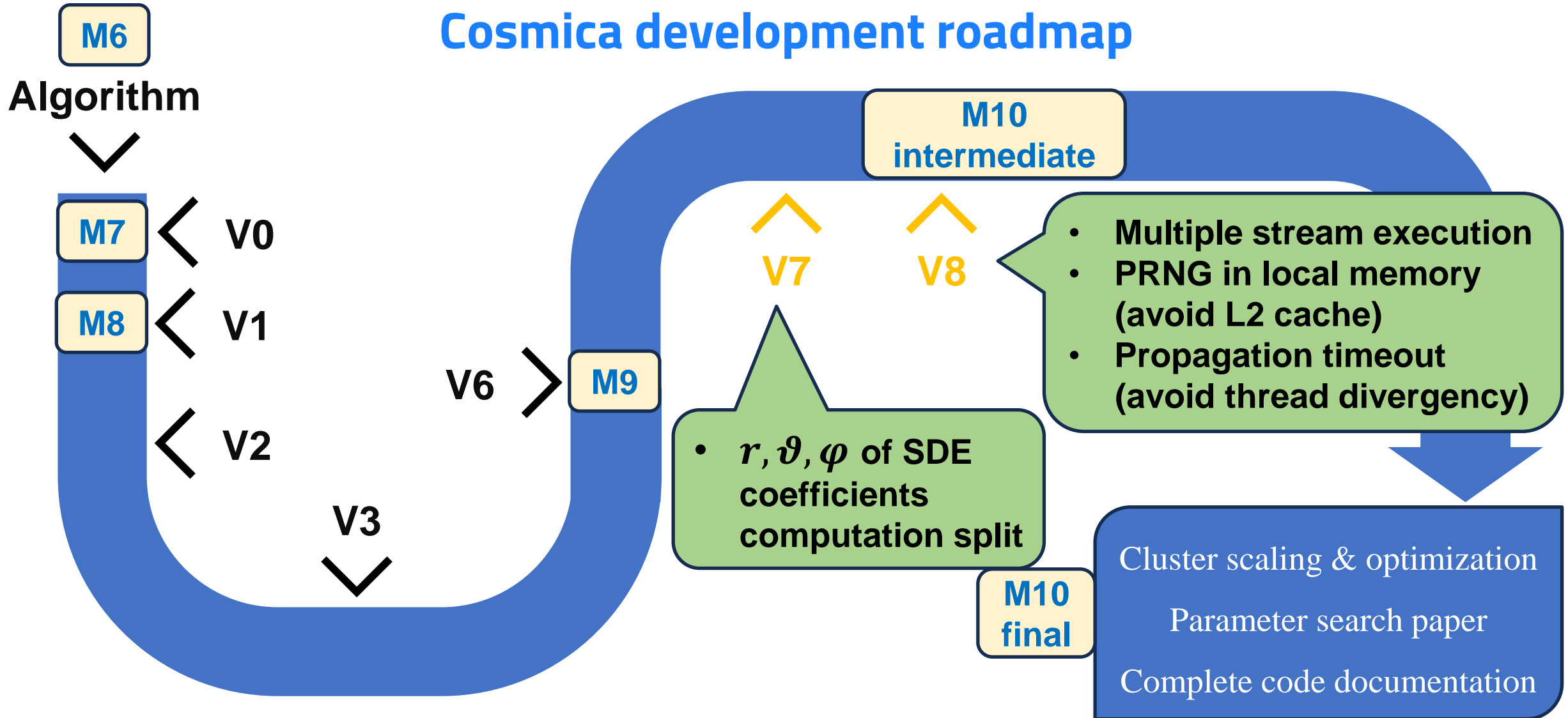
**Preliminary new model agreement test**

**Code modularity**

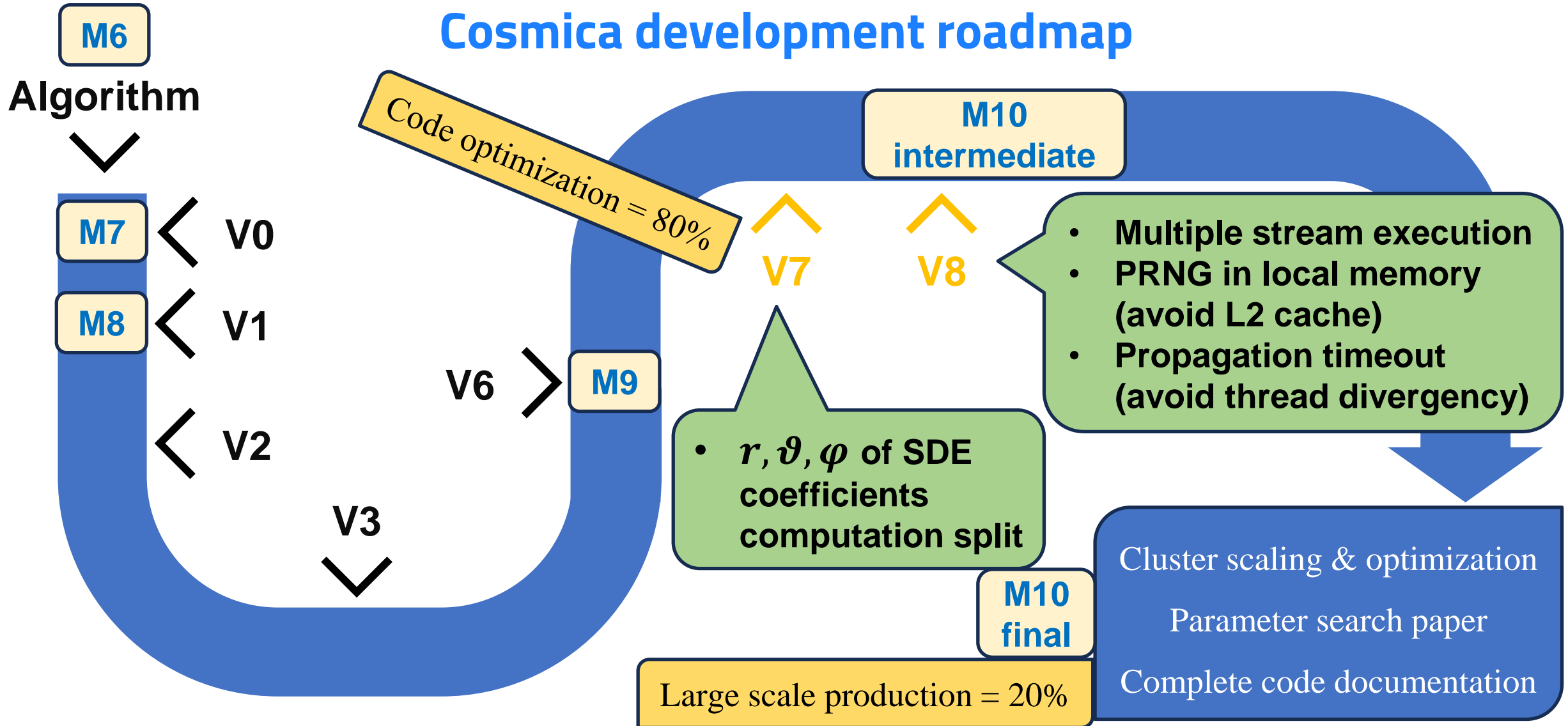
**COSMICA 1D & COSMICA 3D**



# Cosmica development roadmap



# Cosmica development roadmap



## Final Steps

Last optimized version releases



Version 7  
Version 8

Test of different physical models  
and code modularity



Strauss-Moraal model  
Latest tuning of the group

Large scaling and cluster optimization  
Intermediet M10 target



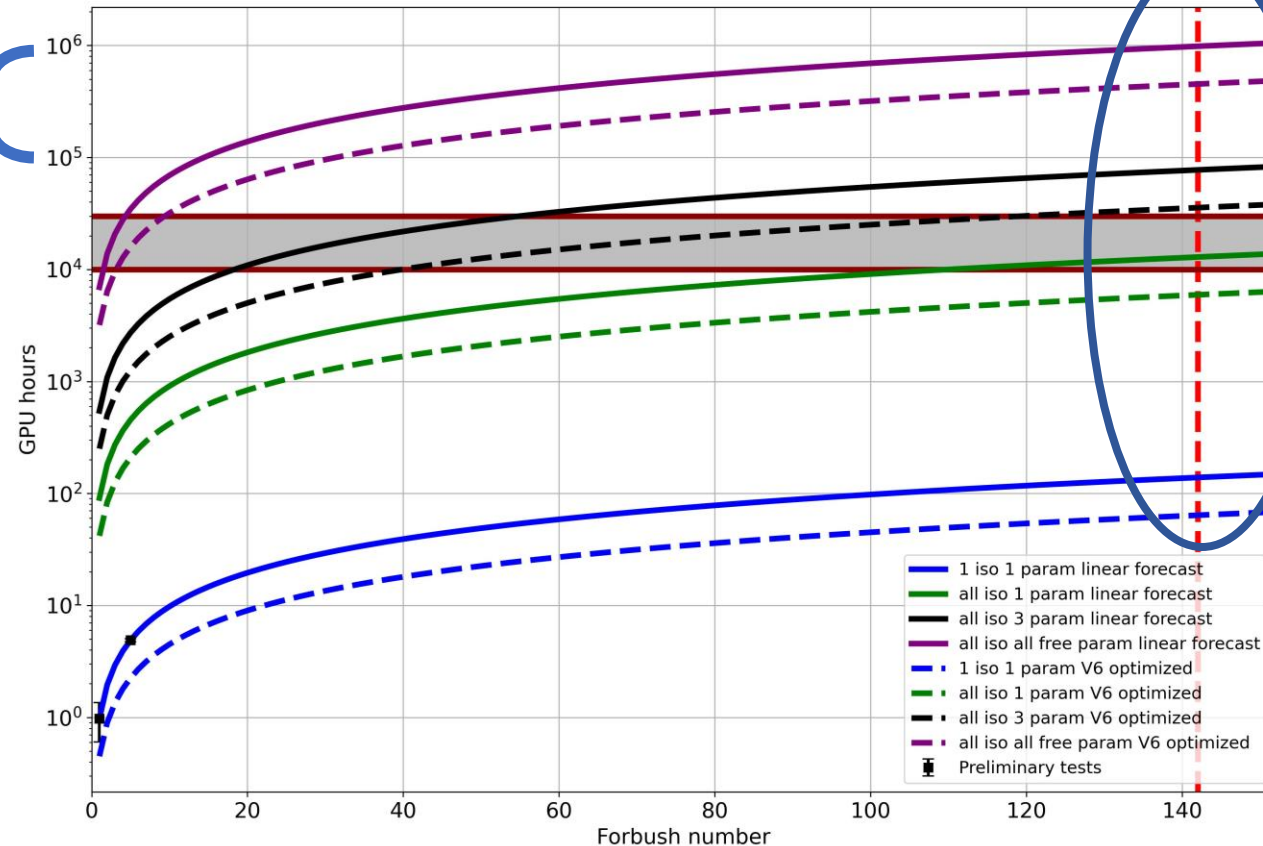
Parameter search paper publication  
Complete code documentation  
M10 final target



# Technical Objectives, Methodologies and Solutions

Execution times for Forbush

Minimal & maximum resources request at CINECA



- 1 parameter not optimized **inside**
- 1 parameter optimized **below**
- 3 parameter not optimized **over**
- 3 parameter optimized **inside**

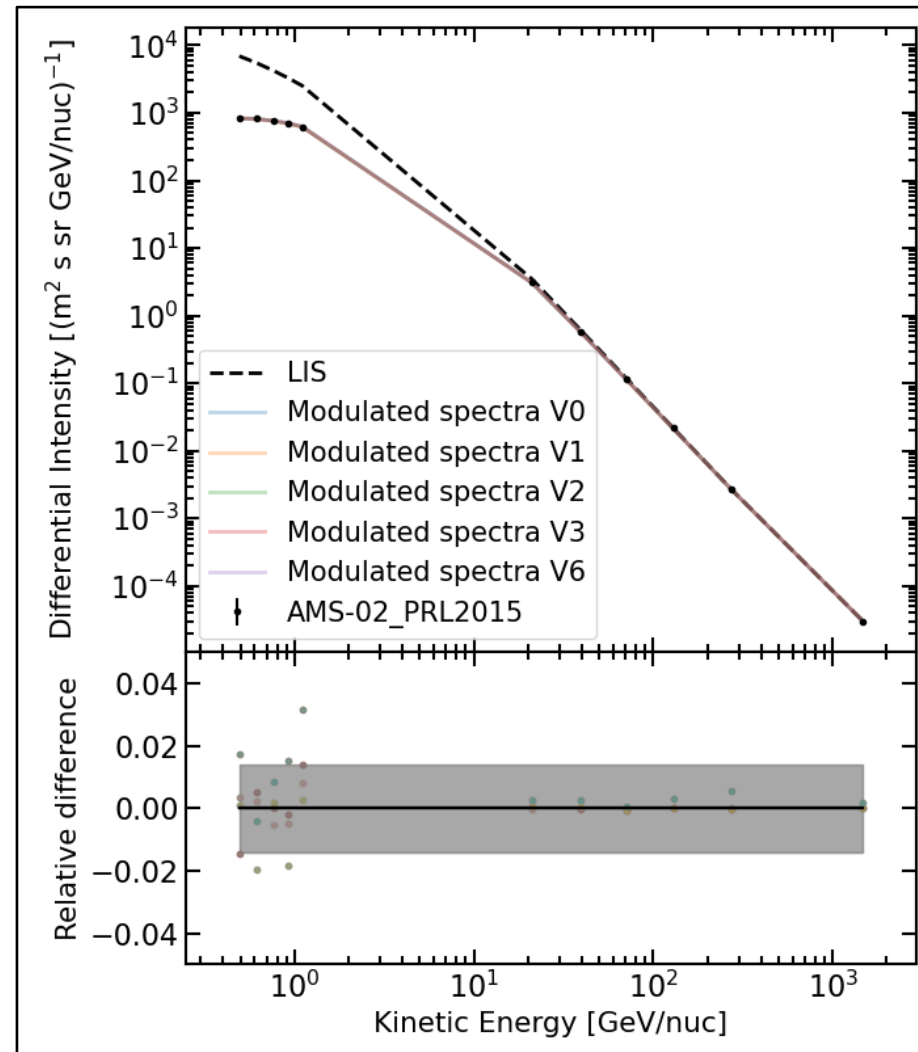
# Cosmica development history

- **V1 Milestone 7 version of the code**
  - Illustrated in the Elba meeting
- **V2 Improving internal structure**
  - Usage of customized compilation flags to reduce register compilation allocation
  - Implementation of the best Warp number per block derived from performance tests executed on A30 and A40 NVIDIA boards
- **V3 Optimization of stochastic computations**
  - Optimization of partial computations of the stochastic differential equation coefficients
  - Reduction of the allocated variables lightening over-structures
- **V6 Use of the rigidity as main variable instead of kinetic energy**
  - Reformulation of SDE in momentum form (one of which becomes trivial)

[Available on the Spoke3 GitHub \(open source\)](#)



# Main Results

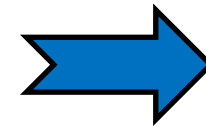




# New propagation computation

## Energy formulation

- $A_r = \frac{2k_{rr}}{r} + \frac{dk_{rr}}{dr} + \frac{dk_{r\theta}}{r d\theta} + \frac{\cot(\theta)k_{r\theta}}{r} + \frac{dk_{r\phi}}{r \sin(\theta) d\phi} - v_{drift,r} - v_{sw}$
- $A_\theta = \frac{\cot(\theta)k_{\theta\theta}}{r^2} + \frac{dk_{\theta\theta}}{r^2 d\theta} + \frac{dk_{r\theta}}{r dr} + \frac{dk_{\theta\phi}}{r^2 \sin(\theta) d\phi} + \frac{k_{r\theta}}{r^2} - \frac{v_{drift,\theta}}{r}$
- $A_\phi = \frac{dk_{\phi\phi}}{r^2 \sin^2(\theta) d\phi} + \frac{k_{r\phi}}{r^2 \sin(\theta)} + \frac{dk_{r\phi}}{r \sin(\theta) dr} + \frac{dk_{\theta\phi}}{r^2 \sin(\theta) d\theta} - \frac{v_{drift,\phi}}{r \sin(\theta)}$
- $A_T = \frac{2Tv_{sw}(T+2T_0)}{3r(T+T_0)}$
- $C = \begin{pmatrix} 2k_{rr} & \frac{2k_{rt}}{r} & \frac{2\csc(\theta)k_{rp}}{r} \\ \frac{2k_{rt}}{r} & \frac{2k_{tt}}{r^2} & \frac{2\csc(\theta)k_{tp}}{r^2} \\ \frac{2\csc(\theta)k_{rp}}{r} & \frac{2\csc(\theta)k_{tp}}{r^2} & \frac{2\csc^2(\theta)k_{pp}}{r^2} \end{pmatrix}$
- $L_B = -\frac{2v_{sw}(2T^2+4TT_0+T_0^2)}{3r(T+T_0)^2}$



## Rigidity formulation

- “
- “
- “
- $A_P = \frac{2}{3} \frac{V_{SW}}{r} P$
- “
- $L = 0$

1 trivial propagation equation

# Energy – Rigidity conversion

Energy unit

$$J = \frac{\beta U}{4\pi}$$

$$J_{mod}(T_{inj}) = \frac{\beta(T_{inj})}{N_{ev}} \sum_{j=1}^{N_{ev}} \frac{J_{LIS}(T_j)}{\beta(T_j)} \cdot \exp(L_j)$$

$$J_{LIS}(T_j) = J_{LIS}(P_j) \left( \frac{dP}{dT} \right)$$

Rigidity unit

$$J = fP^2 \quad \left( p = \frac{ze}{c} P \right)$$

$$J_{mod}(R_{inj}) = \frac{P_{inj}^2}{N_{ev}} \sum_{j=1}^{N_{ev}} \frac{J_{LIS}(P_j)}{P_j^2} \cdot \exp(L_j)$$

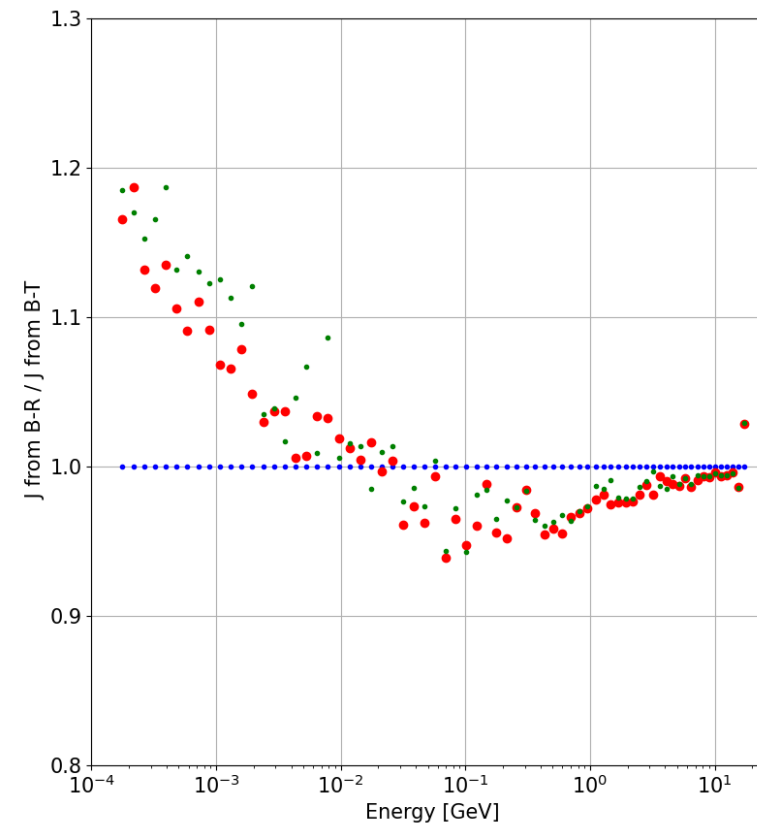
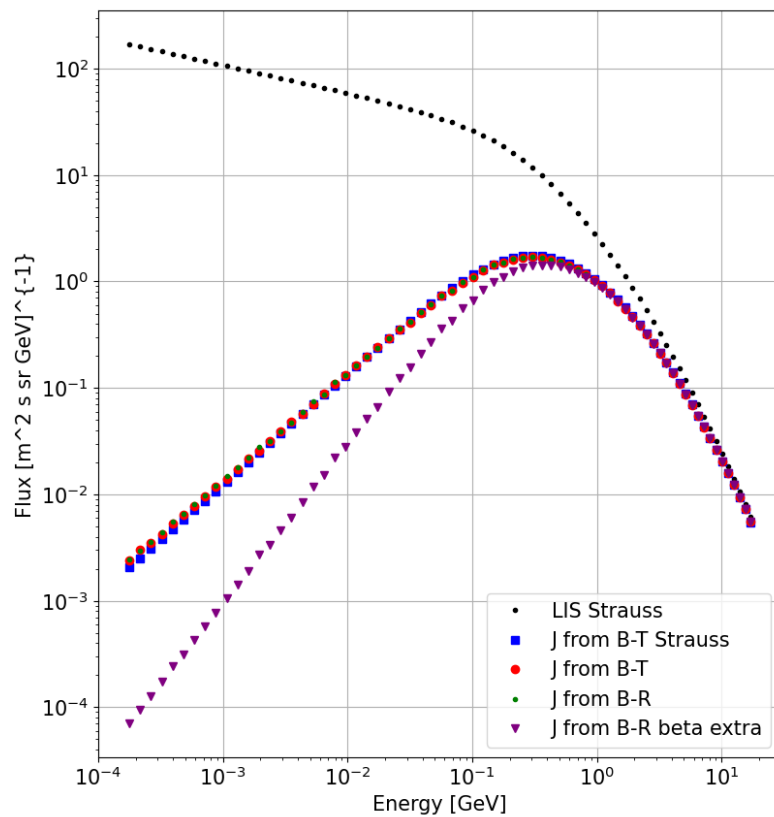
Extra conversion  
not needed

$$\frac{dP}{dT} = \frac{1}{\beta}$$

$\beta$  term cancels with  $\frac{dR}{dT}$

# Energy – Rigidity conversion

Comparison between Strauss-Moraal and Cosmica models



## Main Results



# PDP 2025

33rd Euromicro International Conference on Parallel, Distributed, and Network-Based Processing (PDP 2025)  
12-14 March 2025, Turin (Italy)

Submission of a paper to a HPC and parallel computing conference

### Cosmica: a novel parallel GPU code for Cosmic Rays propagation in heliosphere

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**Abstract**—The complex structure of interplanetary magnetic fields and their variability, due to solar activity, make it necessary to compute the CR modulation with numerical simulations. CODE for a Speedy Monte Carlo Involving Cuda Architecture (COSMICA) is a speedy and high precision Monte Carlo simulator of CR modulation, which solve the system of Stochastic Differential Equations (SDE) associated to the Parker Transport Equation

#### I. INTRODUCTION

Galactic Cosmic Rays (GCR) are charged particles mainly originating from astrophysical sources outside the solar system. These particles enter isotropically the heliosphere, which is the region dominated by the solar wind (SW) and its

## Activated new project paths

- **Sdegno Cascade founding project: parameter search (Cafoscari)**
- **Forbush scale study: CINECA GPU hours allocation (PICA)**

**Project:** CN HPC  
**AccountID:** 1667457  
CRPSF - Spoke3  
**Science Domain:** Astrophysics and Plasma Physics  
**Validity:** Monday, 16 September, 2024 to Thursday, 28 August, 2025  
**Status:**  
Active  
**ExpirationDelay:**  
6  
**Note:**  
16/09/2024: creazione progetto per 2024Q2 (tts#49882 - FDP) 27/09/2024: sostituzione fake user con PI Cavallotto - giovanni.cavallotto@mib.infn.it (tts#49882 - FDP)

### Details

**Hosts:** LEONARDO\_B  
**Budget (standard hours):** 30 000  
**WORK Quote (in GB):** 1 024

### People

**External Supervisor (PI):** Giovanni Cavallotto - INFN - Istituto Nazionale di Fisica Nucleare - 2 43 Dipartimento di Fisica

### Collaborators

**Collaborators:** Stefano Della Torre - Non Italian Universities - INFN Sezione di Milano Bicocca

## Cosmica development plan

- **V7 (under development) Separation of SDE coefficients computation for each coordinate**
  - Instead of using matrices of coefficients they are computed separately to relieve the register pressure
- **V8 (planned) Multi stream execution**
  - Execute multiple streams on the same GPU instead of using a 1 tot 1 coupling CPU-GPU with multiple GPUs
  - Relieve register pressure with architecture compiler flag
  - Change the PRNG (now it is Philox4\_32\_10 of cuRAND) to have more local RNG (not in L2 cache)
  - Avoid thread divergency and execution time differences

