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The OPAL KSP: status update

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1. INAF- OATo; 2. INAF-OATs; 3. INAF-IAPS; 4. ICSC

Spoke 3, II Technical Meeting, Bologna 17-19/12/2024

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OPAL is:

- -A PNRR Key Science Project with two goals:
 - produce a *library of detailed synthetic atmospheric models of giant planets* for the Ariel space mission (Turrini+2018; Edwards+2019)
 - port the $Ar\chi es$ planet formation suite of codes to HPC

Takes advantage of allocated time in: Leonardo (4 million core hours) Pleiadi (2.5 million core hours)

as well as the dedicated cluster GENESIS+.









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Meet the OPAL Team!

We do:

- -Modelling
- -Observations
- -Computational science

We focus on:

- Star formation
- Planet formation
- Solar system
- Exoplanets



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Why do we need the OPAL project?

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02 Jun 202

Mass - Period Distribution

- New observational data highlight our limited understanding of where planets form in disks

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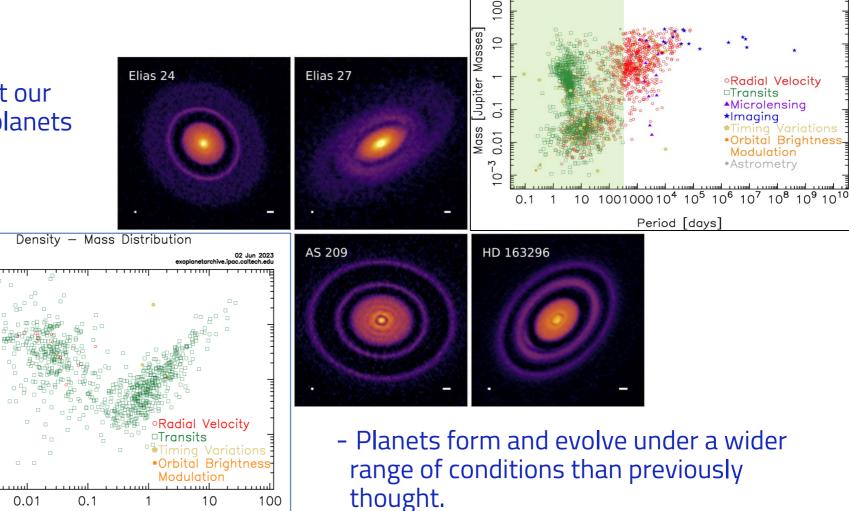
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Density



 10^{-3}

Mass [Jupiter Masses]

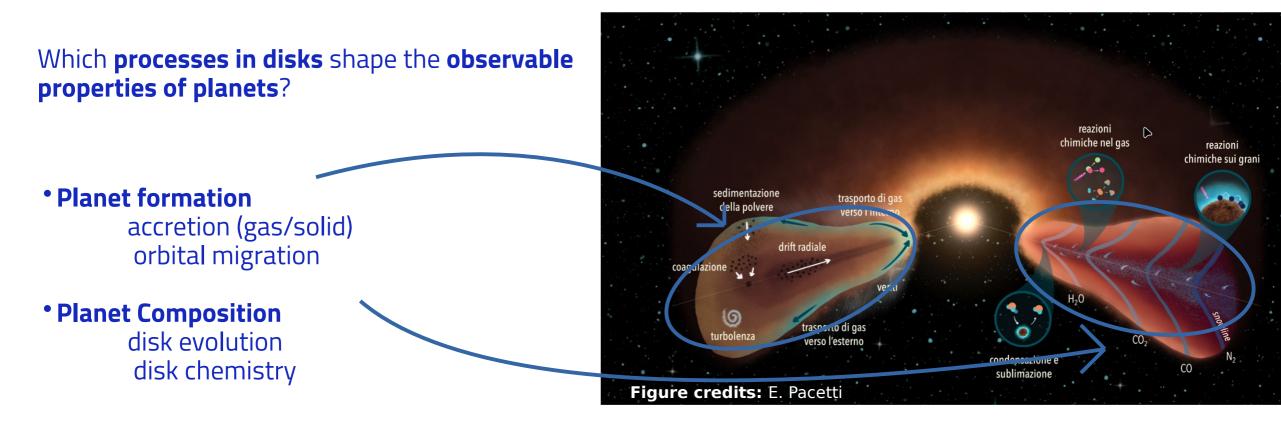








Modern Challenges in Planetary Formation



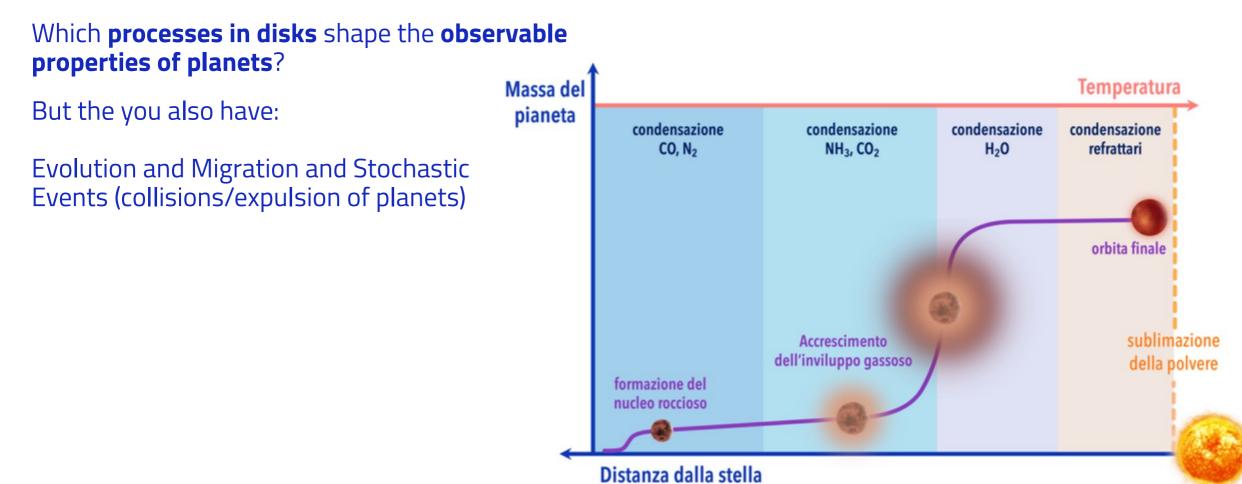








Modern Challenges in Planetary Formation



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Modern Challenges in Planetary Formation

What *can* we observe?

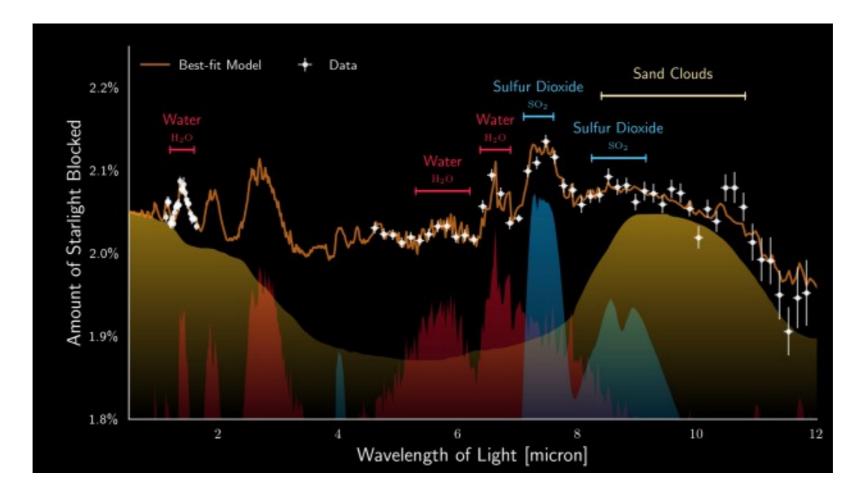








Atmospheres Basically





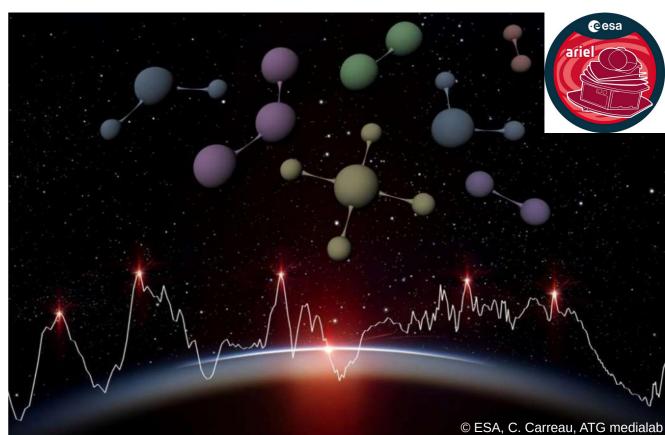






A Library of Synthetic Planetary Atmospheres

OPAL is the ongoing program to create an unprecedented library of end-to-end synthetic (exo)-planetary atmospheres, informed by actual data, using the **Arxes** suite of codes, to predict and consolidate what molecules can be detected in the planetary atmospheres that the **Ariel Space mission** will observe





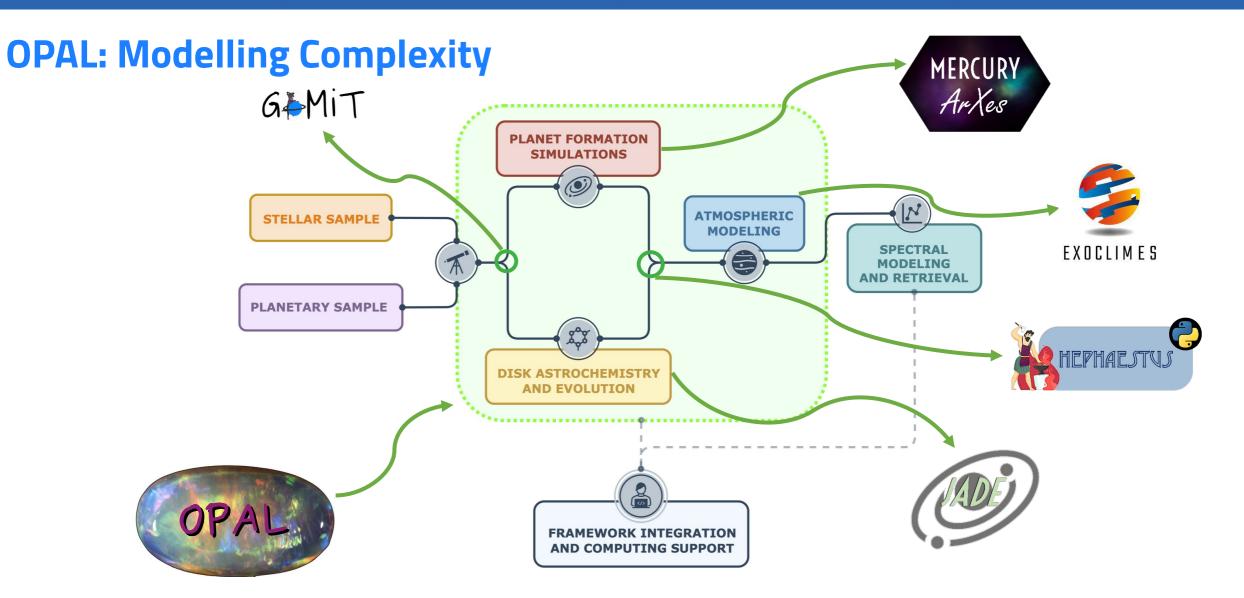
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- -Design Phase is finished:
 - First systems to simulate chosen, their physical parameters 'frozen'
 - Initial conditions are set
- -1st batch of simulations with Mercury-Ar χ es are done
- -Jade simulations on the way
- -The next batch of Mercury-Ar χ es are going to be launched these days
- -Vulcan synthetic atmospheres will be launched once Jade simulations are ready









OPAL: Modelling Complexity



JADE (Pacetti et al. in prep.) is the multi-language (Python + Fortran) data-parallel code to jointly simulate the *physical and chemical evolution of protoplanetary disks* of the Arxes suite.

Mercury-Arxes (Turrini+2019,2021) is the parallel n-body code of the Arxes suite incorporating MERCUR physical libraries to simulate *planet formation in protoplanetary disks* (see tomorrow's presentation).



HEPHAESTUS (Turrini+2021; Pacetti+2022) is the Python compositional post-processing tool of the *Arxes* suite that resolves planetary compositions across 20+ elements.

FastChem and **Vulcan** are atmospheric modelling codes from the University of Bern's *ExoClimes* suite that convert Hephaestus' elemental compositions into atmospheric molecular composition (e.g. Fonte+2023).



Ankes









OPAL: Modelling Complexity



IADE (Dacetti et al. in pren.) is the multi-language (Dython + Fortran) data-parallel code to Arxes An extra word on Vulcan:

Mercuryphysical I presenta

We (*Poolo Simonetti*) are expanding the chemical network of Vulcan to also include the condensation and sedimentation of iron!



(and soon also the silicates!)



sitional

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FXDCLIMES









GroMiT: The new kid in town!

GroMiT, the *Gro*wth and *Mi*gration *T*rack code, tracks how a forming planet accretes pebbles and gas as well as how it migrates depending on the local environment of its native circumstellar disk (e.g. Johansen+2019, Ida+2016, Tanaka+2020).

Born as a plug in for *Mercury-Ar* χ *es* to make it self consistent it found plenty of applications also a stand alone in exploring the formation histories of *observed planets*.

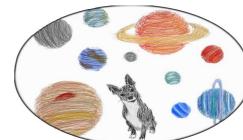
Used in **OPAL** to consolidate the initial conditions of the simulations run with *Mercury-Ar* χ *es* and *Jade*

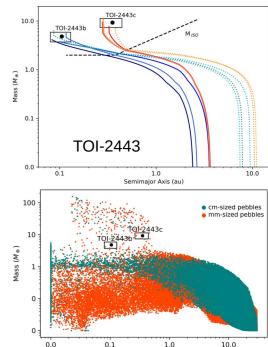
→ Single Planet Mode (Polychroni+2023)

It comes in two flavours!

→ Population Synthesis Mode (Polychroni+ in prep)







Final Semimajor Axis (au)

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Eccentricity



150

100



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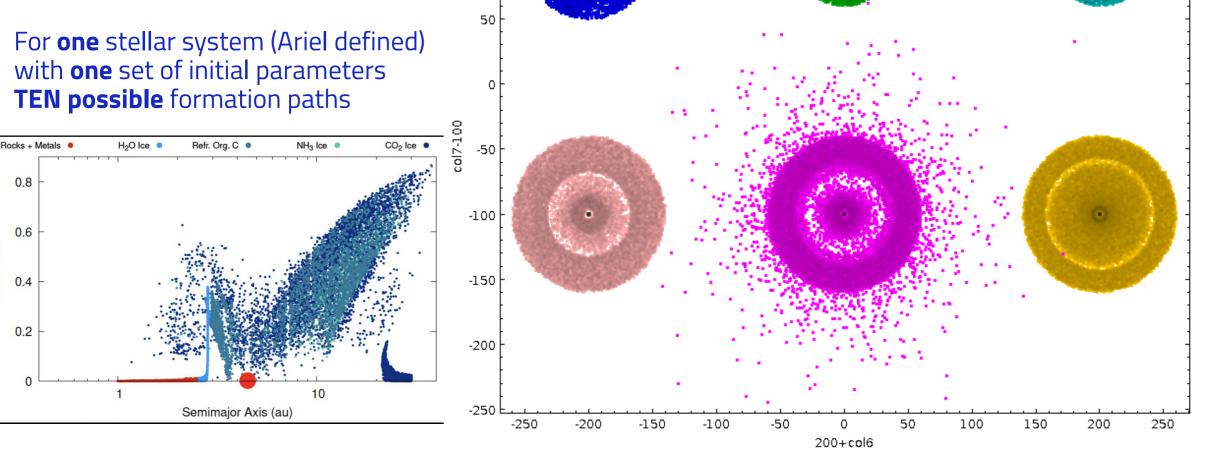
40au initial 30au 5au

> 2au 15au

Where we are now

- First results with Mercury-Ar χ es

with **one** set of initial parameters **TEN possible** formation paths



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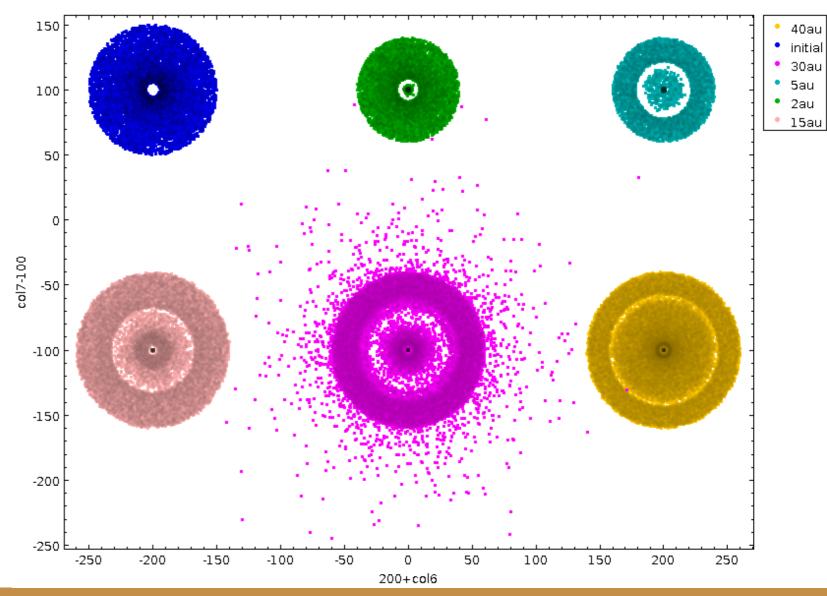


- First results with Mercury-Ar χ es

For **one** stellar system (Ariel defined) with **one** set of initial parameters **TEN possible** formation paths

(Historic overview) one such simulation *before parallelisation* Mercury-Arχes: Duration: ~3-4 months on GENESIS *after parallelisation:* Duration: ~3-4 weeks on GENESIS ~3-4 DAYS(!!) on Leonardo

Remains to be seen the GPU version



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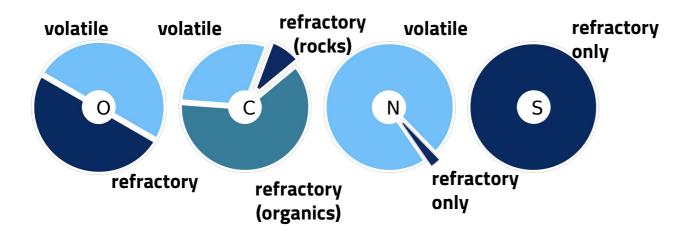


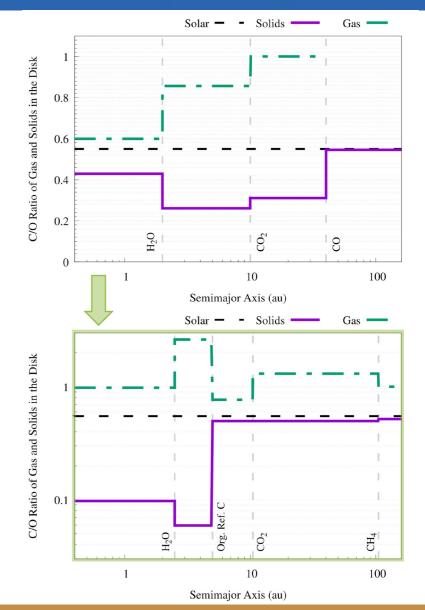






Jade Input for Circumstellar Disks





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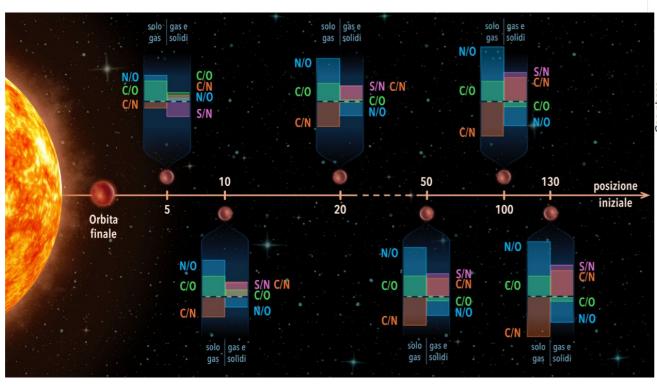


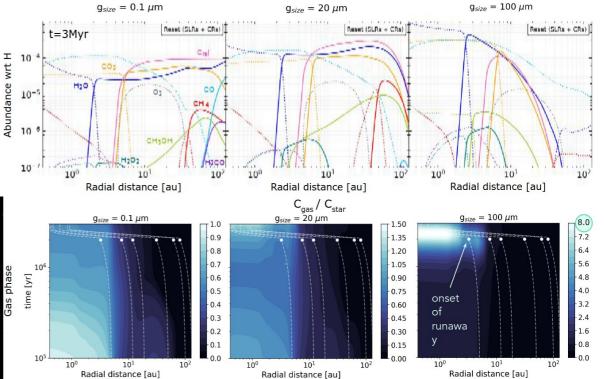




Jade Output:

Gas + Solids output













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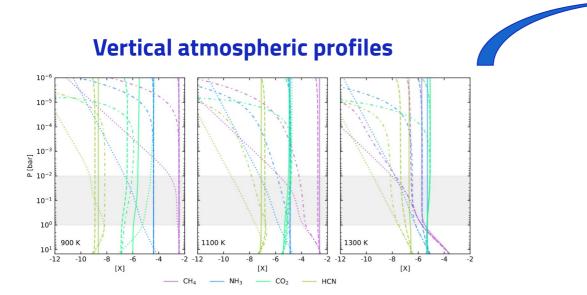


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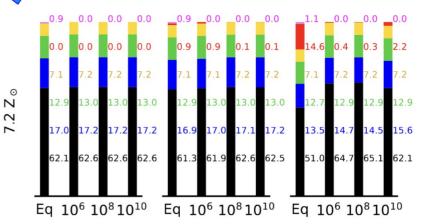
Chemical disequilibrium model with mixing, condensation/sedimentation & photochemistry

- chemical network expanded to include Mg, Si and Fe species
- recipes for the condensation and sedimentation of these refractories

Used to expand investigation of oxygen deficit (Fonte+23) & abundances of S-bearing species



Integrated abundances



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OPAL progress:

GroMiT: completion: ~70%; Code operational, needs to be parallelised & optimised (in progress)

Mercury-Ar es GPU porting: ~75% ; check P. Simonetti's presentation tomorrow

Simulation Campaign: ~10% ; so much more to come.

Still need to define: Repository and database for the completed simulations









OPAL as a Key Science Project

- The **seed of OPAL** is part of the Ariel-IT 2025's activities: its scope and planetary sample are widening as support from *INAF Pleiadi* and *Leonardo* became available.

- OPAL is the **next frontier in computational planetary science**: combining multi-domain and multiphysics *ab initio* simulations to fully understand the information encoded into planetary atmospheres.
- **Our community is currently the most advanced** in this integration of expertise and code in a unified scientific and computational ecosystem in disk-planet-atmosphere formation.

OPAL is a coordinated effort between the ARIEL, INAF and the PNRR