



## The selection effects in the LGRB $E_{p,i} - L_{iso}$ correlation

We simulate a large LGRB population with P, z,  $\alpha$ ,  $E_{p,o}$ , and  $L_{iso}$  to study the impact of peak flux P on the LGRB  $E_{p,i} - L_{iso}$  correlation. The mock z and  $L_{iso}$  are obtained from the redshift and luminosity distribution models (section 1) in previous work. The mock spectral parameters:  $E_{p,o}$  is from the observed bivariate  $(E_{p,i}\{z, E_{p,o}\}, L_{iso})$  distribution, and  $\alpha$  is from the observed bivariate  $(\alpha, E_{p,o})$ distribution. Finally, the mock P can be calculated through the above mock z,  $\alpha$ ,  $E_{p,o}$ , and  $L_{iso}$  data. These allow the mock data of each parameter to closely follow (section 2) the *Swift* observed distribution. However, to make the simulated P distribution consistent with the observed P distribution, the mock  $E_{p,o}$  has to be obtained from the mock  $E_{p,i}$  which is simulated based on the observed  $(E_{p,i}, L_{iso})$ distribution. This means that the joint  $(E_{p,i}, L_{iso})$  distribution is still effective to constrain the LGRB parameters. With this large simulated sample that can well represent Swift results, we find that the  $(E_{p,i}, L_{iso})$  distribution, which will directly affect the best-fitting result of the correlation, is significantly dependent on the value of P. Moreover, the P distribution at low- $E_{p,i} \& L_{iso}$  region is different from at higher- $E_{p,i}$  &  $L_{iso}$  region, which implies that there may be a subgroup of LGRBs in the low- $E_{p,i}$  &  $L_{iso}$  region.

## 1. The basic formula for z and $L_{iso}$ simulation

$$N_{LGRB} \propto \int_{0}^{z_{max}} \int_{\max(L_{lim}(z), L_{min})}^{L_{max}} \theta(P(L, z)) \varphi(z) \phi(L) dL dz$$

Probabilities of trigger and redshift measurement to **Expected LGRB** Intrinsic Intrinsic **Authors: Guangxuan Lan & Jean-Luc Atteia** Email: glan@irap.omp.eu

How to estimate  $\theta(P(L,z))$ ?

correct the intrinsic distribution (number) to the number that can be detected observational distribution (number)

redshift luminosity distribution distribution

Please see:

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## 2. The simulated distribution vs. the observed distribution



## The result of simulated $(E_{p,i}, L_{iso})$ distributions









The simulated data (blue) are well consistent with the observed data (red) in both 1D-parameter and 2D-parameter distributions. The observed sample comes from The Swift/BAT Gamma-Ray Burst Catalog<sup>1</sup> directly, finally including LGRBs with P >259  $0.5 \, ph \, cm^{-2} \, s^{-1} \, (15 - 150 \, keV)$ ,  $15 < E_{p,o} < 9000 \ (keV)$ and measured *z*.

log L<sub>iso</sub> (erg s<sup>-</sup>)

The colour represents the value of  $\log P$ . It is clear that data with different *P* have different  $(E_{p,i}, L_{iso})$  distributions and best-fitting  $E_{p,i} - L_{iso}$  correlations. In addition, the *P* distribution at low- $E_{p,i} \& L_{iso}$  region is significantly different from at higher-  $E_{p,i} \& L_{iso}$  region. This is another factor that affects the slopes of different fitting lines. We suppose that sub-class LGRBs maybe exist in the low- $E_{p,i} \& L_{iso}$  region. However, to clarify it, more data with low  $E_p$  are required, especially data with  $E_p$  lower than Swift's limit.

It can also be seen in this picture that the boundaries for data with  $\log P > 1.5$ are different from the entire simulated data. The colour here represents the value of log z. It is clear that the value of z at the boundaries is not simplex, which means the impact of P is not determined by redshift. Our results can also explain best-fitting why the  $E_{p,i} - L_{iso}$ correlation will change in different observed samples.

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1. https://swift.gsfc.nasa.gov/results/batgrbcat/index\_tables.html