



SEARCH FOR ORPHAN GAMMA-RAY BURST AFTERGLOWS IN RUBIN LSST DATA

HOW TO CREATE A MACHINE LEARNING CLASSIFIER TO IDENTIFY THEM

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ORPHAN GAMMA-RAY BURST (GRB) AFTERGLOWS: DEFINITION

CORE-COLLAPSE SUPERNOVA AFTERGLOW \Rightarrow long GRBs (> 2 seconds) PROMPT EMISSION (GRB) [1] θ_{jet}

 $\theta_{obs} > \theta_{iet} \Rightarrow \text{ORPHAN AFTERGLOW}$

Hard-to-find, faint and slow transients

MOTIVATIONS



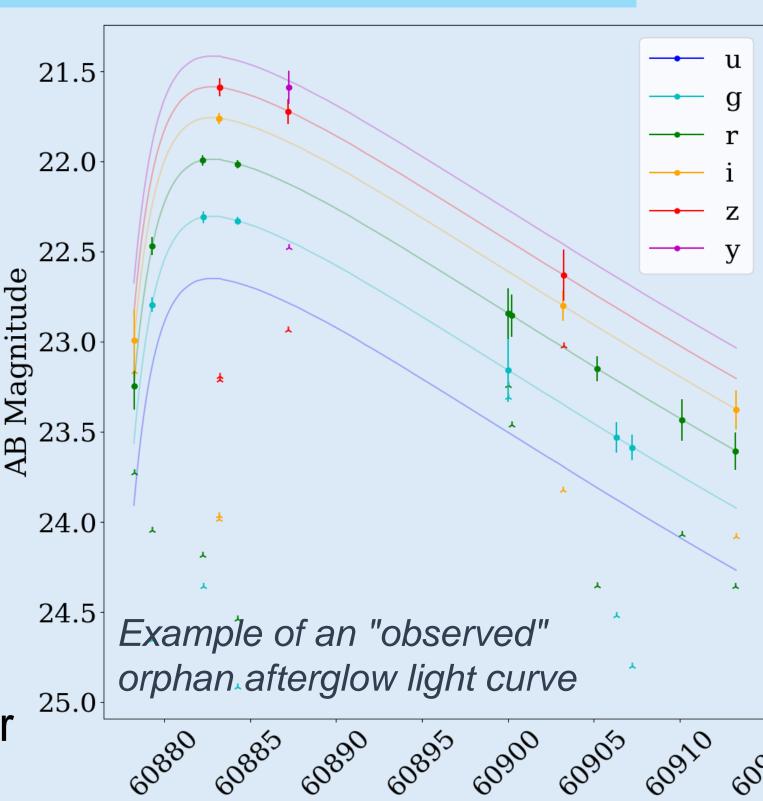
- Learn more about GRB physics and their progenitors,
- Help constrain jet structure,
- Multi-messenger analysis with gravitational waves [3].



- Simulation of a **realistic short GRB population** based on the **SBAT4** catalogue,
- Generation of their off-axis afterglow multi-wavelength light curves with afterglowpy [4]

 \Rightarrow short GRBs (< 2 seconds)

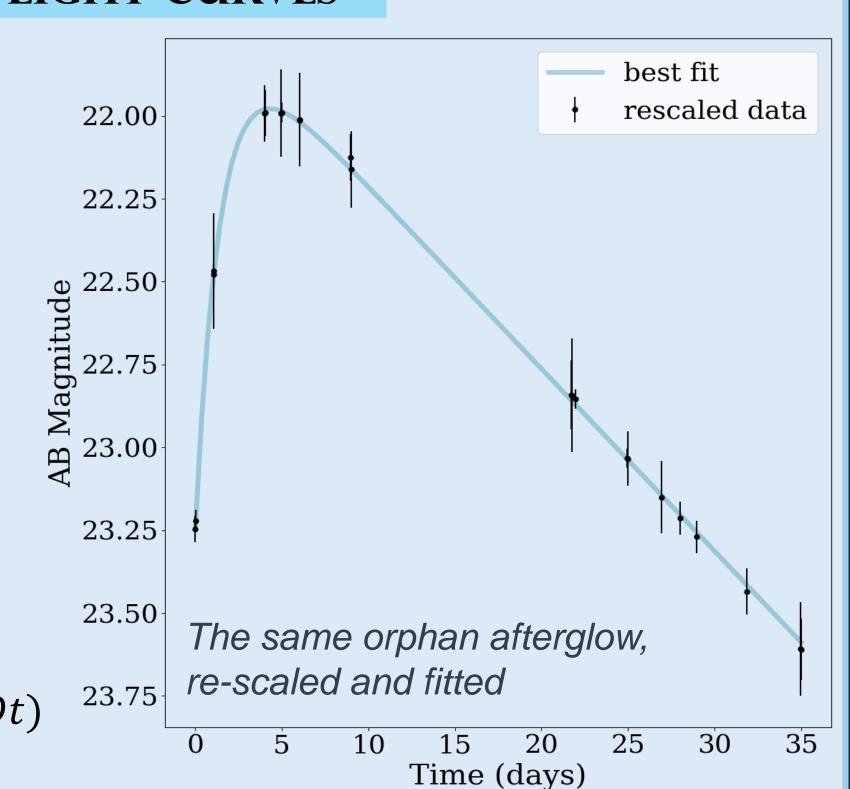
Simulation of their observation with the rubin sim scheduler emulator

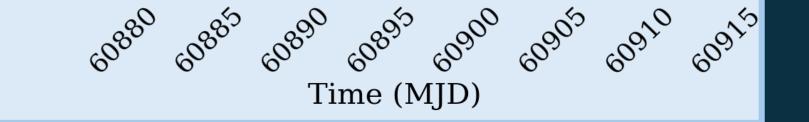


STEP 2: CHARACTERISE LIGHT CURVES best fit 22.00 Straightforward 22.25 features (colours, rise and fall durations...) و 22.50

Re-scaling of the point to the r-band using $F_{\nu} \propto t^{-\beta}$

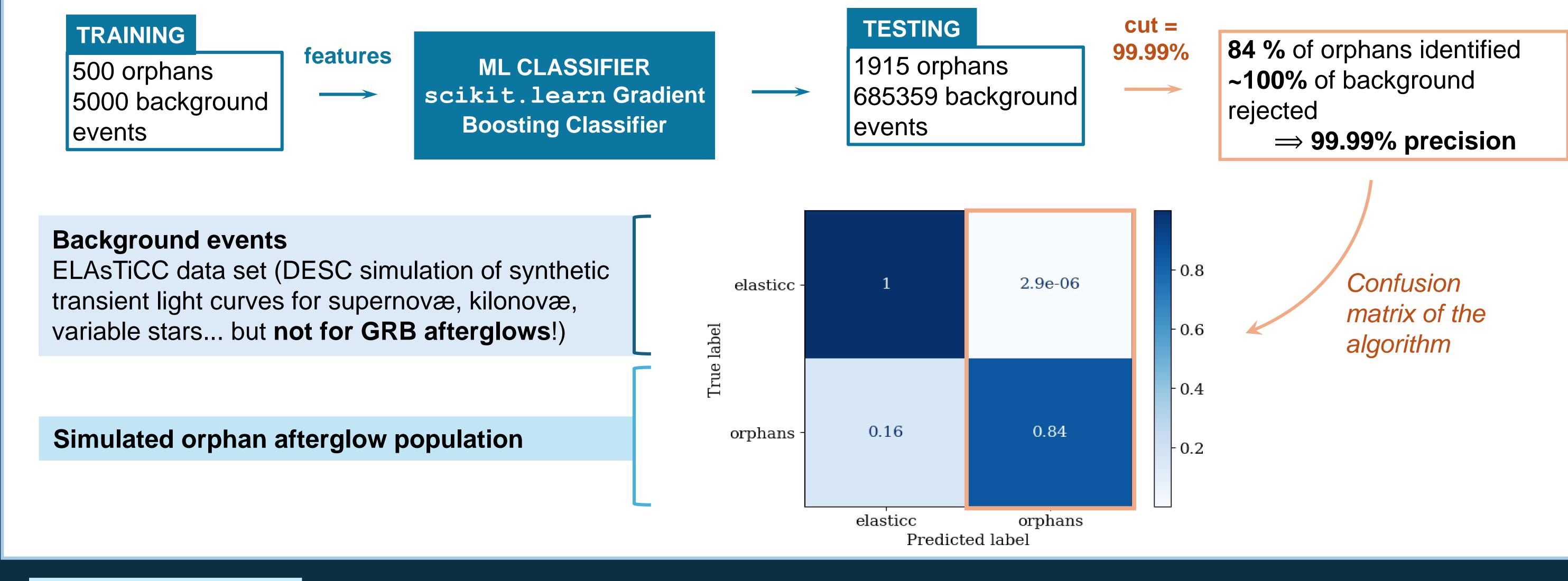
Empirical **fit** with: magAB(t) = At + B + Cexp(-Dt)





STEP 3: DESIGN A MACHINE LEARNING CLASSIFIER AND CHECK ITS PERFORMANCES

Goal: implement a filter in Fink [5] to identify the best orphan afterglow candidates



KEY TAKEAWAYS

- We developed a machine learning algorithm to identify orphan afterglows in Rubin LSST data, and it manages to recognise 84% of our simulated orphan afterglow population and to reject almost all of the background data [6].
- Testing filter on-going with ZTF public data and soon with Rubin ComCam DP1 data release

REFERENCES

[1] Sari, R., and T. Piran, 1997, *Ap. J.*, 485, 270+. [2] Mészáros, P., and M. J. Rees, 1997, *Ap. J.*, 476, 232. [3] Abbott, B. P. et al., 2017. Ap. J., 848(2):L12.!

[4] Ryan, G. et al., 2020, Ap. J., 896, 166. [5] Möller, A., Peloton, J., Ishida, E. E. O., et al., 2020, *MNRAS*, 501, 3272. [6] Masson, M. and J. Bregeon, 2025, *in prep.*