

# A New Sample of Transient Ultraluminous X-Ray Sources Serendipitously Discovered by Swift/XRT

Caltech

Celebrating 20 years of Swift Discoveries

Murray Brightman (Caltech)

Illustration

Brightman, M., Hameury, J.-M., Lasota, J.-P., et al. 2023 ApJ, 951, 51 (arXiv:2305.01693)

# Ultraluminous X-Ray Sources



- X-ray sources with  $L_x > 10^{39}$  erg s<sup>-1</sup> (Eddington luminosity of 10 solar-mass black hole), not including accreting SMBHs.
- Imply more massive black holes (perhaps intermediate mass ones), or super-Eddington accretors.
- Several well known ULXs are now known to be powered by neutron stars (e.g. M82 X-2, Bachetti+14, NGC 5907 ULX1, Israel+17).
- Most well studied ULXs are persistent sources.

# Ultraluminous X-Ray Sources

- M51 has a large and diverse population of ULXs

ULX9

M51b

ULX5

ULX7 (ULX pulsar,  
Rodriguez Castillo+20)

Eclipsing ULXs

ULX8  
(Cyclotron line source,  
Brightman+18)

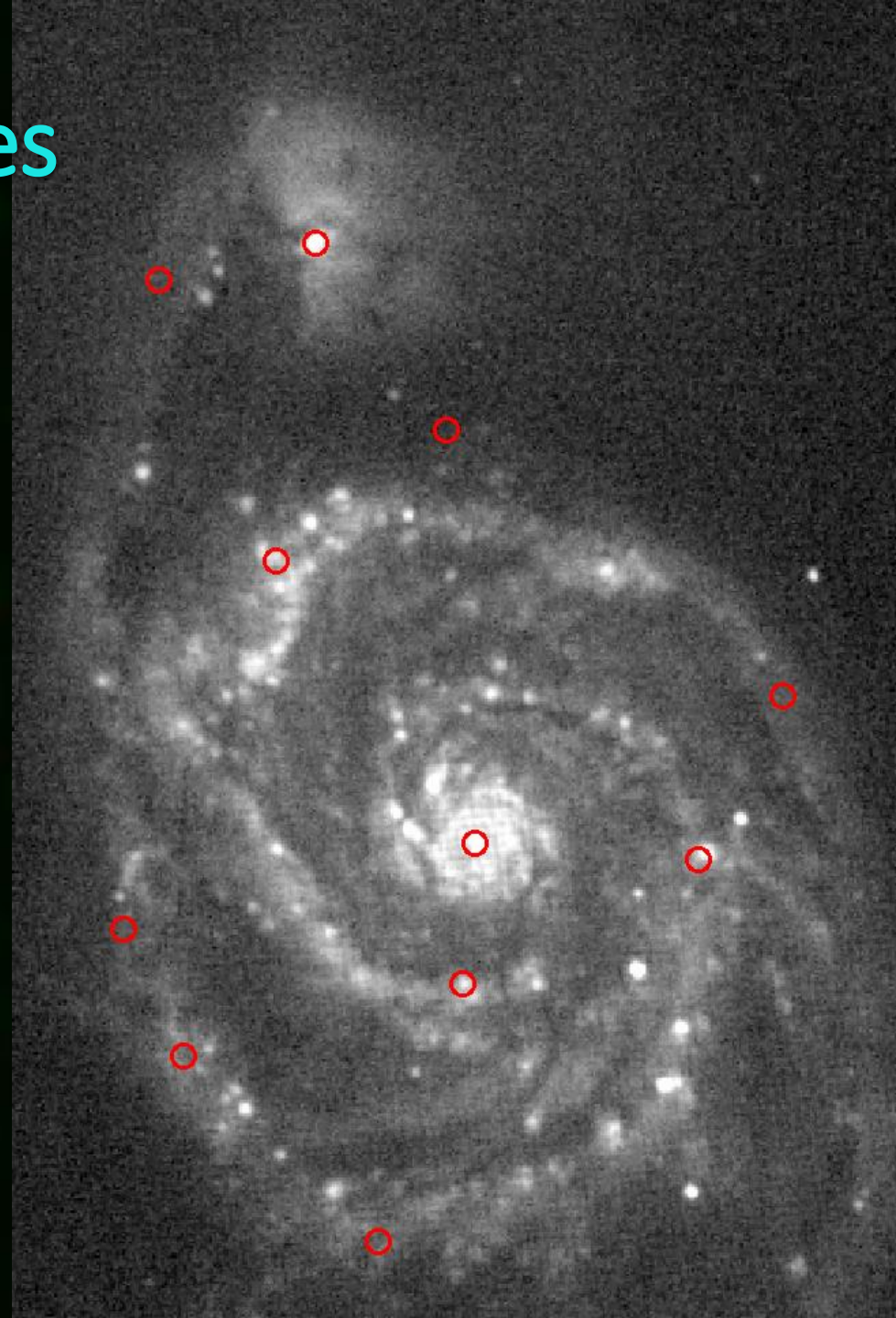
M51a

ULX2

ULX4

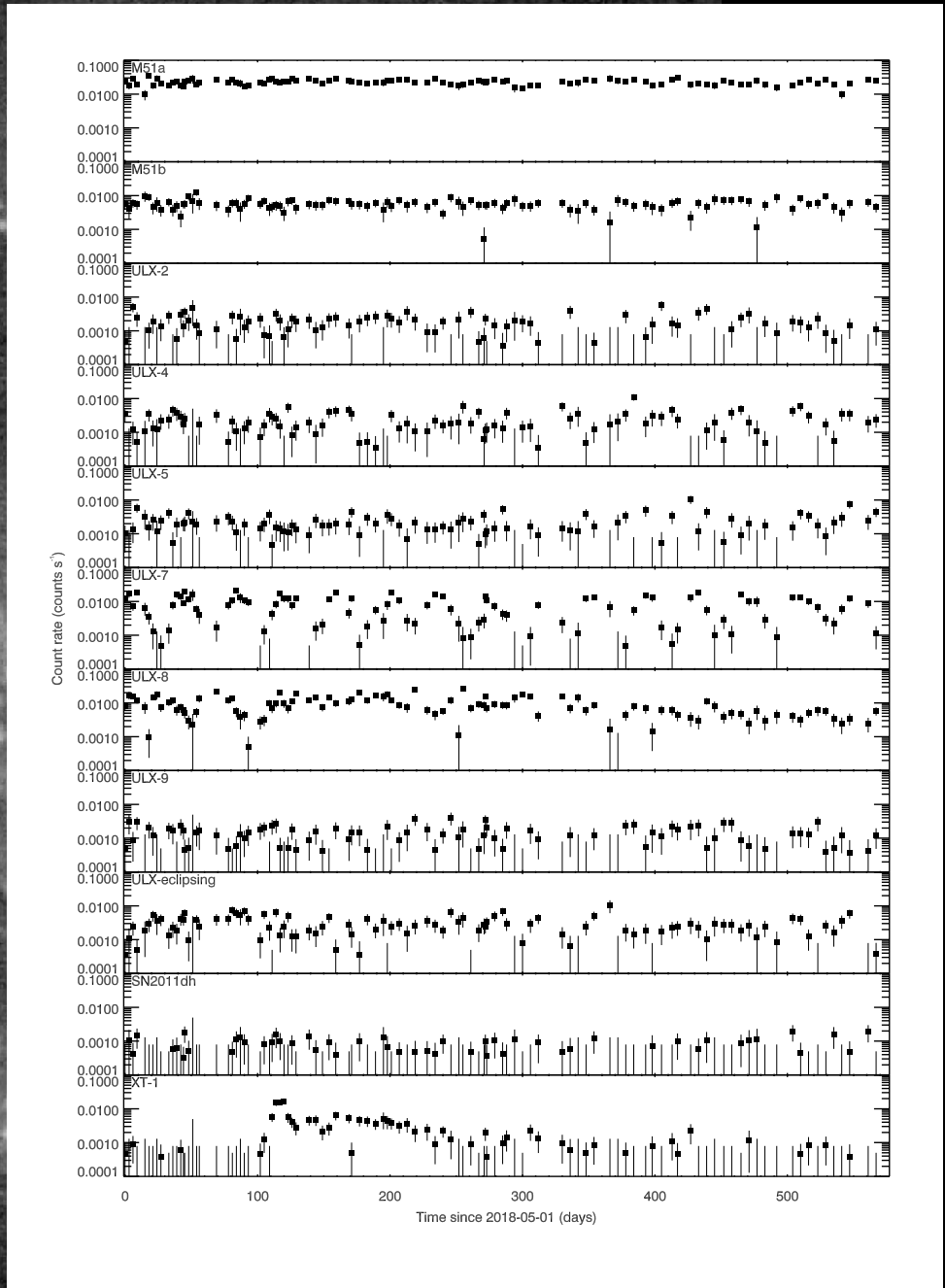
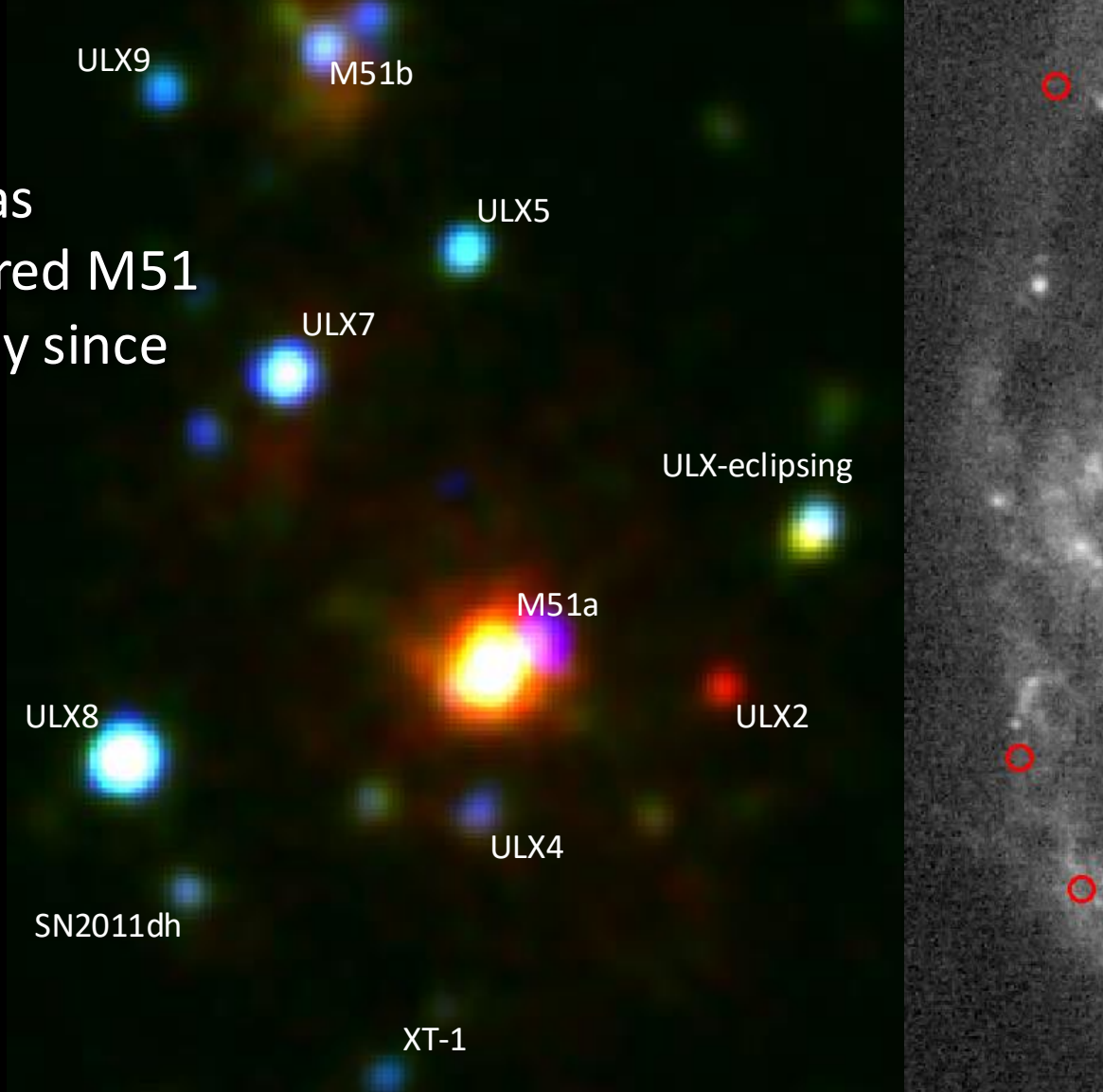
SN2011dh, type IIb  
supernova

XT-1 (transient ULX,  
Brightman+20)



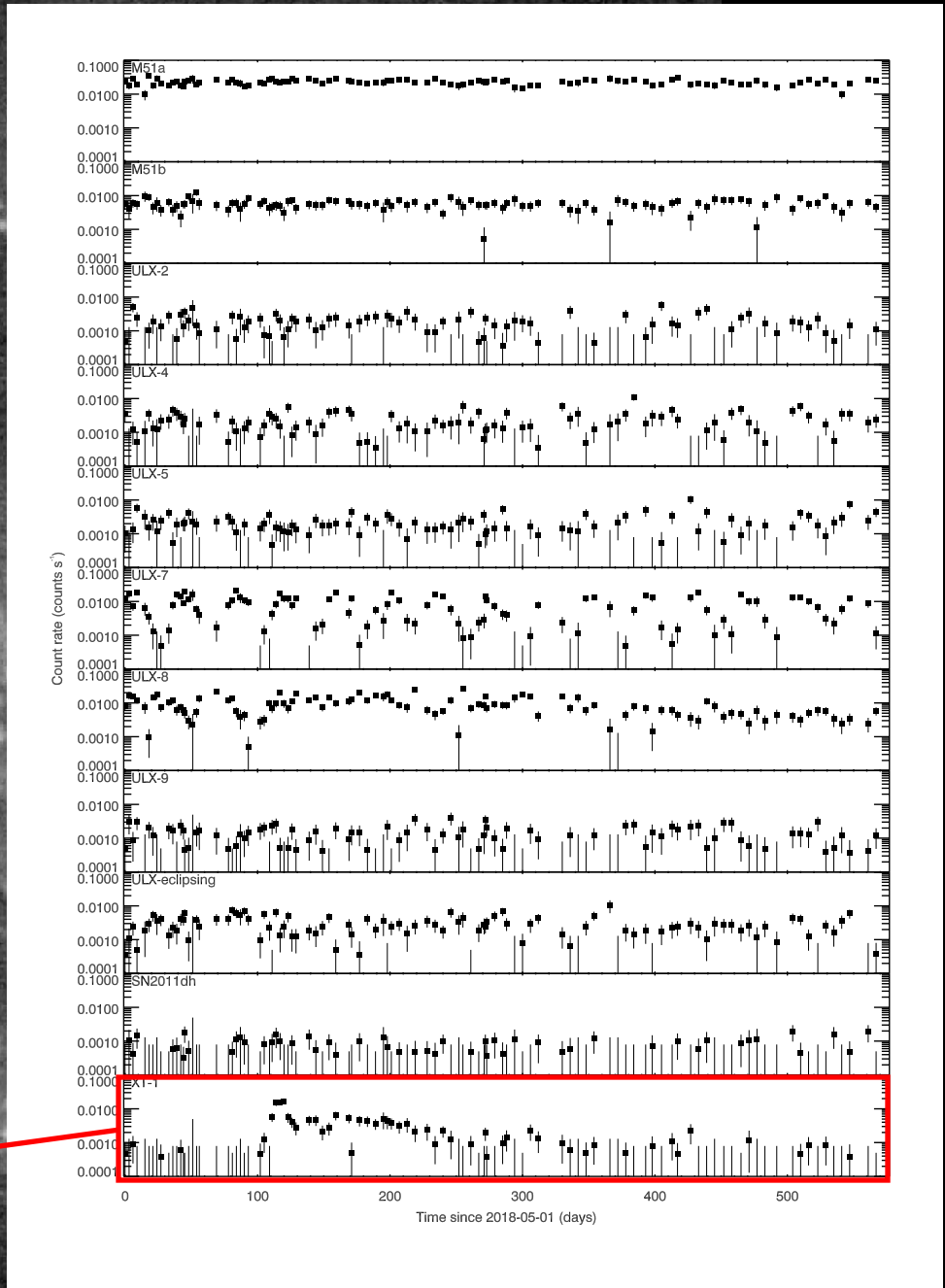
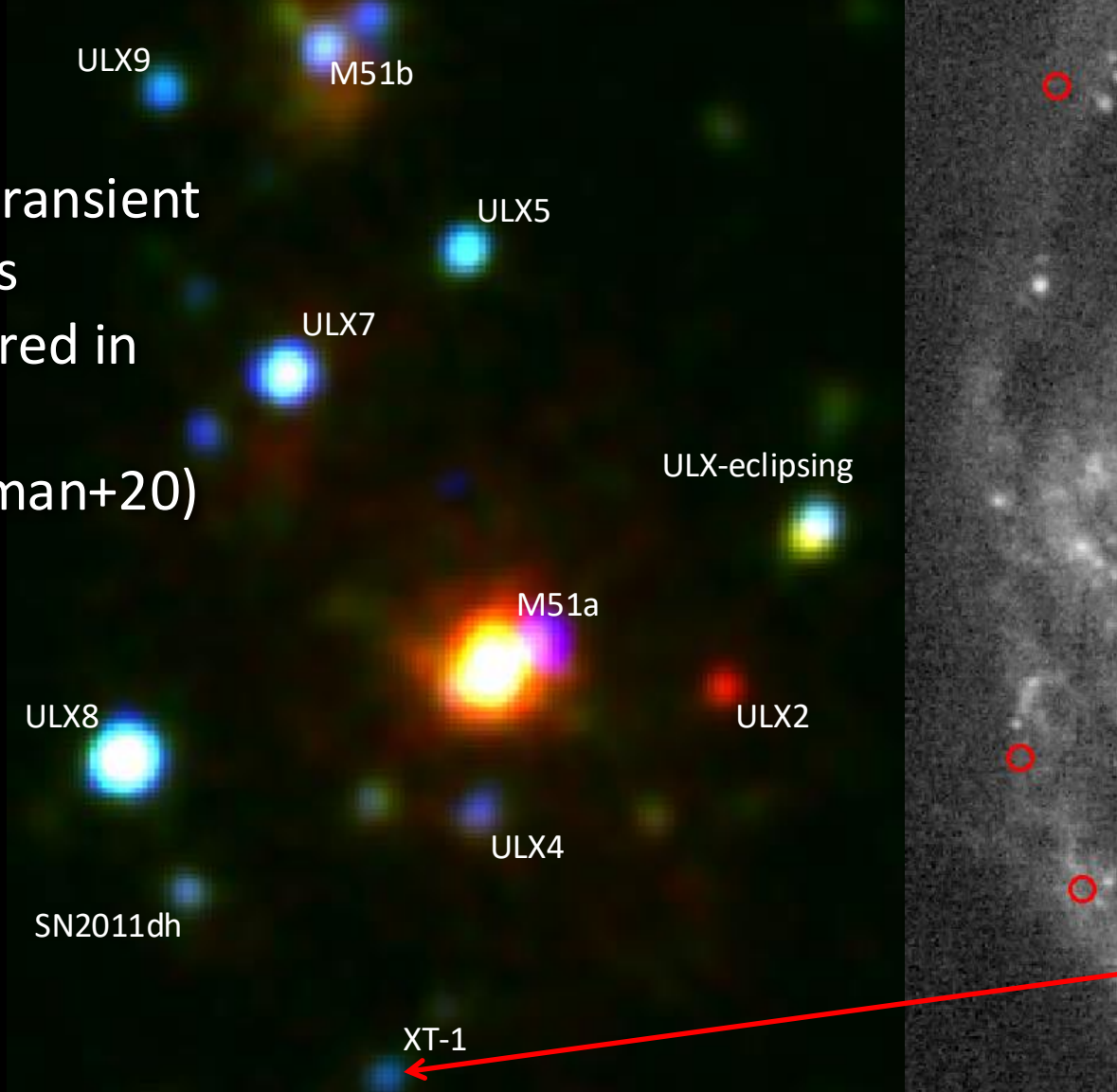
# Ultraluminous X-Ray Sources

- Swift has monitored M51 regularly since 2018

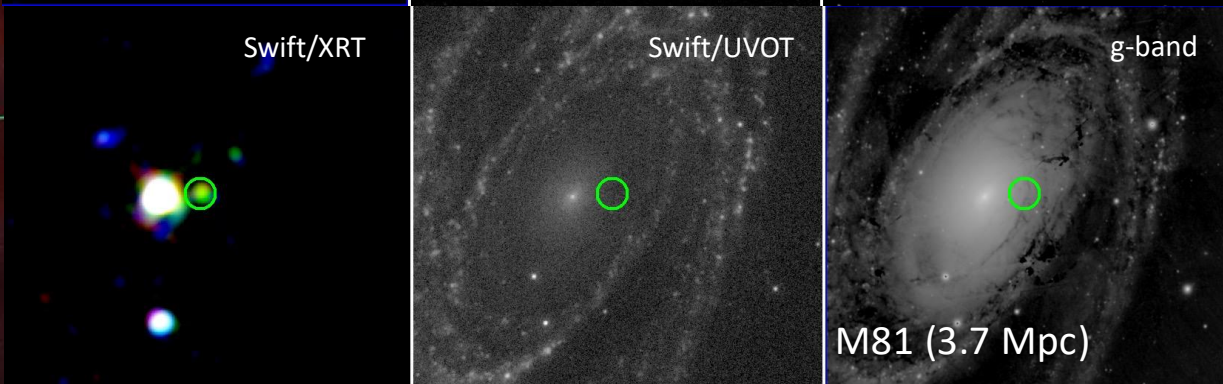
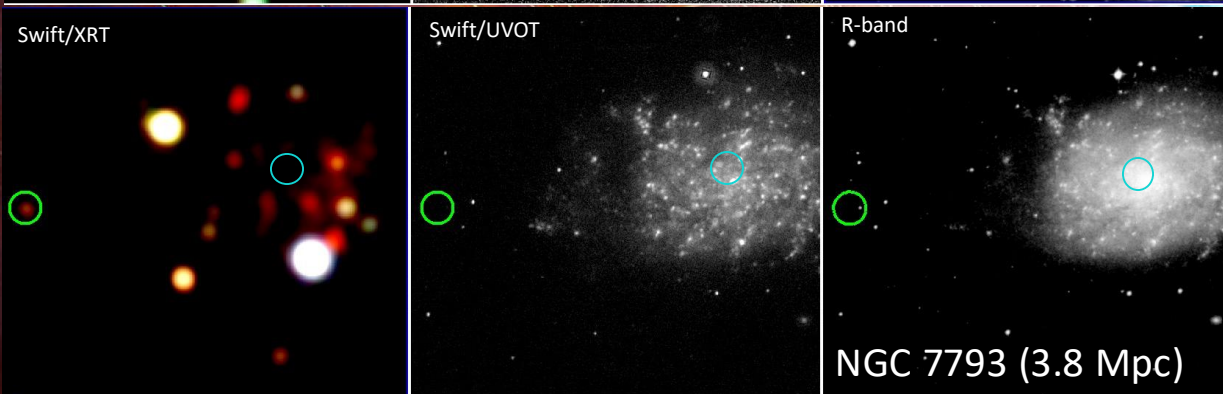
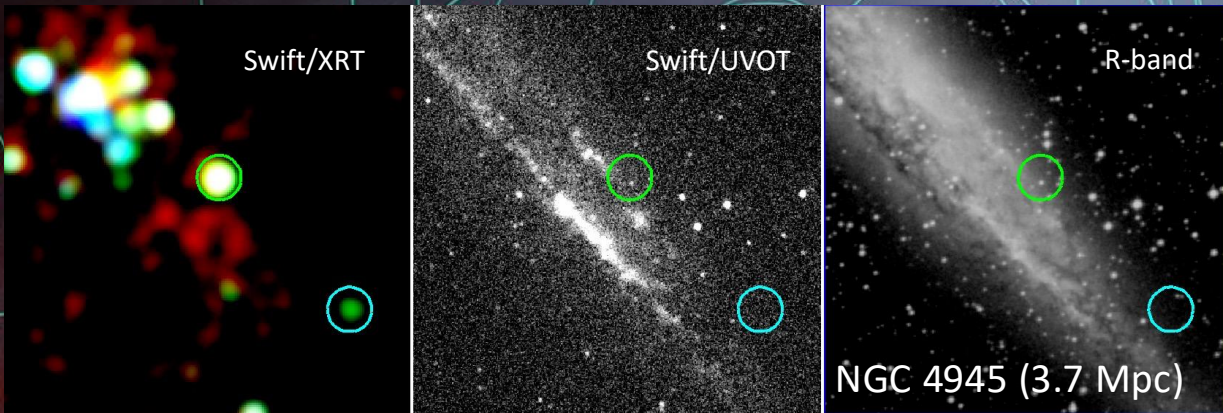


# Ultraluminous X-Ray Sources

- A new transient ULX was discovered in 2018 (Brightman+20)

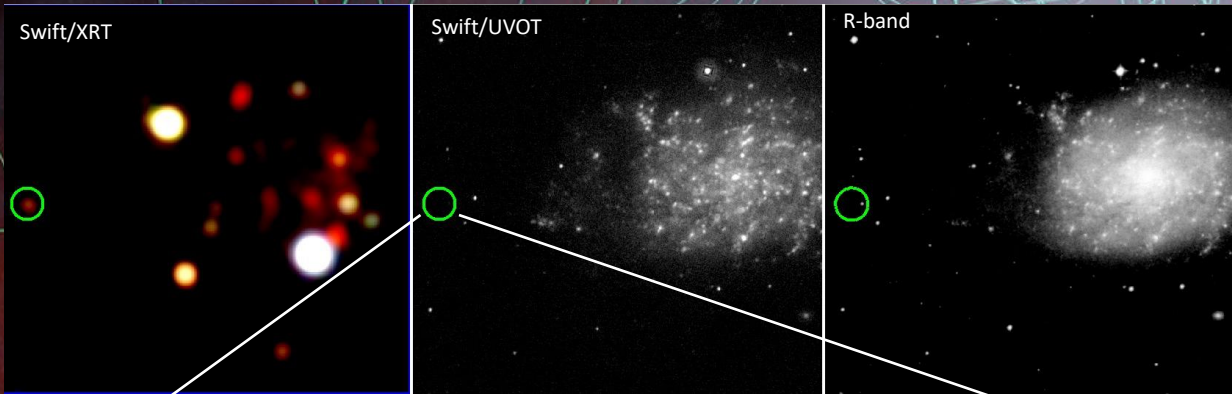


# A New Sample of Transient Ultraluminous X-Ray Sources Serendipitously Discovered by Swift/XRT



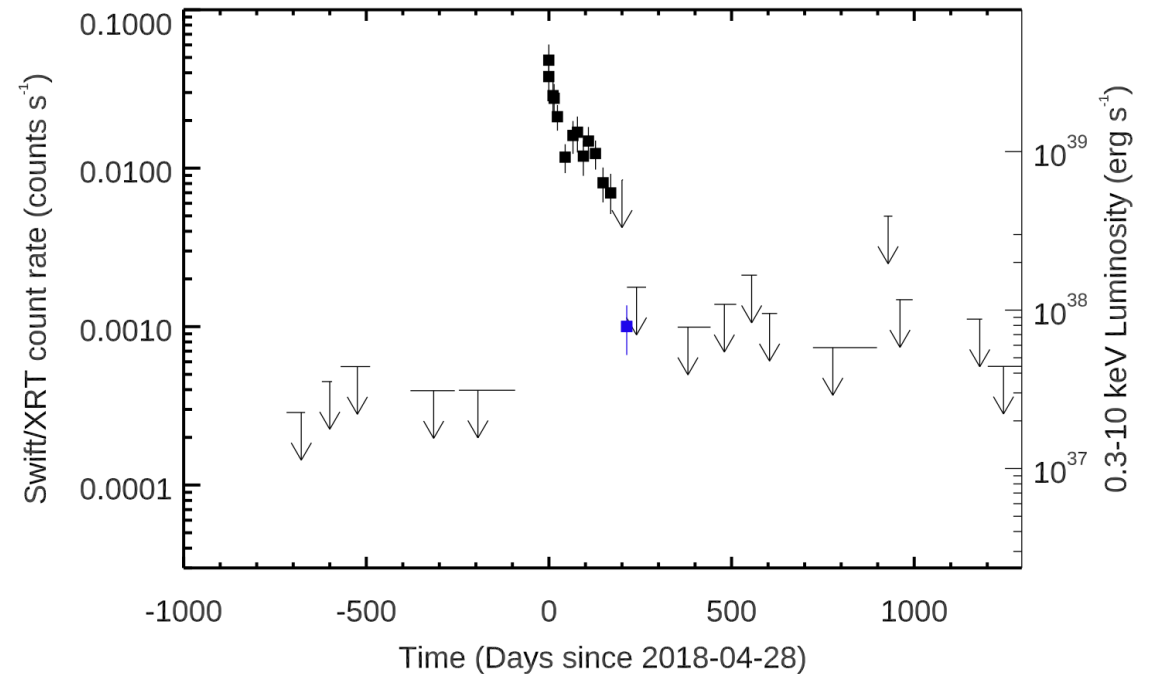
- In  $\sim 3$  years of searches, we found 4 new transient ULXs in nearby galaxies, plus 1 other source in an archival search.
- The targets of the observations were the other ULXs in the galaxies, or the AGN in M81.
- Despite extensive prior X-ray studies of these galaxies, none of these sources had been previously detected.
- No counterpart optical/UV transients identified.

# 2SXPS J235825.7-323609, an X-ray transient in the field of NGC 7793



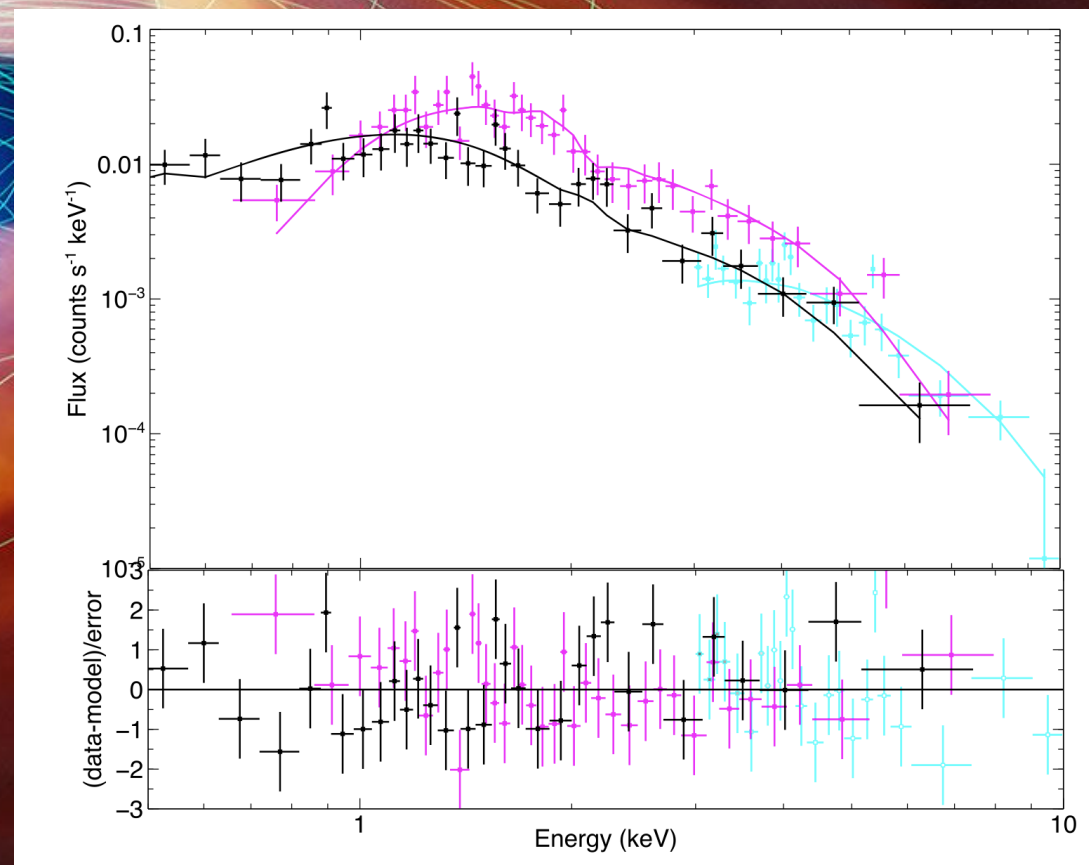
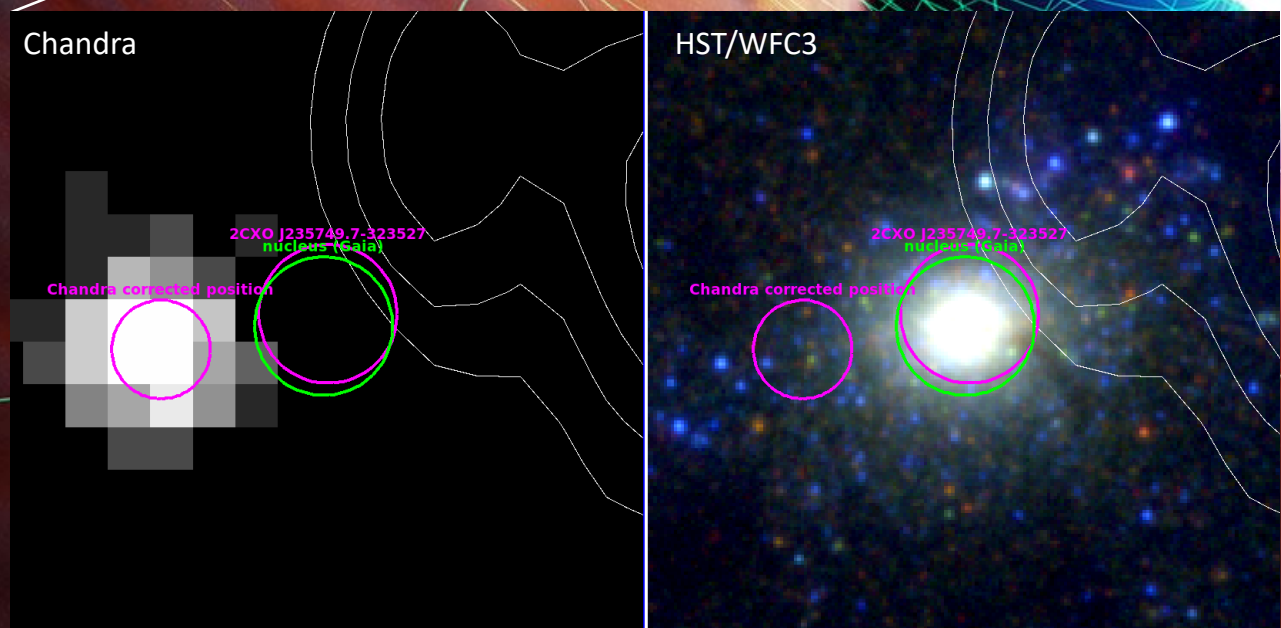
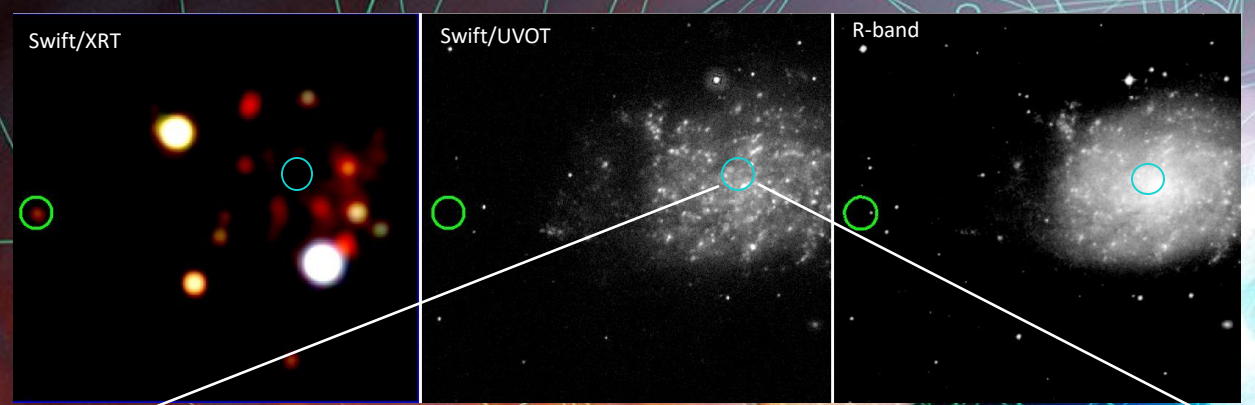
- Source found 7.6 arcmin (8.4 kpc) from the center of NGC 7793.
- No HST sources within X-ray positional error region.
- Outburst lasted for 180 days.

HST ACS F606W



# Swift J235749.9-323526, a second X-Ray transient in NGC 7793

- Source found 2.0 arcsec from the center of NGC 7793.
- Swift, Chandra and NuSTAR spectra are well fitted by a disk-like spectrum (power-law ruled out) with low level of absorption.





# Source type determination

The background of the slide is a dark, reddish-brown image of a galaxy. A bright, glowing blue and white source is visible in the center, emitting a large number of thin, curved lines that represent magnetic field lines or particle trajectories. The lines are more densely packed around the central source and spread out towards the edges of the galaxy.

## Background sources

**( $10^{44}$  erg s $^{-1}$  @  $z \sim 0.1$ )**

- TDEs, AGN flares, gamma-ray burst afterglows are expected to show optical/UV emission (but see EP240408a, O'Connor+25)
- A host galaxy should be seen (also in optical/UV or IR).

## Other type of X-ray transient in the galaxy in question ( $\sim 10^{39}$ erg s $^{-1}$ @ 3-4 Mpc)

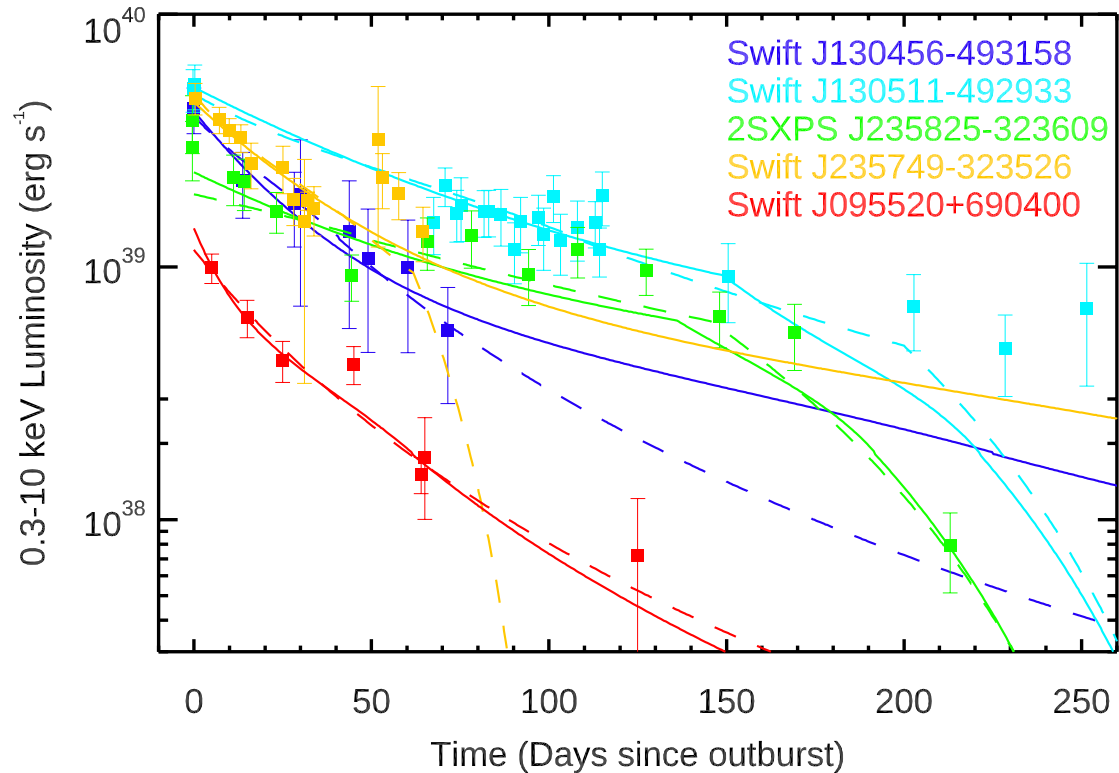
- supernovae also expected to show optical/UV emission.
- supernovae rates much lower.

## Foreground sources

**( $\sim 10^{33}$  erg s $^{-1}$  @ 10 kpc)**

- classical/dwarf novae are also expected to show optical/UV emission.
- Stellar flares are typically much shorter in duration, no Galactic stars in error circle.

# Lightcurve fitting



- Lightcurves modelled with the disk instability model of Hameury & Lasota (2020).
- Model explains that super-Eddington outbursts can be explained by thermal-viscous instabilities in large unstable disks with outer radii greater than  $10^7$  km.

# Conclusions

- We have presented results on five newly found X-ray transients in the fields of nearby galaxies identified in a search of Swift/XRT observations.
- The timescales, fluxes, and lack of bright optical/UV counterparts argue against other types of X-ray transient.
- The rate of transient ULXs for these three galaxies appears high and the number of systems that produce ULX luminosities are likely dominated by transient sources.
- Still much to learn, such as the nature of the accretor (NS/BH?) and accretion mechanisms.