Un optional spettroscopico per PRISMA

Andrea Bianco INAF-Osservatorio Astronomico di Brera

Parco Astronomico La Terra del Sole, October 4, 2024







Diffraction gratings in astronomy

"Let in the sun's light, by a small hole to a darkened house, and at the hole place a feather (the more delicate and white the better for this purpose)...a number of small circles and ovals (appear), one is somewhat white and all the rest severally coloured."

$$R = \frac{\lambda}{\Delta \lambda}$$



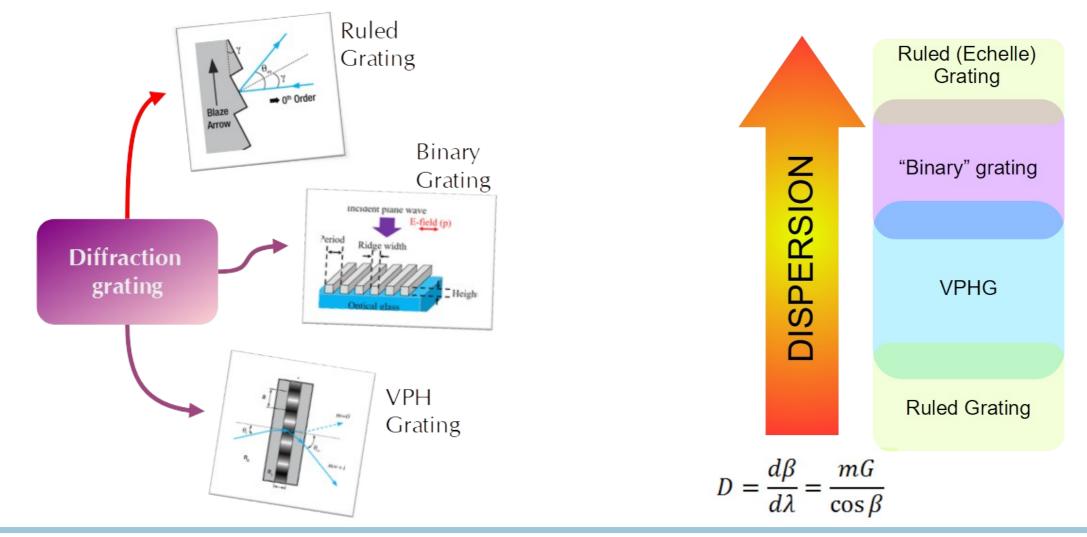


James Gregory (1638-1675)





Diffraction gratings: possibilities







VPHG: principle of work

The diffraction occurs thanks to a periodic modulation of the refractive index in the volume of the material.

Periodic modulation of the refractive index Δn (usually sinusoidal).

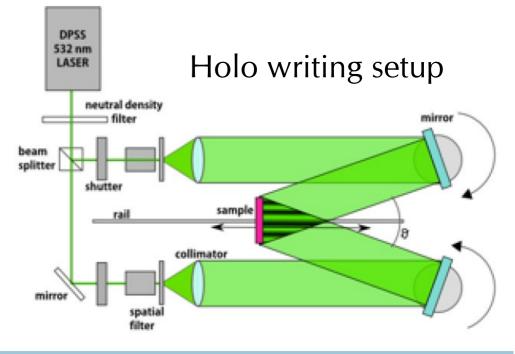
α



VPHG: principle of work

The diffraction occurs thanks to a periodic modulation of the refractive index in the volume of the material.

Periodic modulation of the refractive index Δn (usually sinusoidal).





α

VPHG: main features for astronomy

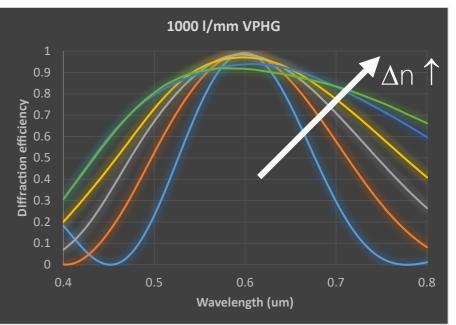
• The peak efficiency can be easily > 90%; **large bandwidth** if we can store a high Δn ;

•The device is **easily customizable** from UV to NIR and multi-functional structures can be obtained (multi order, multiplexed);

- •Especially suitable for **low and medium dispersion** spectrographs;
- •Reliable and repeatable design process;

•Fast production and characterization process suited for multi VPH copies.



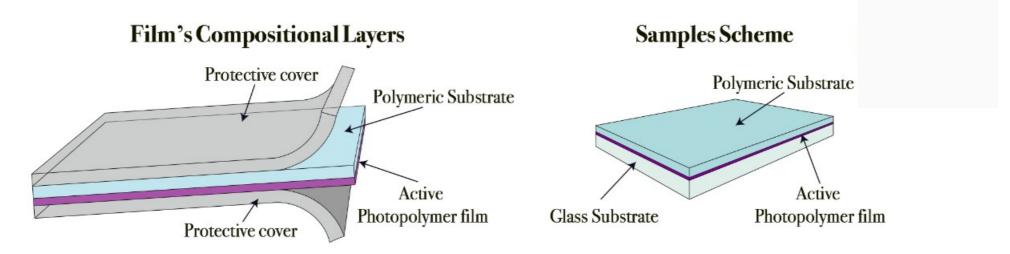






Materials for VPHGs: photopolymers

Photopolymers are available showing a very simple and "easy to use" structure:



•The structure is like a protective layer of smartphone and tablet;

- •They can show different thicknesses and size;
- •The holograms can be removed and attached to different substrates like the protective layers.



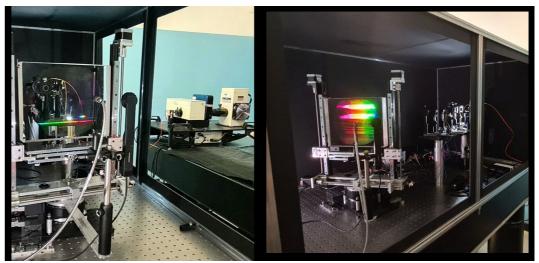


Actual capabilities @ INAF – OABr



Writing capabilities

- RGB laser system;
- Size up to 190 mm x 200 mm;
- Line density: 200 3000 l/mm;
- Spectral range: 0.33 2.5 um;
- VPHGs and GRISMs.

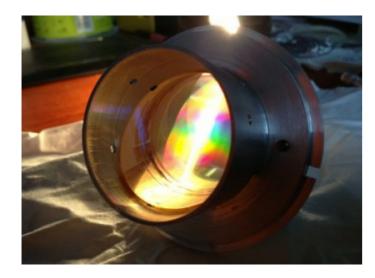


Characterization capabilities

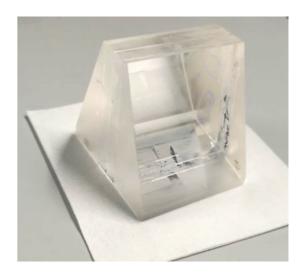
- Fully automatic measurements;
- Diffraction efficiency from UV to NIR;
- Efficiency map (X Y);
- Line density (< 0.5 l/mm) and orientation;
- WFE distortion.



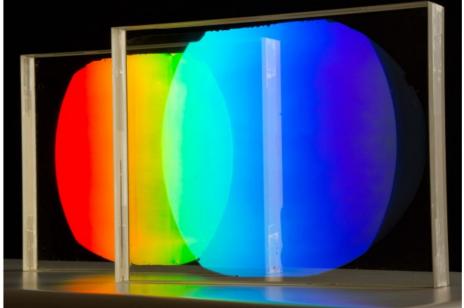










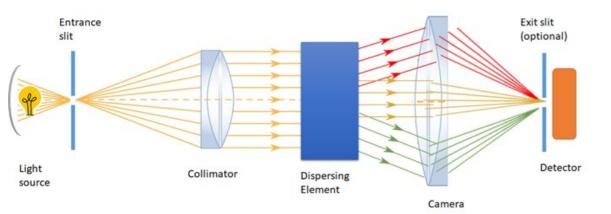






PRISMA Camera with spectral capabilities: issues

• Usually a spectrograph is attached to the main optics:



Not so easy in PRISMA camera, because of the extremely wide field of view and compactness.

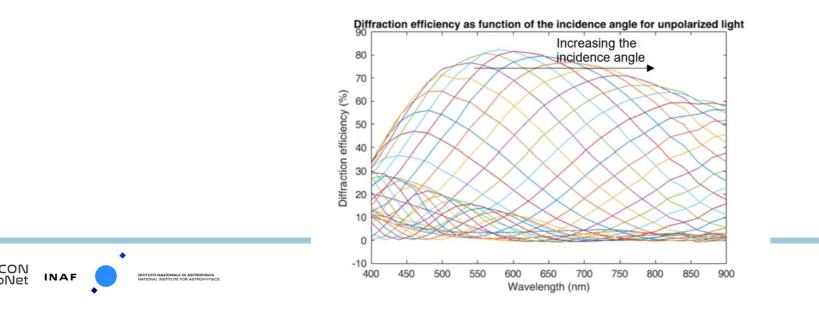
• A different approach is to place a dispersing element in front of the main optics: a collimated beam is coming from the sky and the camera forms the spectrum on the CMOS detector.





Key points of the dispersing element

- What the spectral range of interest?
- The dispersion/resolution provided by the grating cannot be too large;
- The response of the grating depends on the target position/trajectory
- The throughput decreases (reflection losses, diffraction efficiency,...);
- Example of the efficiency curve at t are the following:





Overcoming the direction uncertainty

- The grating is divided in two sections with patterns at 90°: criss-cross;
- The throughput is lower;
- The robustness is much bigger: no matter the object comes, a "good" spectrum is recorded;
- This has been done for AMOS camera (Slovak Video Meteor Network).

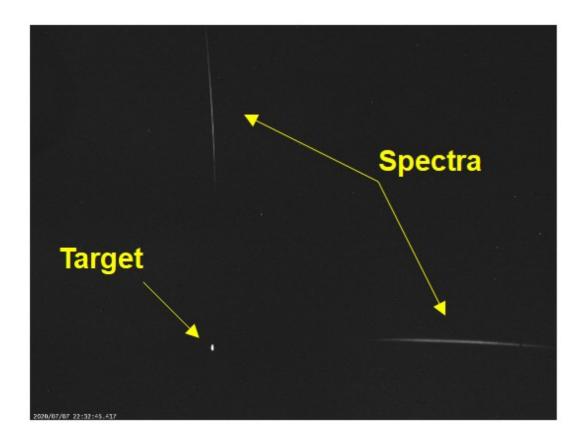


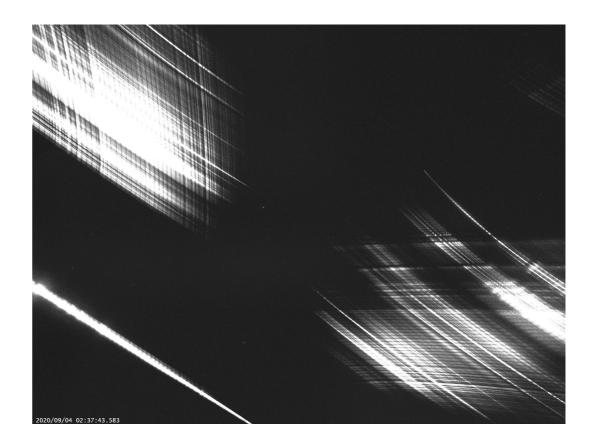
Property	Value
Shape and size	Round, 40 mm in diameter
Field of view	60°
Structure	2 halves, same size
Line density	800 l/mm
Spectral range	400 – 900 nm
Peak diffraction efficiency	80%





Overcoming the direction uncertainty

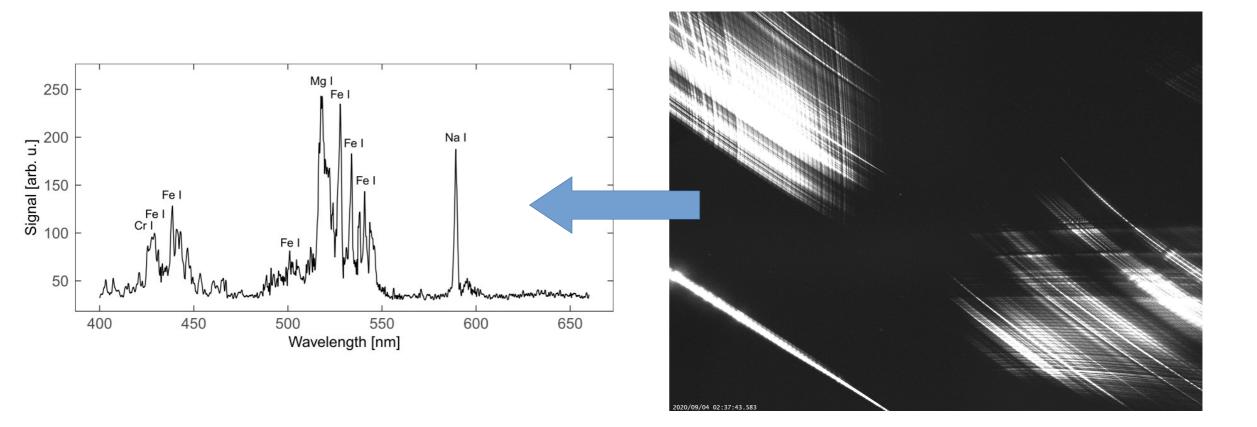








Overcoming the direction uncertainty

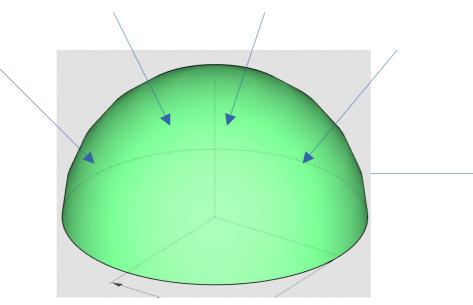






Other possibilities

- A flat grating can be a limit if the FoV is very wide;
- Curved "VPHGs": apply the grating on the hemisphere;
- Necessary to model the system and the response.









Conclusions

- VPHGs are useful dispersing elements for astronomical spectrographs;
- We can produce them in a wide range of parameters exploiting a flexible proces;
- They are suitable for the spectroscopic upgrade of PRISMA camera;
- Know-how developed with the AMOS cameras;
- Some ideas vs some issues.













The gOlem galaxy...25 years of technologies for astronomy



