

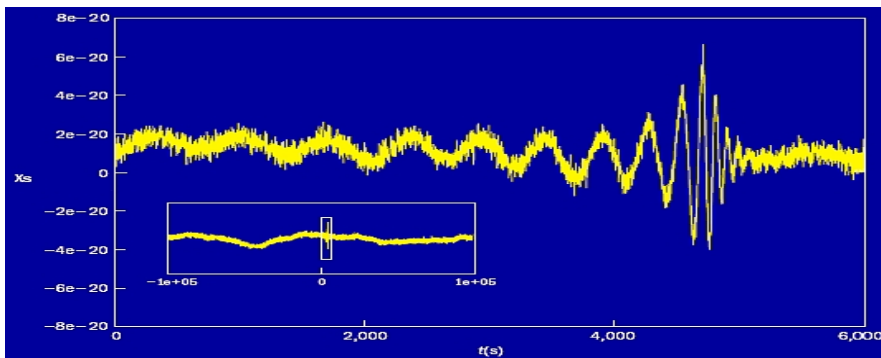
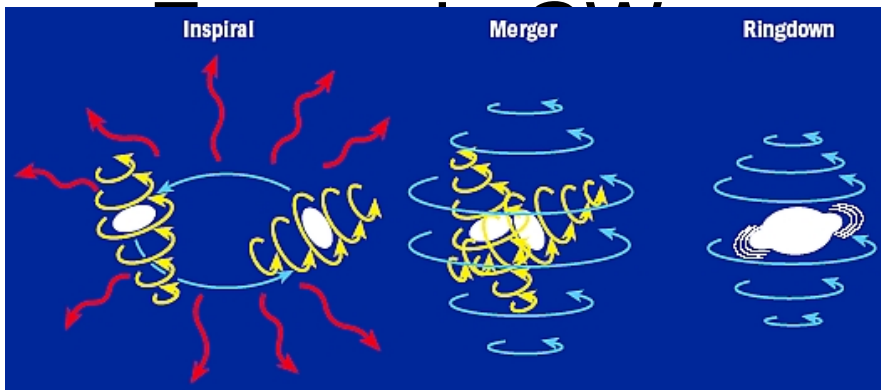
**Science case:
Search of electromagnetic counterparts of
Gravitational Waves**

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on behalf of the INAF GW Collaboration***

Gravitational waves: sources

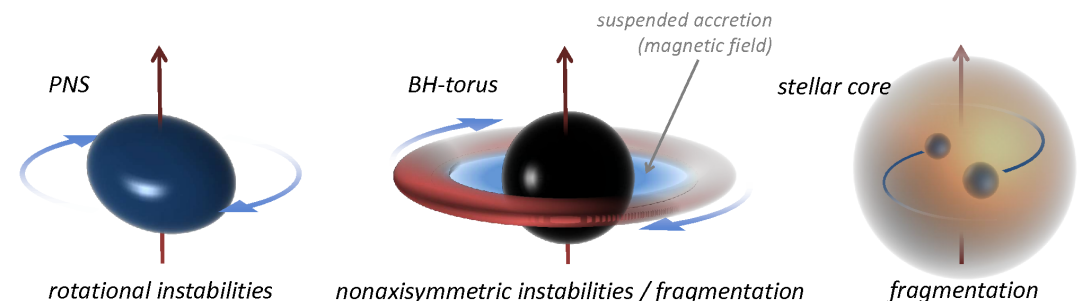
Coalescence of compact objects (CBC)

- Known waveform

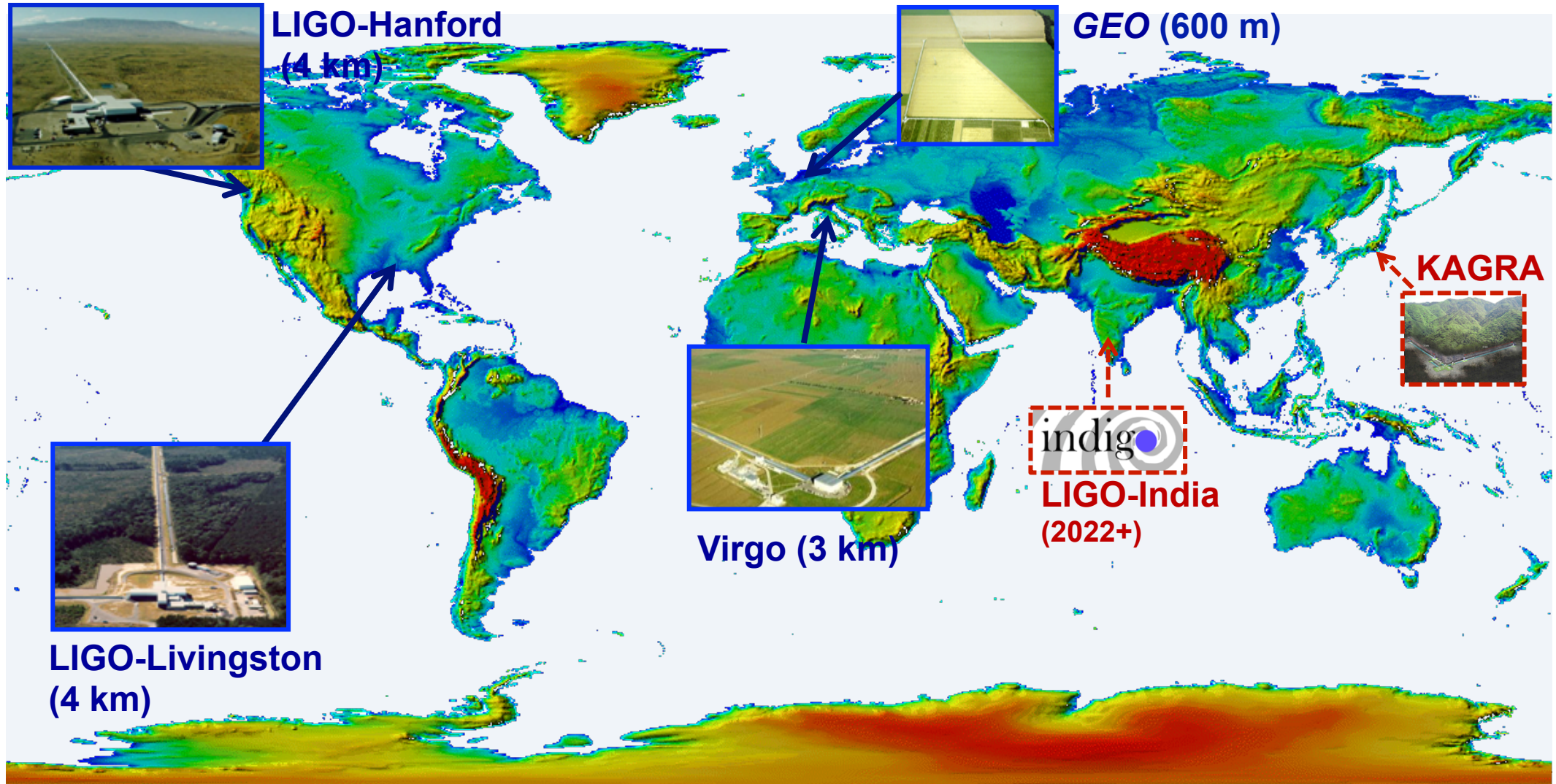


Core-collapse of massive stars

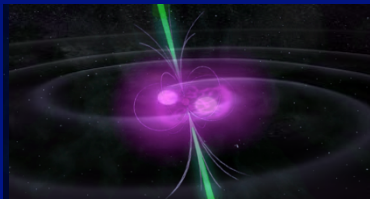
- Waveform and energy emitted in GW uncertain:
- Likely $10^{-8} M_{\odot} c^2$
- Optimistic $10^{-4} M_{\odot} c^2$



Gravitational Waves: Detection



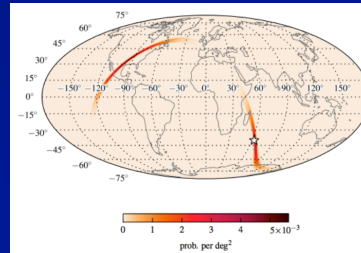
LIGO and Virgo currently under upgrade
They will observe the sky (10-1000 Hz) as a single network
aiming at the first direct detection of GWs



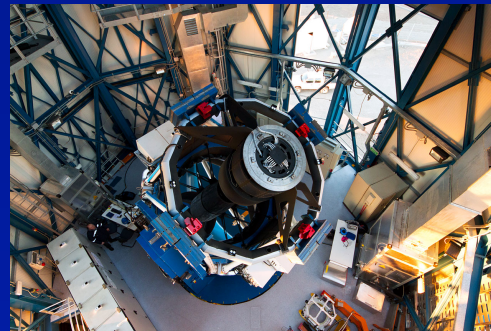
Search of electromagnetic counterparts of Gravitational Waves



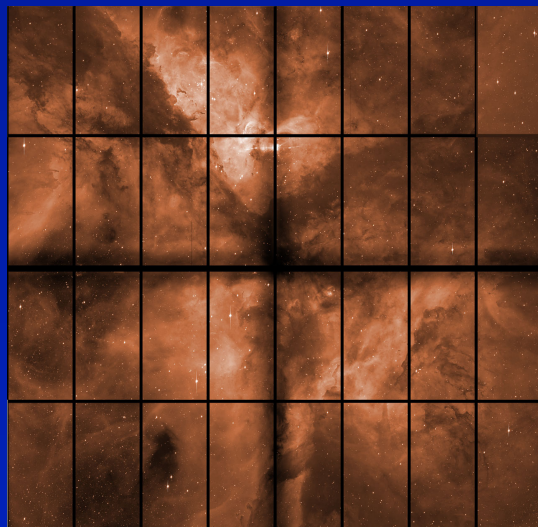
LIGO/Virgo
Interferometers →
error boxes on sky:
 $\geq 200 \text{ deg}^2$



Problem:
→ Hundreds of astronomical
images have to be analyzed to
search transient objects



VST – OMEGA-CAM
Field of View 1 deg^2
Resolution 0.21 arcsec/pixel
1 single image ~ 1.3 Gb



Computing challenge:
Calibrate and analyzes hundreds of wide field astronomical
images and
search for transients objects in few hours (constraints due to
the fading time of the astronomical targets)

Of course, further observational steps (spectroscopy & follow-up) are
needed to characterize the discovered transients and then to validate
the final identification of the counterpart of GW signals

Science case: Search of electromagnetic counterparts of Gravitational Waves

A real example:

After LIGO/Virgo alert, VST images are obtained by INAF GW collaboration

- Number of images : ≥ 200 images
($\sim 18000 \times 18000$ pixels each to map 1 deg^2)
- Image size : $\sim 1.3 \text{ GB}$ / image
- Calibration time: ~ 6.5 hours for a set of ~ 200 astronomical images
(VST center : computer spec. : 256 cpu cores / $\sim 900 \text{ GB}$ RAM / $\sim 300 \text{ TB}$ storage)



Bottleneck :

- each single images has to be analyzed with specific software to identify transients
- human intervention is still required
- small workstations are presently used for this step

Case A (*reference images not available*)

Analysis time of one VST image: **~ 2 hours / image** **1800000 detected objects**

(computer spec.: 12 cpu cores / 128 GB RAM / $\sim 12 \text{ TB}$ storage)

+ time needed to identify transients by using available catalogs in the observed field

Case B (*reference images available*)

Image subtraction (well established technique to search transients when reference images are available)

~ 20 min / image (computer spec.: 32 cores AMD 6328 / 128 GB RAM)

We are currently working to improve software performances and to obtain faster (and large RAM) computing facilities