

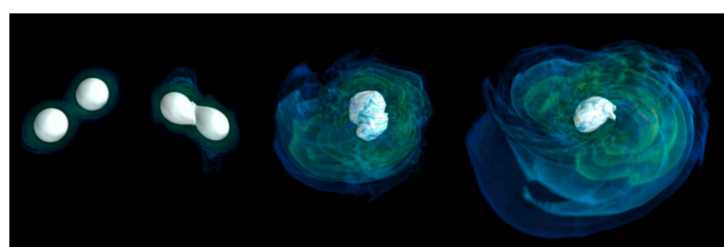
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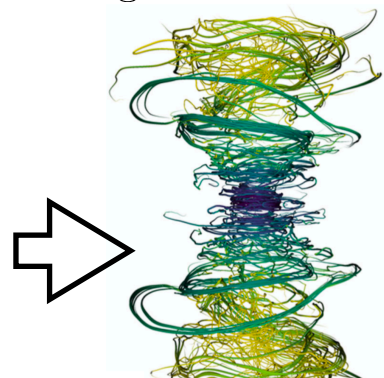
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INTRODUCTION

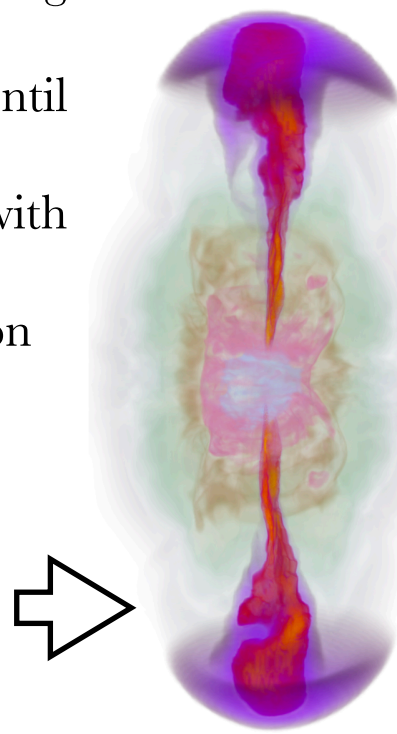
1. Binary neutron star mergers can form a compact object (e.g. accreting black hole) able to launch a **relativistic jet**
2. The jet breaking out of post-merger environment keeps evolving until reaches the **ballistic regime** (saturation of velocity and structure)
3. Powerful jets can produce a gamma-ray burst. Later interaction with the interstellar medium leads to the **afterglow signal**
4. The jet evolution details are imprinted in the electromagnetic emission



Binary neutron star merger



Jet formation

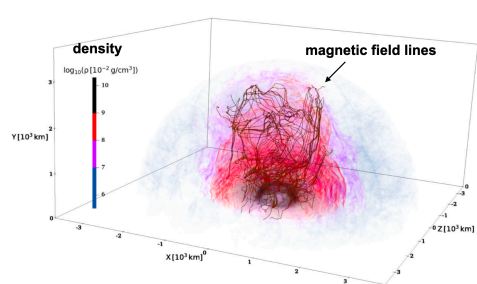


Jet break-out

SCIENTIFIC GOAL

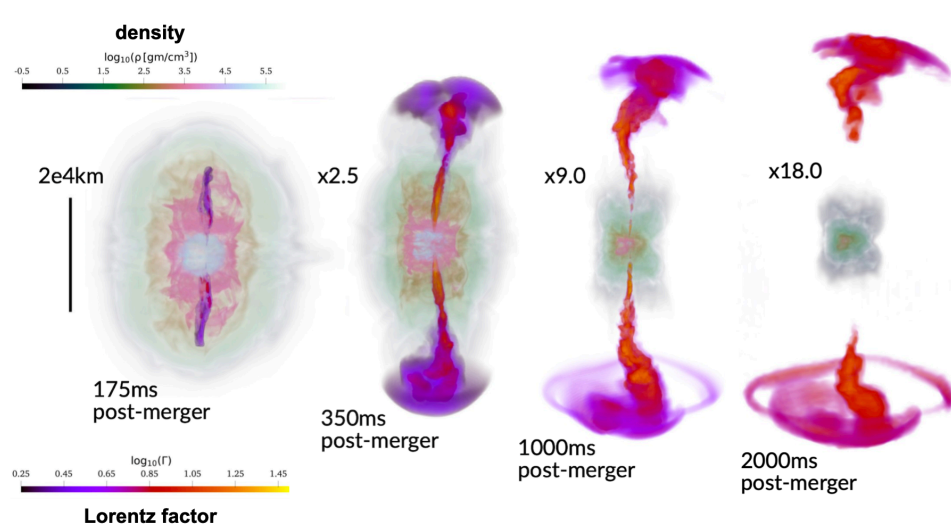
- Combine simulations of the merger process, jet break-out and jet propagation up to a quasi-ballistic regime in the first consistent end-to-end description
- Connect the afterglow emission with the progenitor system and jet injection parameters

JET INJECTION AND BREAKOUT



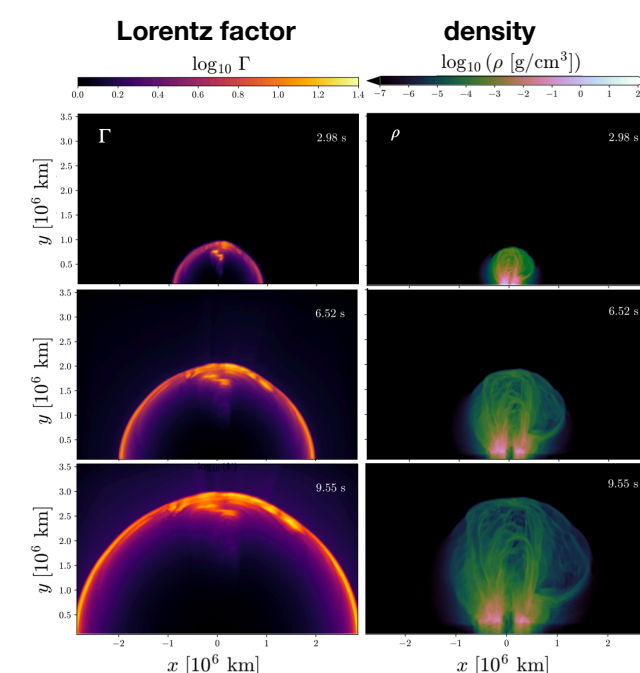
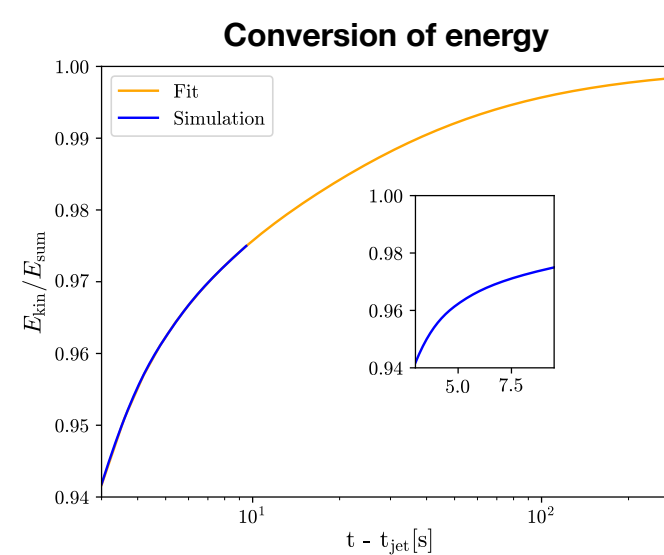
A magnetized jet is launched in a realistic post-merger environment, directly imported from the binary neutron star merger simulation of [1].

The jet is evolved in a spherical grid with the PLUTO code ([2]) for 3 s, with the details of break-out dynamics imprinted in the final angular structure and energetics [3,4].

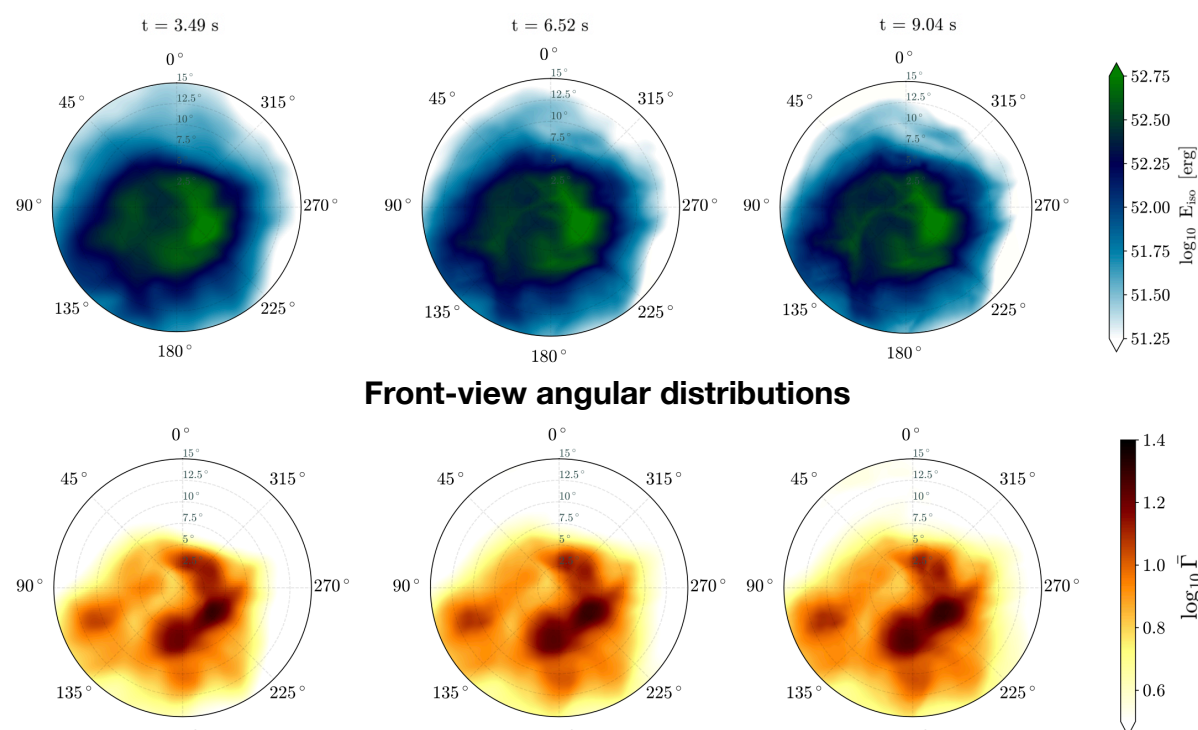


LATE EVOLUTION AND BALLISTIC PHASE

- **Aim:** Follow the subsequent jet evolution without loss of resolution up to a quasi-ballistic regime
- **Method:** Remap the output of the early evolution on a Cartesian grid with uniform cells



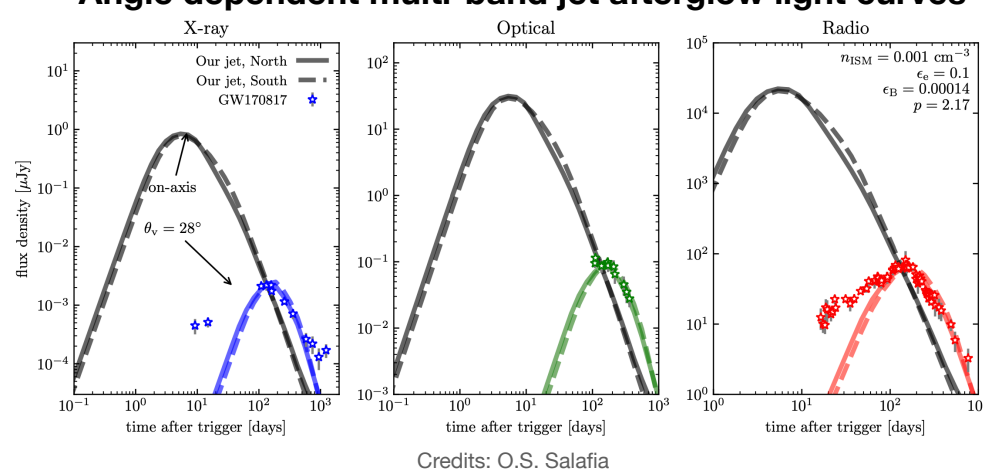
At the end of the simulation (~ 10 s), $\sim 98\%$ of energy is converted into kinetic form and the angular structure is no longer changing. Strong deviations from axisymmetry are observed.



OUTLOOK

- Systematic application to a large set of sGRB jet initial parameters, linking the final jet structure with injection and break-out conditions
- Use the outputs in semi-analytic afterglow models to produce lightcurves and compare with observations (e.g. GRB 170817A)

Angle dependent multi-band afterglow light curves



Credits: O.S. Salafia

REFERENCES

- [1] Ciolfi (2020), MNRAS Lett. 495, L66
- [2] Mignone et al. (2007), ApJS 170, 228M
- [3] Pavan et al. (2021), MNRAS 506, 3483
- [4] Pavan et al. (2023), MNRAS 524, 260