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Low-mass carbon-enhanced metal-poor thermally pulsating AGB star HD112869

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"s, i & r Element Nucleosynthesis (sirEN)" Conference

08.06.2025, Giulianova

Introduction/Relevance

- Observations of metal-poor AGB stars are lacking
- The chemical abundances of carbon-enhanced metal-poor (CEMP) stars reflect the origin of CNO and neutron capture elements in the early Galaxy
- CEMP stars have a wide variety of elemental abundance patterns
- The majority of CEMP stars have enhanced s-process elements

CEMP star HD112869



- Bright ($M_{\text{bol}} = -3.35$, $V = -8.91$), high-latitude ($b = +79^\circ$) extremely metal poor star classified as R type carbon star with enhanced CH bands;
- HD112869 has an extremely low intensity of molecular lines with isotopic ^{13}C
 - A high $^{12}\text{C}/^{13}\text{C}$ ratio is observed rarely among the CEMP stars
 - The majority of CEMP stars have enhanced s-process elements

Observations of HD112869

- Radial velocity monitoring (> 2500 days) - CORAVEL
- Simultaneous monitoring of the brightness (465 days)
- Parallel high-resolution spectroscopy (3600 – 9300 Å)

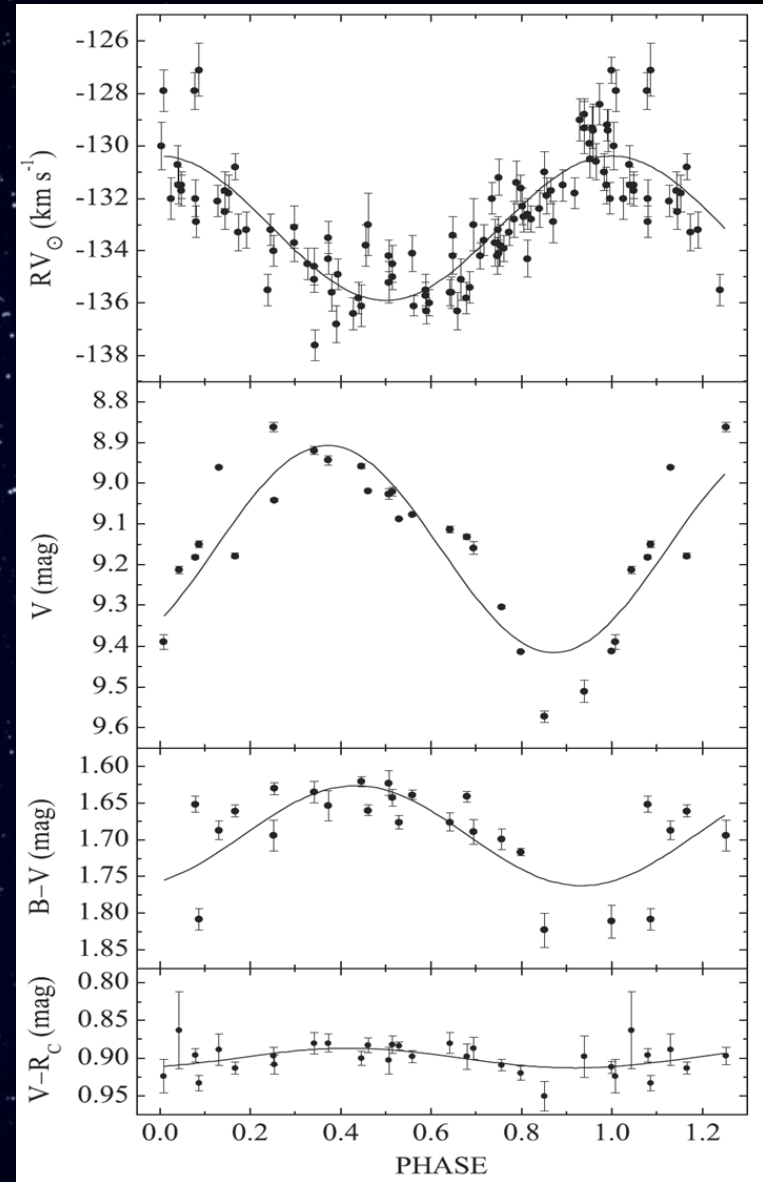
CORAVEL
spectrometer
installed on the 1.65-
m telescope at the
Molėtai Observatory
(Lithuania)



Coudé échelle spectrometer
MAESTRO on the 2 m telescope at
the Observatory on the Terskol
Peak

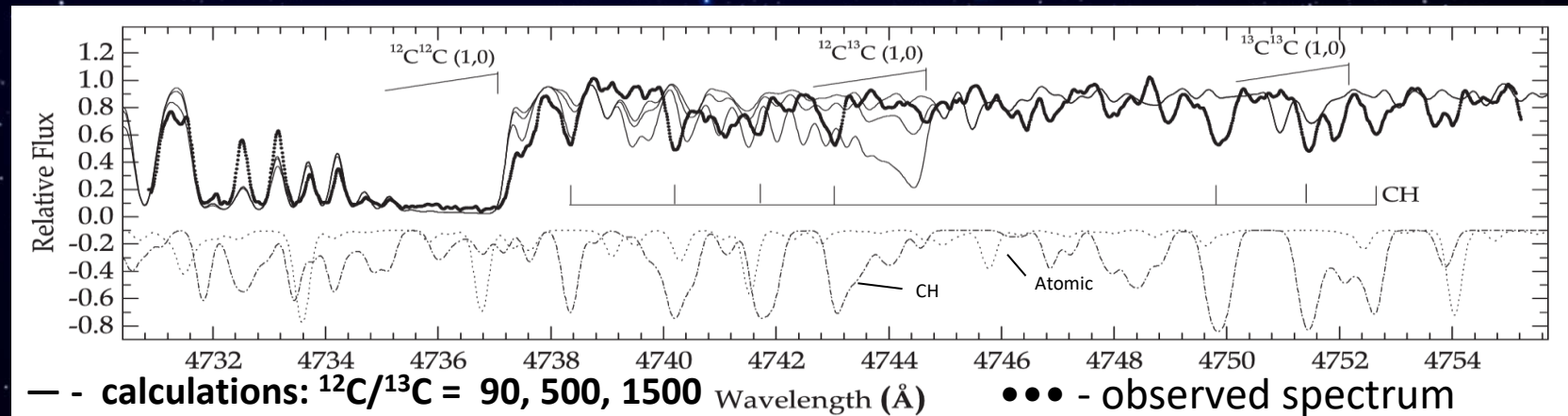
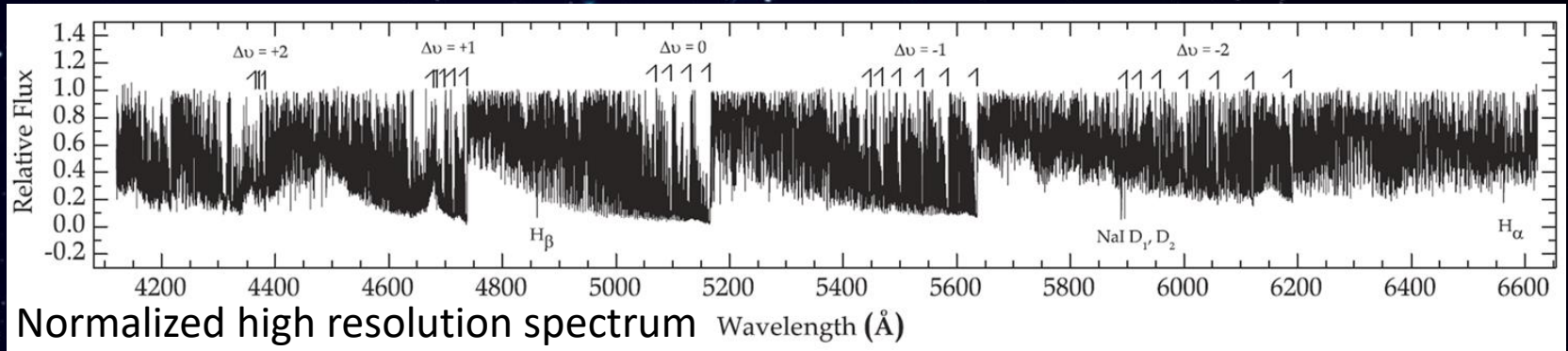
Conclusions

- Radial velocity ~ 114.9 days
 - Peak-to-peak amplitude 10 km s^{-1}
- Light and colour variations are shifted in phase relative to the velocity curve
 - typical for pulsating evolved stars



CORAVEL radial velocities (top panel) and photometry

- The lower limit of isotopic ratio was found to be extremely high, $^{12}\text{C}/^{13}\text{C} \geq 1500$



Observed spectrum of HD112869, along with synthesized C_2 spectra for the final carbon abundance: $\log \epsilon(\text{C}) = 8.3$ dex, and atmospheric model parameters: $T_{\text{eff}} = 3900$ K, $\log g = 0.4$ (cgs), $\zeta_t = 4.0 \text{ km s}^{-1}$

- The abundance of s-process elements Y, Sr and Ba is not enhanced:
 - $[Y/Fe II] \leq 0.0 \text{ dex}$; $[Sr/Fe II] \leq -1.7 \text{ dex}$; $[Ba/Fe II] \leq -0.3 \text{ dex}$
- But Nd, La and Sm is overabundant:
 - $[Nd/Fe II] \leq +0.7 \text{ dex}$; $[La/Fe II] \leq +0.8 \text{ dex}$; $[Sm/Fe II] \leq +0.5 \text{ dex}$
- A strong enhancement of r-process was denied for atmosphere of HD 112869, $[Eu/Fe] \leq 0.8 \text{ dex}$
- According the evolutionary models, HD112869 is a single metal-poor low-mass thermally pulsating AGB star