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s, i & r Element Nucleosynthesis

Lessons from Dwarf Galaxies

Ása Skúladóttir
University of Florence

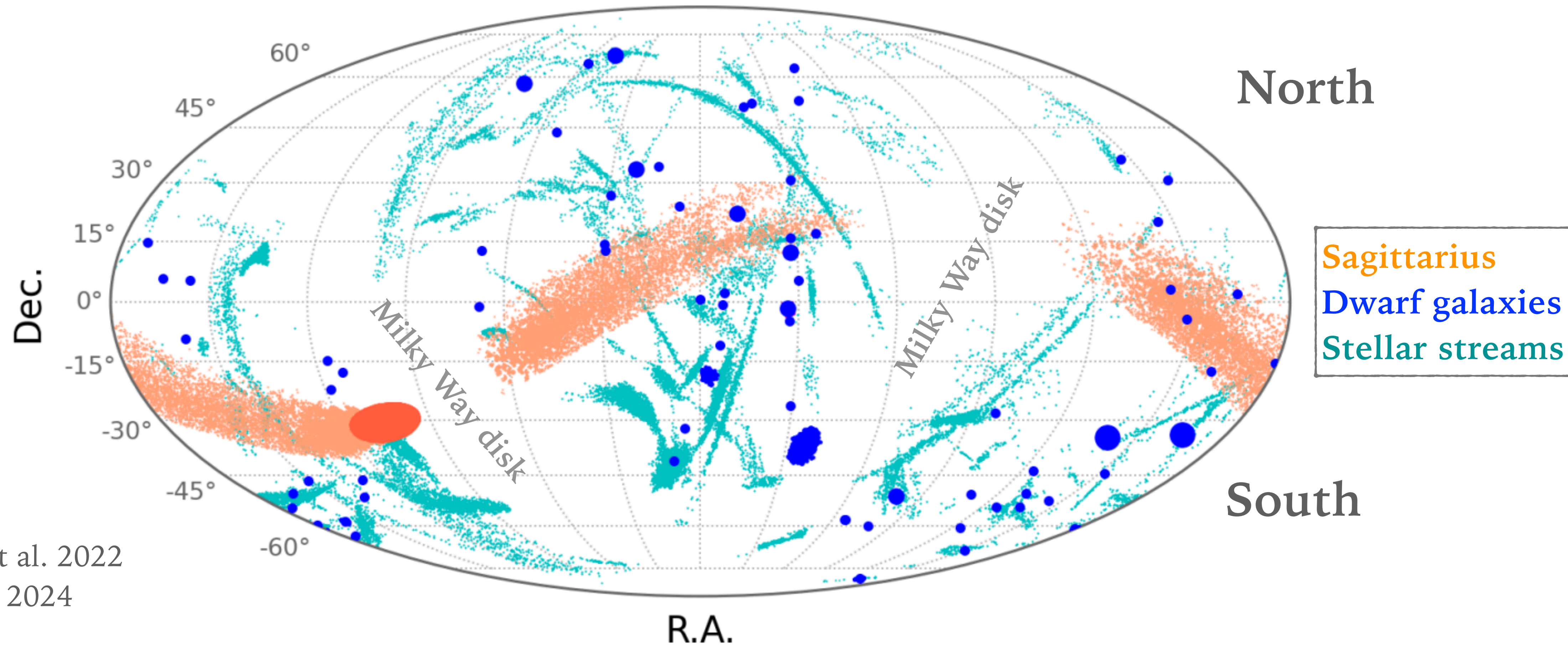


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sirEN conference - Giulianova, Italy - 13th of June 2025

DWARF GALAXIES AROUND THE MILKY WAY

- The **Milky Way** has many satellite **dwarf galaxies**, and **stellar streams**.



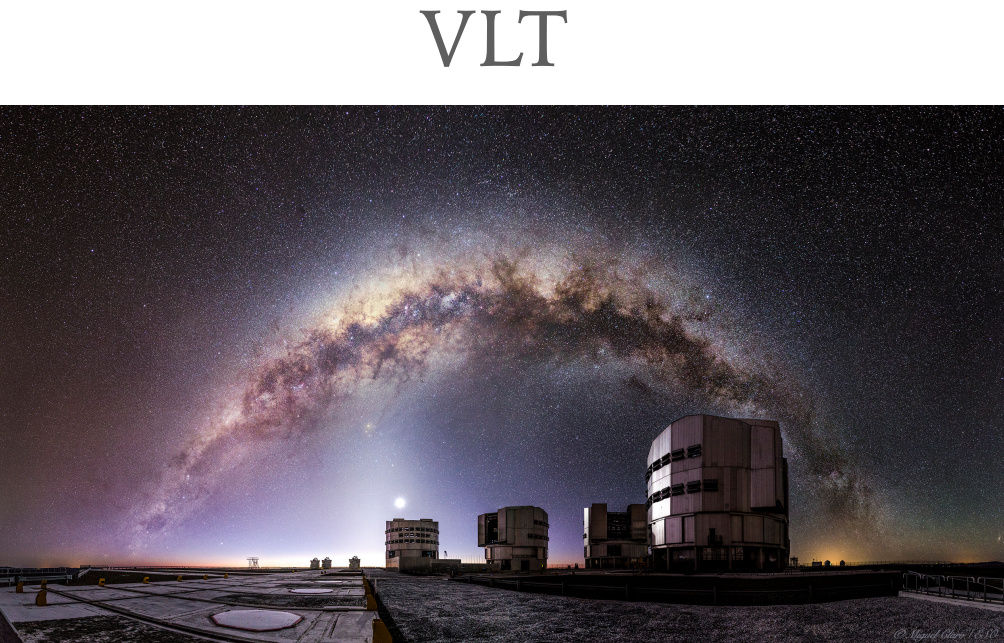
Data from:
Battaglia et al. 2022
Ibata et al. 2024

DWARF GALAXIES AROUND THE MILKY WAY

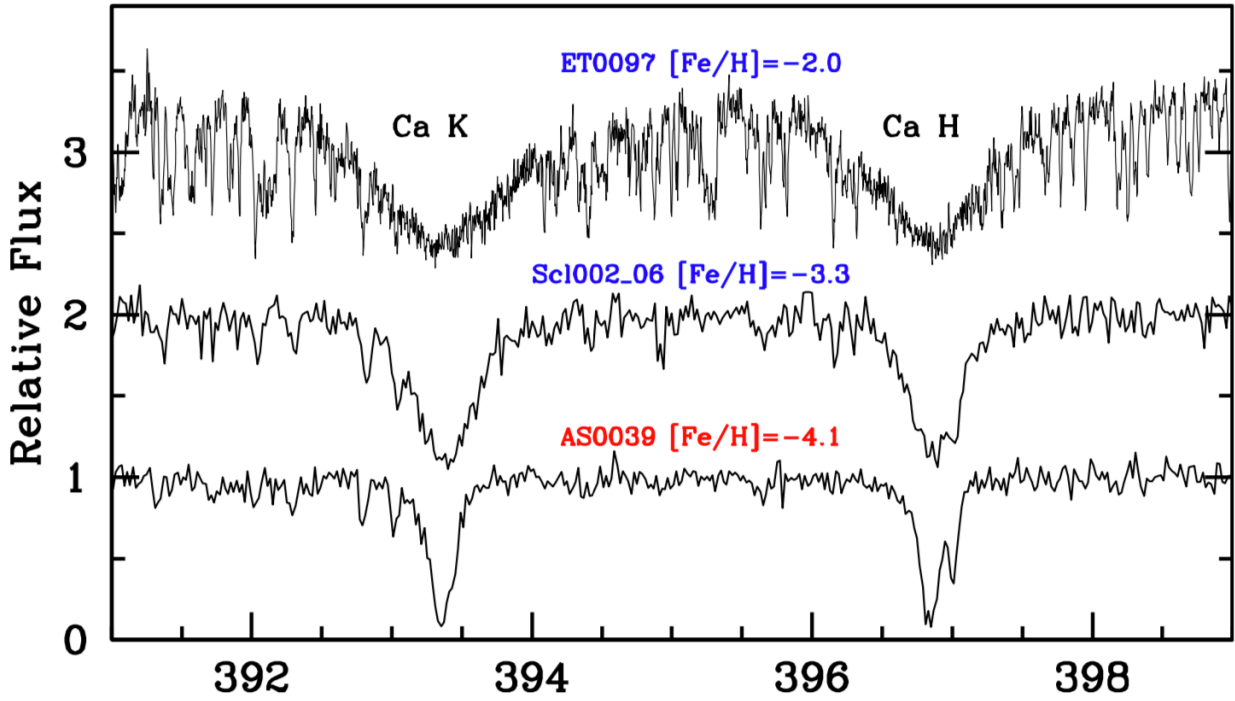


Gaia

Positions
Proper motions
Photometry

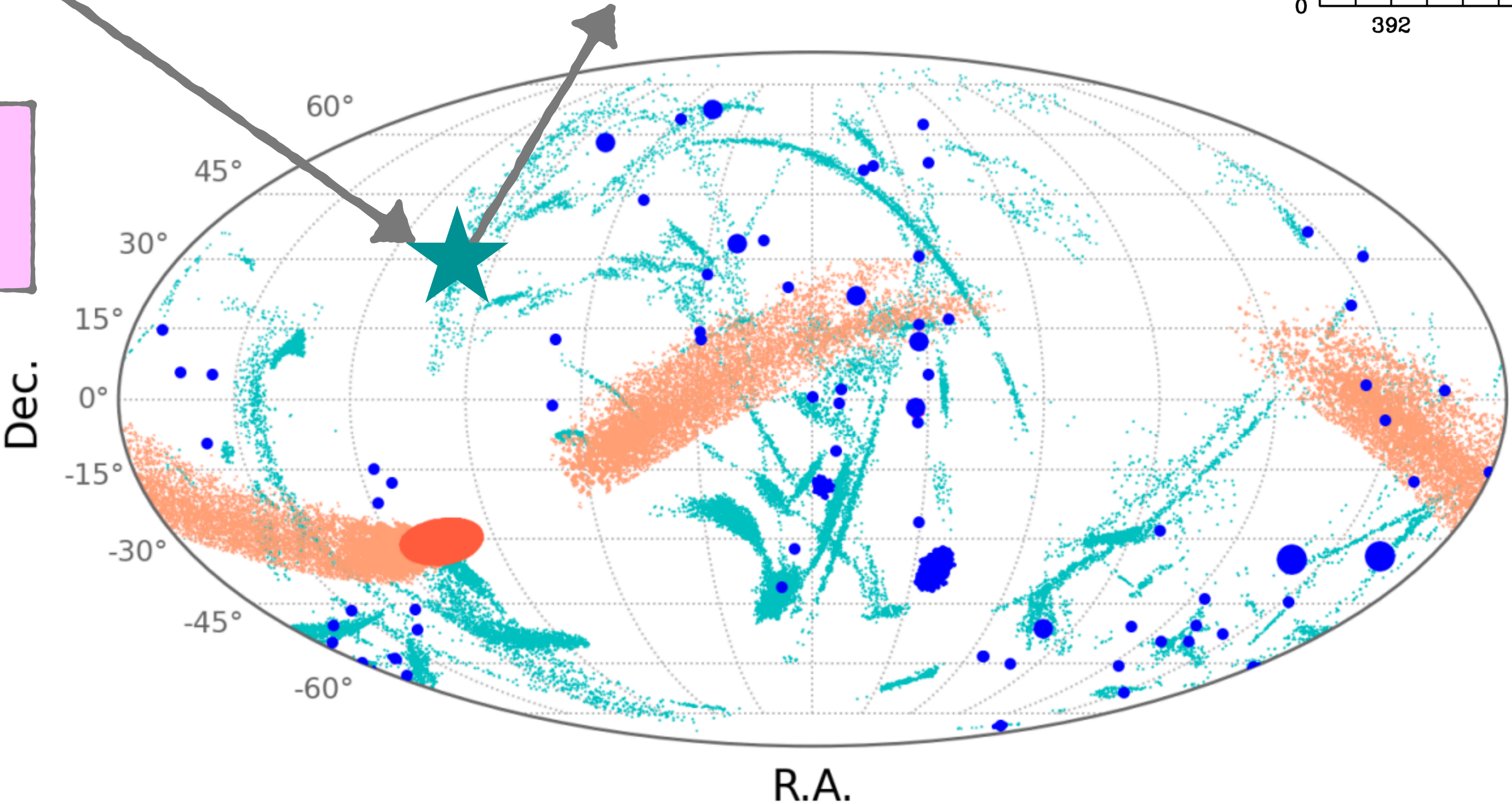


VLT



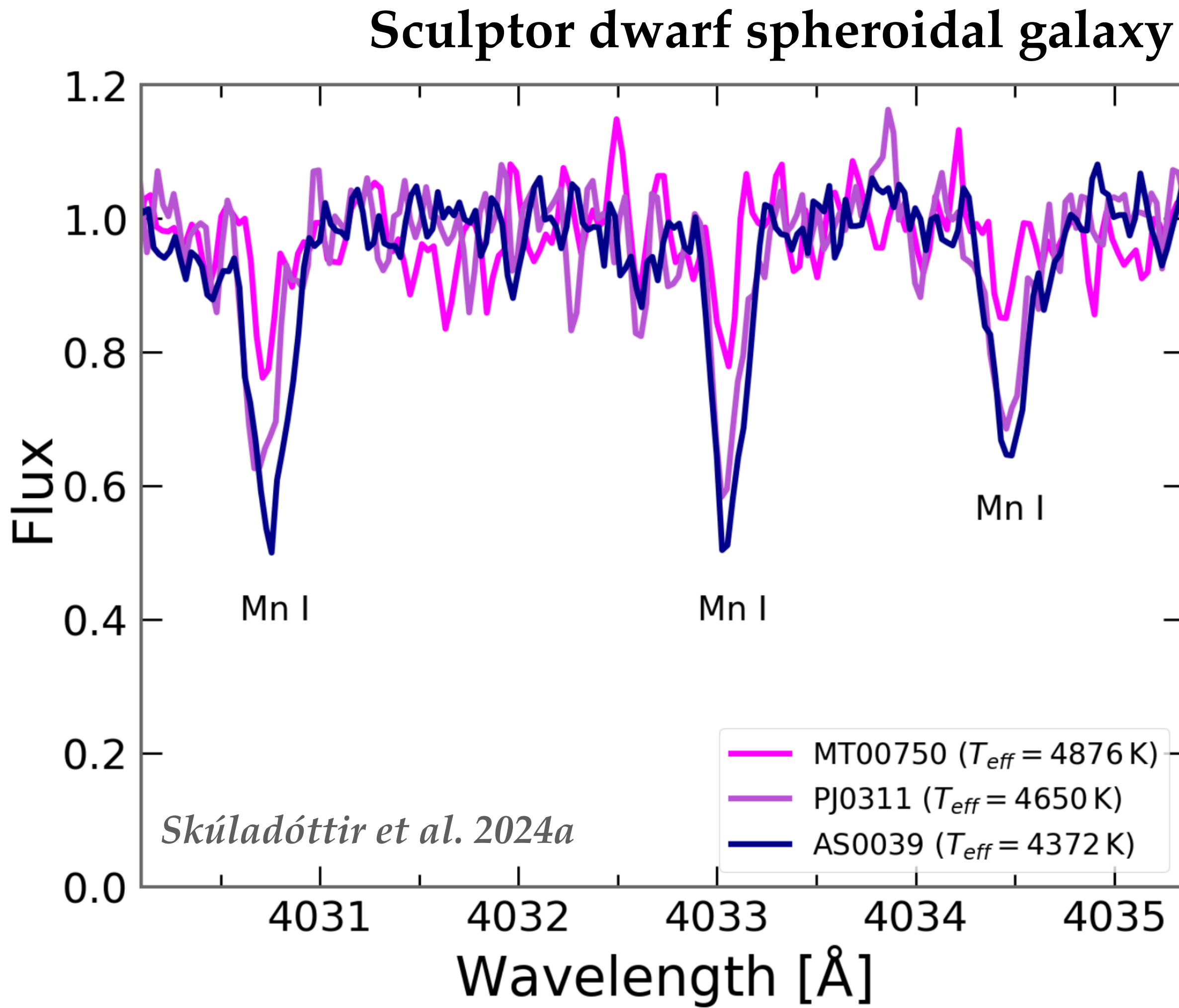
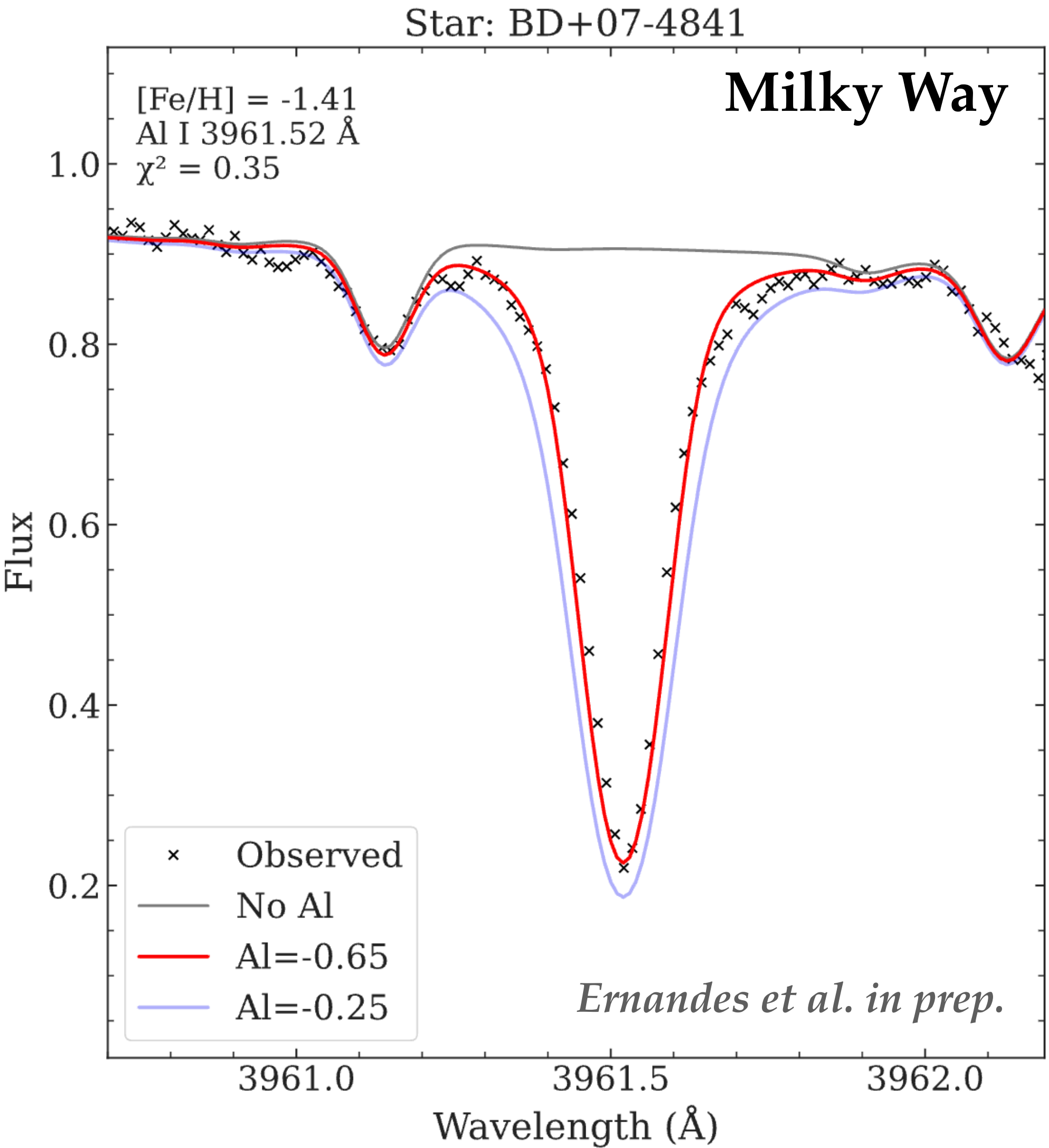
Spectra

Chemical abundances
Line-of-sight velocities
Stellar ages



Data from:
Battaglia et al. 2022
Ibata et al. 2024

CHALLENGES OF GETTING HIGH-QUALITY SPECTRA





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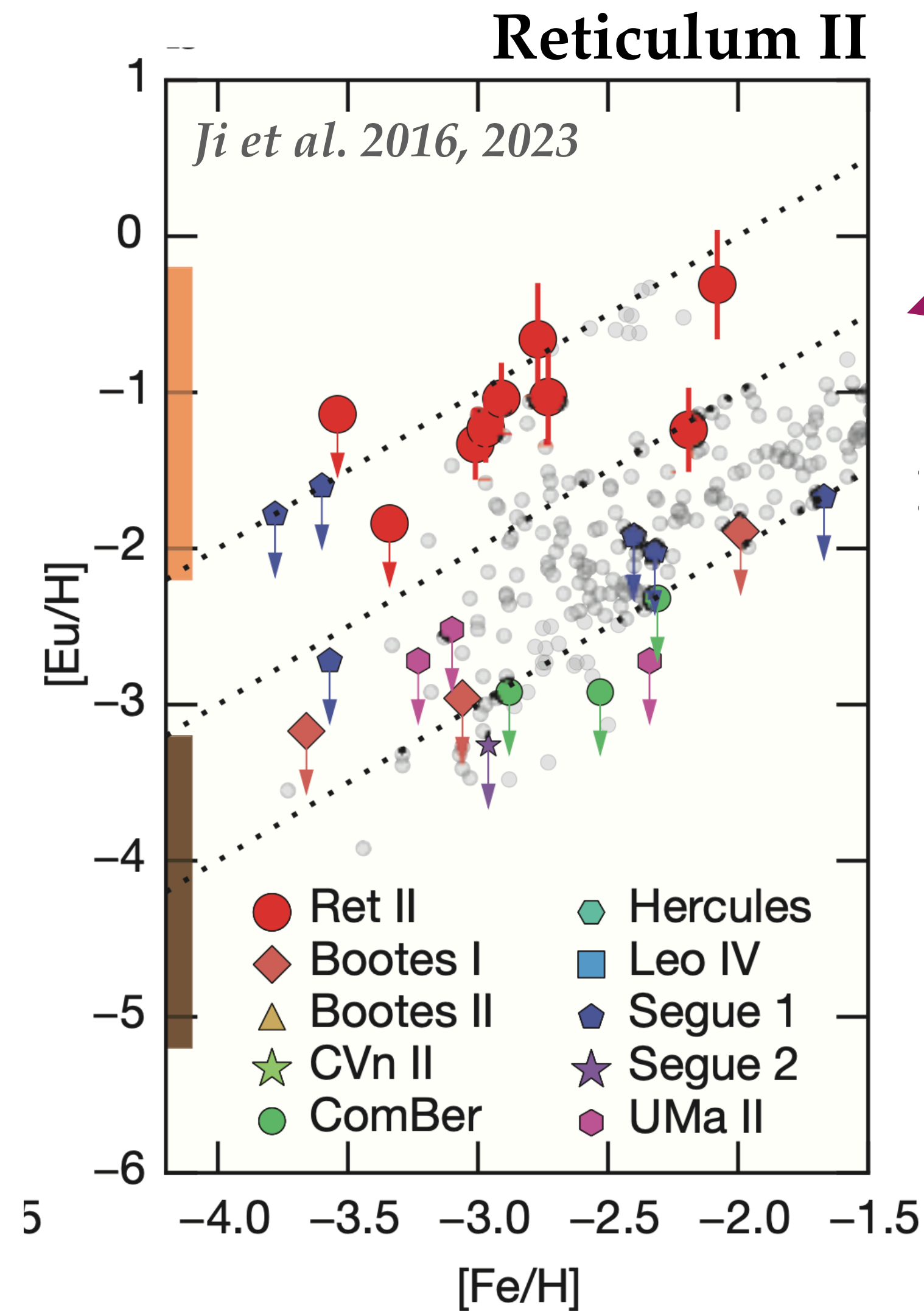
Tracing individual events



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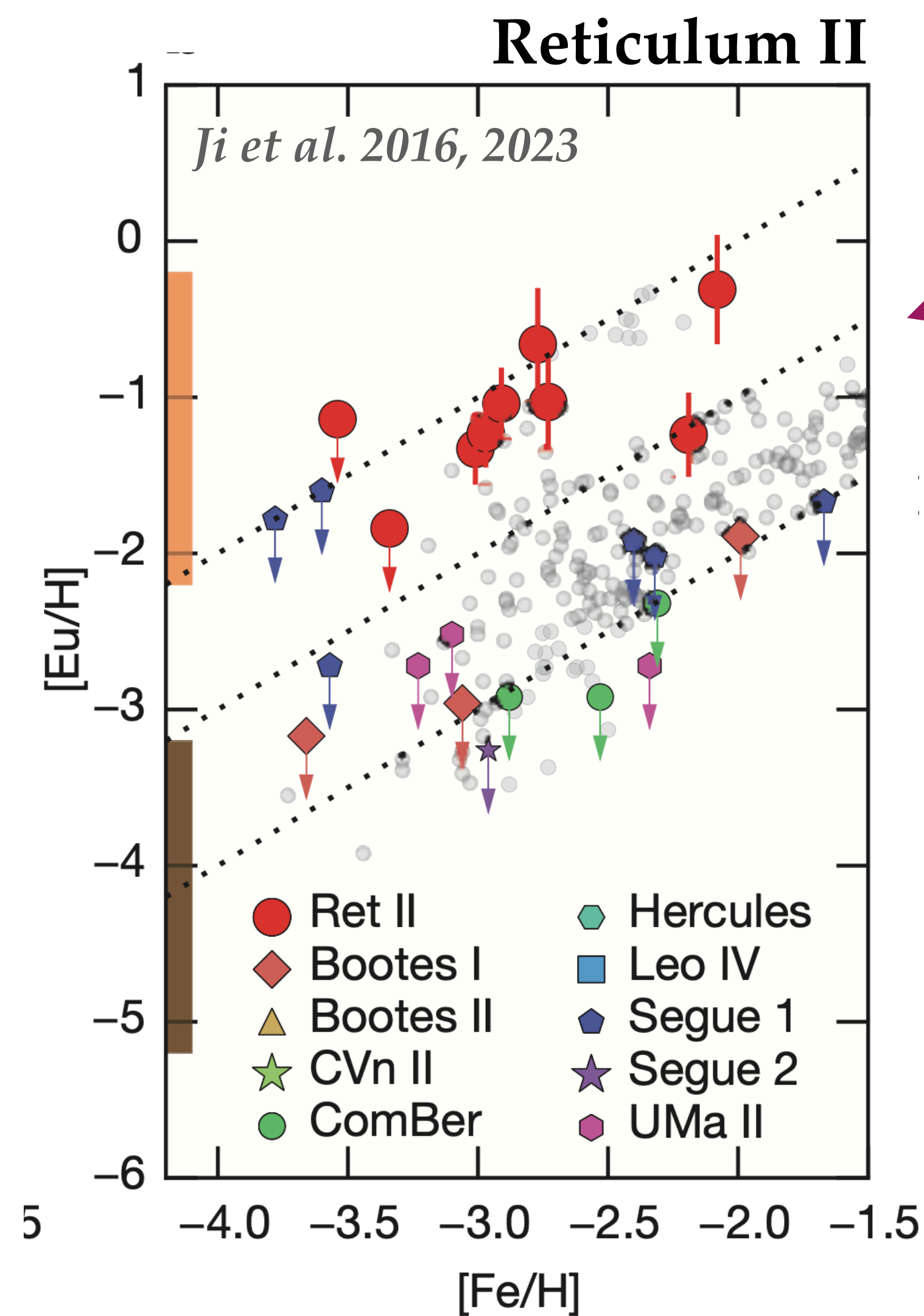
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R-PROCESS RICH STARS



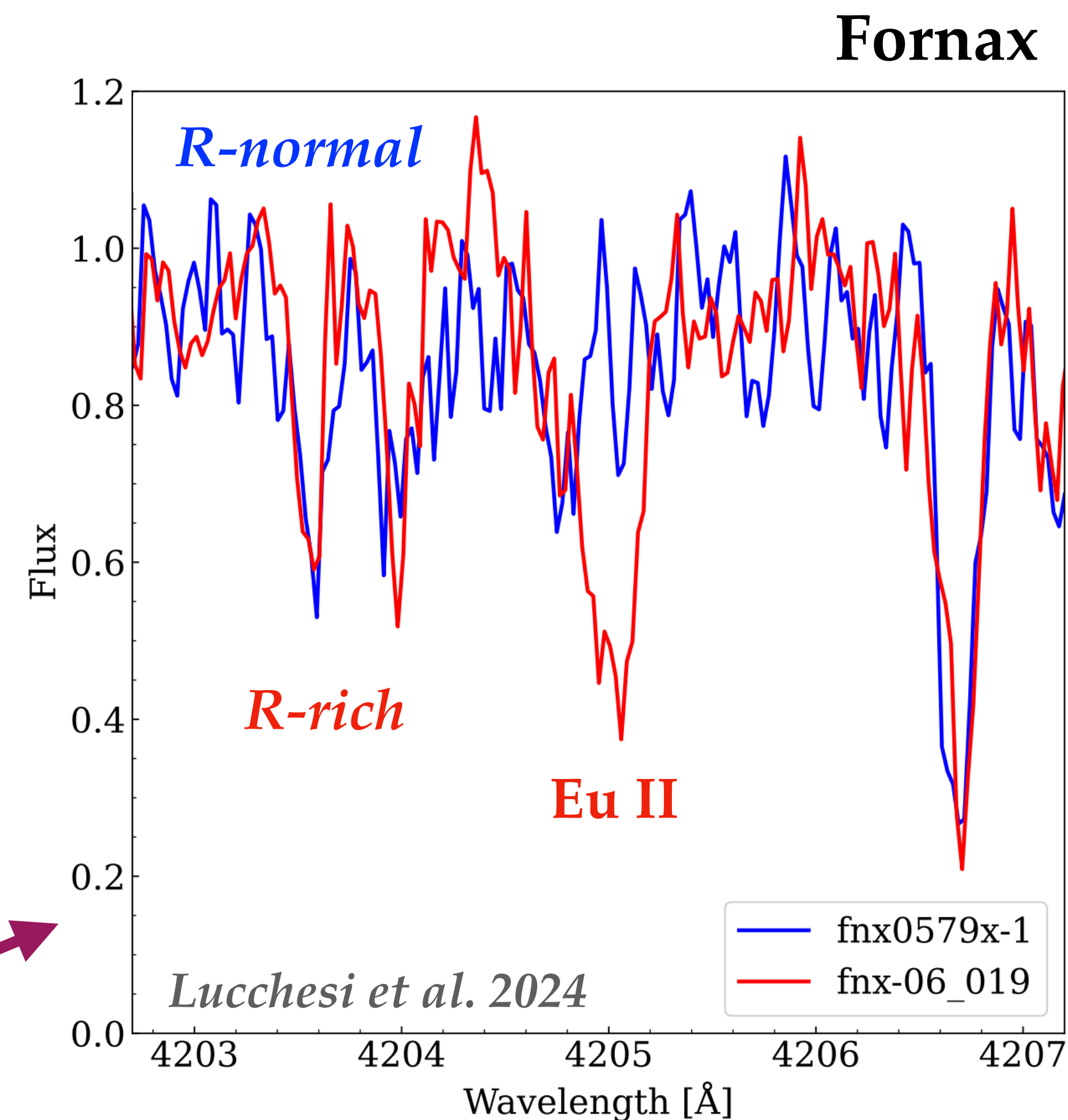
Ultra faint dwarf galaxy
(UFD) enriched in r-
process!

R-PROCESS RICH STARS

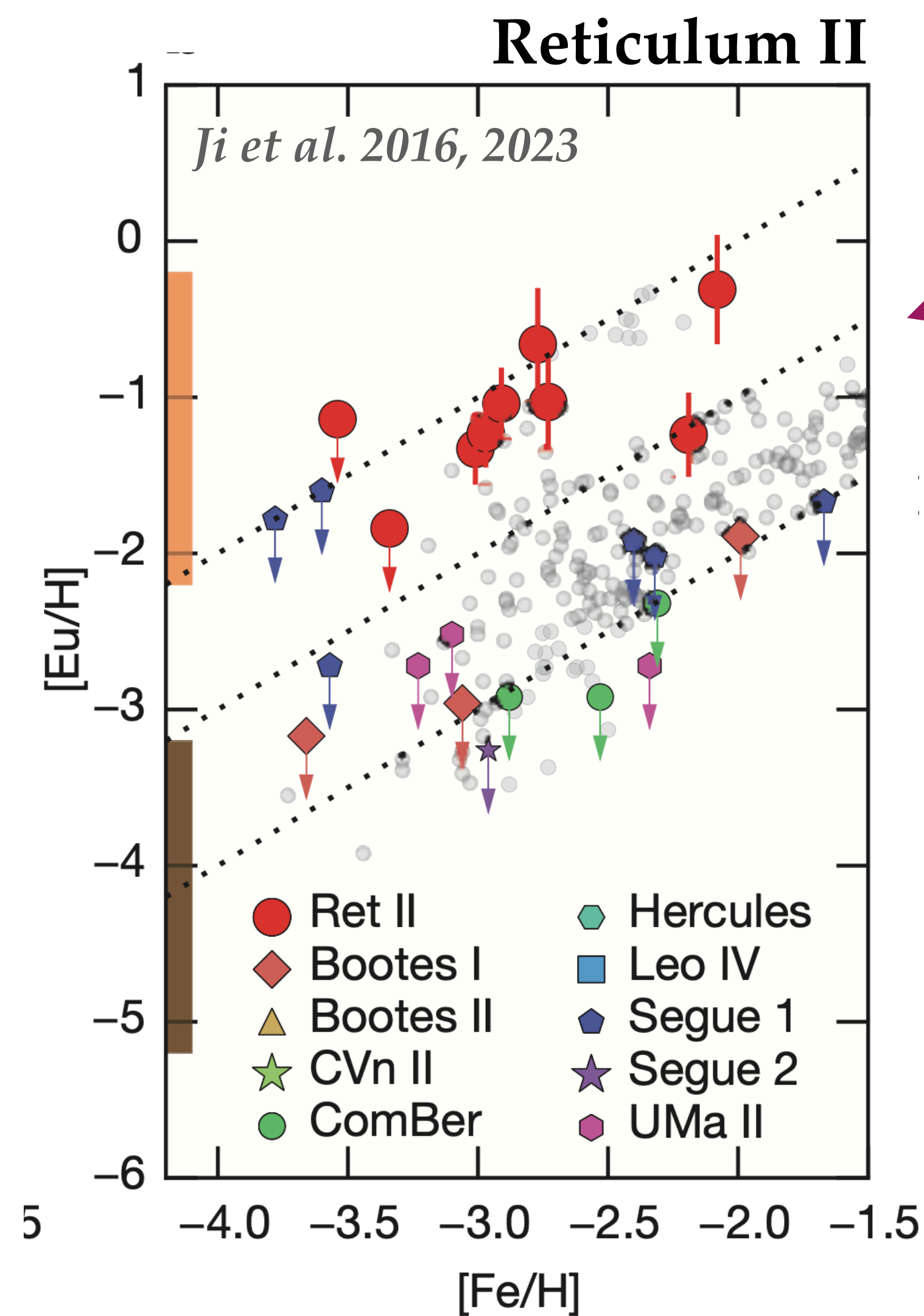


Ultra faint dwarf galaxy
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The most metal-poor r-
rich star in a dwarf
spheroidal galaxy (dSph),
 $[Fe/H]=-3$
 $[Eu/Fe]=+0.8$

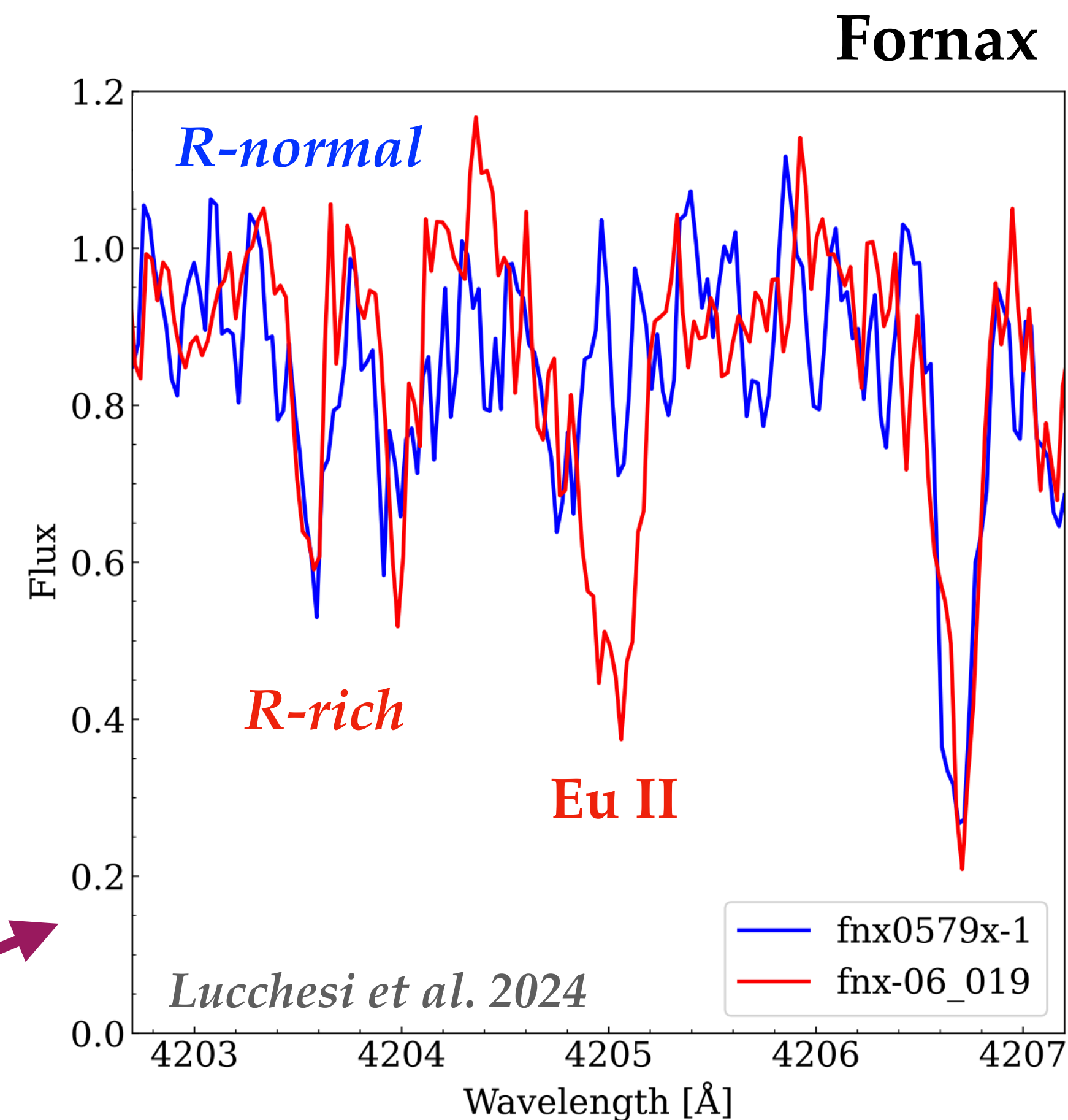


R-PROCESS RICH STARS



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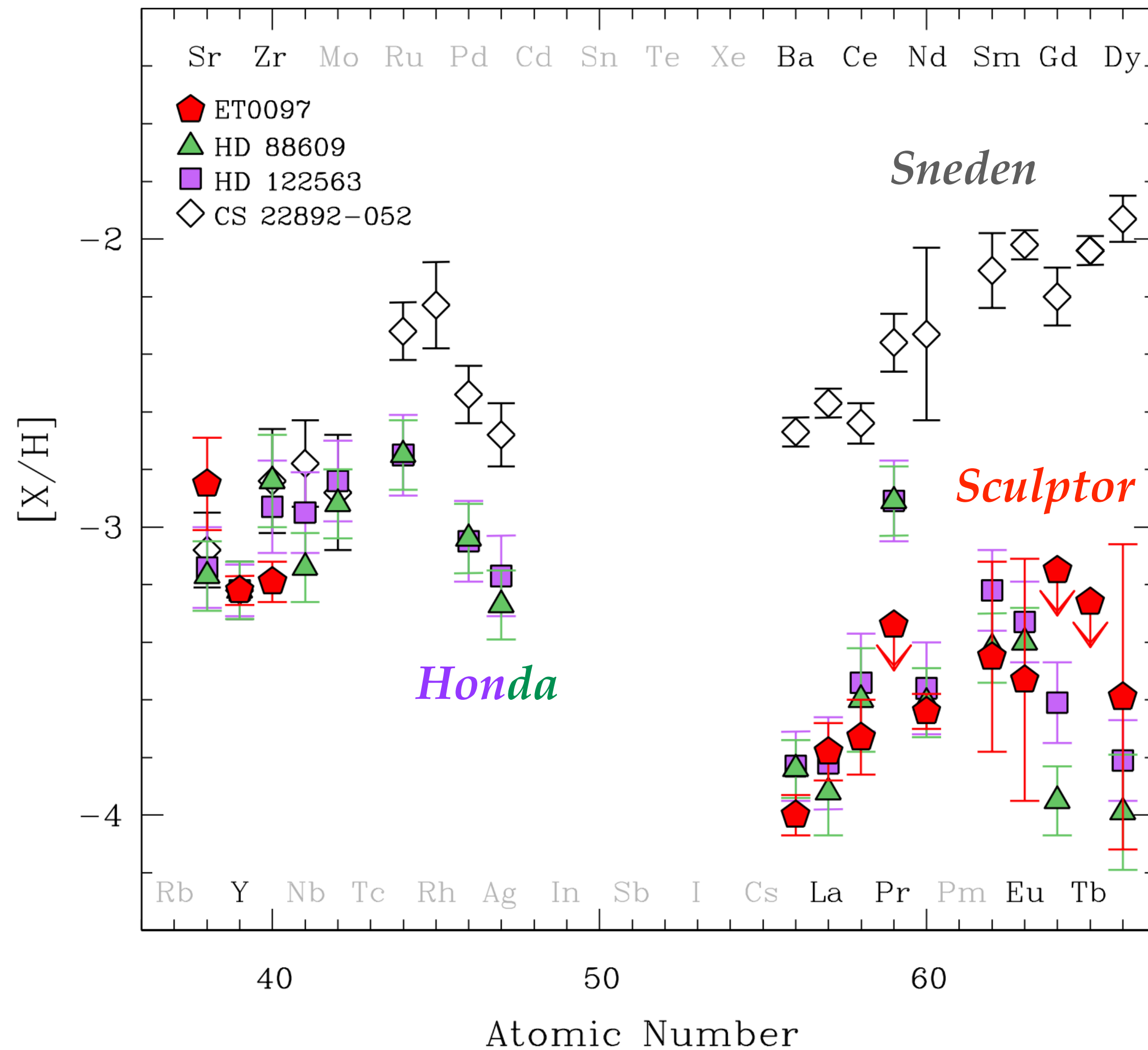
The most metal-poor r-
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 $[Fe/H]=-3$
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R-process events are rare and prolific.
They can happen in any galaxy
Present in the early Universe (few ~100 Myrs)

HONDA STARS IN DWARF GALAXIES!

Skúladóttir et al. 2015



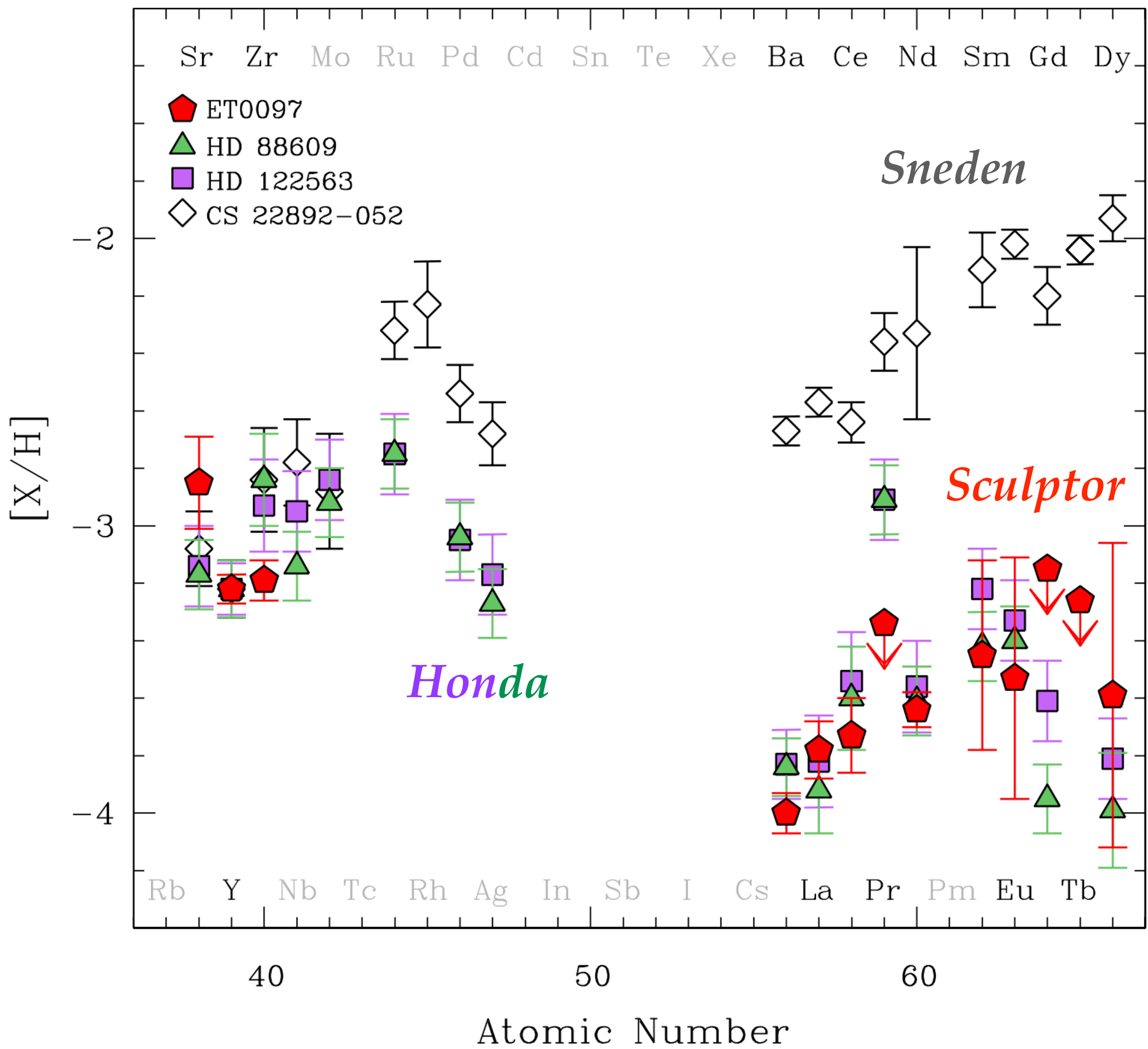
Honda stars (rich in the first peak) can
be found in dwarf galaxies...

...but they are typically Carbon-rich!

(See also: *Susmitha et al. 2017, Spite et al. 2018*)

HONDA STARS IN DWARF GALAXIES!

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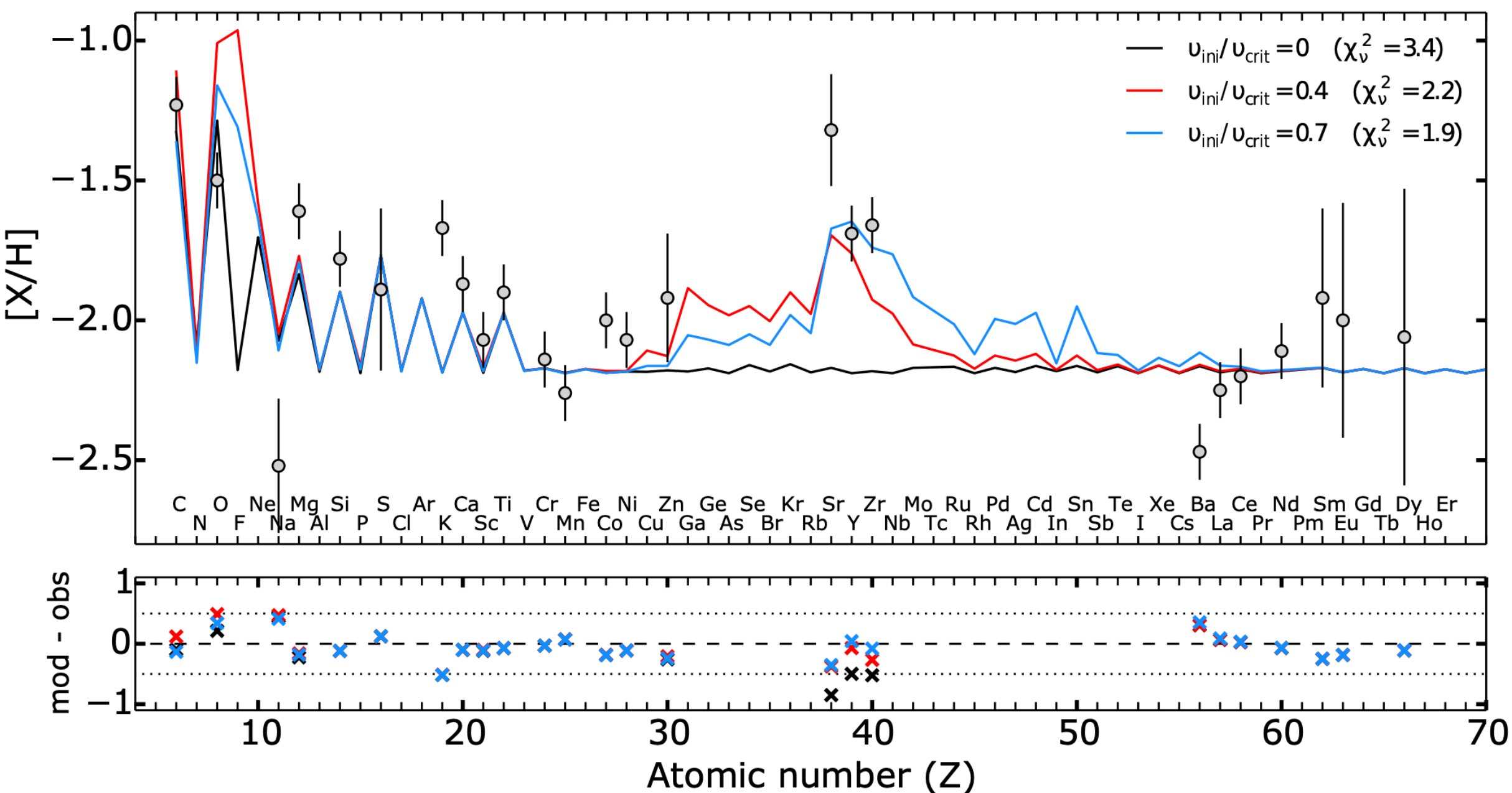


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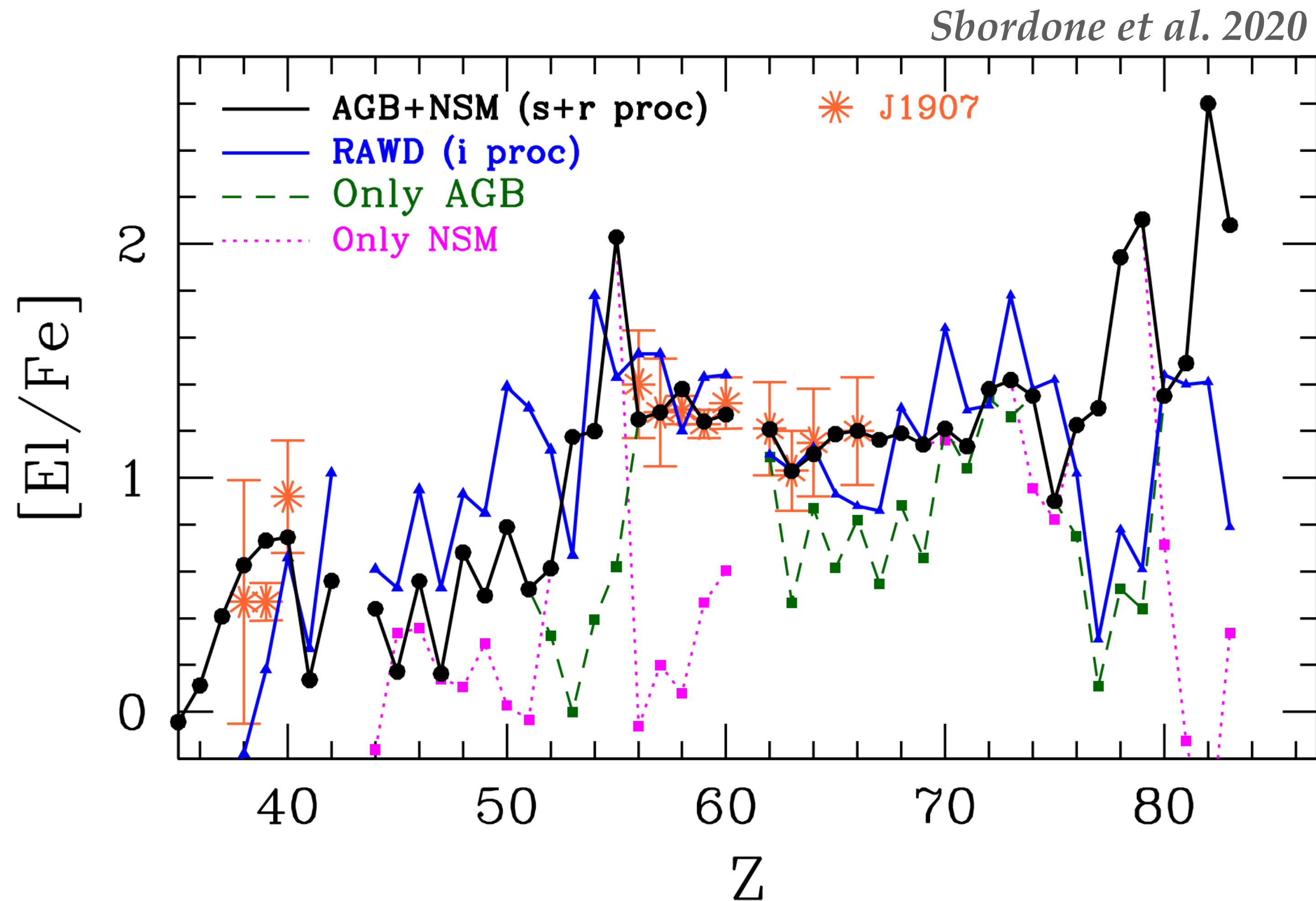
...but they are typically Carbon-rich!

(See also: *Susmitha et al. 2017, Spite et al. 2018*)

Massive star models of A. Choplin (Skúladóttir et al. 2020)



CEMP-R/S STAR IN SAGITTARIUS



Signature of the
i-process?

Are the CEMP-s
and r/s stars the
same as in the
Milky Way?

Population studies of
CEMP-s and r/s star
severely lacking!



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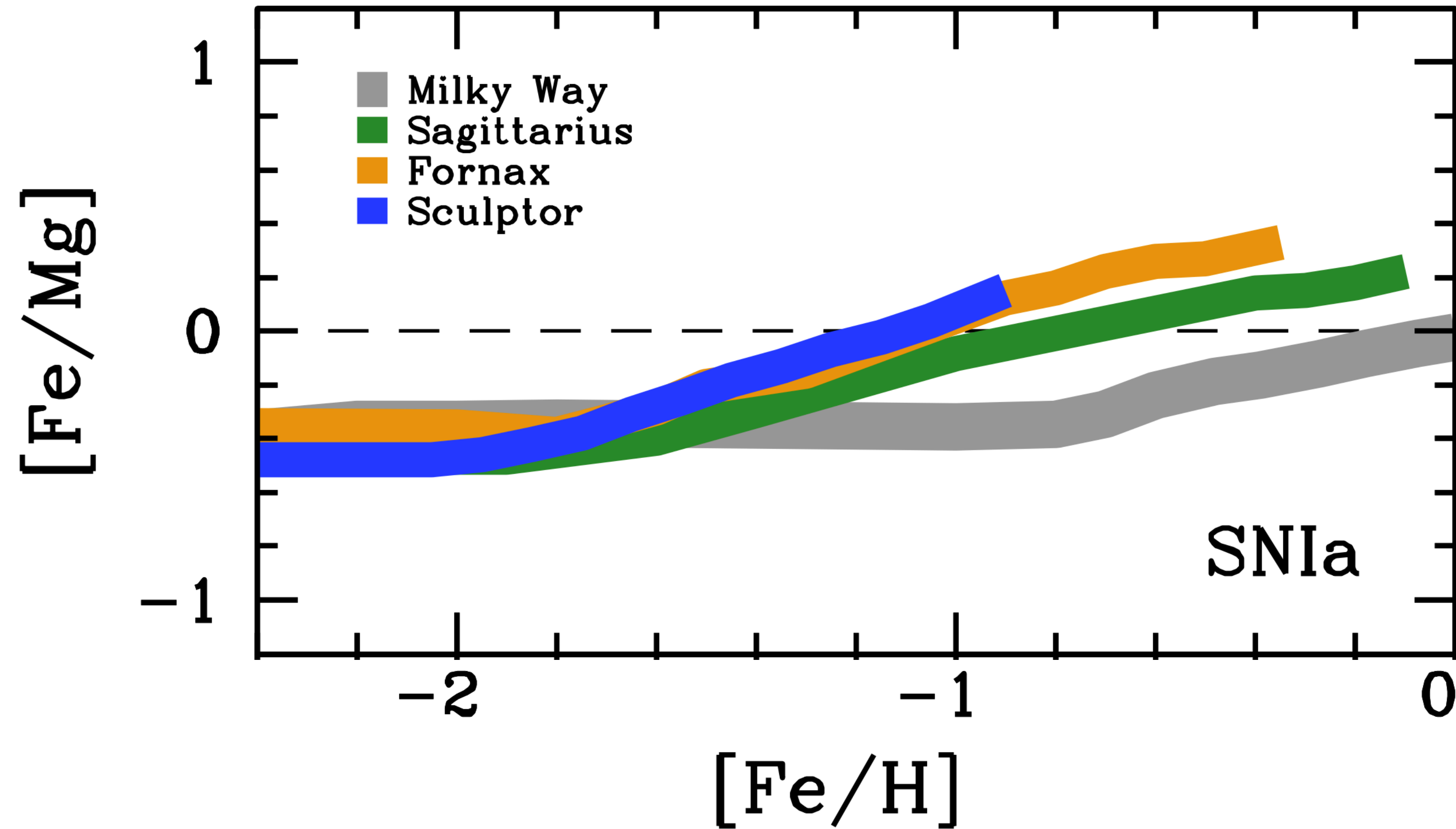
The r-process



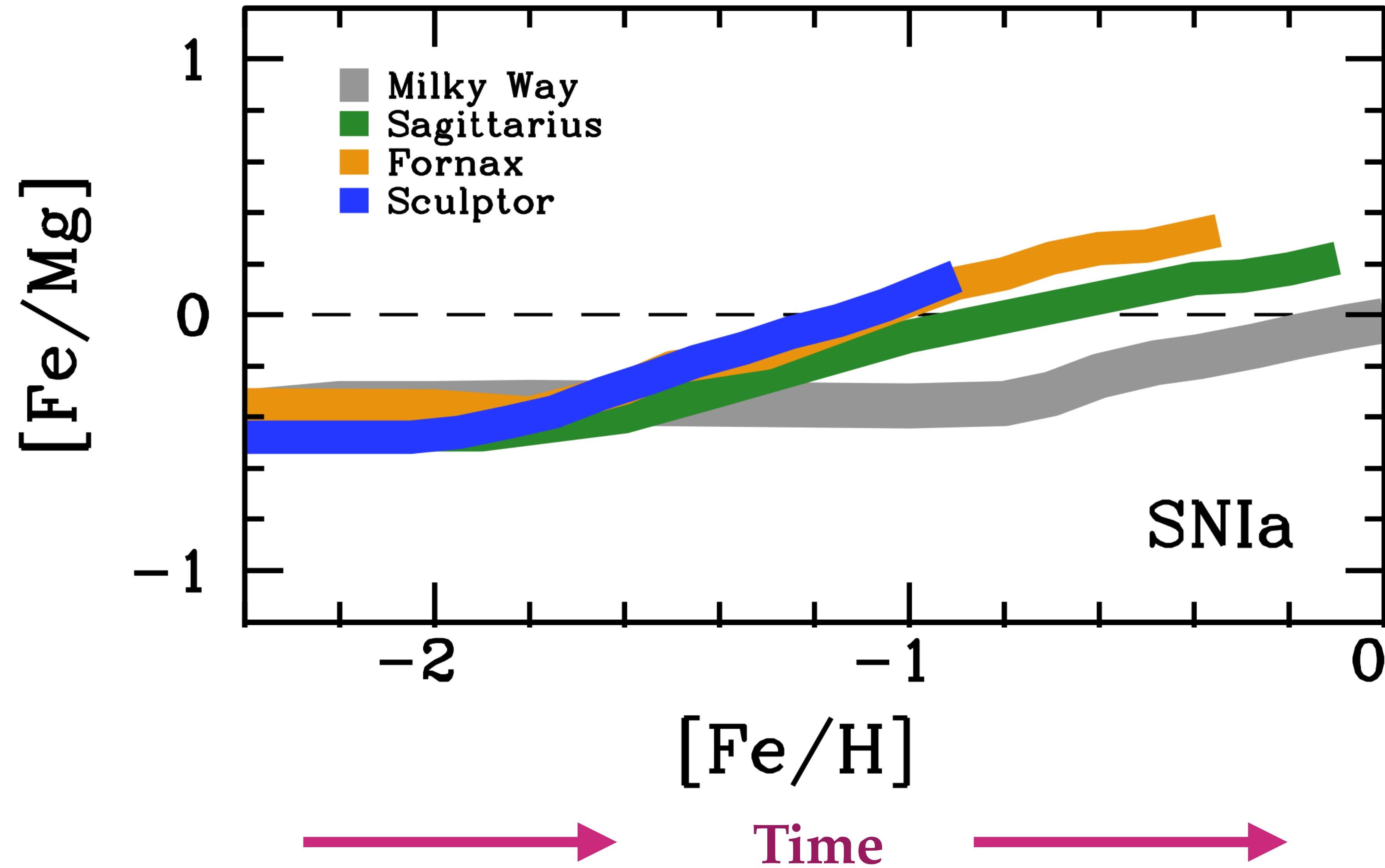
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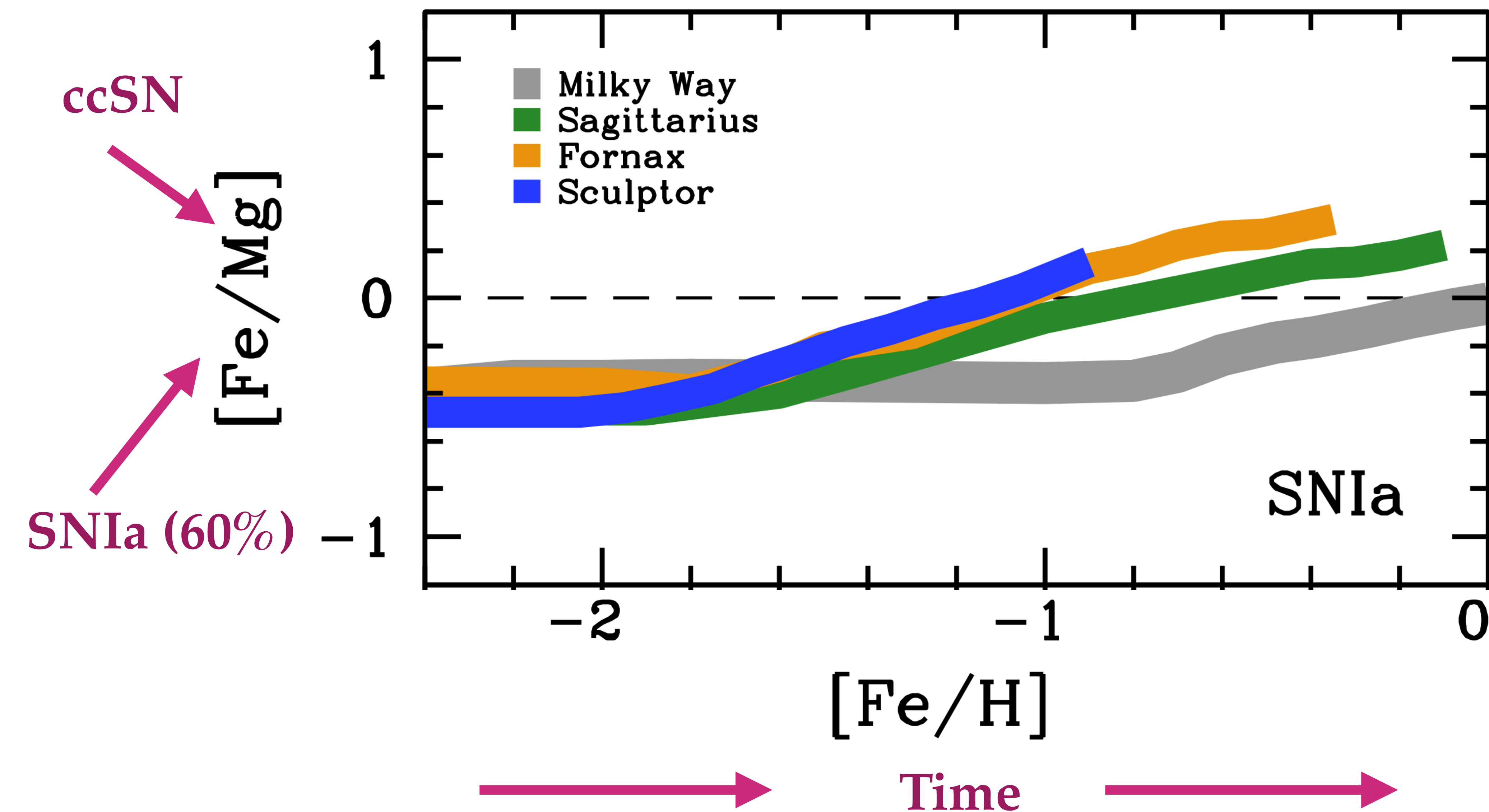
DELAYED PROCESSES – SUPERNOVA TYPE IA



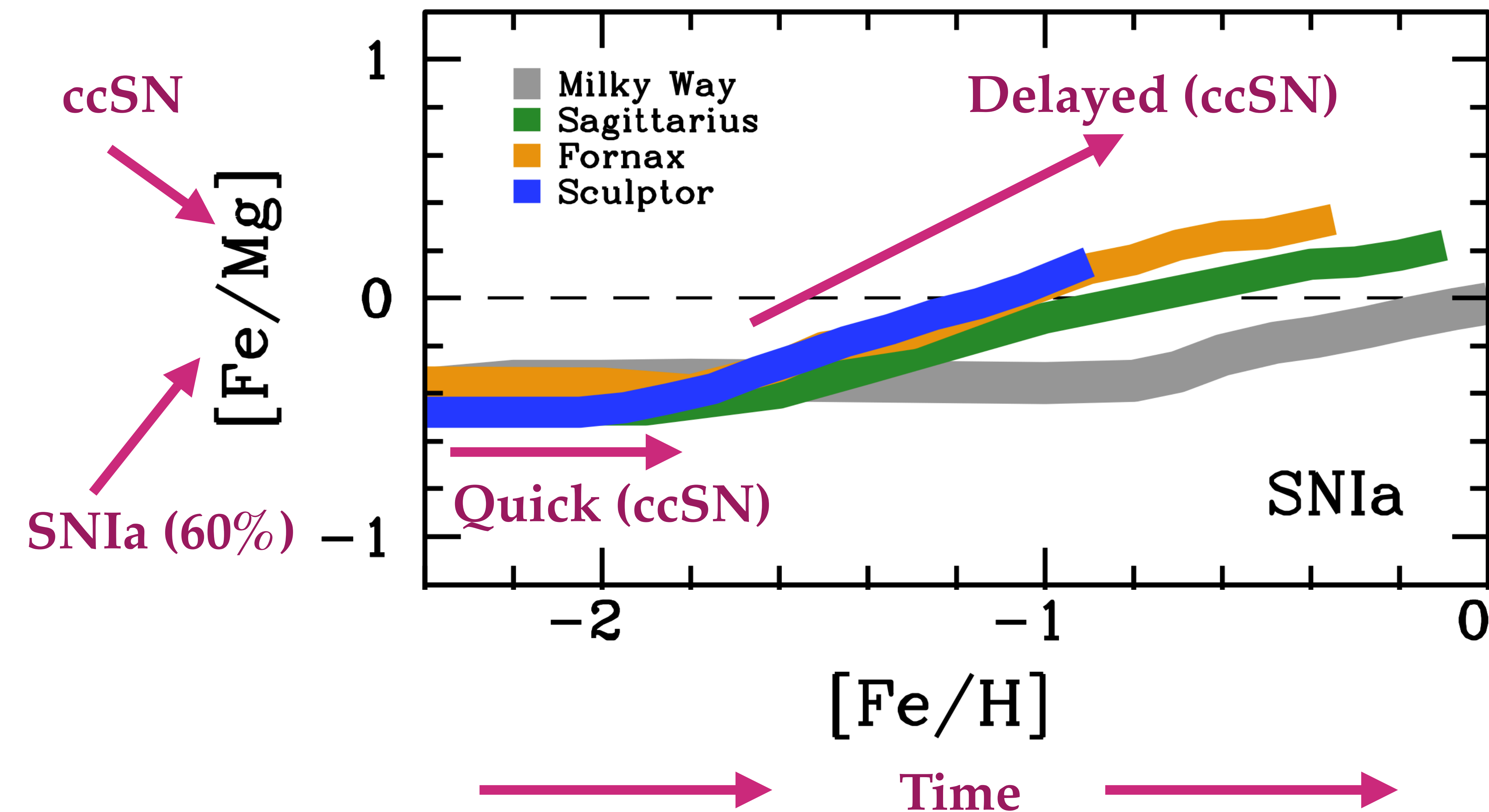
DELAYED PROCESSES – SUPERNOVA TYPE IA



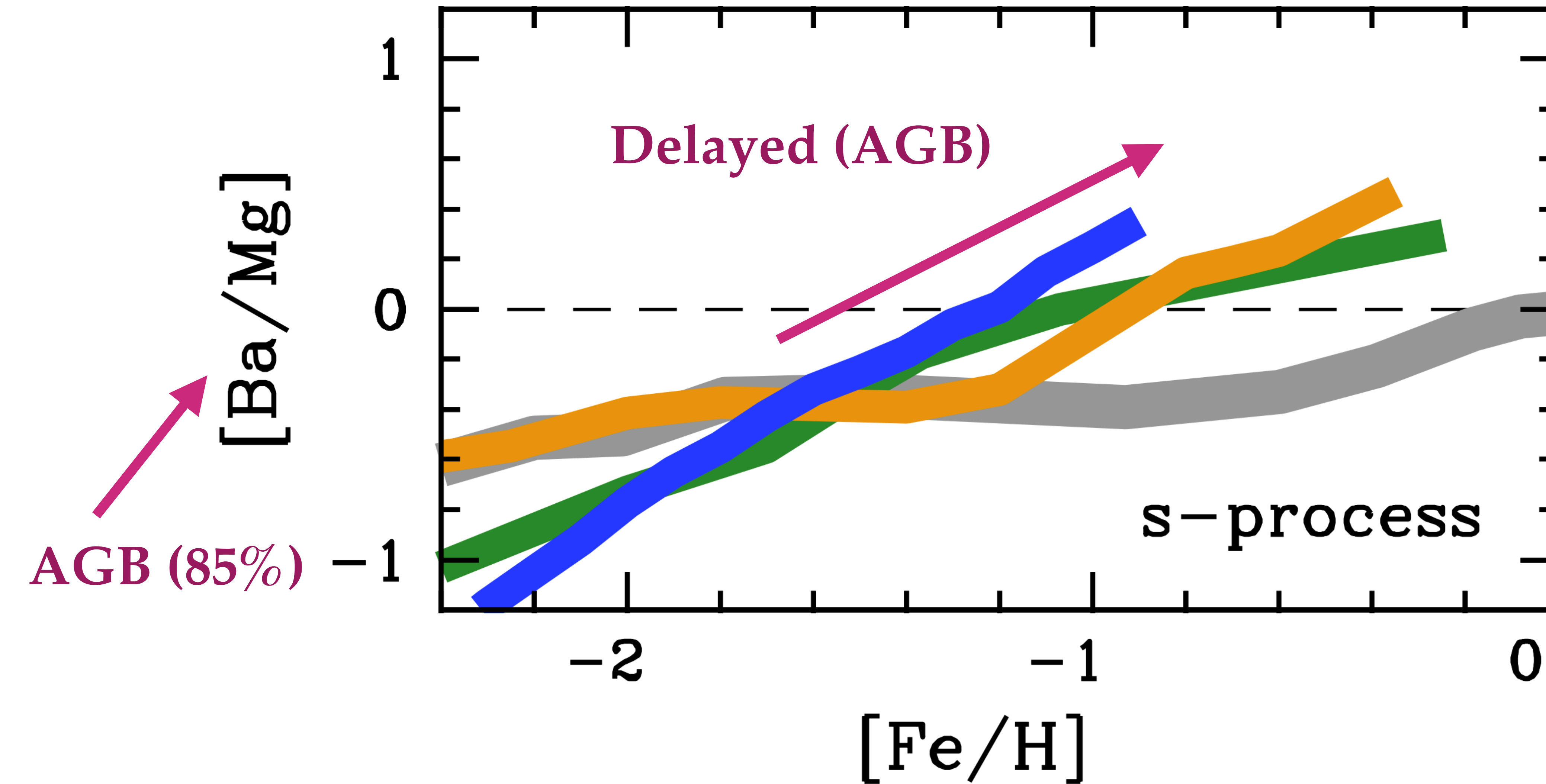
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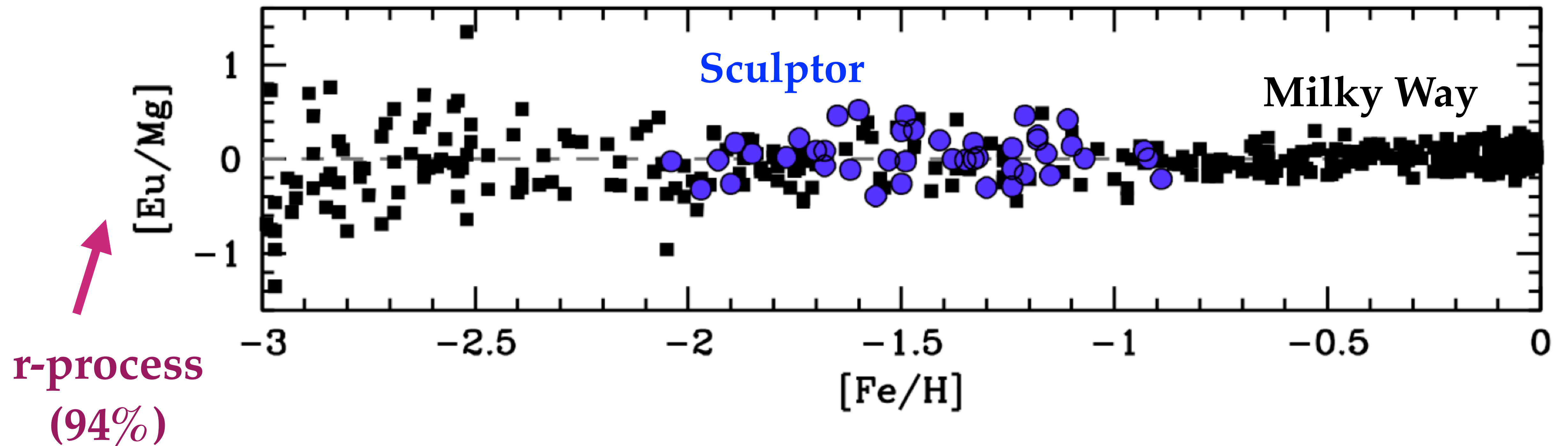


DELAYED PROCESSES – AGB STARS



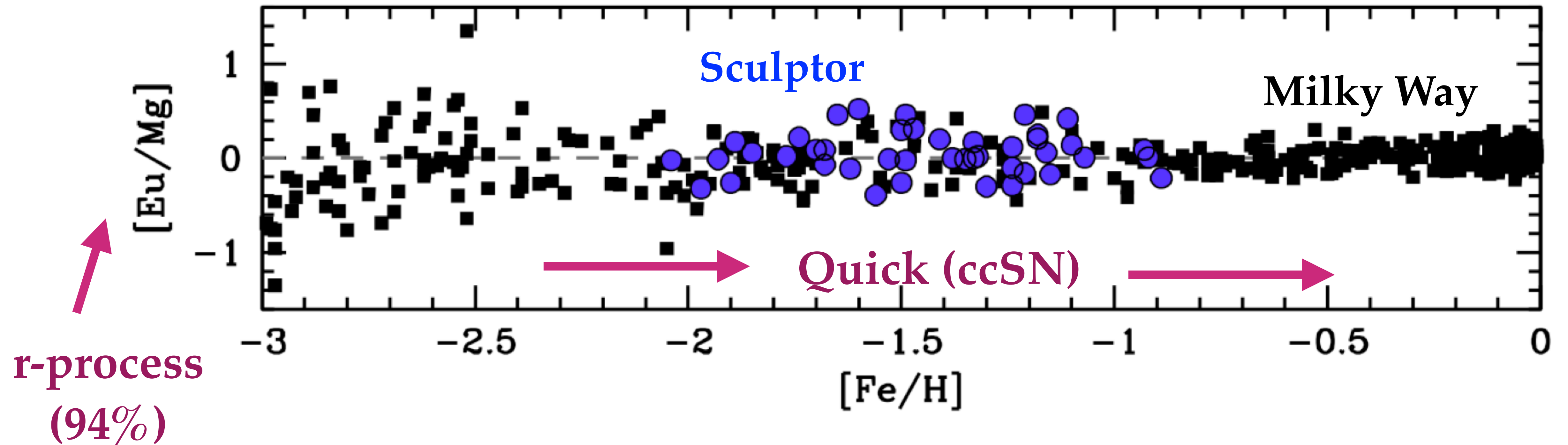
DELAYED PROCESSES – R-PROCESS

Skúladóttir et al. 2019



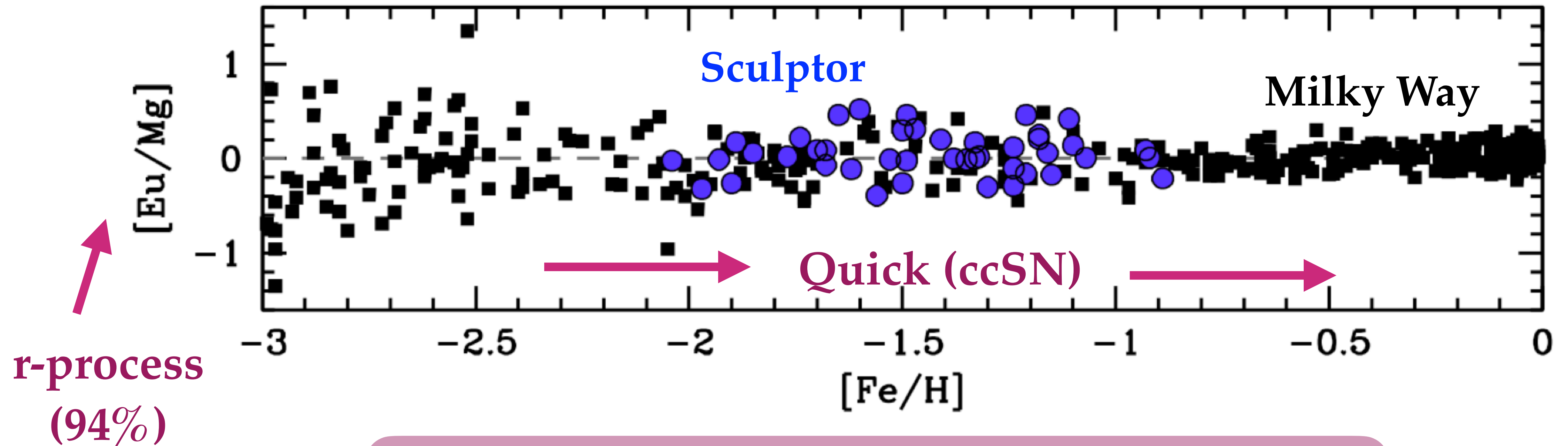
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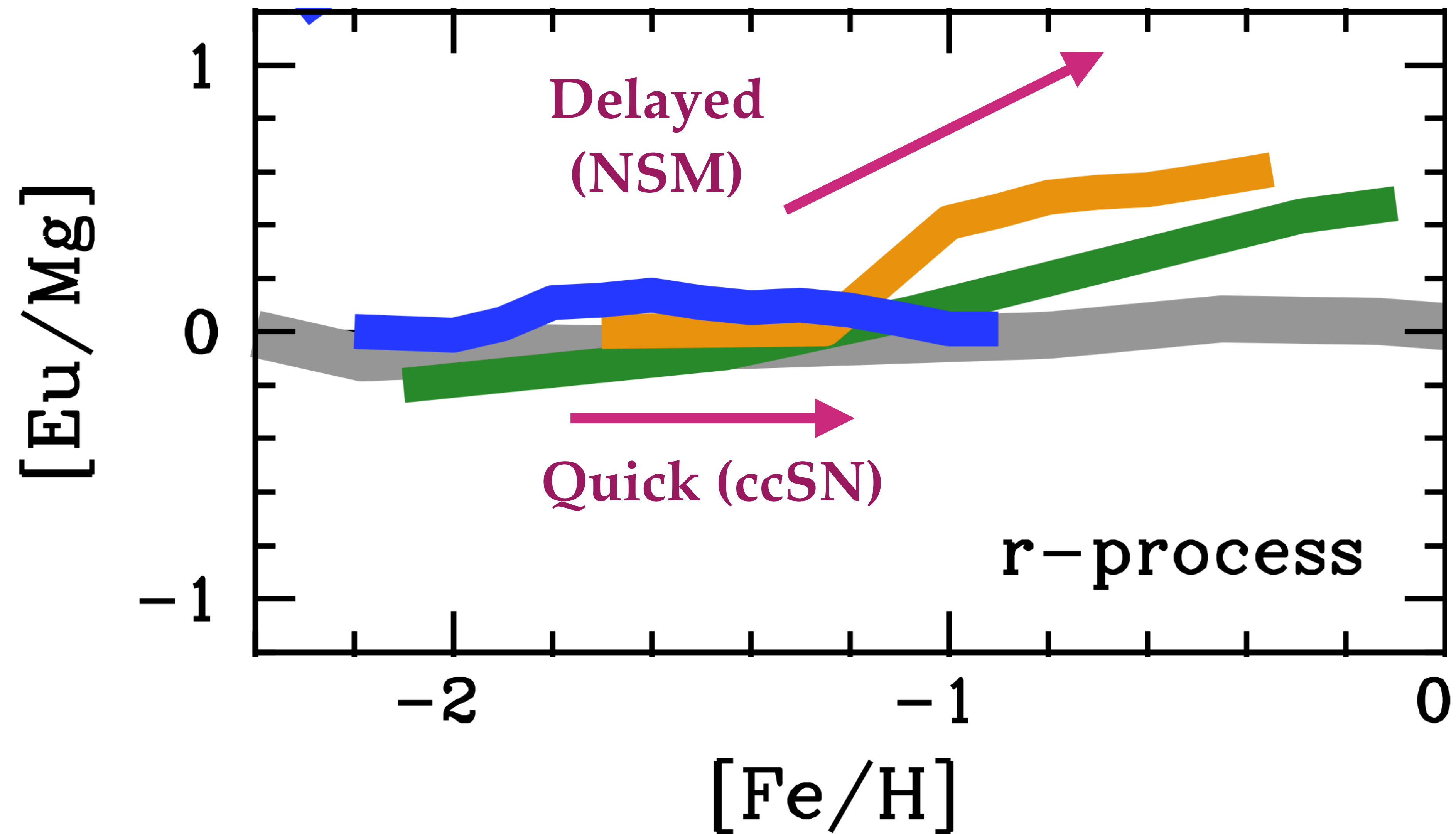
DELAYED PROCESSES – R-PROCESS

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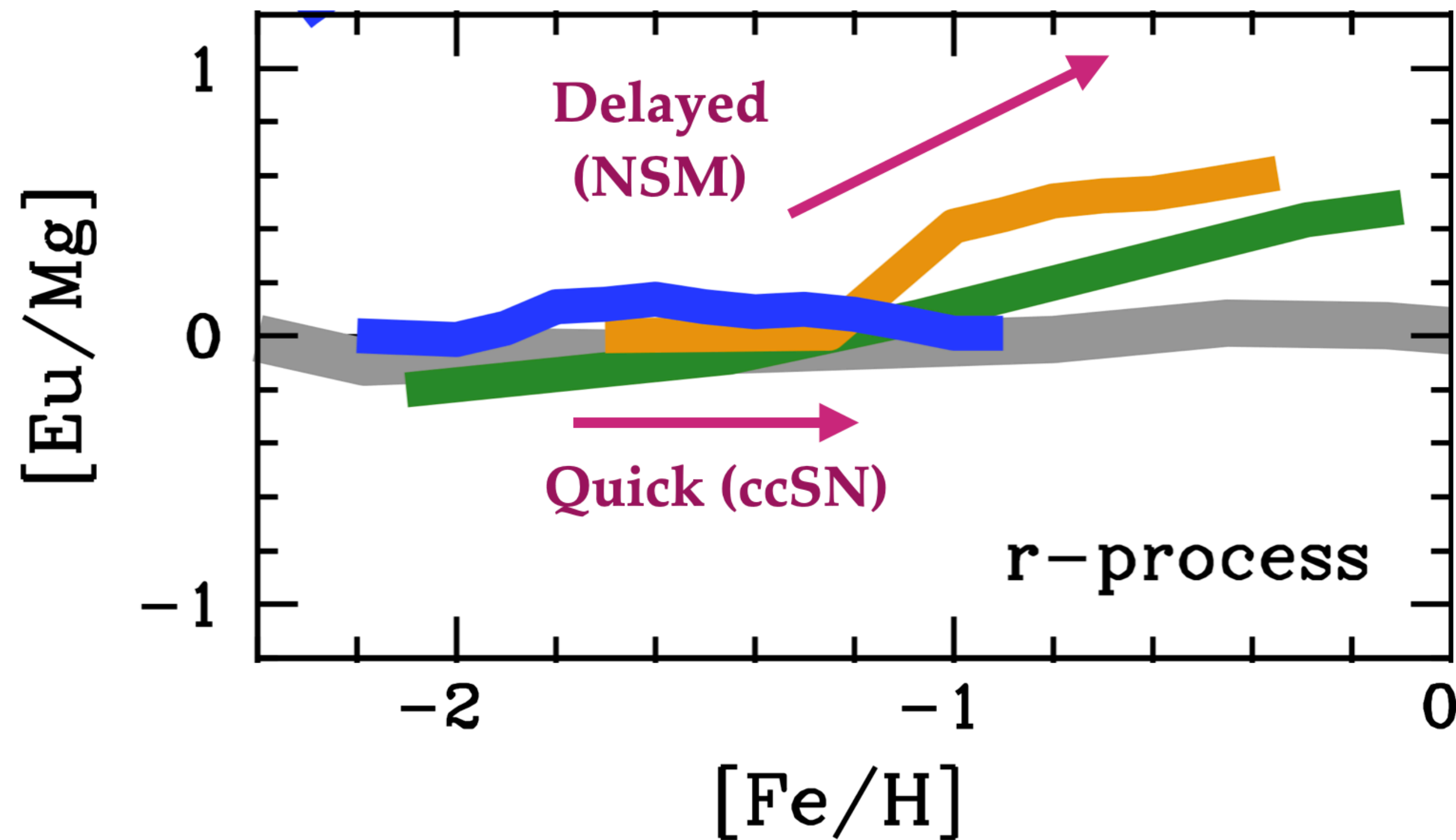


The r-process in Sculptor and the Milky Way is not delayed

DELAYED PROCESSES - R-PROCESS

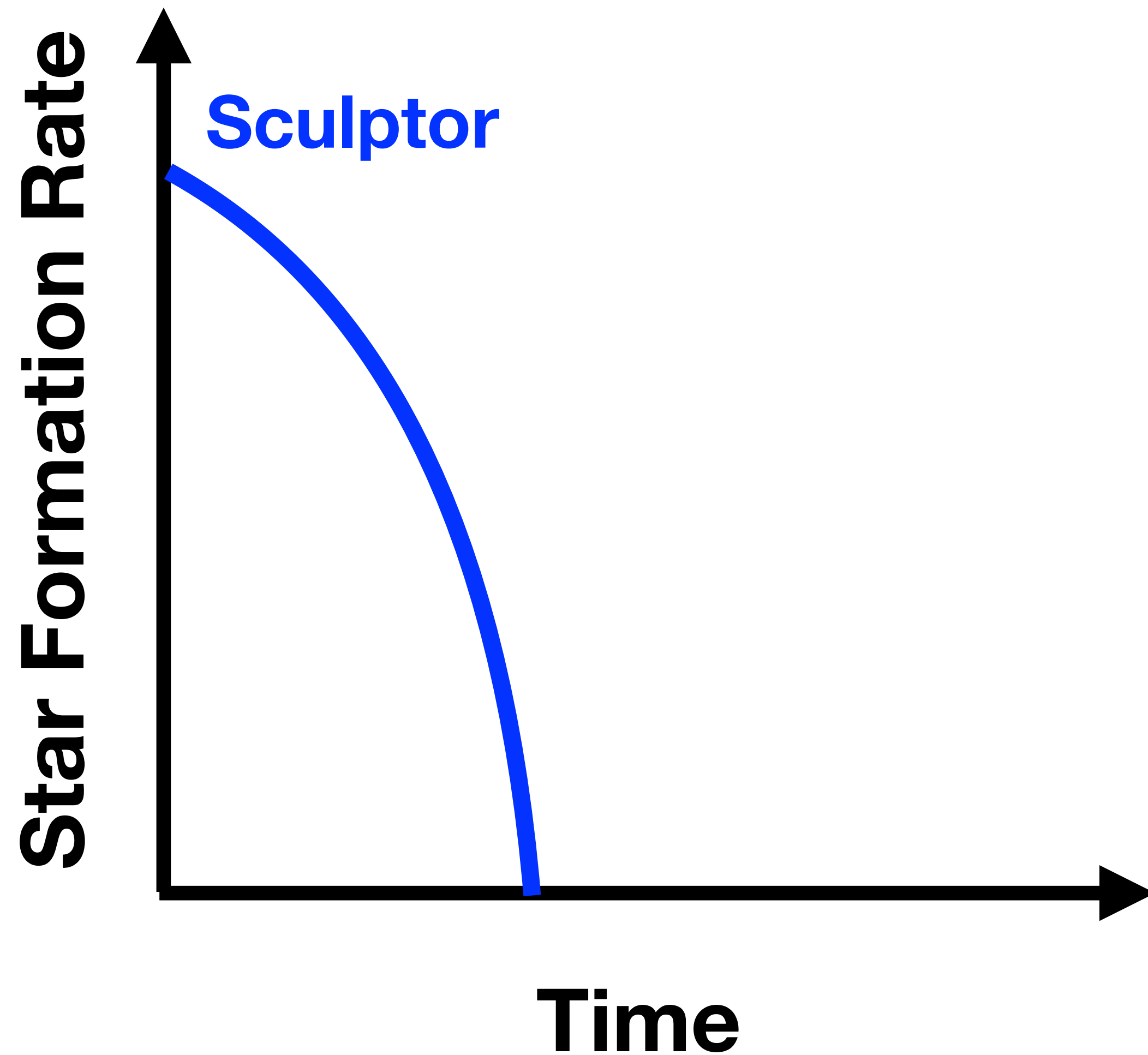


DELAYED PROCESSES – R-PROCESS



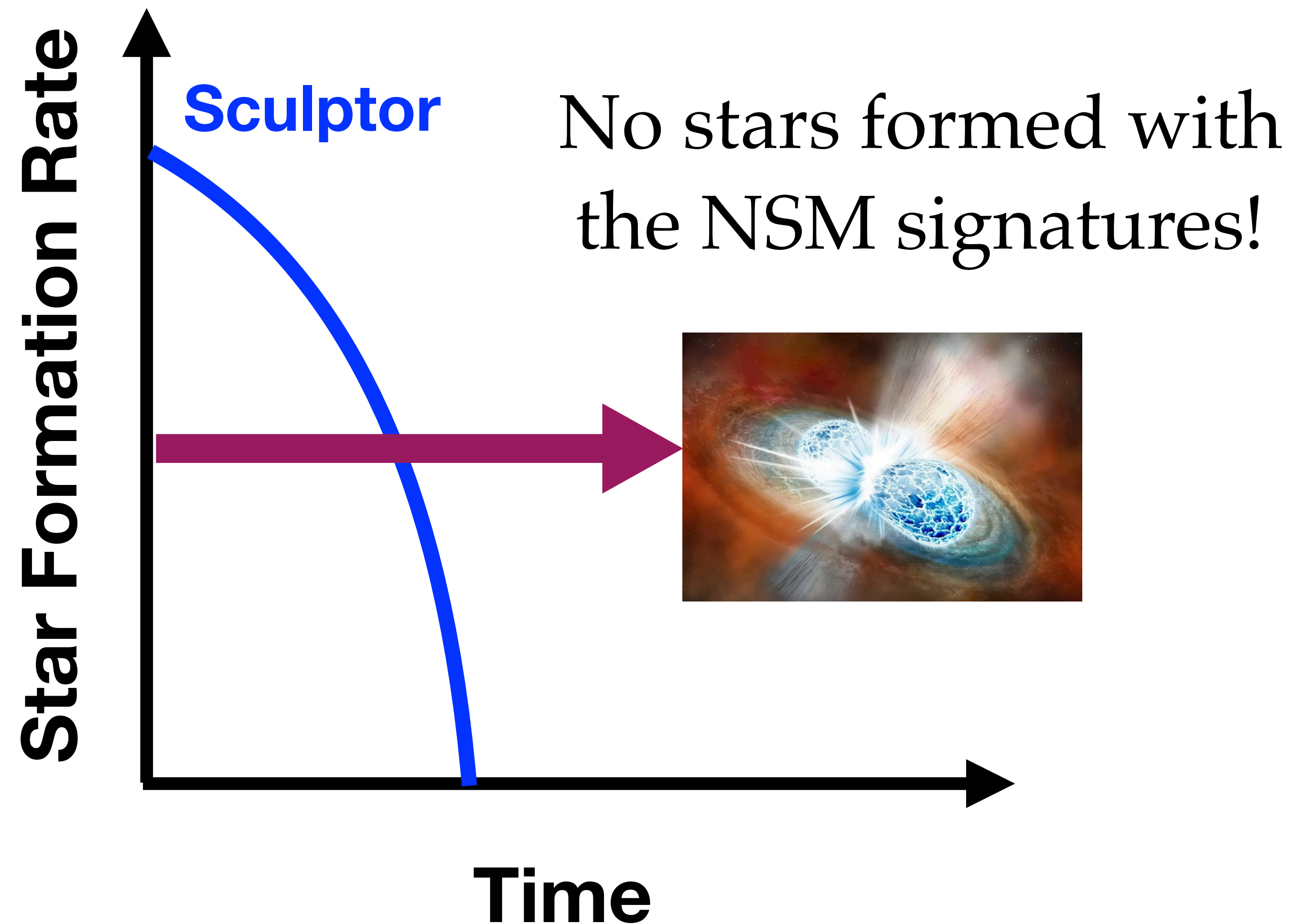
Two processes are needed to explain the data: Quick (ccSN) and Delayed (NSM)

WHY IS SCULPTOR DOMINATED BY THE QUICK R-PROCESS?



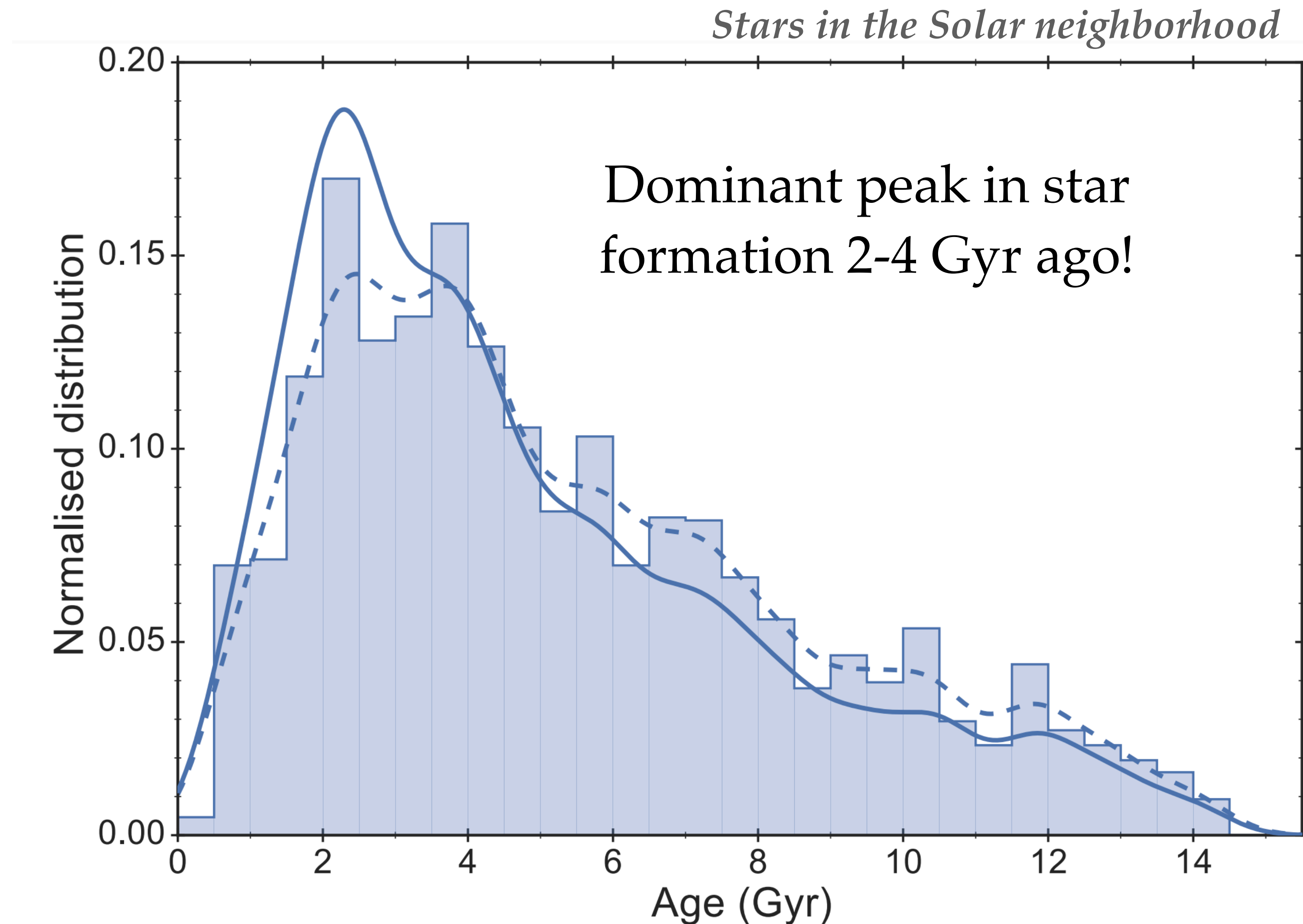
Star formation in
Sculptor is too short!

WHY IS SCULPTOR DOMINATED BY THE QUICK R-PROCESS?



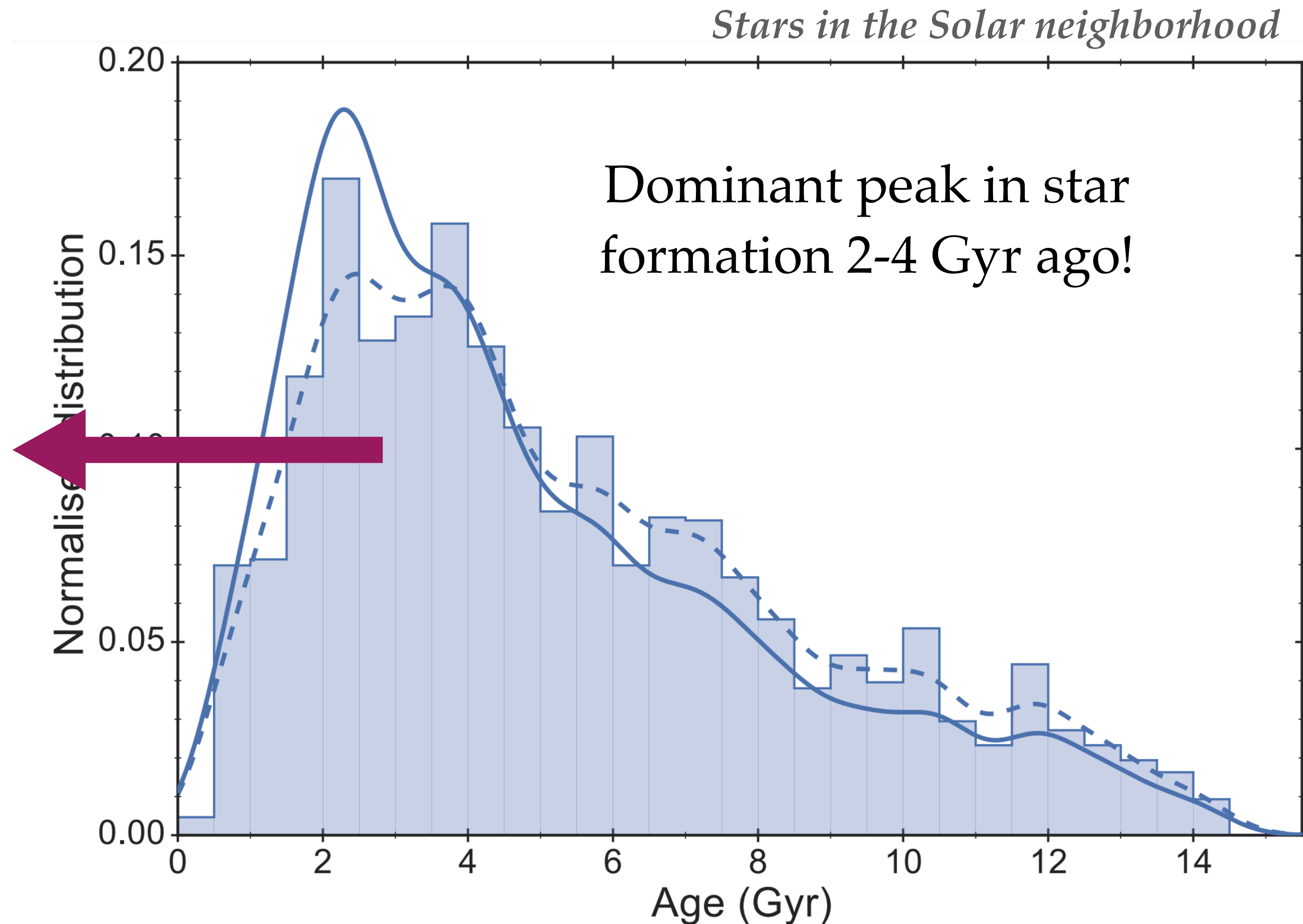
Star formation in Sculptor is too short!

WHY IS THE MILKY WAY DOMINATED BY THE QUICK R-PROCESS?



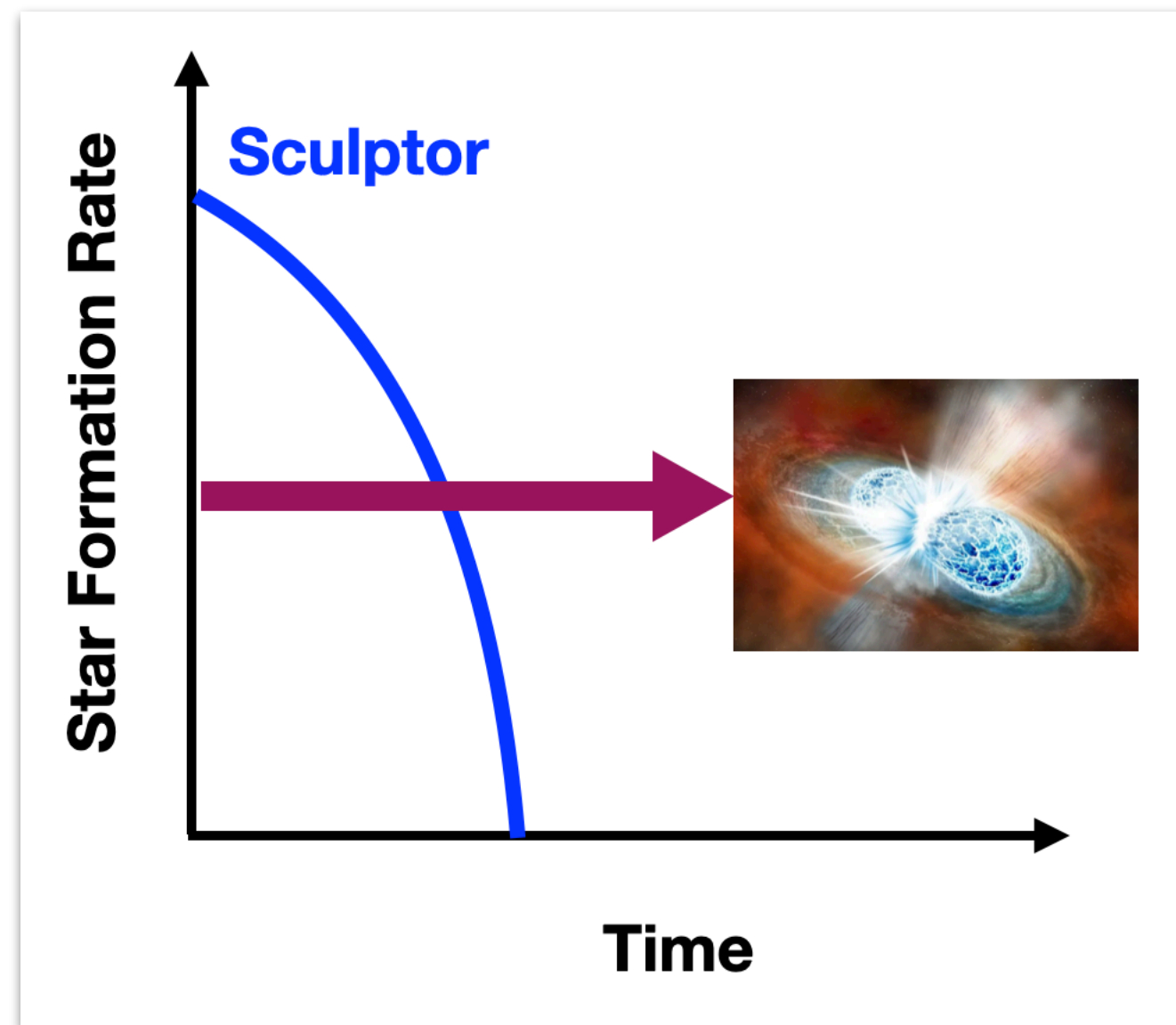
Star formation
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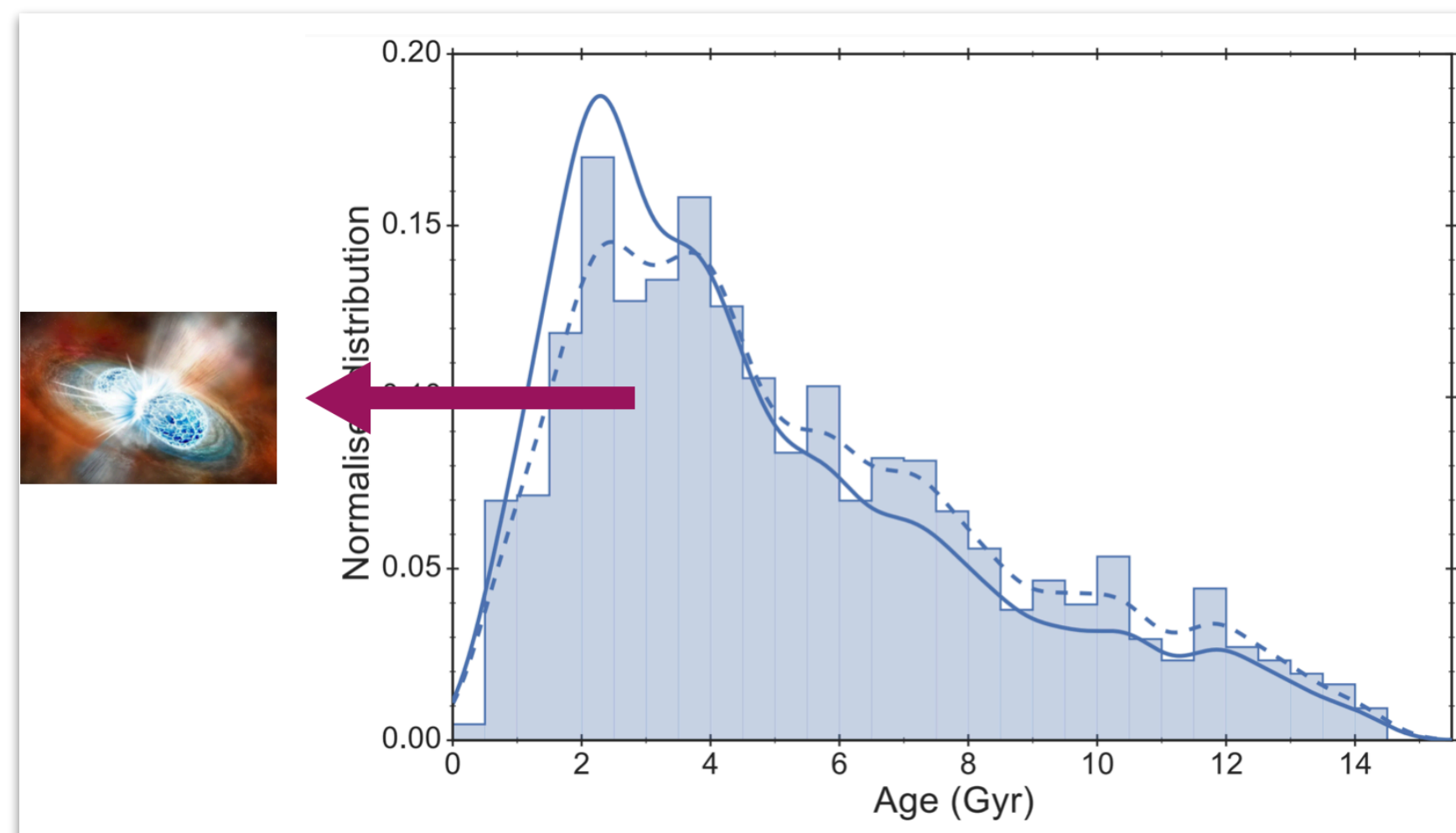
Star formation
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TIMESCALE OF NEUTRON STAR MERGERS



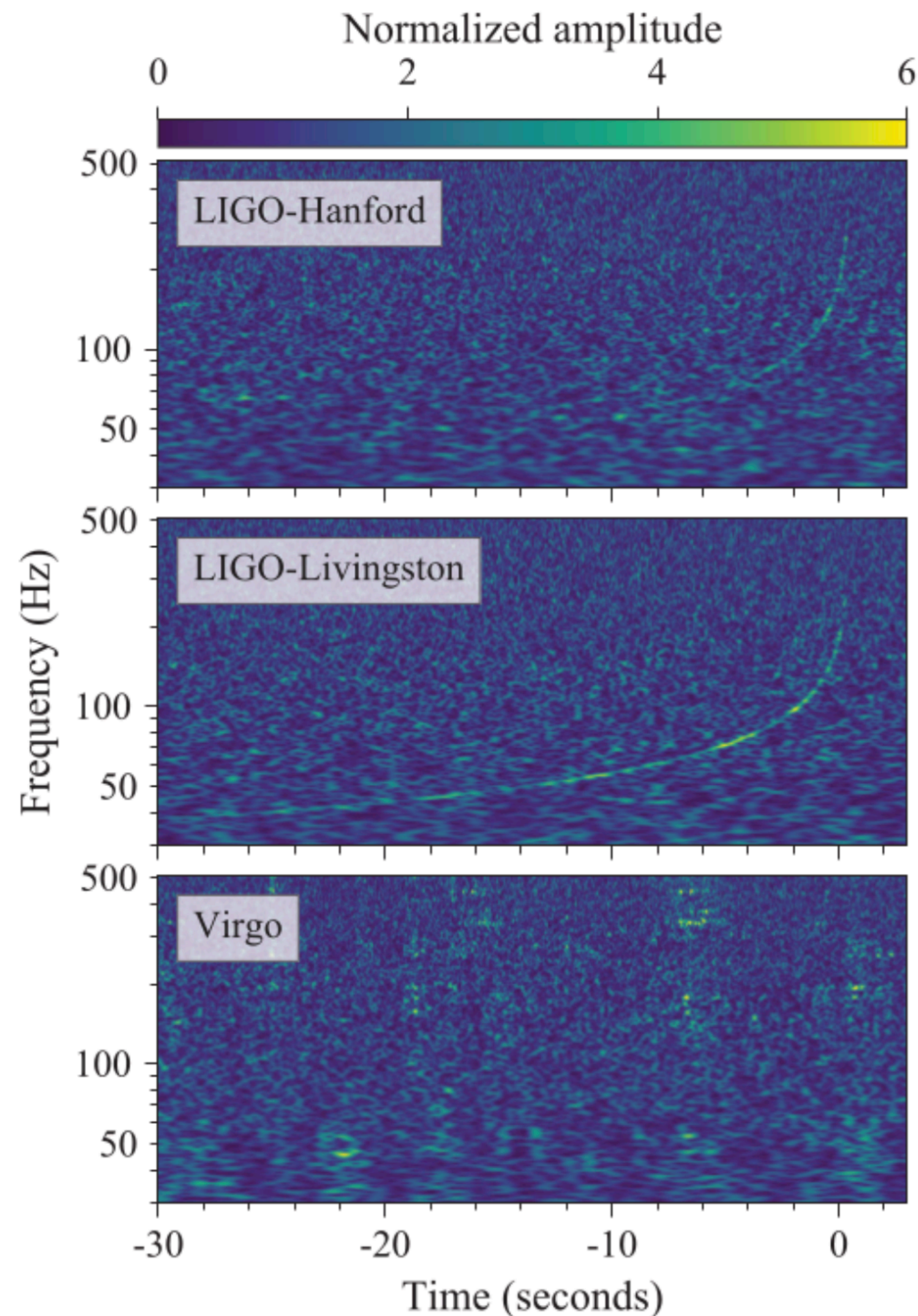
Typical timescales of NSM are ≈ 2 Gyr

This is the timescale that it takes NSM to rival ccSN in production of r-process elements*



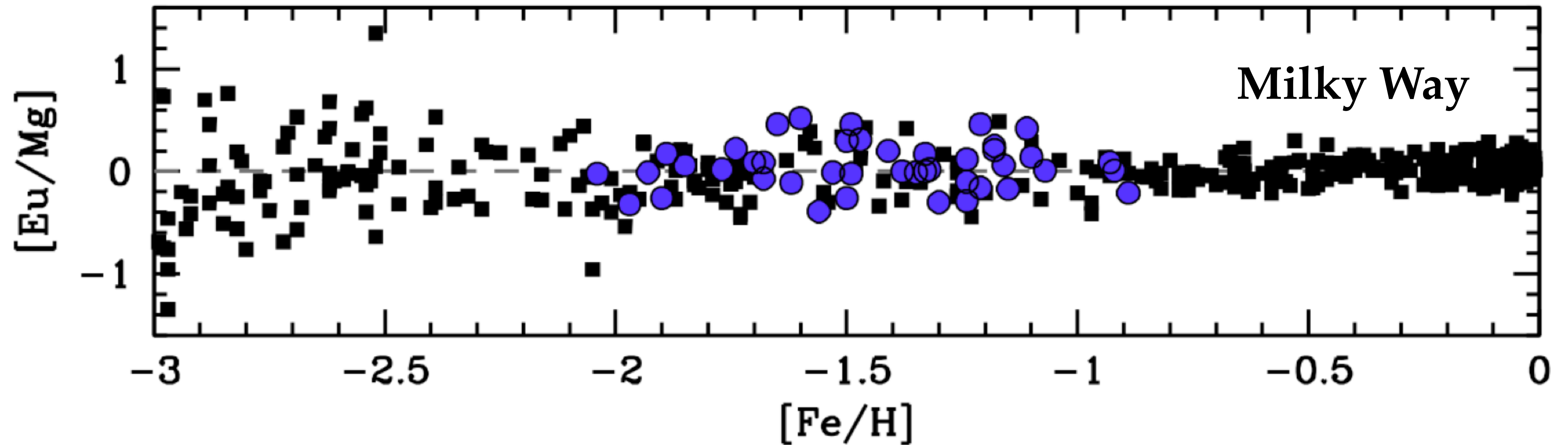
**In Skúladóttir & Salvadori 2020 we estimated ≈ 4 Gyr, but taking into account more data I recommend a revised more conservative ≈ 2 Gyr timescale*

TIMESCALE OF NEUTRON STAR MERGERS – OTHER EVIDENCE



- GW170817 from a non-star forming region! Estimated delay time >6.8 Gyr (Blanchard et al. 2017).
- Other Kilonovae also found far from star formation (e.g. talk by Hotokezaka).
- Significant fraction of short gamma ray bursts found in early type galaxies (Berger 2014).
- Different intrinsic distribution of GRBS
 - Short GRBs: peak at $z \sim 1.5$, 9.5 Gyr ago (Ghirlanda et al. 2022).
 - Long GRBs: peak at $z \sim 3.5$, 12 Gyr ago. (Salvaterra et al. 2012).

R-PROCESS



The quick source (ccSN) is the dominant r-process source at **all** metallicities in the Milky Way disk



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The s/i -processes

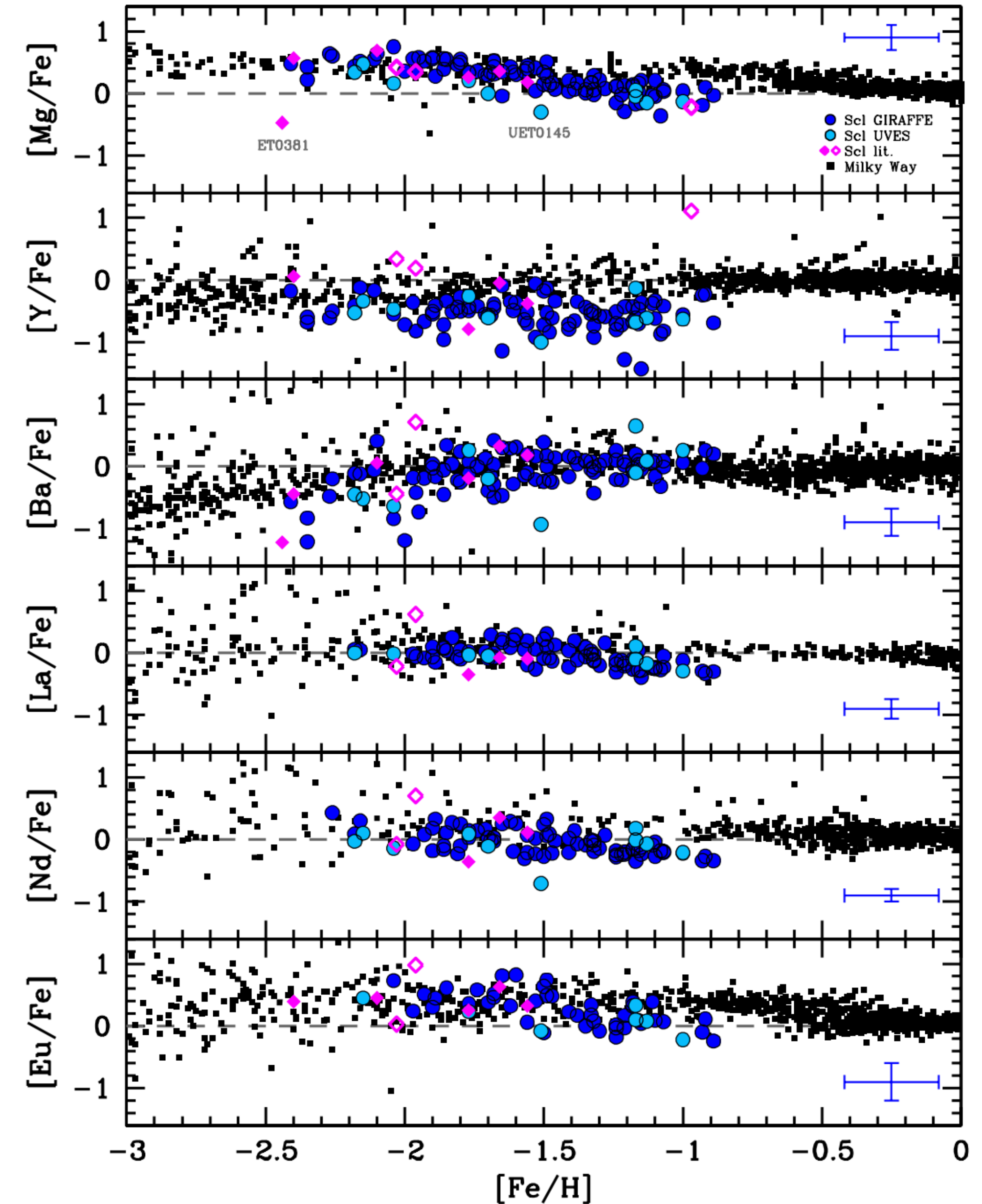


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N-CAPTURE ELEMENTS IN DWARF GALAXIES

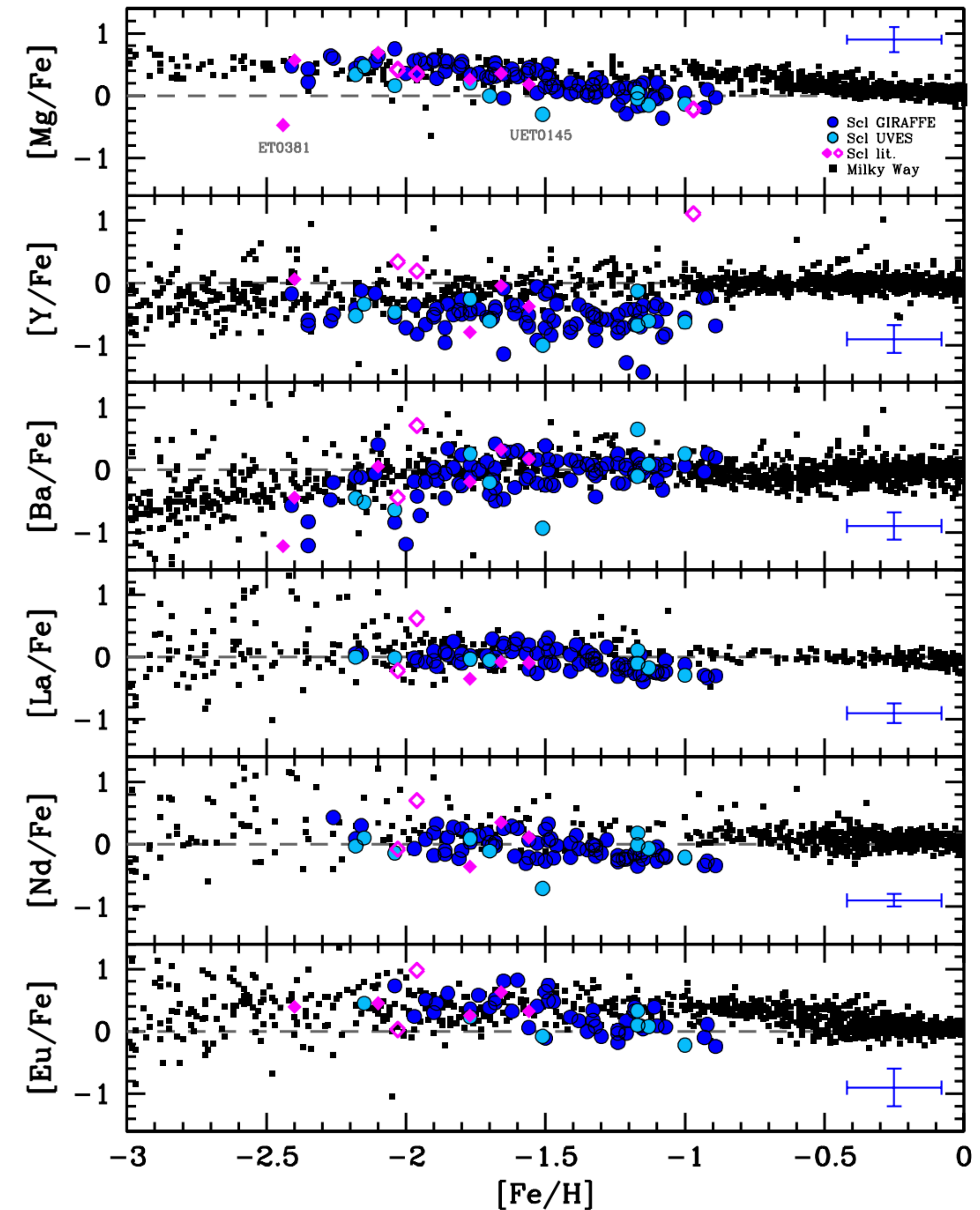
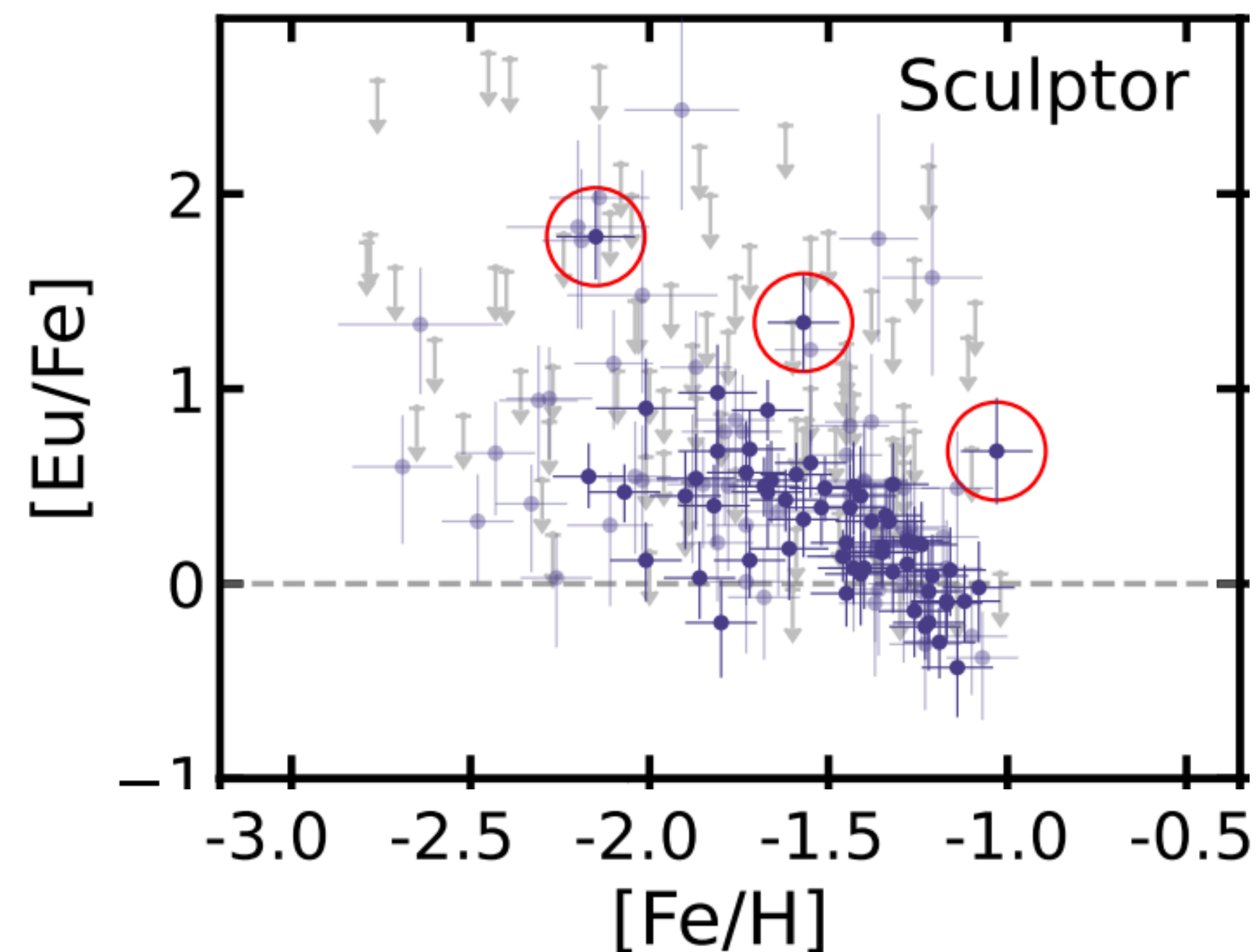
N-capture elements measured in
100 stars in the Sculptor dwarf
spheroidal galaxy



N-CAPTURE ELEMENTS IN DWARF GALAXIES

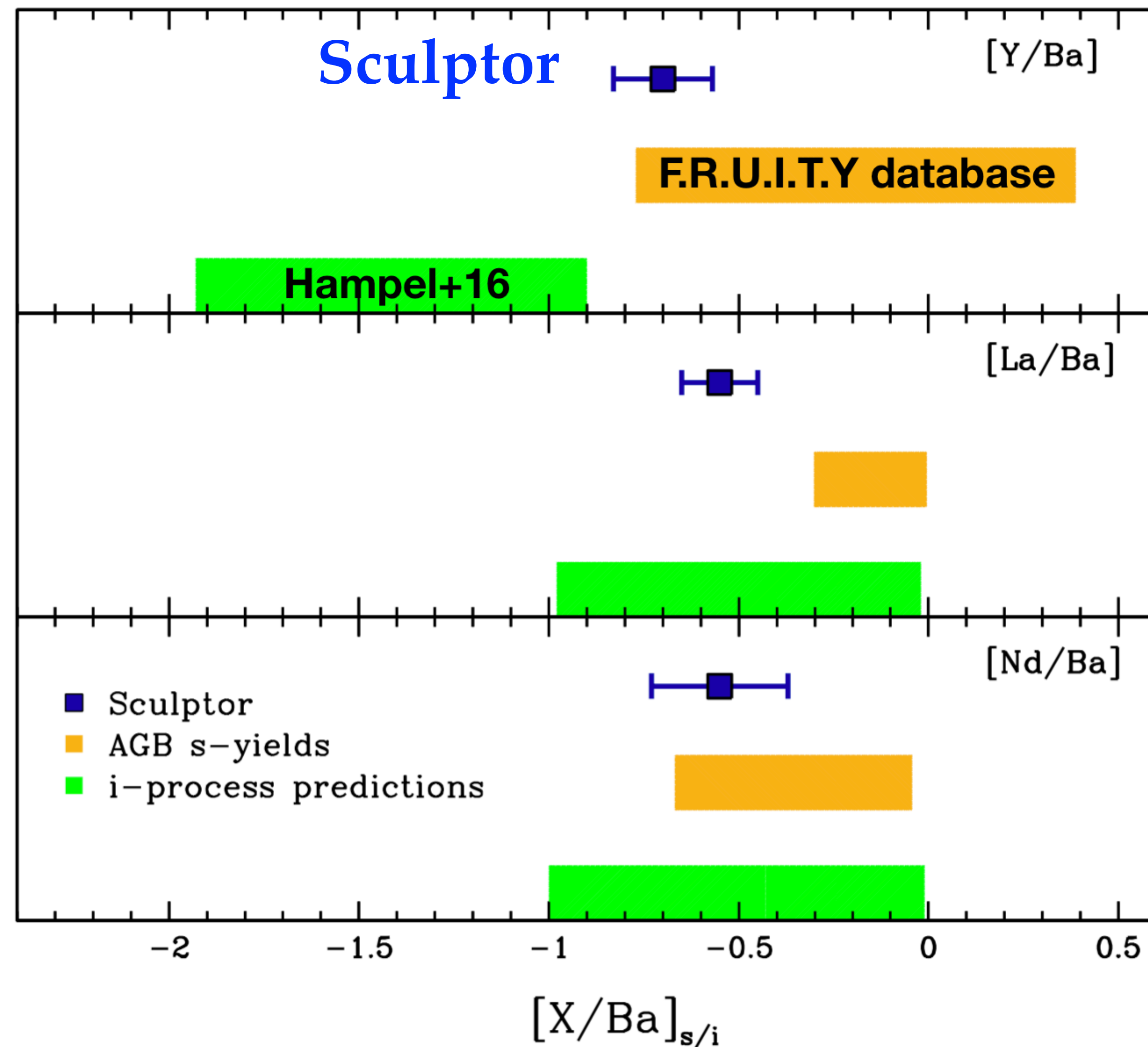
N-capture elements measured in
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See also Henderson et al. 2025



THE S/I PROCESSES IN SCULPTOR

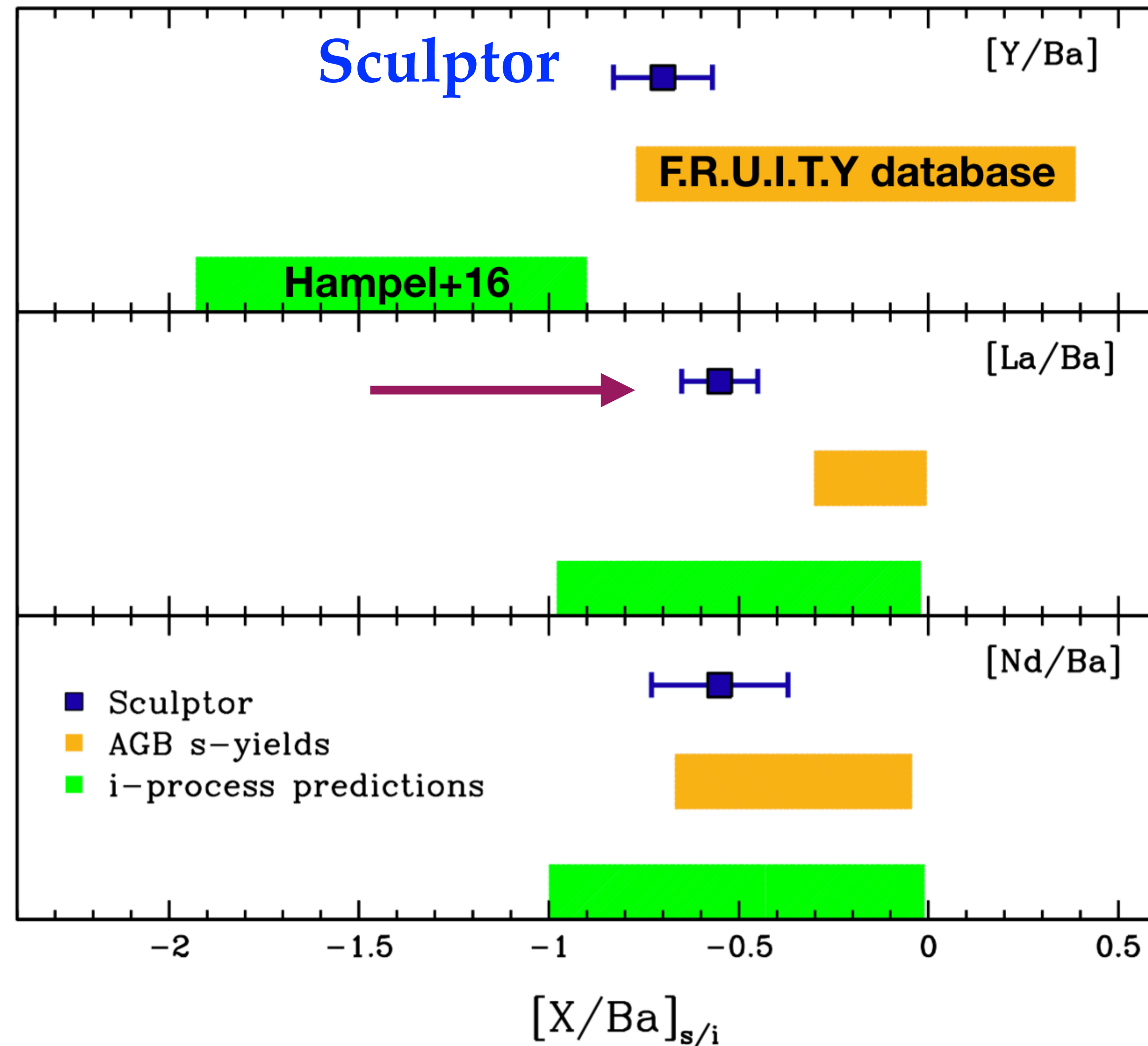
Skúladóttir et al. 2020



Taking all the n-capture elements formed in Sculptor and removing the r-process gave us the $[X/Ba]_{s/i}$ values.

THE S/I PROCESSES IN SCULPTOR

Skúladóttir et al. 2020



Taking all the n-capture elements formed in Sculptor and removing the r-process gave us the $[X/Ba]_{s/i}$ values.

The i-process is **needed** to explain the chemical abundances in Sculptor!

THE I-PROCESS SOURCE IN SCULPTOR

$$[\text{Fe}/\text{H}] \approx -2$$

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Not massive stars
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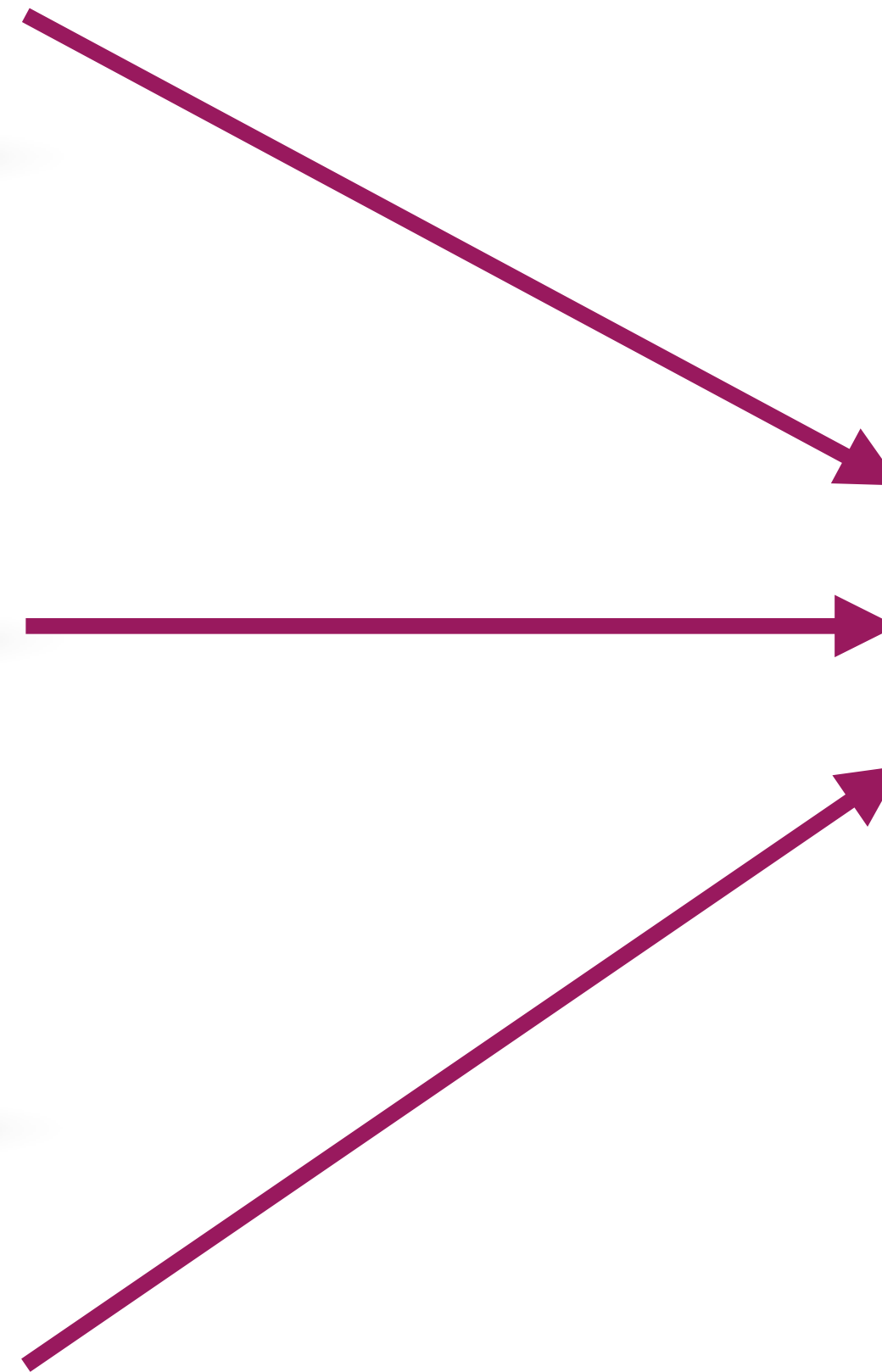
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Metal-poor AGB stars
with $M > 1.5 M_{\odot}$?



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Metal-poor AGB stars
with $M > 1.5 M_{\odot}$?

Accreting white
dwarf?



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Chemical clocks

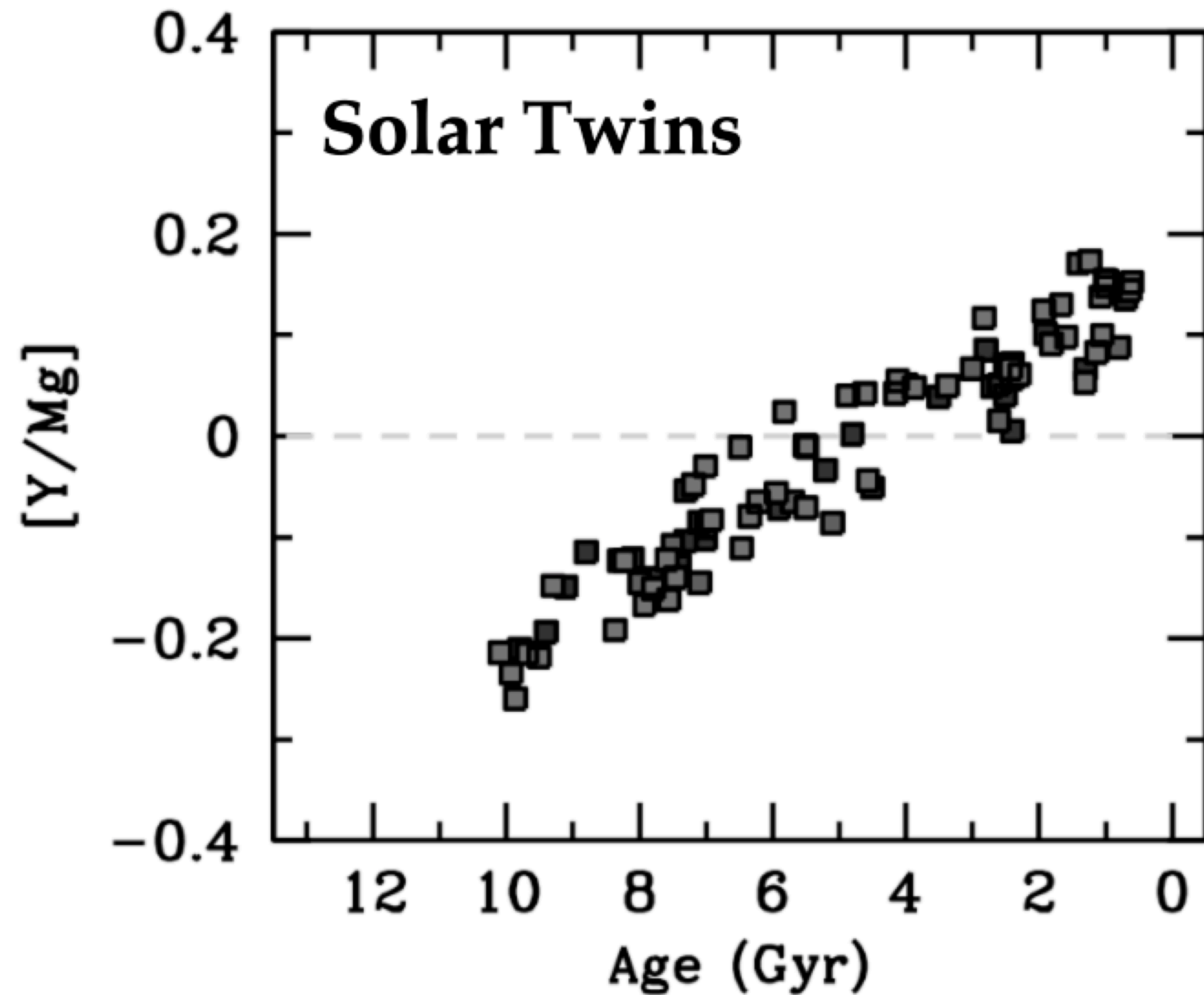


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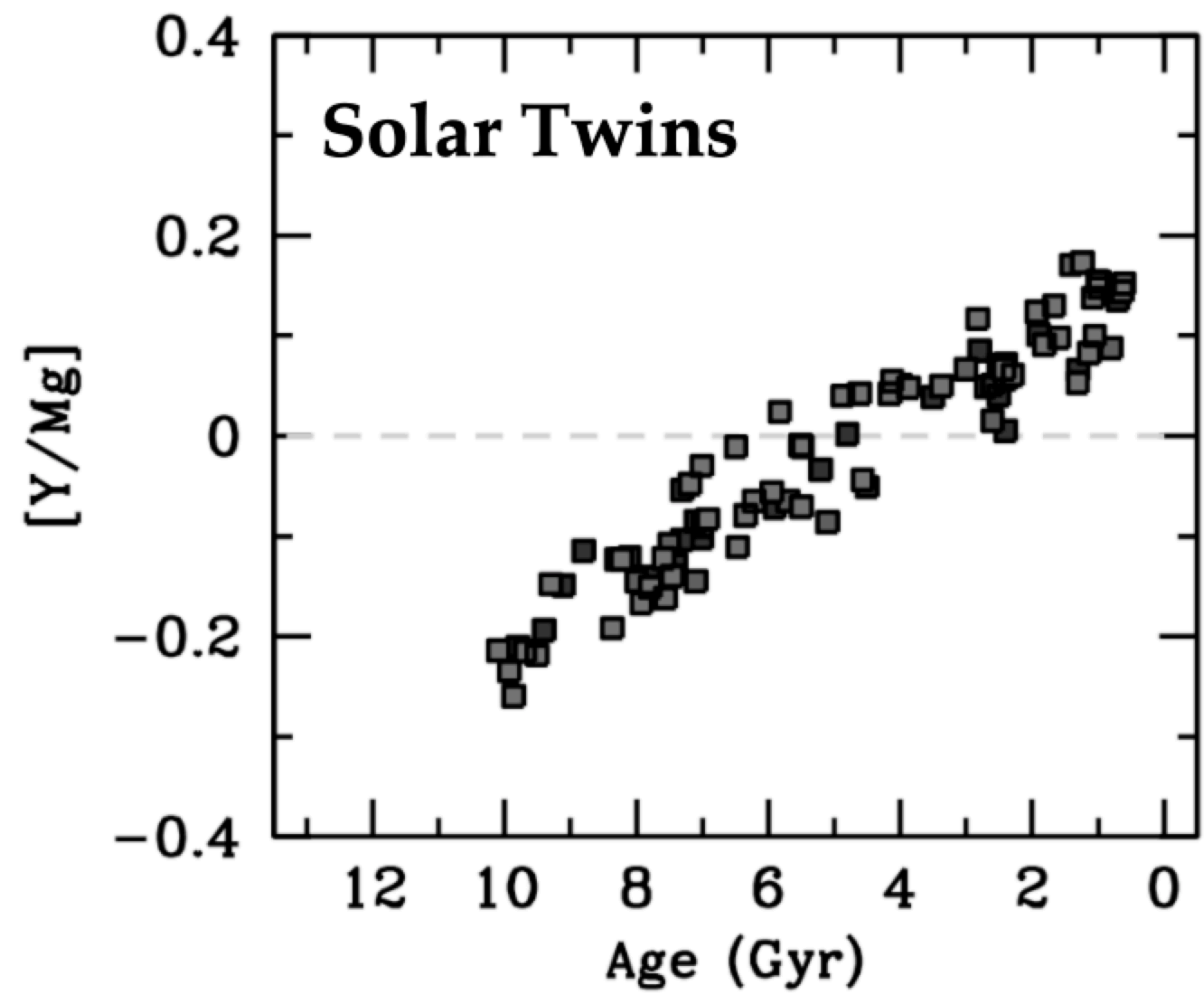
CHEMICAL CLOCKS

*Nissen 2015, 2017, Spina et al. 2016,
Tucci-Maia et al. 2016*

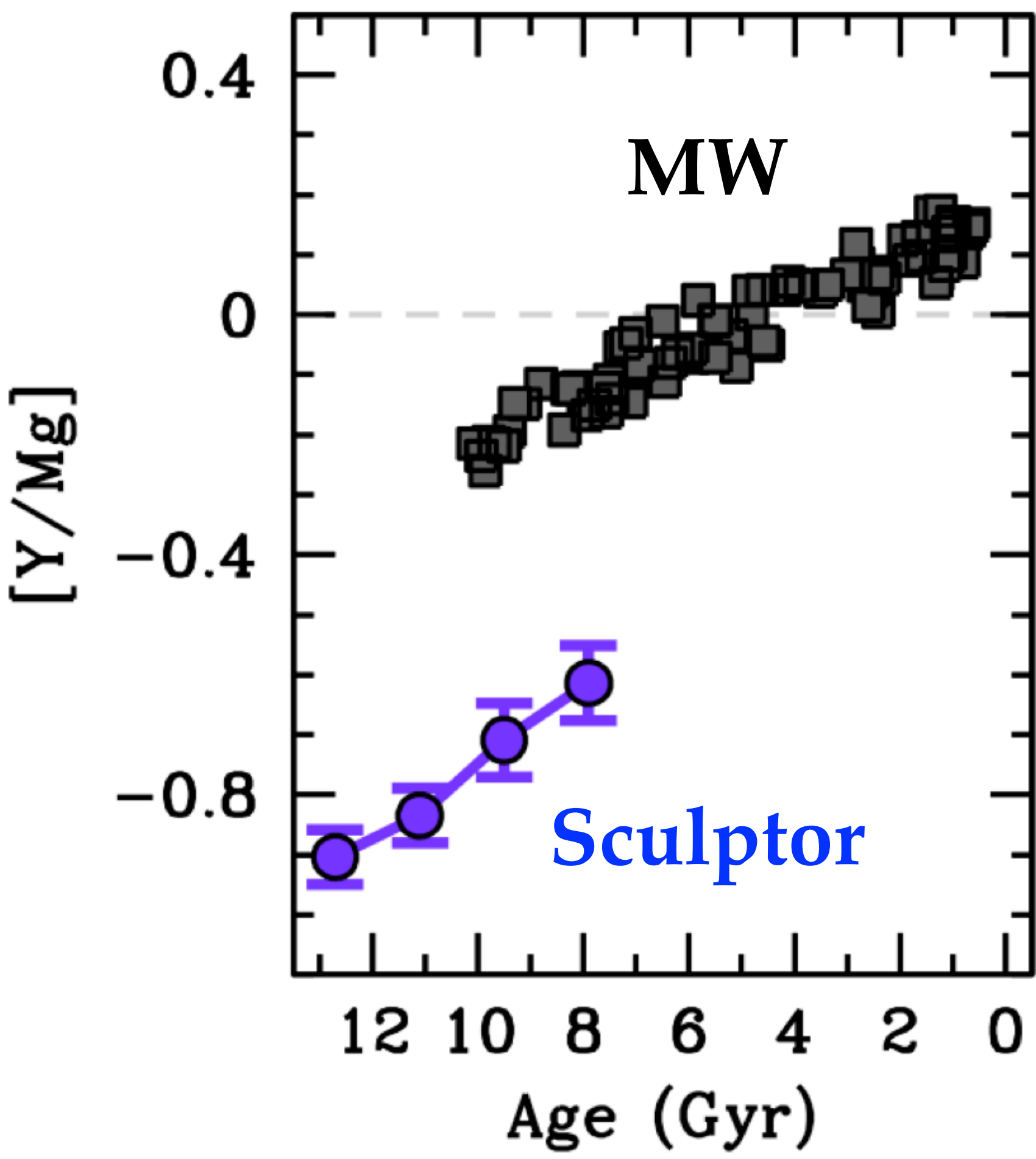


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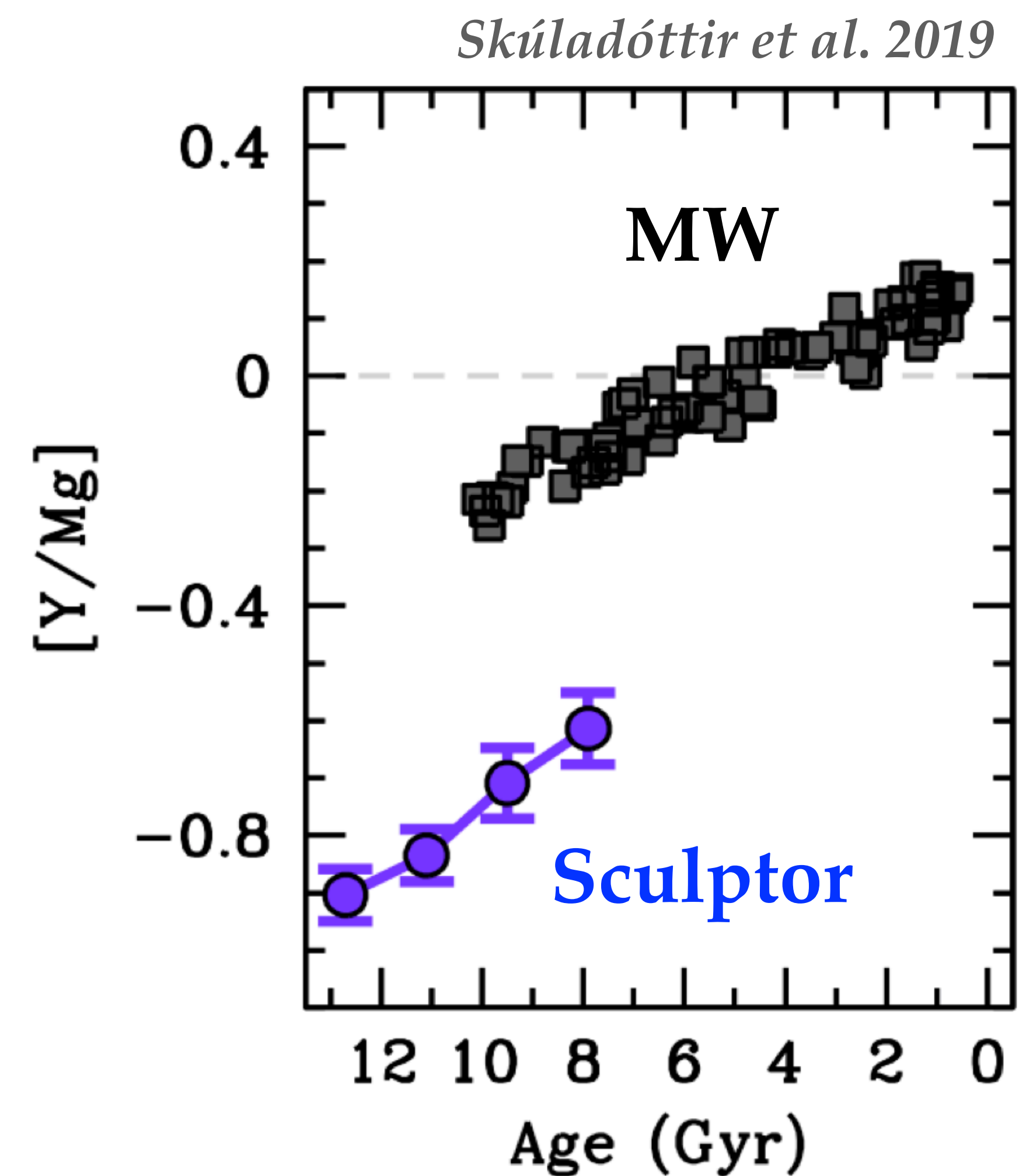
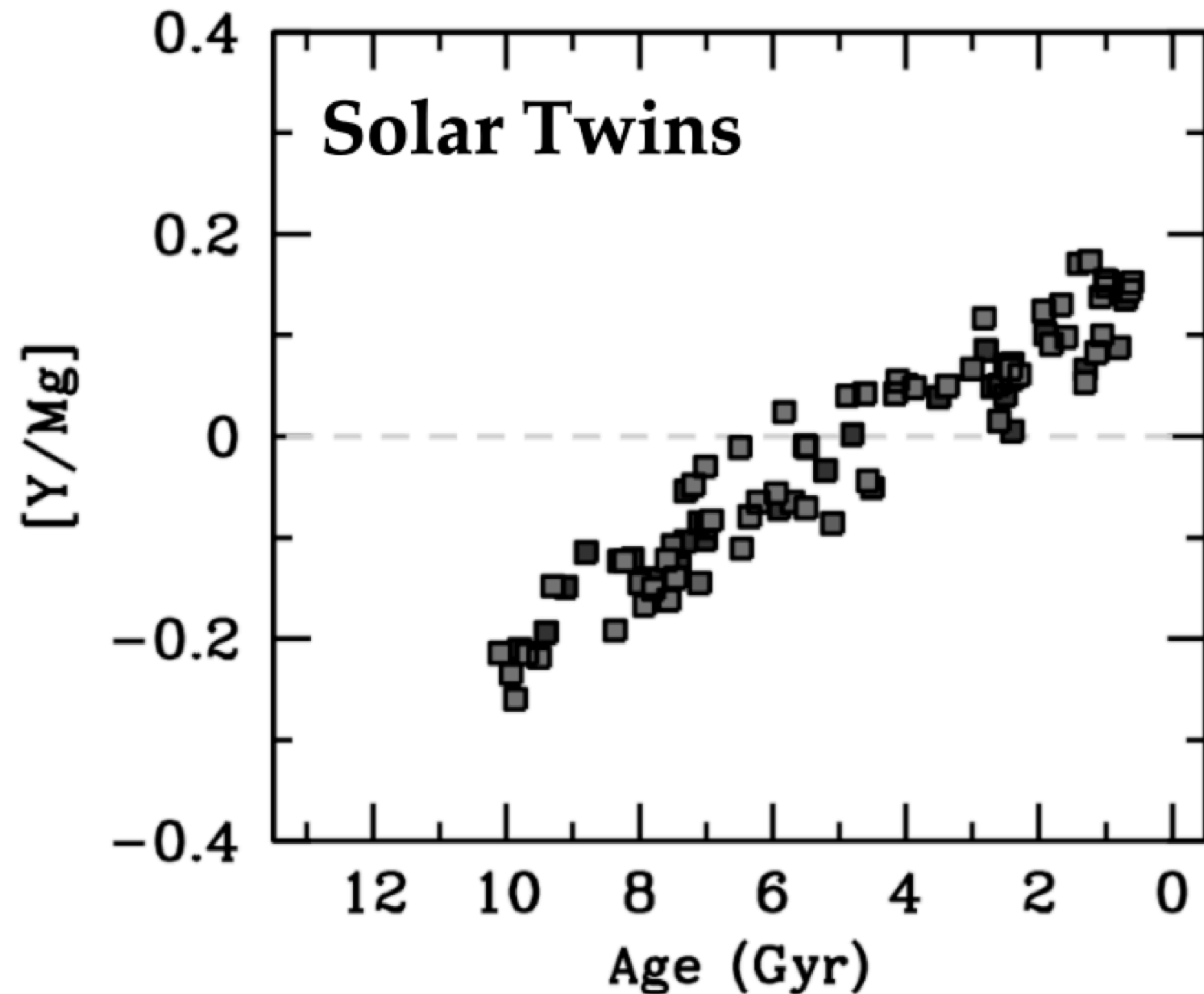


Skúladóttir et al. 2019



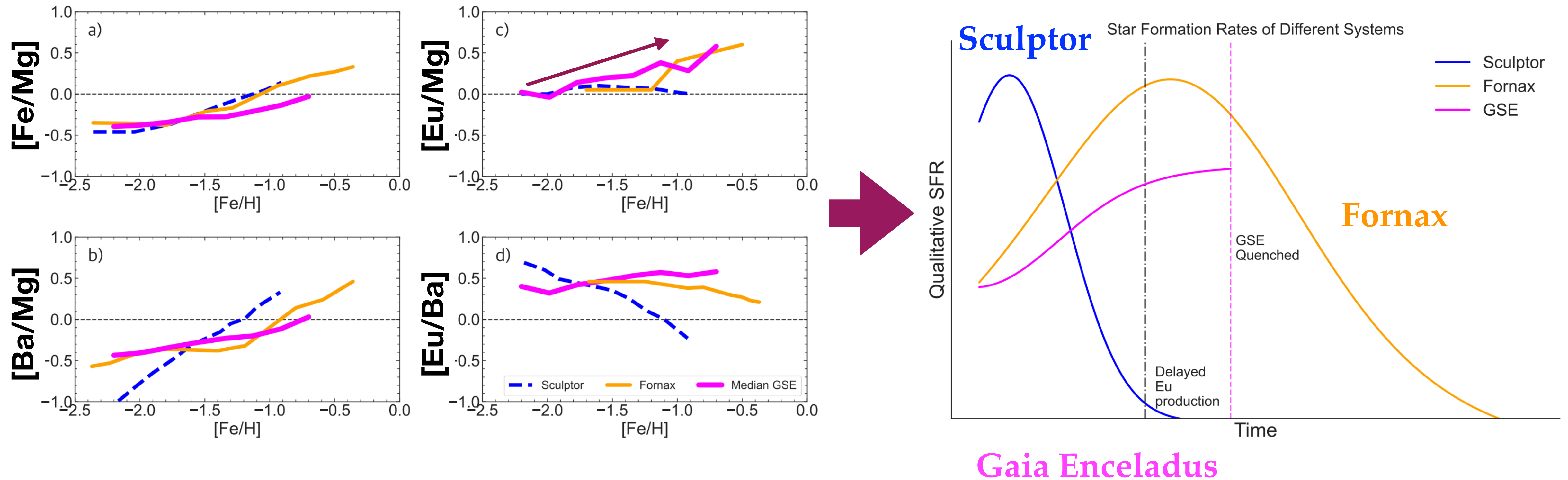
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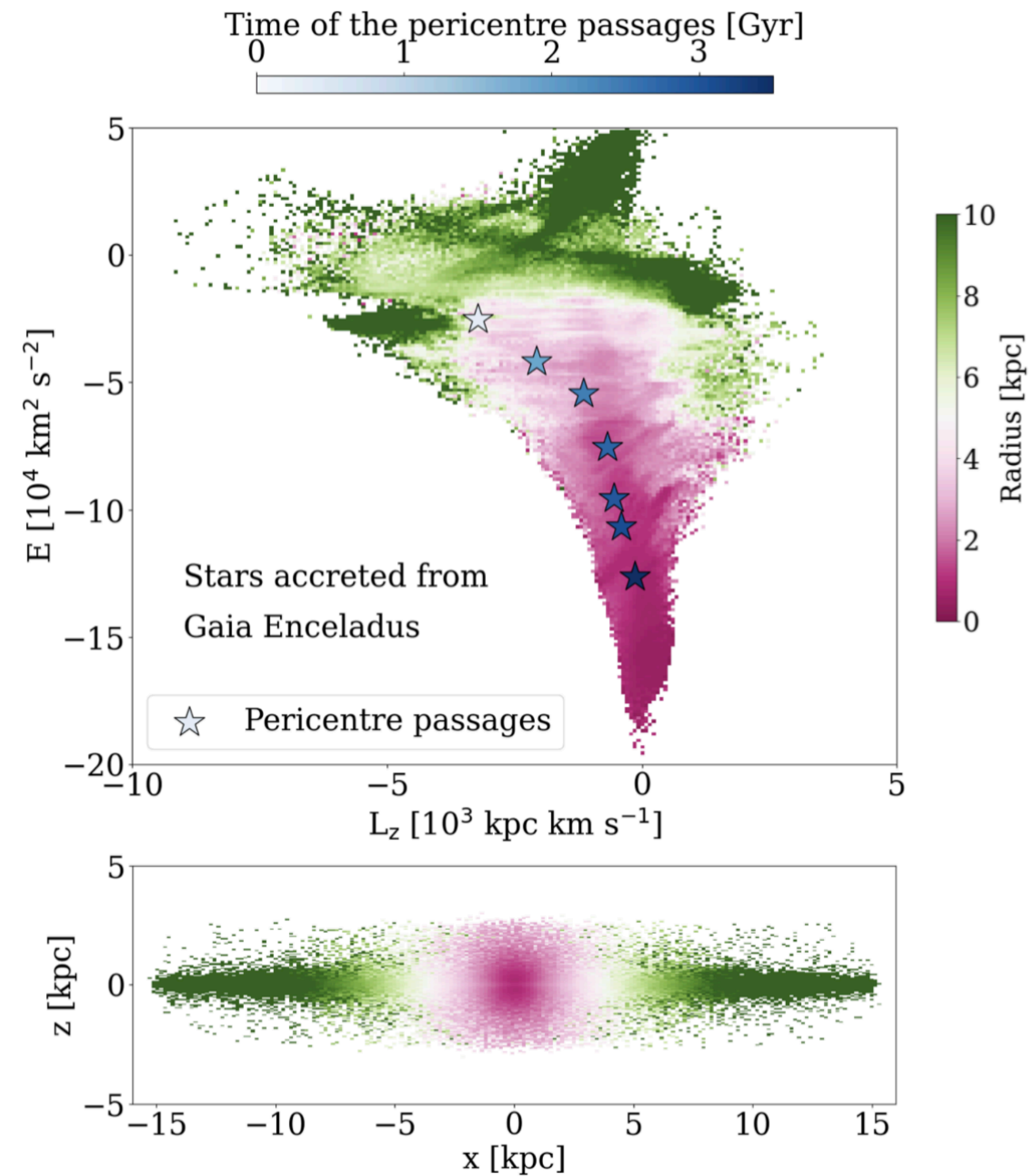
Chemical clocks are not universal

USING “CHEMICAL CLOCKS”

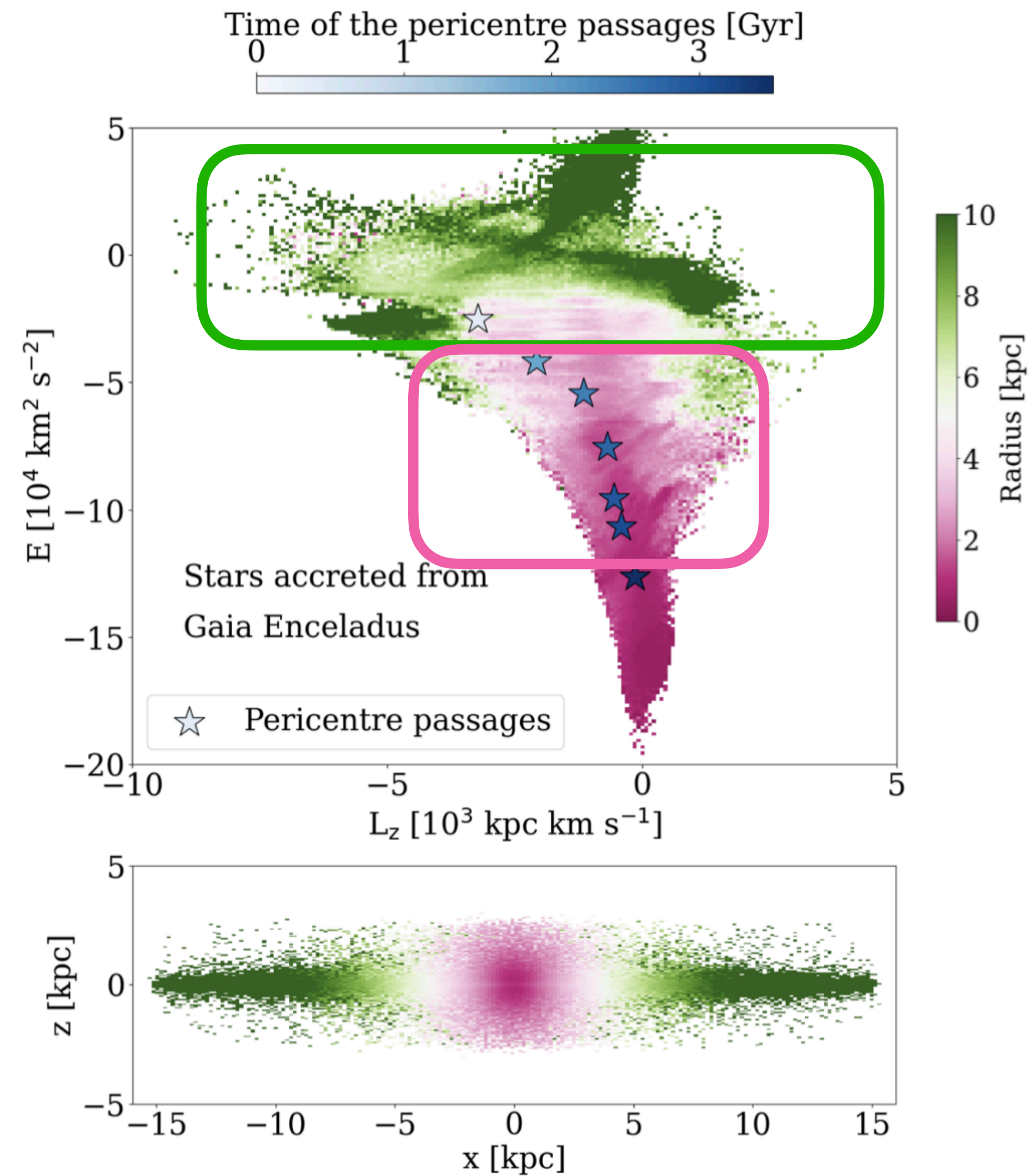


Gaia Enceladus was the last major merger event of the Milky Way
Comparing the abundance trends reveals its star formation history

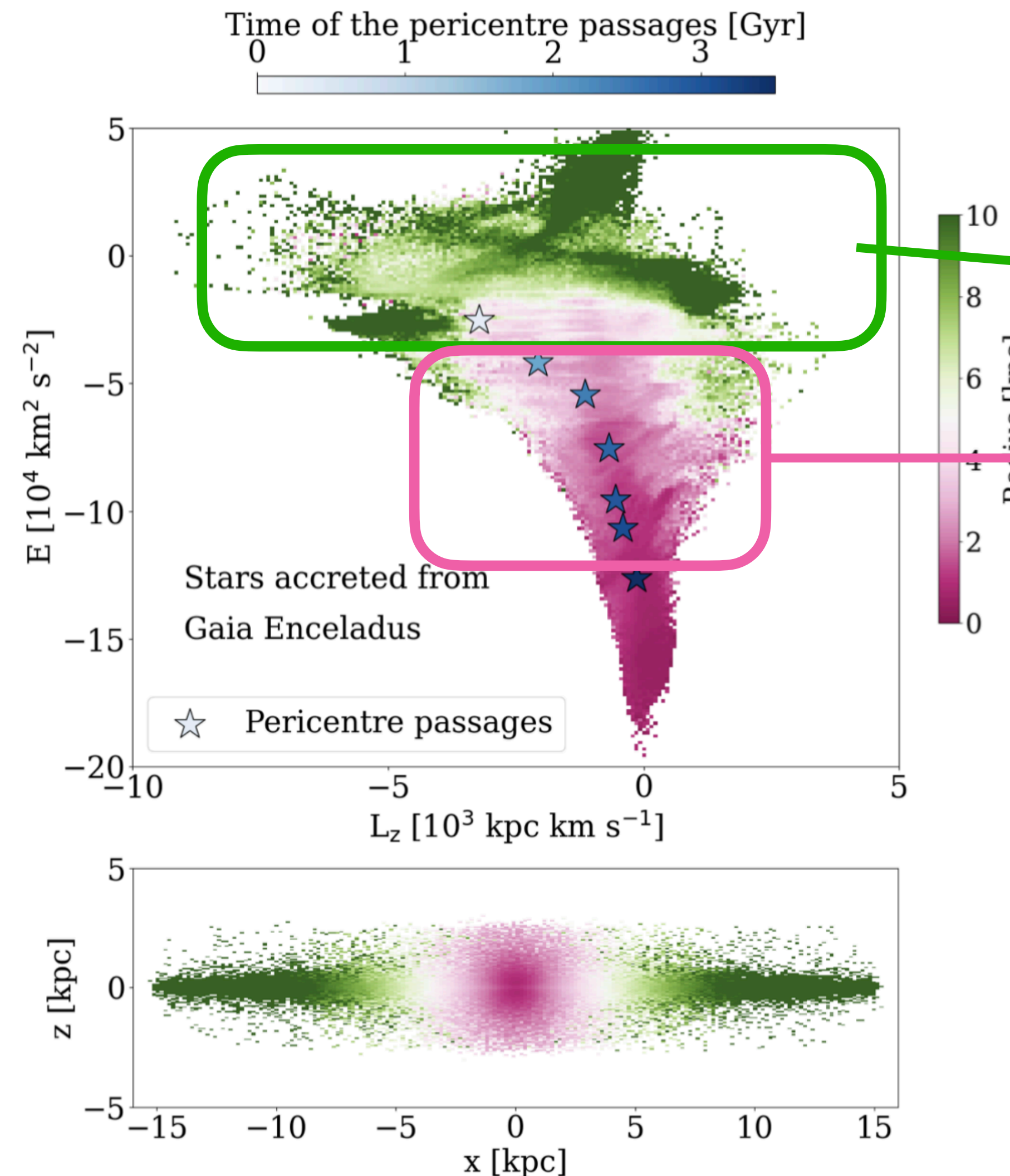
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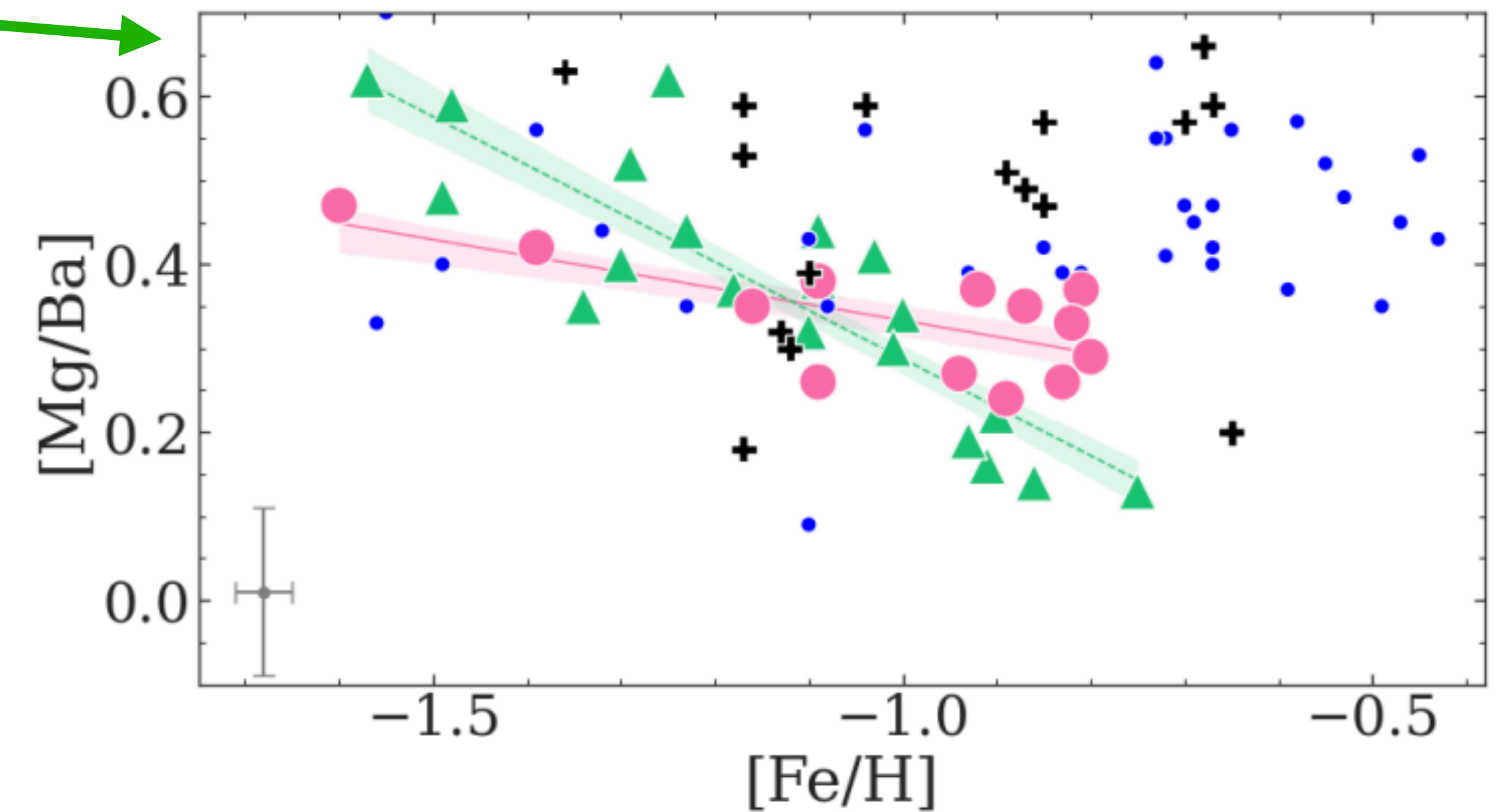
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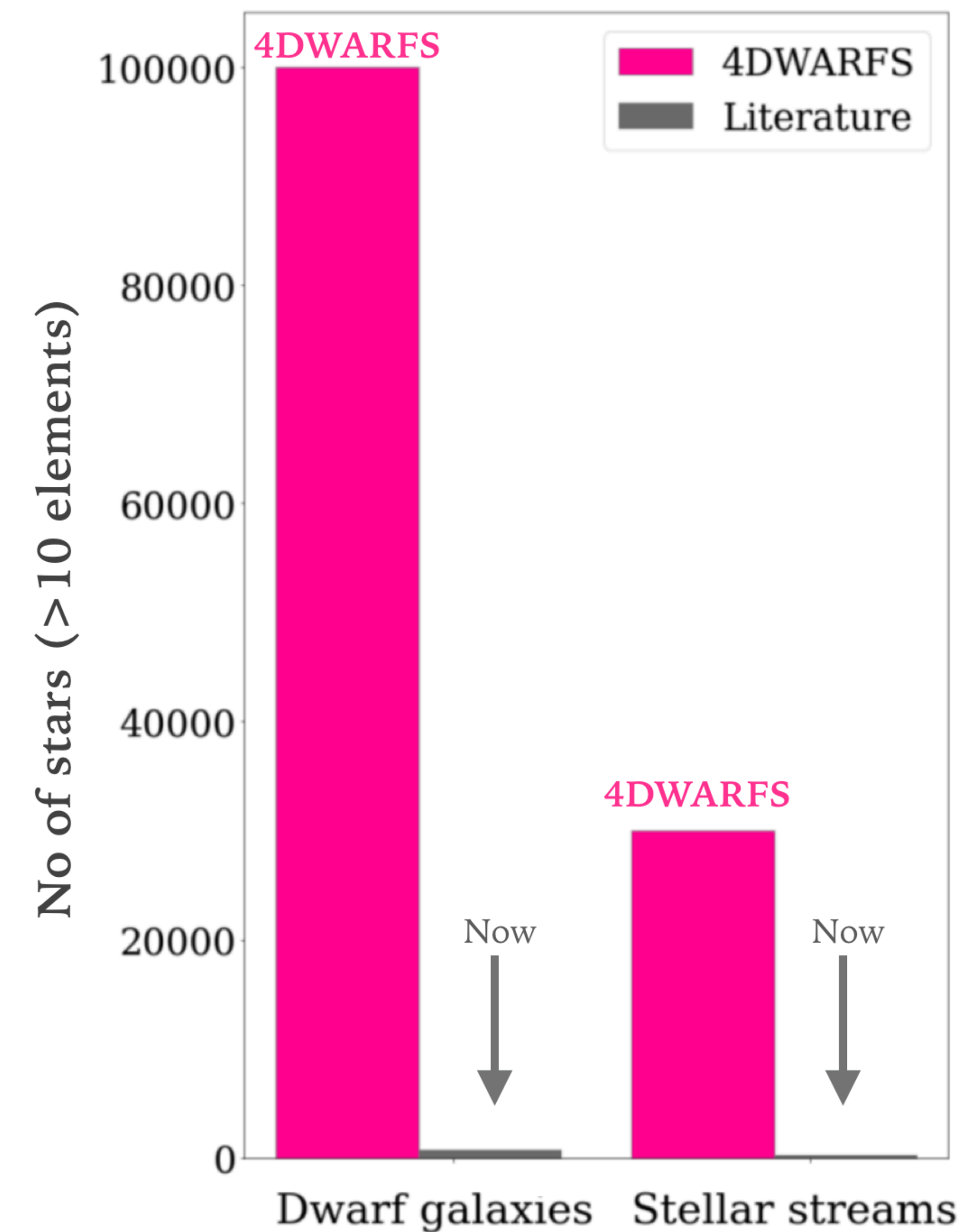
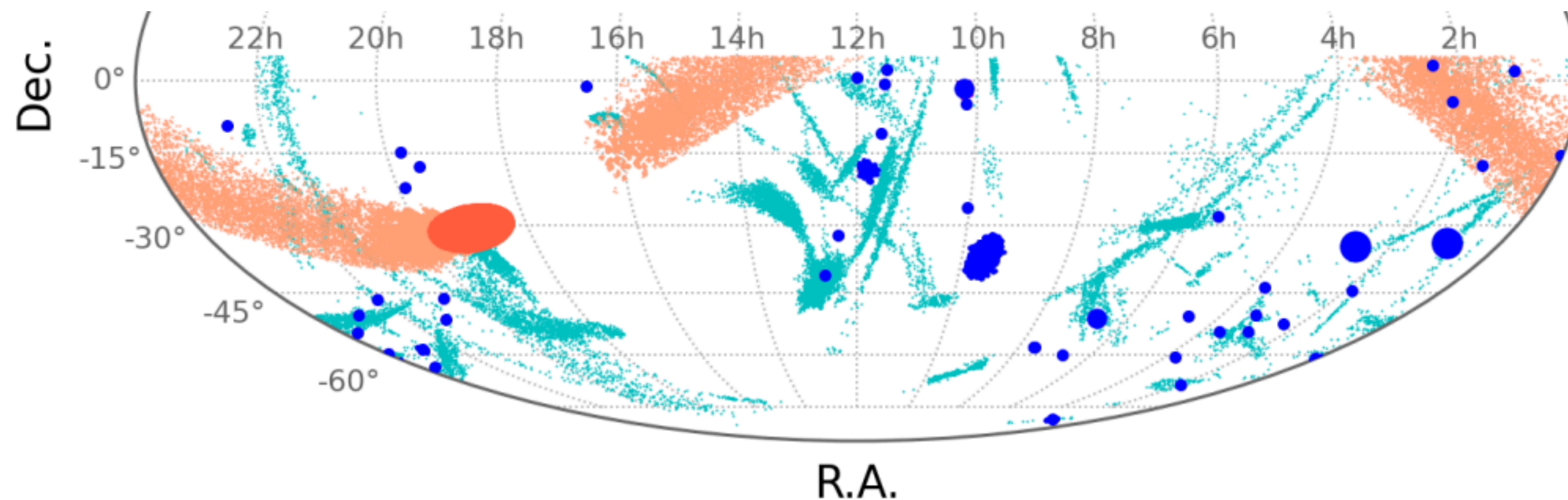
Stars accreted first from the outer regions of Gaia Enceladus



Stars accreted later from the inner regions of Gaia Enceladus

4DWARFS SURVEY

- 4DWARFS - Survey of Dwarf Galaxies and their Stellar streams (PI: Skúladóttir).
- All dwarf galaxies and streams in the south.
- Large discovery space of unexpected findings.
- First light in 2025



LESSONS FROM DWARF GALAXIES



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- There are (at least) two r-process sources - quick (ccSN) and delayed (NSM).
- Neutron star mergers have typical timescales of ≈ 2 Gyr.
- The i-process is needed to explain the n-capture elements in the Sculptor dwarf spheroidal galaxy.
 - Metal-poor AGB stars with $M > 1.5 M_{\odot}$?
- Chemical clocks are not universal - but they are useful!
- Data are lacking - but Data are coming!

This project has received funding from the European Research Council (ERC) under the European Union's Horizon 2020 research and innovation programme (TREASURES - grant agreement No. 101117455, PI: Skúladóttir).



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